

## 1.1. Kepler Problem and Action-Angle Variables (6 points)

Consider a mass  $m$  in a centrally symmetric potential  $V(r)$ ,  $r = \sqrt{x_1^2 + x_2^2 + x_3^2}$ . In the Kepler problem,  $V(r) = C/r$ , but we can keep  $V(r)$  general. The Hamiltonian reads

$$H = \sum_{i=1}^3 \frac{p_i^2}{2m} + V(r), \quad (1.1)$$

where  $p_i$  is the momentum conjugate to  $x_i$ .

- The phase space has six dimensions. Hence the system is integrable if there are three integrals of motion in involution. Show that the components  $J_i = \varepsilon_{ijk} x_j p_k$  of the angular momentum vector  $J = (J_1, J_2, J_3)$  are conserved.
- Compute the Poisson bracket  $\{J_i, J_j\}$  to show that the components  $J_i$  are *not* in involution. Do you recognize their algebra?
- The Hamiltonian  $H$  itself is another conserved quantity. Show that the combination  $J^2 = J_1^2 + J_2^2 + J_3^2$  commutes with both  $H$  and  $J_i$ . Hence  $P := (H, J^2, J_3)$  are three integrals of motion in involution.
- Due to the rotational symmetry, it is useful to employ spherical coordinates  $r, \vartheta, \varphi$ ,

$$x_1 = r \sin \vartheta \cos \varphi, \quad x_2 = r \sin \vartheta \sin \varphi, \quad x_3 = r \cos \vartheta. \quad (1.2)$$

What are the momenta  $p := (p_r, p_\vartheta, p_\varphi)$  conjugate to the positions  $q := (r, \vartheta, \varphi)$ ?

- Rewrite the conserved quantities  $P$  in terms of spherical coordinates  $(q, p)$ . Invert these relations to express the momenta  $p$  in terms of  $P$  and  $q$ . Observe that the variables  $q$  are *separated*:  $p_r$  only depends on  $r$ ,  $p_\vartheta$  only depends on  $\vartheta$ , and  $p_\varphi$  only depends on  $\varphi$ .
- The generating function for a canonical transformation from  $(q, p)$  to  $(Q, P)$  is

$$S(q, P) = \int^{r'} p_r(q', P) dr' + \int^{\vartheta'} p_\vartheta(q', P) d\vartheta' + \int^{\varphi'} p_\varphi(q', P) d\varphi' \quad (1.3)$$

The positions conjugate to the momenta  $P_i$  are  $Q_i = \partial S / \partial P_i$ . Verify that  $\dot{Q}_{J^2} = 0$ ,  $\dot{Q}_{J_3} = 0$ , and  $\dot{Q}_H = 1$ . Use the last equation to derive the standard solution of the Kepler problem

$$t - t_0 = \int_{r_0}^r \frac{m dr'}{\sqrt{2m(H - V(r')) - J^2/r'^2}}. \quad (1.4)$$

*Hint:* Write the time derivative as  $d/dt = \dot{r} d/dr + \dot{\vartheta} d/d\vartheta + \dot{\varphi} d/d\varphi$ , and use Hamilton's equation of motion  $\dot{r} = \partial H / \partial p_r$  etc.

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## 1.2. Euler Top and Lax Pair (6 points)

Consider a rotating body attached to a fixed point without external forces. In the co-moving frame, where the coordinate axes are aligned with the principal moments of inertia  $I_i$  of the body, the Hamiltonian reads

$$H = \sum_{i=1}^3 \frac{J_i^2}{2I_i}, \quad (1.5)$$

where  $J_i = I_i \omega_i$  are the components of the angular momentum  $J = (J_1, J_2, J_3)$ , with  $\omega = (\omega_1, \omega_2, \omega_3)$  the rotation vector of the comoving frame.

a) Using the Poisson structure  $\{J_i, J_j\} = \varepsilon_{ijk} J_k$ , derive the equation of motion for  $J$ :

$$\frac{dJ_i}{dt} = \varepsilon_{ijk} \omega_j J_k. \quad (1.6)$$

*Hint:* Use the Hamilton equation of motion  $dF/dt = -\{H, F\}$ .

b) Show that the square of the angular momentum  $J^2 = \sum_i J_i^2$  is conserved. Since the system has only two degrees of freedom, and the Hamiltonian  $H$  itself is also conserved, the system is integrable. Express  $J_2$  and  $J_3$  in terms of  $H$ ,  $J^2$ , and  $J_1$ .

c) As a candidate Lax pair, consider the matrices  $L_{ij} = \varepsilon_{ijk} J_k$  and  $M_{ij} = -\varepsilon_{ijk} \omega_k$ . Show that the equation of motion (1.6) can be written as

$$\frac{dL}{dt} = [M, L]. \quad (1.7)$$

d) Calculate the first few conserved quantities  $\text{tr}(L^n)$ , and observe that the Hamiltonian is not among them. Why not?

e) Show that there exist rescaled variables  $\mathcal{J}_i := \alpha_i J_i$  such that the equations of motion take the form

$$\frac{d\mathcal{J}_i}{dt} = 2\mathcal{J}_j \mathcal{J}_k, \quad (1.8)$$

where  $(i, j, k)$  is any cyclic permutation of  $(1, 2, 3)$ .

f) Show that the Lax equation  $dL(z)/dt = [M(z), L(z)]$  for the following  $2 \times 2$  matrices  $L(z)$  and  $M(z)$  is also equivalent to the equations of motion:

$$\begin{aligned} L(z) &= (1 - z^2)\mathcal{J}_1\sigma_1 + (1 + z^2)\mathcal{J}_2 i\sigma_2 - 2z\mathcal{J}_3\sigma_3, \\ M(z) &= z\mathcal{J}_1\sigma_1 - z\mathcal{J}_2 i\sigma_2 + \mathcal{J}_3\sigma_3, \end{aligned} \quad (1.9)$$

where  $\sigma_i$  are the Pauli matrices.

g) Compute all the independent integrals of motion  $\text{tr}(L(z)^n)$ . Verify that they are indeed conserved, and express them in terms of  $H$  and  $J^2$ . This shows that  $L(z)$ ,  $M(z)$  form a sufficient Lax pair for the system, with spectral parameter  $z$ .

*Hint:* Make use of the Pauli matrix algebra  $\sigma_k \sigma_\ell = i\varepsilon_{klm} \sigma_m + \delta_{k\ell}$ .