

The case for self-seeded beams at MID



MID overview



X-ray scattering and imaging: SAXS, WAXS, XPCS, phase contrast imaging and holography, CXDI, nano focusing, fs laser pump - X-ray probe

European XFEL

Anders Madsen, European XFEL

MHz area detector, 10^6 pix of 200 µm size (AGIPD) ePix, Gotthard detector, CCD cameras,... Versatile setup, multi-purpose interaction chamber Windowless (in-vacuum setup) or sample in air Sample - detector dist: 0.2 m (LFOV) to 8 m (HiRes) 20 up to ~50°, 5 - 24 keV (7-18 keV used so far)



A. Madsen *et al.*, JSR (2021) **28**, 637 <u>https://scripts.iucr.org/cgi-bin/paper?S1600577521001302</u>



Beam coherence

Example: Young's double slit experiment (Thomas Young, 1801) [wave-character of quantum mechanical particles (photons)]



 $P=|\Sigma_{j}\Phi_{j}|^{2}$ $\Phi: \text{ probability amplitude}$ $\Phi_{j} \sim \exp[-i(\omega t-kl_{j})]$ $\omega=ck, \ k=2\pi/\lambda, \ l_{j}(L,y)$

 $\begin{array}{l} P(y) \sim \cos^2(\pi y d/\lambda L) \\ \Delta y = \lambda L/d \end{array}$



Thomas Young (1773-1829)

Spatial coherence Transverse coherence length

Beam coherence

Albert Michelson, 1852-1931



scienceworld.wolfram.com

Longitudinal coherence length

$$\sim \lambda^2/2\Delta\lambda$$

Michelson-Morley experiment (1887)

The experiment failed but the Michelson interferometer is useful to demonstrate the longitudinal (temporal) coherence of light



Courtesy Uni. Freiburg



Edward Morley, 1838-1923



Coherence length



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Example of sunlight:

$$\begin{split} \lambda &= 0.5 \ \mu m \\ L &= 1.5 \ 10^8 \ km \\ d &= 1.4 \ 10^6 \ km \end{split}$$

$l_{\rm h,v} = \lambda L/2d \approx 25 \ \mu m$



http://micro.magnet.fsu.edu/primer/java/i nterference/doubleslit/index.html

X-ray Coherence

Towards full spatial X-ray coherence

Novel MBA storage rings: beam size ~ transverse coherence length (h and v), up to several keV XFELs: close to one transverse coherence mode

Longitudinal coherence (temporal coherence) is more difficult

All accelerator based X-ray sources need monochromators for wide-angle coherent scattering Small beam size (d) and small sample thickness (w) also required in exp

$\lambda/(2\Delta\lambda/\lambda)$

Si reflection	Δλ/λ	long coh (12.4 keV)
(111)	1.4e-4 (1.7 eV)	0.4 um
(220)	6.1e-5 (0.8 eV)	0.8 um
(333)	1.1e-5 (0.1 eV)	4.5 um



Dispersive diamond spectrometer for single pulse spectra

DES: Diagnostics End-Station

Spectral analysis, pulse duration, intensity, beam pointing

X-ray beam



Dispersive spectrometer: Gotthard 1d detector, 50 µm pix 500 kHz, 120 images/train

U. Boesenberg *et al*. Optics Express **25**, 2852 (2017)

Dispersive diamond spectrometer for single pulse spectra



bent diamond crystal R~ 10cm, C*(220)

HPHT type IIa, 20 µm thick made by TISNCM, Russia (S. Terentev et al.)

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Data from MID: 9.1 keV SASE, 30 pulses/train at 500 kHz, Gotthard detector at 1 m, resolution ~ 0.4 eV/pix.

Anders Madsen, European XFEL SASE beam + monochromator Si(111) Δλ/λ ~ 1.4e-4 Si(220) Δλ/λ ~ 6.1e-5 Si(511) Δλ/λ ~ 1.1e-5 spike $\Delta\lambda/\lambda \sim 5e-6$ envelope $\Delta\lambda/\lambda \sim 1e-3$ 0.13814 0.13816 0.13818 0.13803 0.13816 0.13828 Wavelength [nm]

About 15-20 spikes will pass through a Si(111) mono

Coherence time (and hence longitudinal coherence length) not well defined for Si(111) and Si(220)

SASE beam + monochromator



MHz SASE beam + monochromator at MID



2.25 MHz, 0.5 x 0.5 mm (FWHM) beam size, Si(111) reflection, crystal at 100 K

X. Dong et al., AIP Conf. Proc. 1741, 040027 (2016)

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I. Petrov et al. (unpublished)



The case for self-seeded beams at MID

- 1e-4 1e-5 BW is required for many coherent scattering experiments
- Monochromators reduce intensity and increase intensity fluctuations
- Temporal coherence of reflected beam varies from pulse to pulse
- Cryocooled Si monochromators cannot transmit more than ~ 50 mJ during MHz train (adiabatic heating)
- HXRSS increases spectral brightness (Brilliance) \rightarrow more coherent flux than SASE!
- HXRSS intensity fluctuates less and is more stable over train than beam from mono
- Need to better understand spectral drifts and origin (electrons, seeding xtals, spectrometer,...)

Thanks to the HXRSS commissioning team and everyone involved in this project !