
An Introduction to the Physics and Technology of e⁺e⁻ Linear Colliders

Lecture 7: Beam-Beam Effects

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Introduction

- Beam-beam interaction in a linear collider is basically the same Coulomb interaction as in a storage ring collider. But:
 - Interaction occurs only once for each bunch (single pass); hence very large bunch deformations permissible.
 - Extremely high charge densities at IP lead to very intense fields; hence *quantum* behaviour becomes important
- Consequently, can divide LC beam-beam phenomena into two categories:
 - *classical*
 - *quantum*

Introduction (continued)

- Beam-Beam Effects

- Electric field from a “flat” charge bunch
- Equation of motion of an electron in flat bunch
- The Disruption Parameter (D_y)
- Crossing Angle and Kink Instability
- Beamstrahlung
- Pair production (background)

	E_{beam}	β_x	β_y	σ_x	σ_y	σ_z	N_e
	GeV	mm	mm	nm	nm	μm	$\times 10^{10}$
NLC	250	8	0.1	245	2.7	110	0.75
TESLA	250	14	0.4	550	5	300	2.0
CLIC	1500	8	0.15	43	1	30	0.42

Storage Ring Collider Comparison

Linear beam-beam tune shift $\xi_{x,y} = \frac{r_e N_b}{2\pi\gamma} \frac{\beta_{x,y}}{\sigma_{x,y}(\sigma_x + \sigma_y)}$

Putting in some typical numbers (see previous table) gives:

$$\xi_x = 0.54$$

$$\xi_y = 1.44$$

Storage ring colliders try to keep $\xi_{x,y} < 0.05$

Electric Field from a Relativistic Flat Beam

- Highly relativistic beam $\mathbf{E} + \mathbf{v} \times \mathbf{B} \approx 2E$
- Flat beam $\sigma_x \gg \sigma_y$ (*cf Beamstrahlung*)
- Assume
 - infinitely wide beam with constant density per unit length in x ($\equiv \rho(x)$)
 - Gaussian charge distribution in y :
 - for now, leave $\rho(z)$ unspecified

$$\rho(y) \approx \frac{1}{\sqrt{2\pi}\sigma_y} \exp\left[-\frac{1}{2}\left(\frac{y}{\sigma_y}\right)^2\right] \quad \rho(x) \approx \frac{1}{\sqrt{2\pi}\sigma_x}$$

Electric Field from a Relativistic Flat Beam

Use Gauss' theorem: $\oiint_s \mathbf{E} \cdot d\mathbf{s} = \frac{q}{\epsilon_0}$

$$E_y(y, z) \Delta x \Delta z \approx \frac{qN\rho(x)\rho(z)\Delta x \Delta z}{\epsilon_0} \int_{y'=0}^y \rho(y') dy'$$

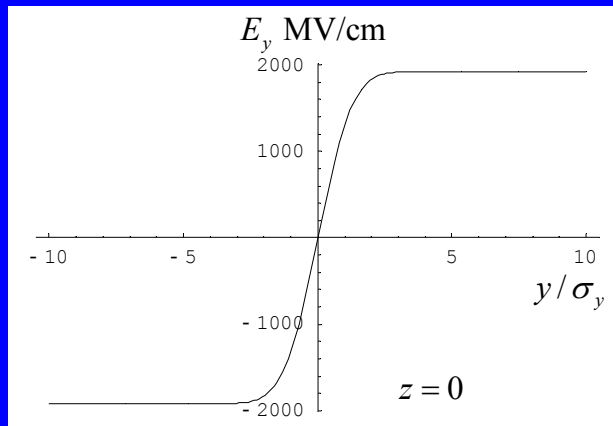
$$E_y(y, z) = \frac{qN}{2\sqrt{2\pi}\epsilon_0\sigma_x} \operatorname{Erf}\left(\frac{y}{\sqrt{2}\sigma_y}\right) \rho(z)$$

Assuming Gaussian distribution for z , the peak field is given by

$$\hat{E}_y = \frac{qN}{4\pi\epsilon_0\sigma_x\sigma_z}$$

Electric Field from a Relativistic Flat Beam

$$\begin{aligned}\sigma_x &= 500 \text{ nm} \\ \sigma_y &= 5 \text{ nm} \\ \sigma_z &= 300 \text{ }\mu\text{m} \\ N &= 10^{10}\end{aligned}$$

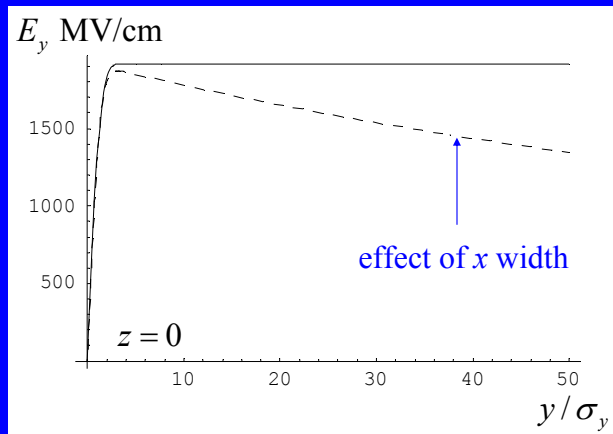


Note: $2 \times E_y$ plotted

Assuming a Gaussian distribution for $\rho(z)$

Electric Field from a Relativistic Flat Beam

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Assuming a Gaussian distribution for $\rho(z)$

Equation Of Motion

$$F = ma: \quad \ddot{y}(t) = -\frac{2qE_y(y,t)}{\gamma m_0}$$

Changing variable to z : $\ddot{y}(t) = c^2 y''(z)$

$$y''(z) = -\frac{2qE_y(y,z)}{\gamma m_0 c^2}$$

$$= -\frac{q^2 N \operatorname{Erf}\left(\frac{y(z)}{\sqrt{2}\sigma_y}\right) \rho_z(2z)}{\sqrt{2\pi} \gamma \epsilon_0 \sigma_x m_0 c^2} \quad \text{why } \rho_z(2z)?$$

$$r_e = \frac{q^2}{4\pi\epsilon_0 m_0 c^2}$$

$$y''(z) = -\frac{2\sqrt{2\pi} N r_e}{\gamma \sigma_x} \operatorname{Erf}\left(\frac{y(z)}{\sqrt{2}\sigma_y}\right) \rho_z(2z)$$

Linear Approximation and the Disruption Parameter

Taking only the linear part of the electric field: $y''(z) \approx -\underbrace{\frac{4Nr_e \rho(2z)}{\gamma \sigma_x \sigma_y}}_{k^2(z)} y(z)$

Take 'weak' approximation: $y(z)$ does not change during interaction $y(z) = y_0$

$$\Delta y' \approx -\frac{4Nr_e}{\gamma \sigma_x \sigma_y} y_0 \int_{-\infty}^{\infty} \rho_z(2z) dz = -\frac{2Nr_e}{\gamma \sigma_x \sigma_y} y_0 \equiv -\frac{1}{f} y_0$$

Thin-lens focal length: $\frac{1}{f} = \frac{2Nr_e}{\gamma \sigma_x \sigma_y}$

Define Vertical *Disruption*

Parameter

$$D_y \equiv \frac{\sigma_z}{f} = \frac{2Nr_e \sigma_z}{\gamma \sigma_x \sigma_y}$$

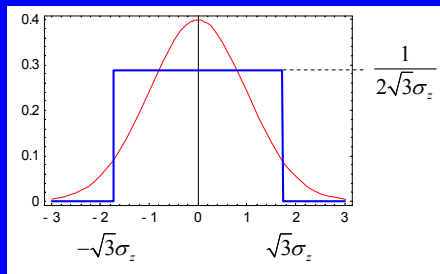
exact: $D_y = \frac{2Nr_e \sigma_z}{\gamma (\sigma_x + \sigma_y) \sigma_y}$

Number of Oscillations

Equation of motion re-visited:

$$y''(z) = -\frac{2D_y}{\sigma_z} \rho_z(2z)y(z)$$

Approximate $\rho(z)$ by rectangular distribution with same RMS as equivalent Gaussian distribution (σ_z)



half-length!

$$y''(z) = -\frac{D_y}{\sqrt{3}\sigma_z^2} y(z); \quad |z| \leq \frac{\sqrt{3}}{2} \sigma_y$$

$$k^2 = \frac{D_y}{\sqrt{3}\sigma_z^2} \quad \nu = \frac{\sqrt{3}k\sigma_z}{2\pi}$$

$$\Rightarrow \nu = \frac{3^{3/4} \sqrt{D_y}}{2\pi} \approx 0.21 \sqrt{D_y}$$

Example of Numerical Solution

$$N = 2 \times 10^{10}$$

$$\sigma_x = 500 \text{ nm}$$

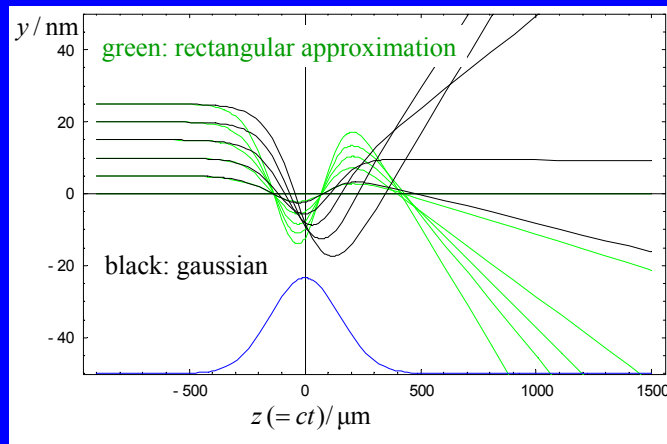
$$\sigma_y = 5 \text{ nm}$$

$$\sigma_z = 300 \text{ } \mu\text{m}$$

$$E = 250 \text{ GeV}$$

$$D_y = 27.7$$

$$\nu = 1.1$$



Pinch Enhancement

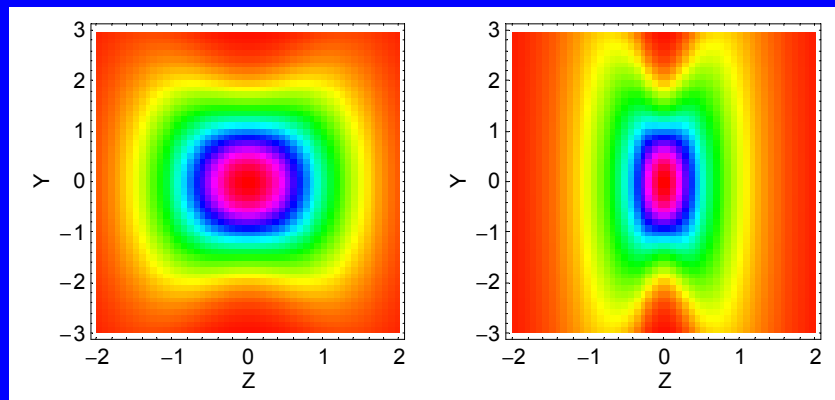
- Self-focusing (pinch) leads to higher luminosity for a head-on collision.

$$H_{D_{x,y}} = 1 + D_{x,y}^{1/4} \left(\frac{D_{x,y}^3}{1 + D_{x,y}^3} \right) \left[\ln(\sqrt{D_{x,y}} + 1) + \underbrace{2 \ln\left(\frac{0.8\beta_{x,y}}{\sigma_z}\right)}_{\text{'hour glass' effect}} \right]$$

Empirical fit to beam-beam simulation results

Only a function of disruption parameter $D_{x,y}$

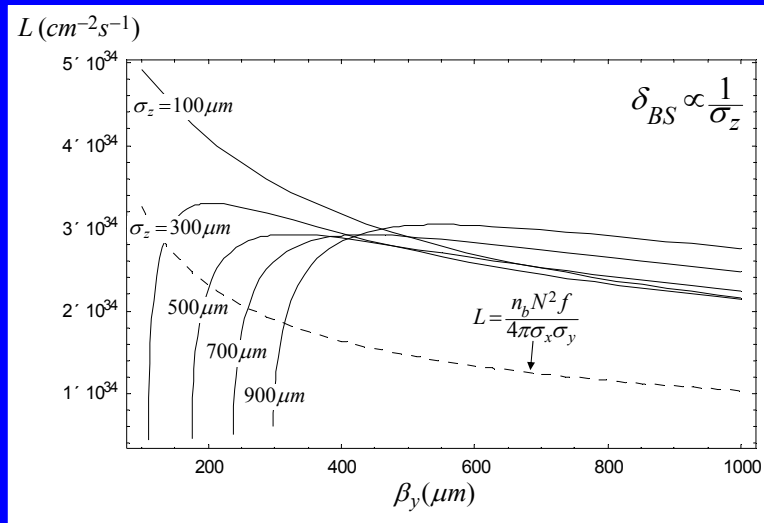
The *Luminosity* Issue: Hour-Glass



β = "depth of focus"

reasonable lower limit for
 β is bunch length σ_z

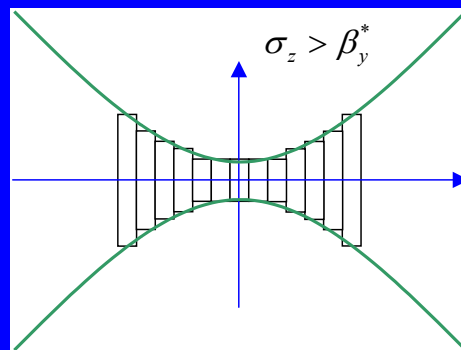
Luminosity as a function of β_y



Beating the hour glass effect

Travelling focus (Balakin)

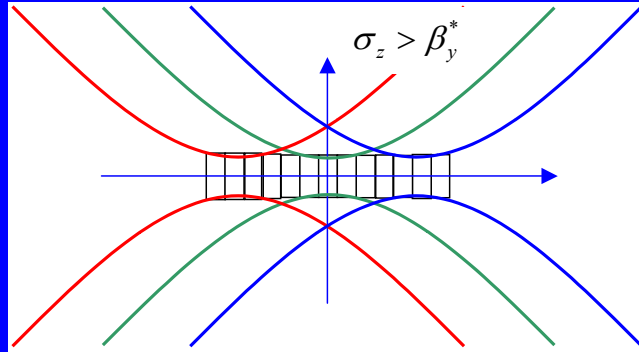
- Arrange for finite chromaticity at IP (how?)
- Create z -correlated energy spread along the bunch (how?)



Beating the hour glass effect

Travelling focus

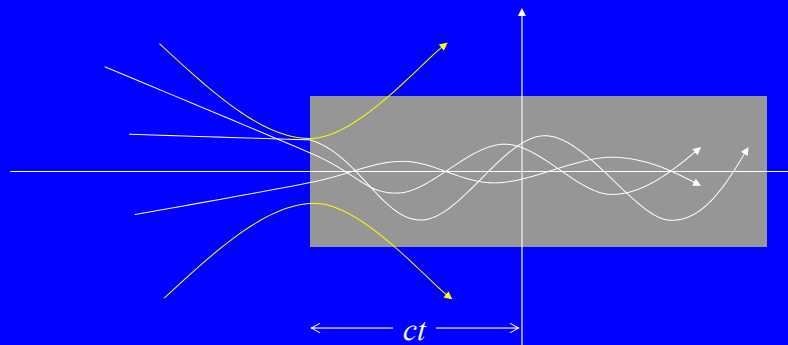
- Arrange for finite chromaticity at IP (how?)
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Beating the hour glass effect

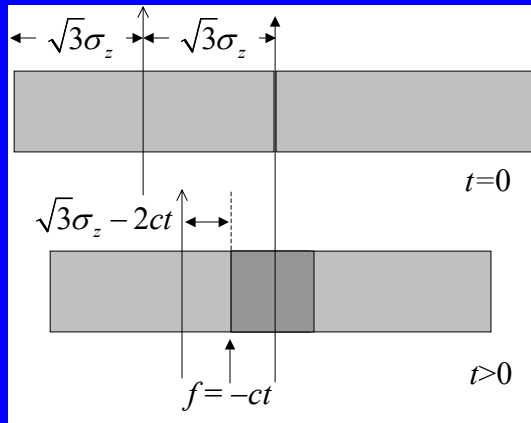
Travelling focus

- Arrange for finite chromaticity at IP (how?)
- Create z -correlated energy spread along the bunch (how?)



Beating the hour glass effect

foci 'travel' from $z = 0$ to $z = \pm\sqrt{3}\sigma_z$



chromaticity:

$$\Delta f_y = \xi_y \delta$$

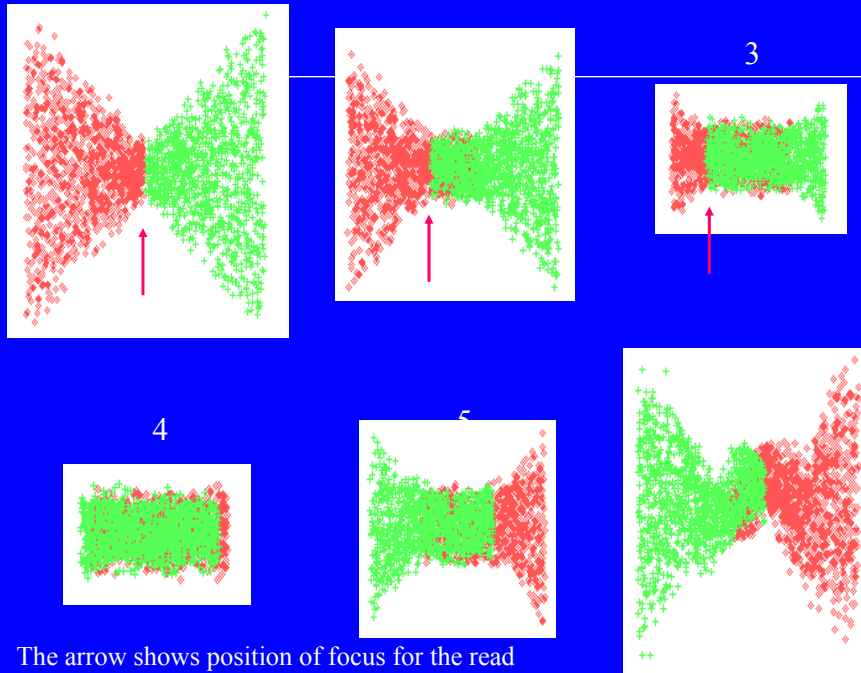
travelling focus:

NB: z correlated!

$$\delta_{\text{RMS}} \approx \frac{\sigma_z}{2\xi_y}$$

$$\langle z\delta \rangle \approx 2\xi_y \delta_{\text{RMS}}^2$$

$$\approx \frac{\sigma_z^2}{2\xi_y}$$



The arrow shows position of focus for the read beam during travelling focus collision

Kink Instability

Simple model: ‘sheet’ beams with: $\rho_{x,z} = \frac{1}{2\sqrt{3}\sigma_{x,z}}$

Linear equation of motion becomes

$$y''(z) = \frac{\sqrt{2\pi}}{6} \frac{D_y}{\sigma_z^2} y(z); \quad |z| \leq \frac{\sqrt{3}}{2} \sigma_z$$

Need to consider relative motion of *both* beams in t and z :

$$\begin{aligned} \left(\frac{\partial}{\partial t} + c \frac{\partial}{\partial z} \right)^2 y_1(t, z) &= -\omega_0^2 (y_1 - y_2) \\ \left(\frac{\partial}{\partial t} - c \frac{\partial}{\partial z} \right)^2 y_2(t, z) &= +\omega_0^2 (y_1 - y_2) \end{aligned} \quad \omega_0^2 = \frac{\sqrt{2\pi}}{6} c^2 \frac{D_y}{\sigma_z^2}$$

Classic coupled EoM.

Kink Instability

Assuming solutions of the form

$$y_{1(2)} = a_{1(2)} \exp[i(kz - \omega t)]$$

and substituting into EoM leads to the dispersion relation:

$$\omega^2 = c^2 k^2 + \omega_0^2 \pm \sqrt{4\omega_0^2 c^2 k^2 + \omega_0^4}$$

Motion becomes unstable when $\omega^2 \leq 0$, which occurs when

$$|k| \leq \frac{\sqrt{2}\omega_0}{c}$$

Kink Instability

Exponential growth rate:

$$\left[\sqrt{4\omega_0^2 c^2 k^2 + \omega_0^4} - \omega_0^2 + c^2 k^2 \right]^{1/2}$$

Maximum growth rate when $k = \pm \frac{\sqrt{3}}{2} \frac{\omega_0}{c}$; $\omega = i \frac{\omega_0}{2}$

Remember! $-\frac{\sqrt{3}}{2} \frac{\sigma_z}{c} \leq t \leq \frac{\sqrt{3}}{2} \frac{\sigma_z}{c} \Rightarrow \Delta t = \sqrt{3} \sigma_z / c$

Thus ‘amplification’ factor of an initial offset is:

$$\exp\left(\frac{\omega_0 \Delta t}{2}\right) = \exp\left[\frac{1}{2} \left(\frac{\pi}{2}\right)^{1/4} \sqrt{D_y}\right] \approx e^{0.6 \sqrt{D_y}}$$

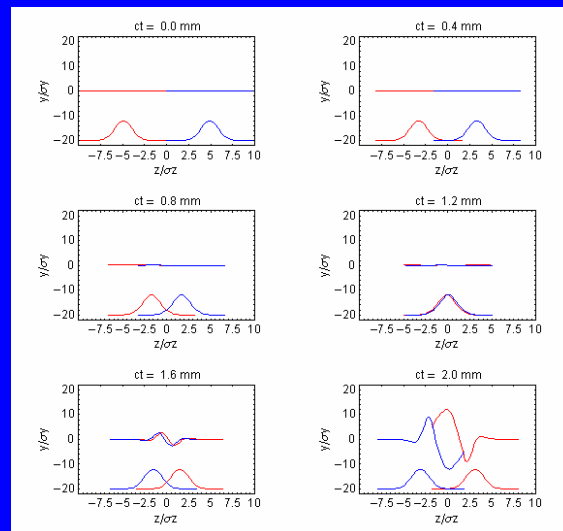
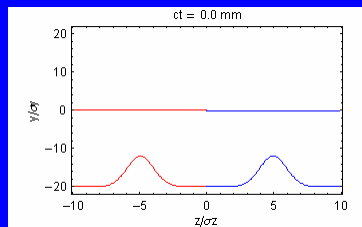
For our previous example with $D_y \sim 28$, factor ~ 3

Kink Instability

$$\Delta y = 0.1 \sigma_y$$

$$D_y \approx 20$$

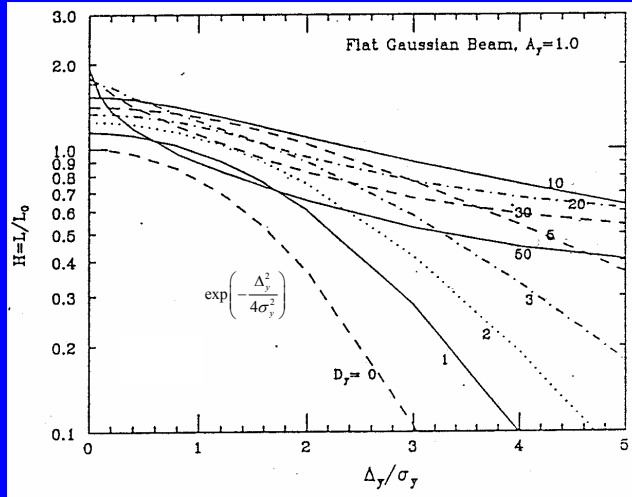
$$\sigma_z = 300 \mu\text{m}$$



Pinch Enhancement

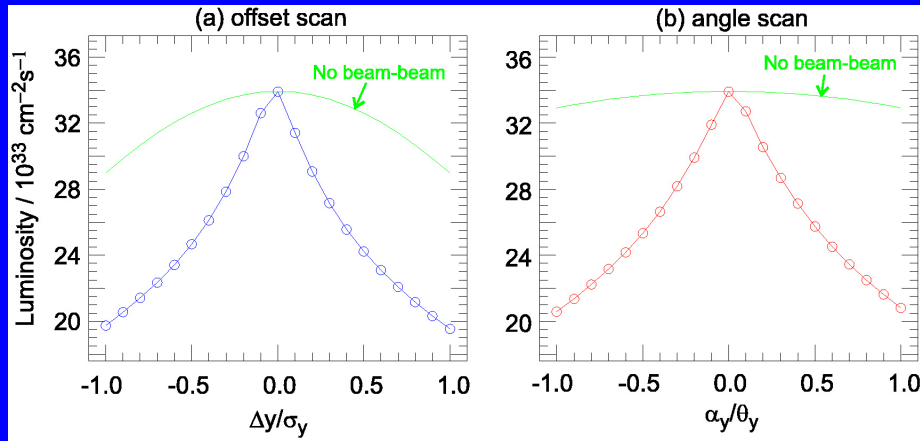
$$L = L_{geom} H \quad H = \text{enhancement factor}$$

results of simulations:



High D_y example: TESLA 500

$$D_y \approx 24$$



Disruption Angle

Remembering definition of D_y $D_y = \frac{\sigma_z}{f}$

The angles after collision are characterised by

$$\theta_0 \equiv \frac{D_x \sigma_x}{\sigma_z} = \frac{D_y \sigma_y}{\sigma_z} = \frac{2Nr_e}{\gamma(\sigma_x + \sigma_y)} \approx \frac{2Nr_e}{\gamma\sigma_x}$$

Numbers from our previous example give $\theta_0 \approx 467 \mu\text{rad}$

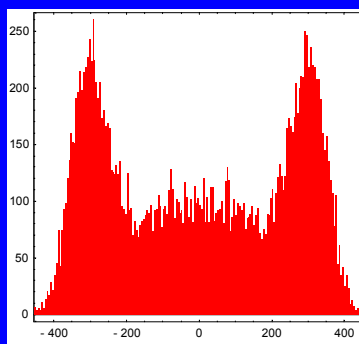
OK for horizontal plane where $D_x < 1$

For vertical plane (strong focusing $D_y > 1$), particles oscillate:

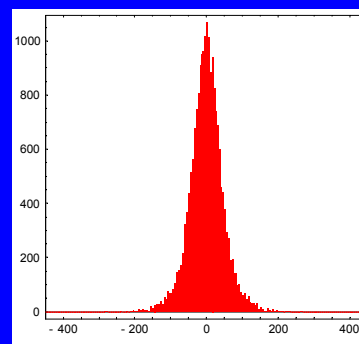
previous linear approximation:

$$y''(z) = -\frac{D_y}{\sqrt{3}\sigma_z^2} y(z); \quad |z| \leq \frac{\sqrt{3}}{2} \sigma_y, \quad \theta \approx \frac{\sqrt{D_y}}{3^{3/4} \sigma_z} \sigma_y = \frac{\theta_0}{3^{3/4} D_y^{1/2}} \approx 67 \mu\text{rad}$$

Disruption Angle: simulation results



horizontal angle (μrad)



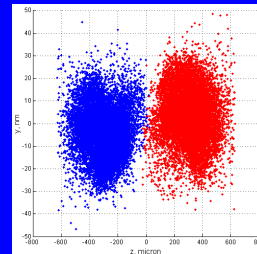
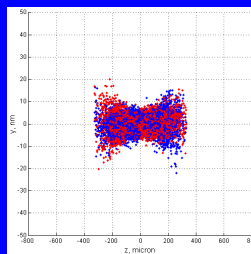
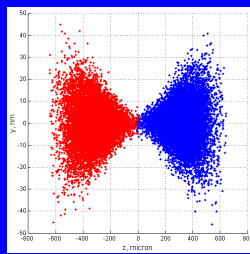
vertical angle (μrad)

Important in designing IR (spent-beam extraction)

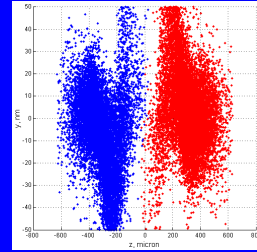
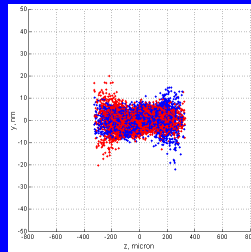
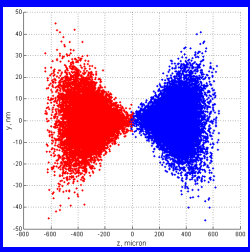
Beam-Beam Animation Wonderland

Animations produced by A. Seryi using the **GUINEAPIG** beam-beam simulation code (written by D. Schulte, CERN).

Examples of GUINEAPIG Simulations

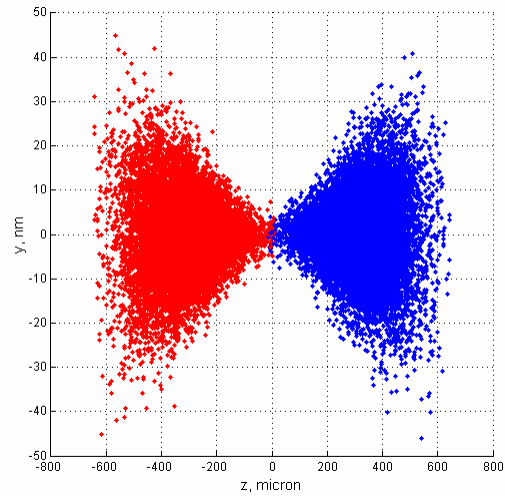


NLC
parameters
 $D_y \sim 12$



Nx2
 $D_y \sim 24$

Examples of GUINEAPIG Simulations



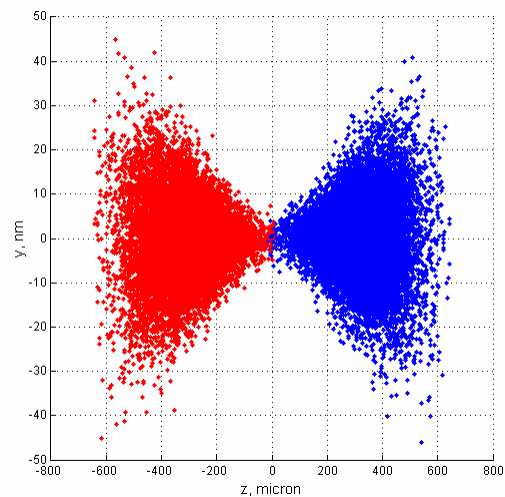
NLC parameters
 $D_y \sim 12$

Luminosity
enhancement
 $H_D \sim 1.4$

Not much of an
instability



Examples of GUINEAPIG Simulations



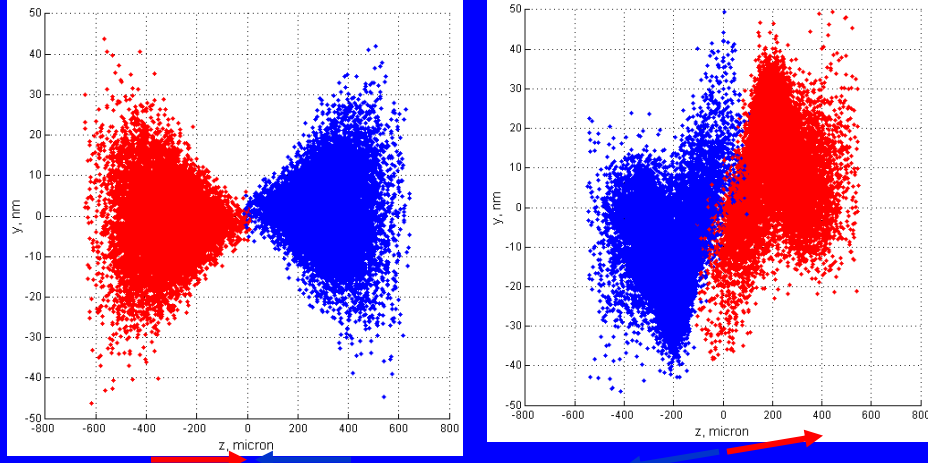
$N_x 2$
 $D_y \sim 24$

Beam-beam
instability is
clearly pronounced

Luminosity
enhancement is
compromised by
higher sensitivity
to initial offsets

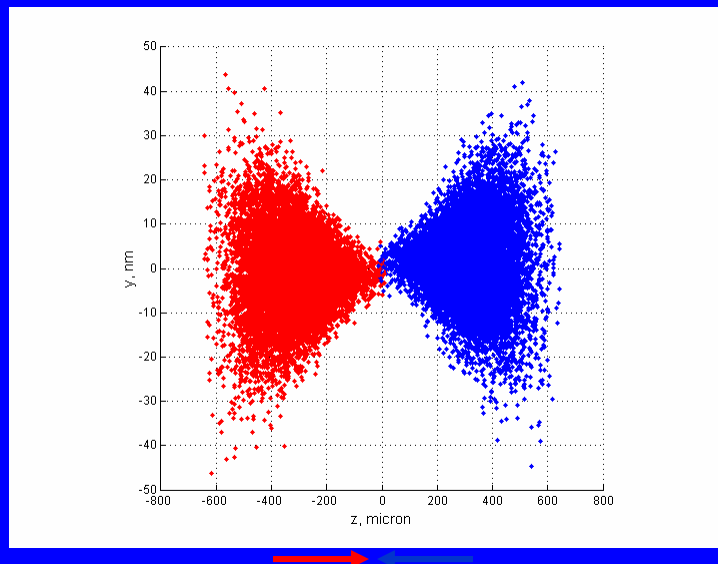


Beam-beam deflection

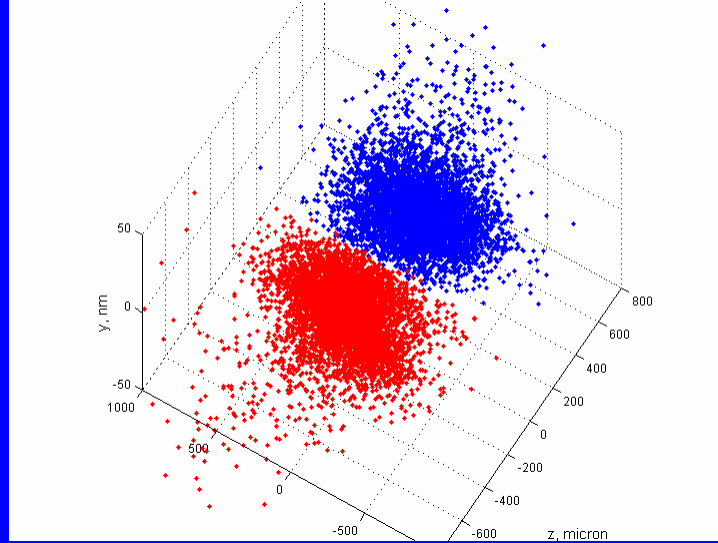


Sub nm offsets at IP cause large well detectable offsets (micron scale) of the beam a few meters downstream

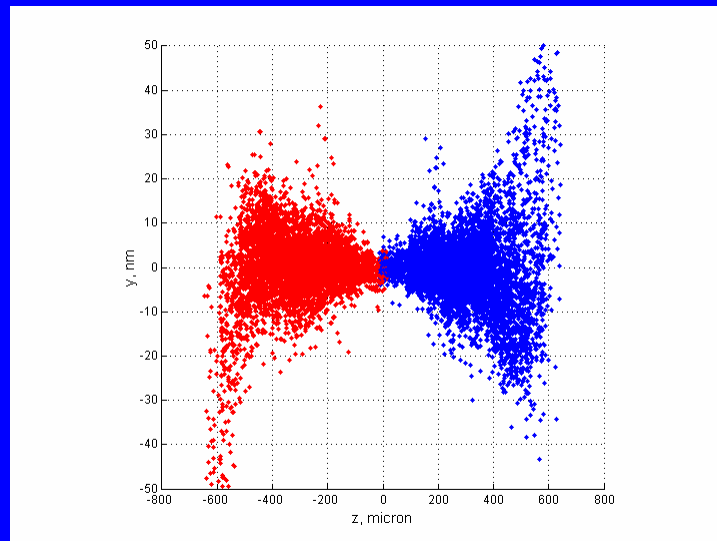
Beam-beam deflection allow to control collisions



Examples of GUINEAPIG Simulations



Examples of GUINEAPIG Simulations



Beam-Beam Kick

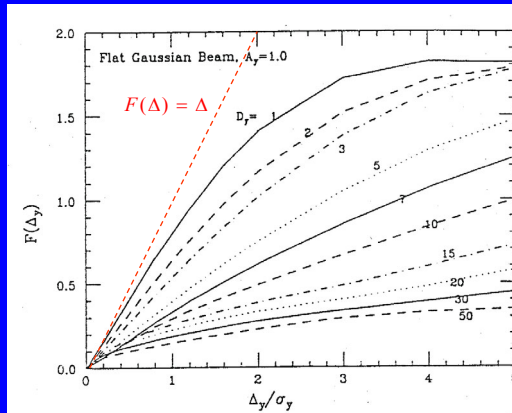
Beam-beam offset gives rise to an equal and opposite mean kick to the bunches – *important signal for feedback!*

For small disruption ($D \ll 1$) and offset ($y/\sigma_y \ll 1$)

$$\theta_{bb} = \frac{1}{2} \theta_0 \frac{y}{\sigma_y}$$

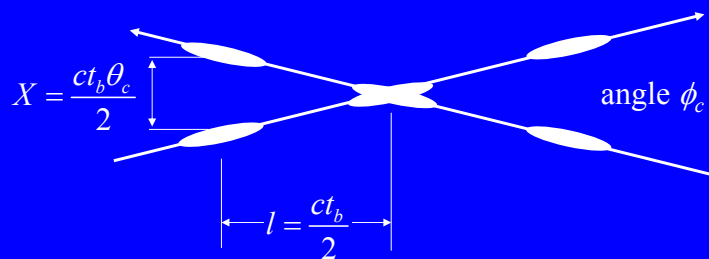
For large disruption or offset, we introduce the form factor F :

$$\theta_{bb} = \frac{1}{2} \theta_0 F(y/\sigma_y)$$



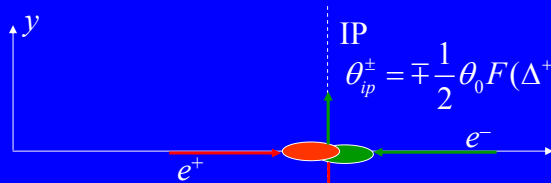
Long Range Kink Instability & Crossing Angle

To avoid parasitic bunch interactions in the IR, a horizontal crossing angle is introduced:



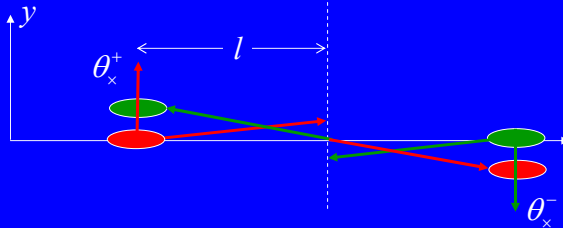
parasitic beam-beam kick: $\Delta r' = \frac{2Nr_e}{\gamma r} r \square \sigma_{x,y}$

Long Range Kink Instability & Crossing Angle



$$\theta_{ip}^{\pm} = \mp \frac{1}{2} \theta_0 F(\Delta^+ - \Delta^-)$$

small vertical offset
($\Delta^{\pm} = \delta y^{\pm} / \sigma_y$) gives
rise to beam-beam
kick



resulting vertical
offset at parasitic
crossing gives next
incoming bunches
additional vertical
kick: IP offset
increases

$$\theta_x^{\mp} = \pm \frac{2N r_e}{\gamma} \frac{l \theta_{ip}^{\pm}}{X^2} = \mp \frac{N r_e}{\gamma} \frac{\theta_0 F(\Delta^+ - \Delta^-)}{l \phi_c^2}$$

\Rightarrow instability

Long Range Kink Instability & Crossing Angle

offset at IP of k -th bunch: $\Delta_k \equiv \Delta_k^+ - \Delta_k^-$

distance from IP to encounter with l -th bunch ($l < k$) $l_{lk} = \frac{(k-l)t_b c}{2}$

contribution from encounter with l -th bunch:

$$\begin{aligned} \Delta_k^{\pm} &= \Delta_{k,0}^{\pm} + l_{lk} \theta_{x,lk}^{\pm} / \sigma_y \\ &= \Delta_{k,0}^{\pm} \pm \frac{N r_e}{\gamma} \frac{\theta_0}{\phi_c^2 \sigma_y} F(\Delta_l) \end{aligned}$$

$$\begin{aligned} \Delta_k &= \Delta_k^+ - \Delta_k^- \\ &= \Delta_{k,0} + \frac{2N r_e \theta_0}{\gamma \phi_c^2 \sigma_y} F(\Delta_l) \end{aligned} \quad \begin{array}{l} \text{NB. independent} \\ \text{of } l_{lk} \end{array}$$

total offset: $\Delta_k = \Delta_{k,0} + C \sum_{i=1}^{k-1} F(\Delta_i); \quad C = \frac{2N r_e \theta_0}{\gamma \phi_c^2 \sigma_y} = D_x D_y \left[\frac{\sigma_x / \sigma_z}{\phi_c} \right]^2$

Long Range Kink Instability & Crossing Angle

Assuming all bunches have same initial offset: $\Delta_{k,0} = \Delta_0$

For small disruption and small offsets, $F(\Delta_i) \approx \Delta_i$

$$\Delta_k = (1 + C)^{k-1} \Delta_0$$

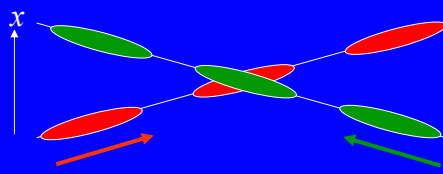
$$\text{since } C = D_x D_y \left[\frac{\sigma_x / \sigma_z}{\phi_c} \right]^2 \ll 1 \quad \frac{\Delta_k}{\Delta_0} \approx 1 + (k-1)C$$

thus $(m_b - 1)C \leq 1$ where m_b = number of interacting bunches

example:

$$\left. \begin{array}{l} D_{x,y} = 0.1, 10 \\ \sigma_x = 220 \text{ nm} \\ \sigma_z = 100 \text{ } \mu\text{m} \\ \phi_c = 20 \text{ mrad} \end{array} \right\} C = 0.012 \Rightarrow m_b < 80 \text{ for factor of 2 increase}$$

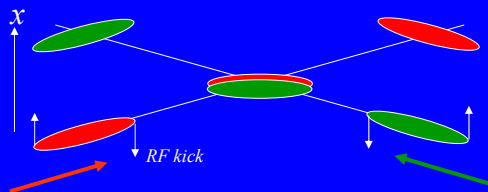
Crab Crossing



$$\begin{aligned} \sigma_{x, \text{projected}} &\approx \sqrt{\sigma_x^2 + \phi_c^2 \sigma_z^2} \\ &\approx \phi_c \sigma_z \\ &= 20 \text{ mrad} \times 100 \mu\text{m} \approx 2 \mu\text{m} \end{aligned}$$

factor 10 reduction in L !

use transverse (crab) RF cavity to 'tilt' the bunch at IP



$$V(z) = \hat{V}_{RF} \sin\left(\frac{2\pi z}{\lambda}\right) \approx \frac{2\pi \hat{V}_{RF} z}{\lambda}$$

$$\phi_c \approx 4\pi \sqrt{\beta_{cav} \beta_{ip}} \frac{\hat{V}_{RF}}{\lambda_{RF}} \sigma_z$$

Beamstrahlung

Magnetic field of bunch $B = E/c$

Peak field:
$$B_{\max} = \frac{2E_{\max}}{c} = \frac{qN}{2\pi\epsilon_0 c \sigma_x \sigma_z} = 1160 \text{ Tesla}$$

From classical theory, power radiated is given by

$$P_\gamma = \frac{cC_\gamma}{2\pi} \frac{E^4}{\rho^2}; \quad C_\gamma \approx 8.85 \times 10^{-5} \text{ m} \cdot \text{GeV}^{-3}$$

For $E = 250 \text{ GeV}$ and $\sigma_z = 300 \mu\text{m}$:

$$\frac{1}{\rho} \approx \frac{cB}{E} \approx 1.4 \text{ m}^{-1}$$

$$\Delta E_{\text{rad}} \approx P_\gamma \Delta t \approx P_\gamma \frac{\sigma_z}{c} \approx 32 \text{ GeV} \quad \Delta E/E \sim 12\%$$

note: $\sigma_z / \rho \approx 200 \mu\text{rad} \square 1/\gamma \approx 2 \mu\text{rad}$

Beamstrahlung

Most important parameter is Υ

$$\Upsilon = \frac{2}{3} \frac{\hbar \omega_c}{E} = \frac{\lambda_e \gamma^2}{\rho} = \gamma \frac{2B}{B_s} = \frac{e}{m_0^2} \sqrt{\left| (F_{\mu\nu} p^\nu)^2 \right|}$$

ω_c critical photon frequency

λ_e Compton wavelength

ρ local bending radius

B beam magnetic field

B_s Schwinger's critical field (= 4.4 GTesla)

$F_{\mu\nu}$ em field tensor

p^ν electron 4-momentum

Beamstrahlung: photon spectrum

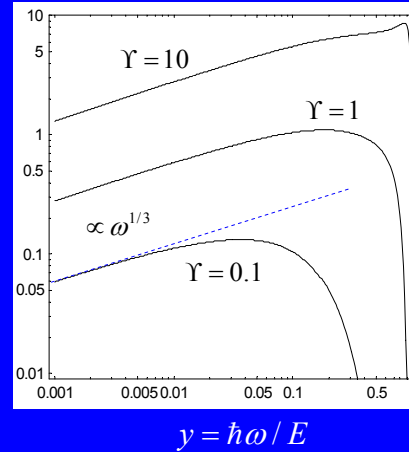
NOT classical synchrotron radiation spectrum!

Need to use Sokolov-Ternov formula:

$$F_{BS} = \underbrace{\int_{\xi}^{\infty} K_{\frac{2}{3}}(\xi') d\xi'}_{\text{classical}} + \underbrace{\frac{y^2}{1-y} K_{\frac{2}{3}}(\xi)}_{\text{quantum theoretical}}$$

$$y = \frac{\hbar\omega}{E}$$

$$\xi = \frac{2}{3\Upsilon} \left(\frac{y}{1-y} \right) = \left(\frac{\omega}{\omega_c} \right) \left(\frac{1}{1-y} \right)$$



Beamstrahlung Numbers

average and maximum Υ

$$\Upsilon_{avr} = \frac{5}{6} \frac{Nr_e \lambda_e \gamma}{\sigma_z (\sigma_x + \sigma_y)}$$

$$\Upsilon_{max} \approx 2.4 \Upsilon_{avr}$$

photons per electron:

$$n_\gamma \approx 2.54 \left[\frac{\alpha \sigma_z}{\lambda_e \gamma} \right] \frac{\Upsilon_{avr}}{\sqrt{1 + \Upsilon_{avr}^{2/3}}}$$

average energy loss: $\delta_{BS} = \left\langle -\frac{\Delta E}{E} \right\rangle \approx 1.24 \left[\frac{\alpha \sigma_z}{\lambda_e \gamma} \right] \frac{\Upsilon_{avr}^2}{[1 + (1.5 \Upsilon_{avr})^{2/3}]^2}$

δ_{BS} and n_γ (and hence Υ) talk directly to *luminosity spectrum* and *backgrounds*

Beamstrahlung energy loss

In lecture 1, we used the following equation to derive our luminosity scaling law

$$\delta_{BS} \approx 0.86 \frac{er_e^3}{2m_0c^2} \left(\frac{E_{cm}}{\sigma_z} \right) \frac{N^2}{(\sigma_x + \sigma_y)^2}$$

Now we have this:

$$\delta_{BS} \approx 1.24 \left[\frac{\alpha \sigma_z}{\lambda_e \gamma} \right] \frac{\Upsilon_{avr}^2}{\left[1 + (1.5 \Upsilon_{avr})^{2/3} \right]^2}$$

with
$$\Upsilon_{avr} = \frac{5}{6} \frac{Nr_e \lambda_e \gamma}{\sigma_z (\sigma_x + \sigma_y)}$$

under what regime is our original expression valid?

Luminosity scaling revisited

$$\beta_y^* \approx \sigma_z$$

low beamstrahlung regime $\Upsilon \ll 1$:
$$L \propto P_{\text{beam}} \frac{\sqrt{\delta_{BS}}}{\sqrt{\epsilon_{y,n}}}$$

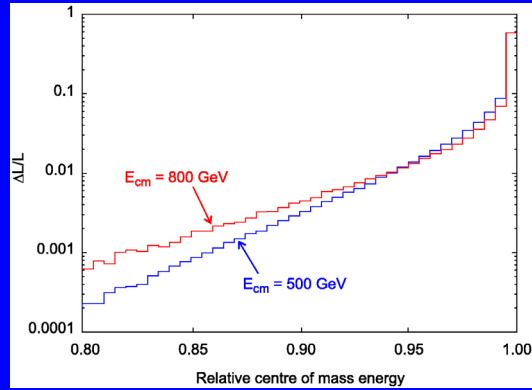
high beamstrahlung regime $\Upsilon \gg 1$:
$$L \propto P_{\text{beam}} \frac{\delta_{BS}^{3/2}}{\sigma_z \sqrt{\epsilon_{y,n}}}$$

homework: derive high BS scaling law

Beamstrahlung Numbers: example

$E = 500 \text{ GeV}$
 $\sigma_x = 220 \text{ nm}$
 $\sigma_y = 2 \text{ nm}$
 $\sigma_x = 110 \mu\text{m}$
 $N = 0.75 \times 10^{10}$

$\Upsilon_{avr} \approx 0.28$ ($E_{\gamma,c} \approx 210 \text{ GeV}$)
 $\Upsilon_{max} \approx 0.66$
 $n_\gamma \approx 1.3$
 $\delta_{BS} \approx 7.5\%$



TESLA

Why Beamstrahlung is bad

- Large number of high-energy photons interact with electron (positron) beam and generate e^+e^- pairs
 - Low energies ($\Upsilon < 0.6$), pairs made by *incoherent* process (photons interact directly with individual beam particles)
 - High energies ($0.6 < \Upsilon < 100$), *coherent* pairs are generated by interaction of photons with macroscopic field of bunch.
 - Very high energies ($\Upsilon > 100$), *coherent direct trident production*
 $e^\pm \rightarrow e^\pm e^+ e^-$

TESLA	0.5 TeV	$\Upsilon \sim 0.06$
NLC	1 TeV	$\Upsilon \sim 0.28$
CLIC	3 TeV	$\Upsilon \sim 9$

- Beamstrahlung degrades *Luminosity Spectrum*

Pair Production

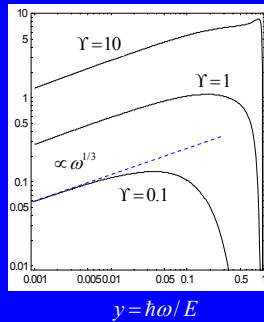
- Incoherent e^+e^- pairs $\Upsilon < 0.6$
 - Breit-Wheeler: $\gamma\gamma \rightarrow e^+e^-$
 - Bethe-Heitler: $e^+\gamma \rightarrow e^+e^+e^-$
 - Landau-Lifshitz: $e^+e^- \rightarrow e^+e^-e^+e^-$

- Coherent e^+e^- pairs $0.6 < \Upsilon < 100$

– threshold defined by

$$\chi \equiv \frac{\hbar\omega}{m_0c^2} \frac{2B}{B_s}$$

$$= \frac{\hbar\omega}{E} \Upsilon \geq 1$$



for $\Upsilon > 1$ $\hbar\omega/E \sim O(1) \Rightarrow \chi \geq 1$

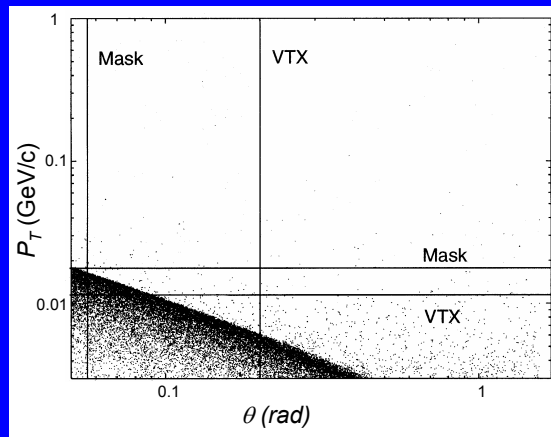
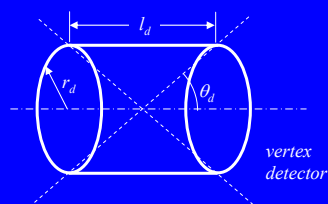
for intermediate colliders ($E_{cm} < 1\text{TeV}$),
incoherent pairs dominate

Pair Production

e^+e^- pairs are a
potential major source
of background

Most important: angle
with beam axis (θ) and
transverse momentum

P_T .



pairs curl-up (spiral) in
solenoid field of
detector

$$r = \frac{P_T}{cB_z} < r_d$$

Summary

- Single pass collider allows us to use very strong beam-beam to increase luminosity
- beam-beam is characterised by following important parameters:
 - $D_y = \sigma_z/f$ defines pinch effect (HD), kink instability, dynamics
 - Υ QM effects, backgrounds,
 - $\delta_{BS} [=f(\Upsilon_{av})]$ energy loss, lumi spectrum
- strong-strong regime requires simulation (e.g. GUINEAPIG). Analytical treatments limited.