



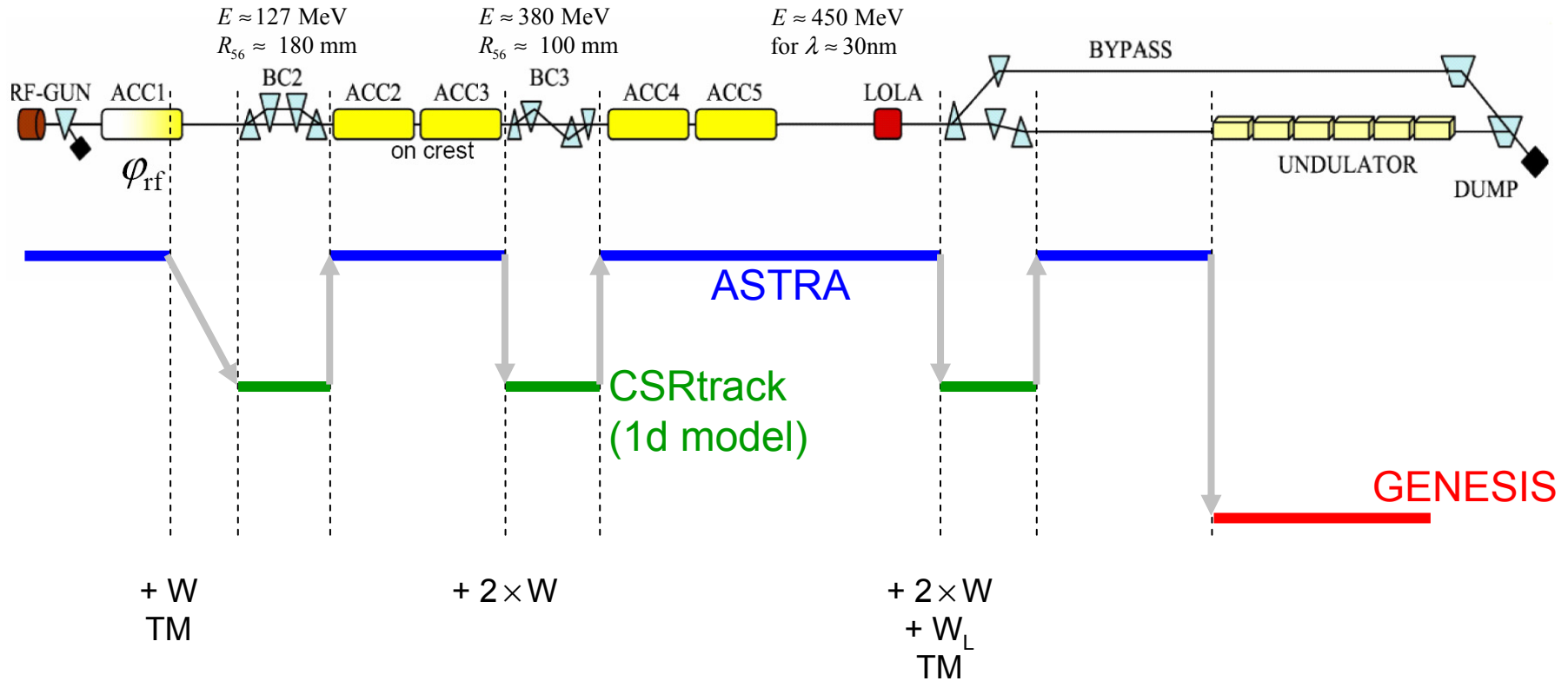
Statistical properties of radiation in TTF2 (Genesis simulations)

Igor Zagorodnov

BDGM, DESY

5.12.05

TTF2 – s2e

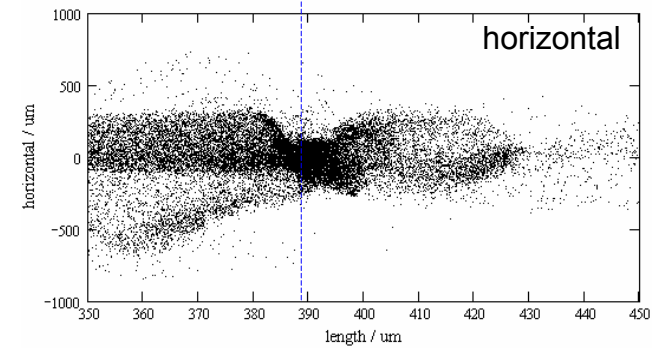
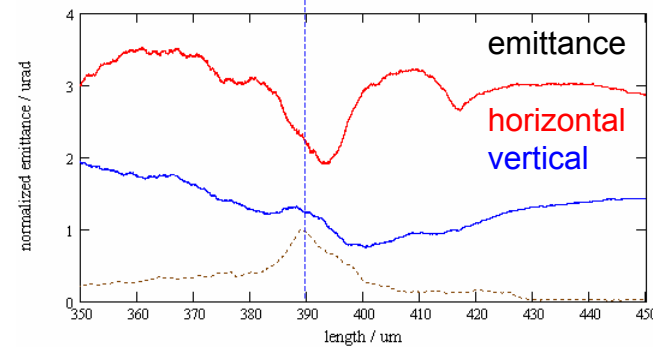
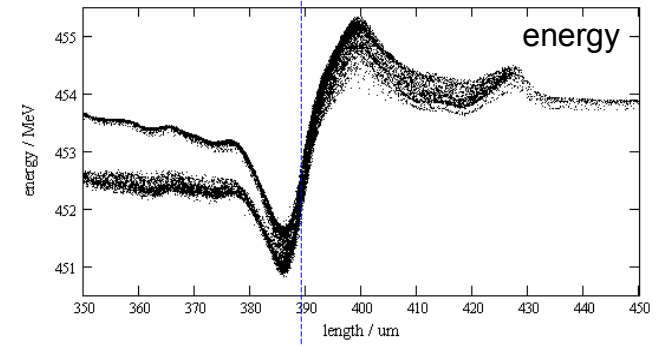
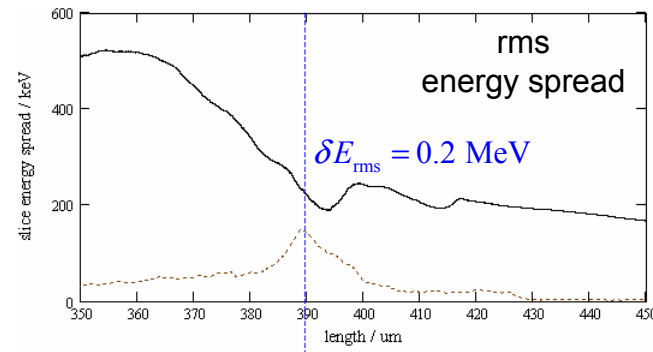
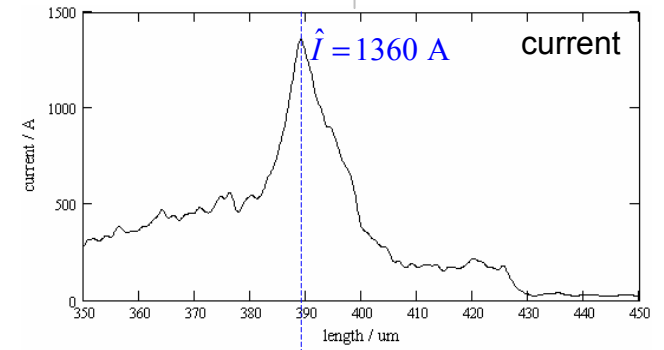
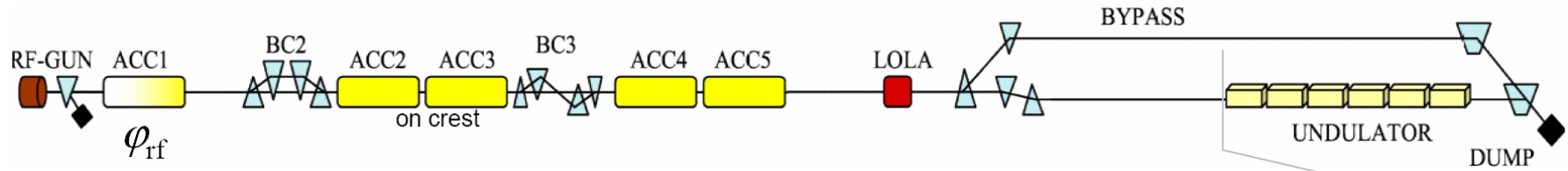


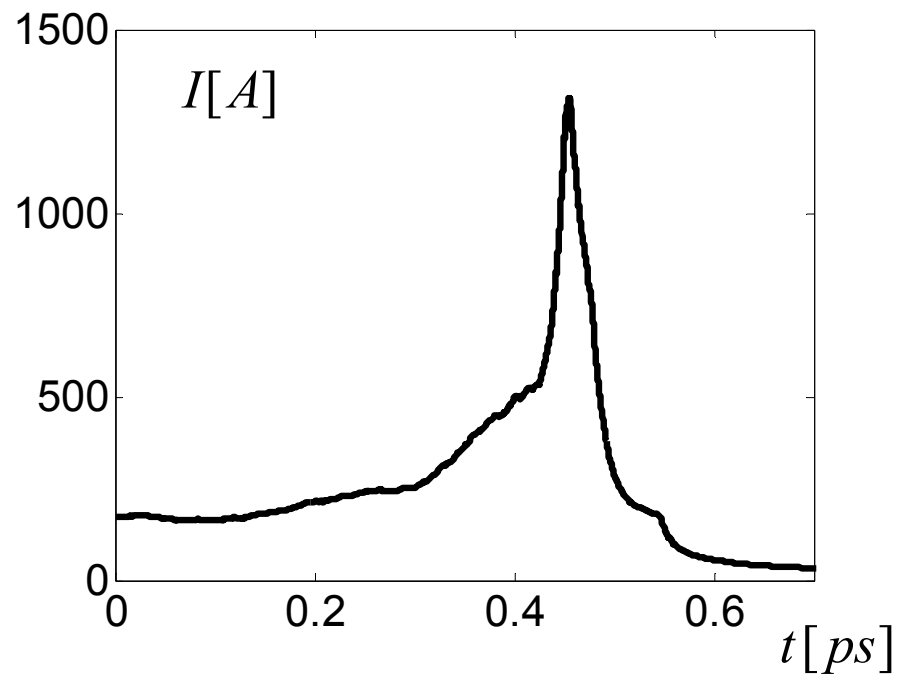
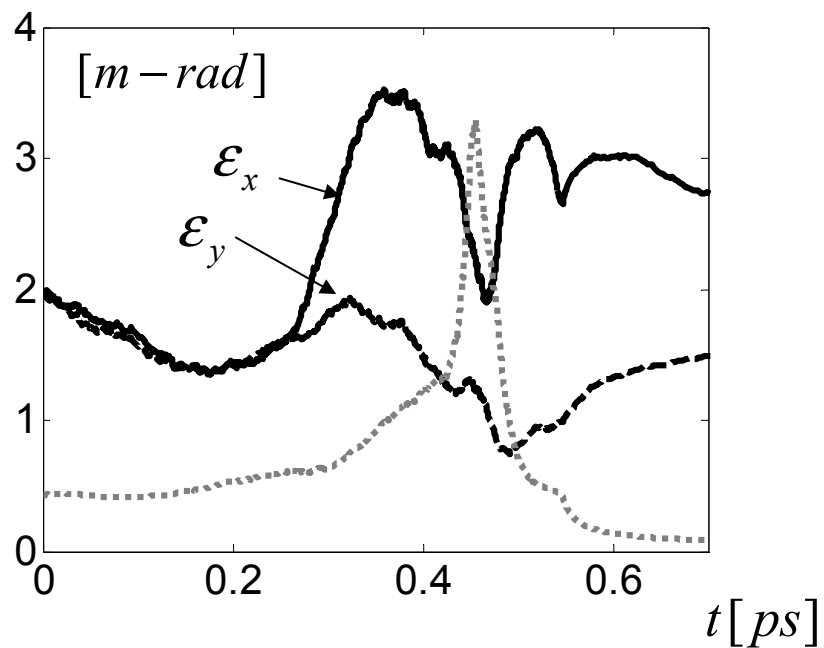
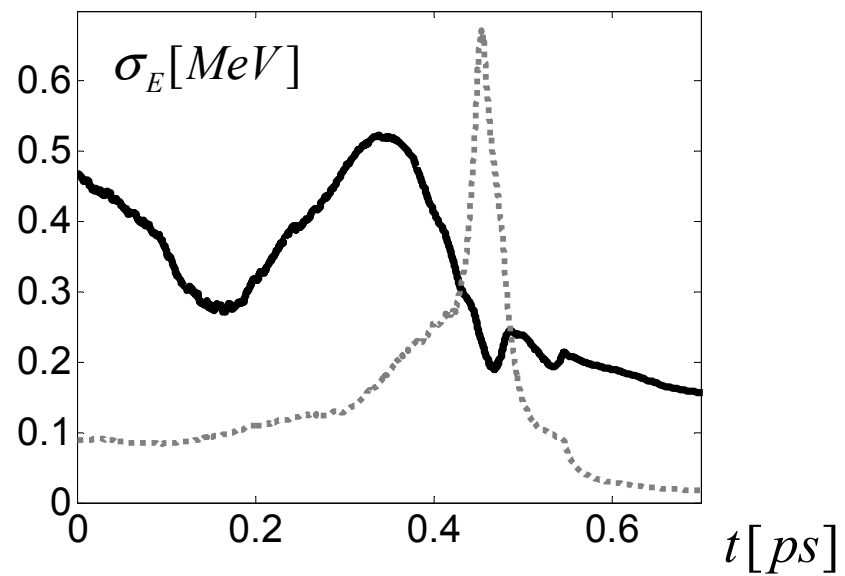
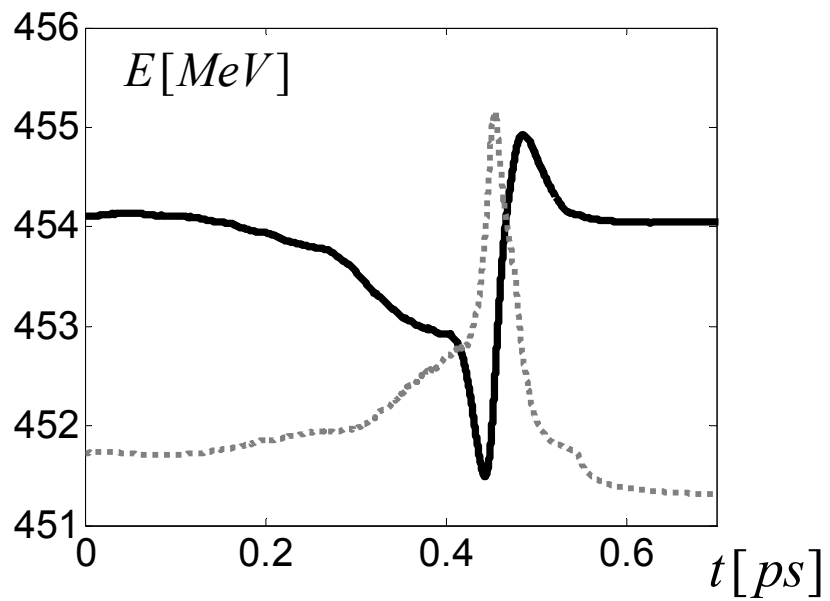
W = wake of one TTF module

W_L = wake of LOLA structure

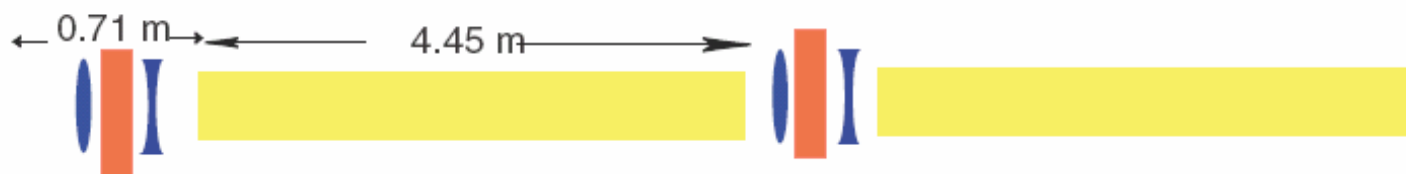
TM = transverse matching to design optic

M.Dohlus





TTF2 Undulator layout



 Undulator segment 4.45 m/163.5 periods, undulator period: 27.3 mm

 Focusing  Defocusing quadrupole 37 T/m with optical length 82 mm

 Diagnostic block 327 mm

Total effective (magnetic) undulator length: 26.78 m (6 segments)

Total geometrical length: 31.16 m

Doublet period: 5.17 m

Separation between quadrupole centroids in doublet: 385 mm

parameters

name	symbol	unit	value
energy	γ		886
energy spread	σ_γ		0.47
emittance	ϵ_x	π mm-mrad	2.4
	ϵ_y		1.14
bunch charge	Q	nC	1
peak current	I_P	kA	1.3
undulator period	λ_u	cm	2.73
undulator parameter	a_u		0.896
quadrupole length	L_Q	cm	12.716
quadrupole gradient	G_F	T/m	10.6
	G_D		-9.1
section length	L_{sect}	m	$4.45+0.75=5.2$
total length	L_{total}	m	31.2
beta function (waist)	$\beta_x,$	m	3.4
	$\beta_y,$		5.63

Parameters of the FEL theory

$$\lambda_s = \frac{\lambda_u}{2\gamma^2} (1 + a_u^2) = 31.348 \text{e [nm]}$$

$$z_R = \frac{\pi w_0^2}{\lambda_s} = \frac{2\pi\sigma_r^2}{\lambda_s} = 1.64 \text{ [m]}$$

Parameters of the FEL theory

Gain parameter

$$\Gamma_3 = \left[\left(\frac{A_{JJ} \omega_s \theta_l}{c \gamma_l} \right)^2 \frac{I_P}{2 \gamma I_A} \right]^{1/2} = 2.5 \quad \Gamma_1 = \Gamma_3 B^{-1/3} = 1.56$$

Efficiency parameter

$$\rho_3 = \frac{c \gamma_l^2 \Gamma_3}{\omega} = 5.4 \text{e-}3 \quad \rho_1 = \frac{c \gamma_l^2 \Gamma_1}{\omega} = 3.4 \text{e-}3$$

Diffraction parameter

$$B = \Gamma_3 \sigma_r^2 \frac{\omega_s}{c} = 4.1$$

Effective power of the input signal

$$P_{sh} = 3 \rho_1 \frac{W_b}{N_c \sqrt{\pi \ln N_c}} = 20.3 [W]$$

Gain length*

$$L_g = L_{g0}(1+\delta) \approx 0.767 \text{ [m]}$$

Optimal beta-function

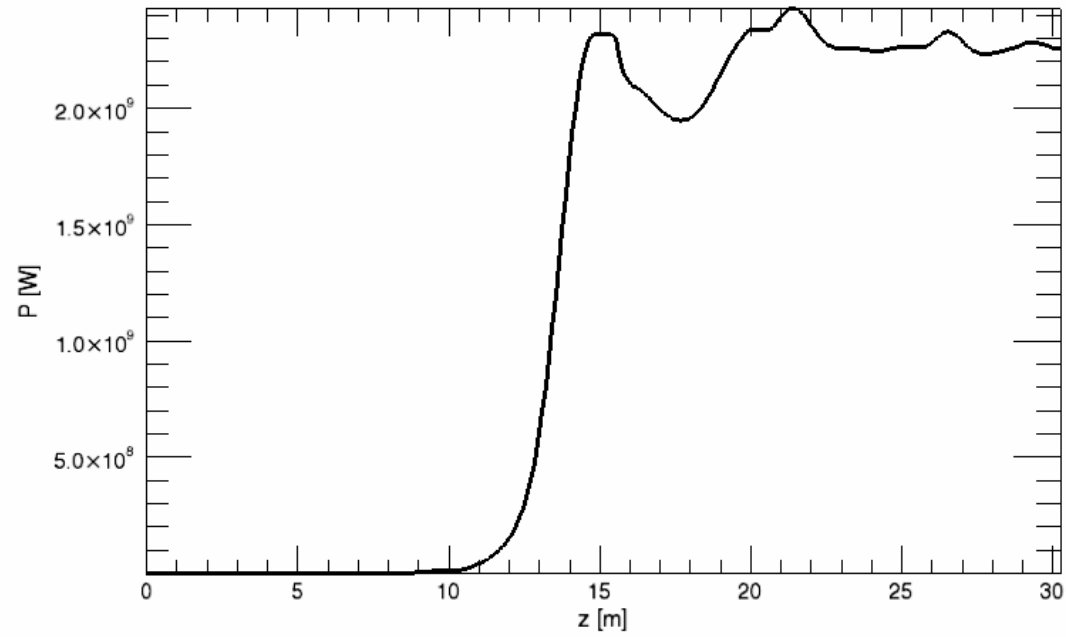
$$\beta_{opt} \approx 0.6 \text{ [m]}$$

Saturation length

$$L_{sat} \approx 10 \div 20 L_{g0} = 8 \div 15 \text{ [m]}$$

* E.L.Saldin et al./Optics Communications 235 (2004) 415-420

Genesis steady state simulation



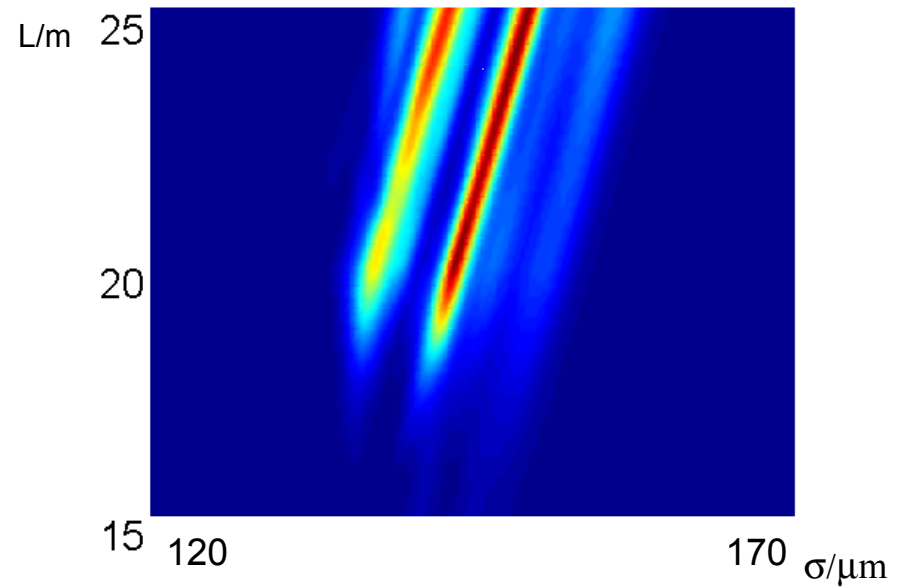
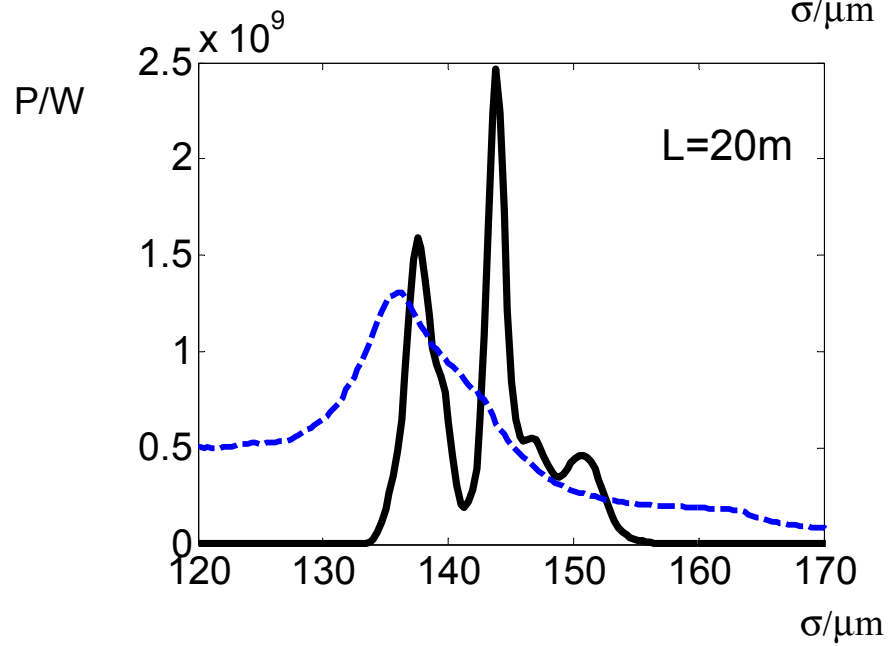
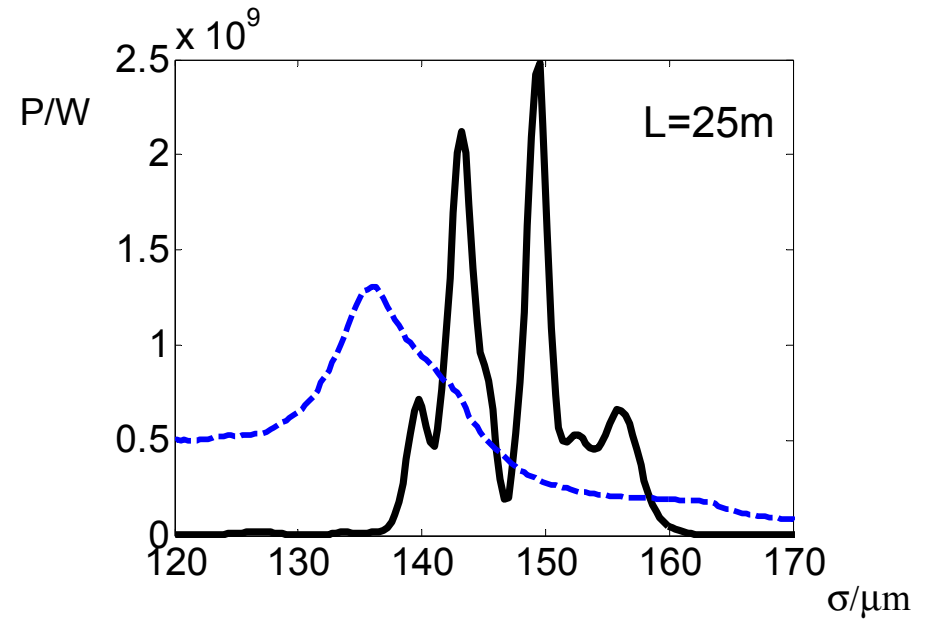
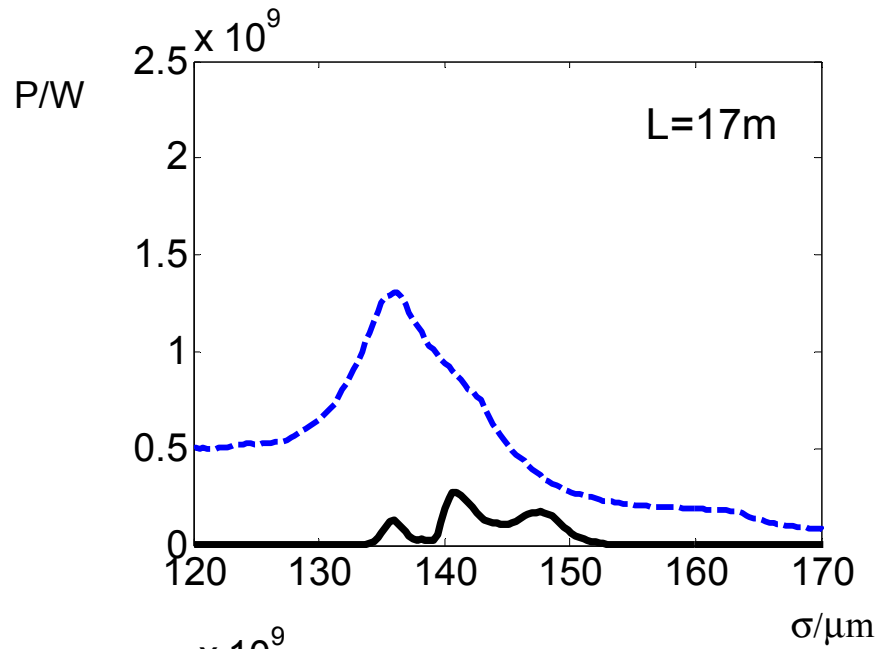
$$\lambda_s^{num} = 31.54 \text{ [nm]}$$

$$L_{sat} = 15 \text{ [m]}$$

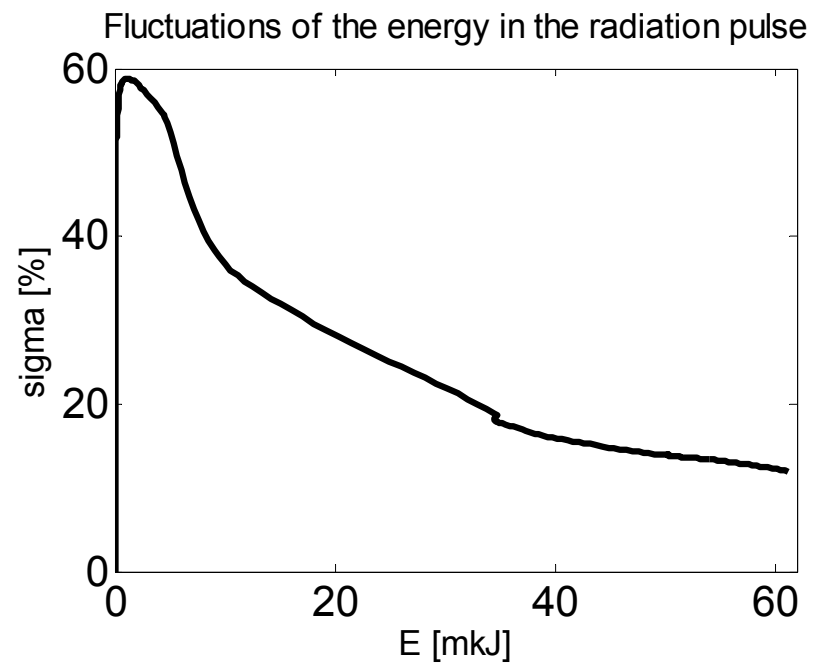
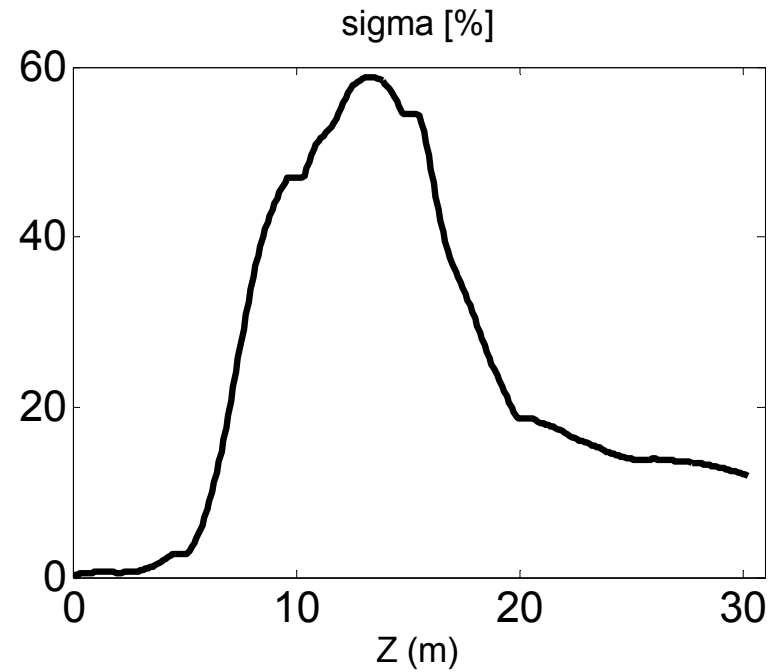
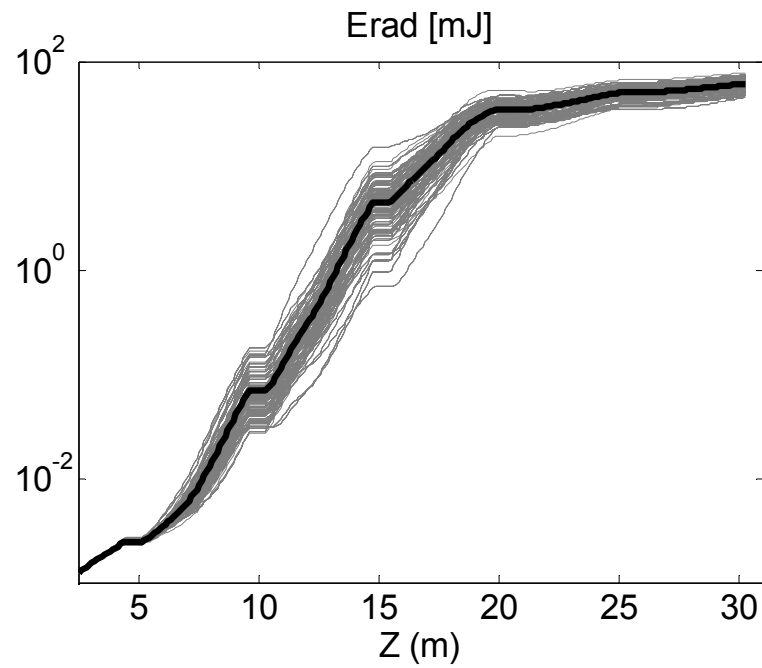
$$L_{slippage} = 37 \mu m$$

$$L_{cooperation} ; 2L_g \lambda_s / \lambda_u = 1.76 \mu m$$

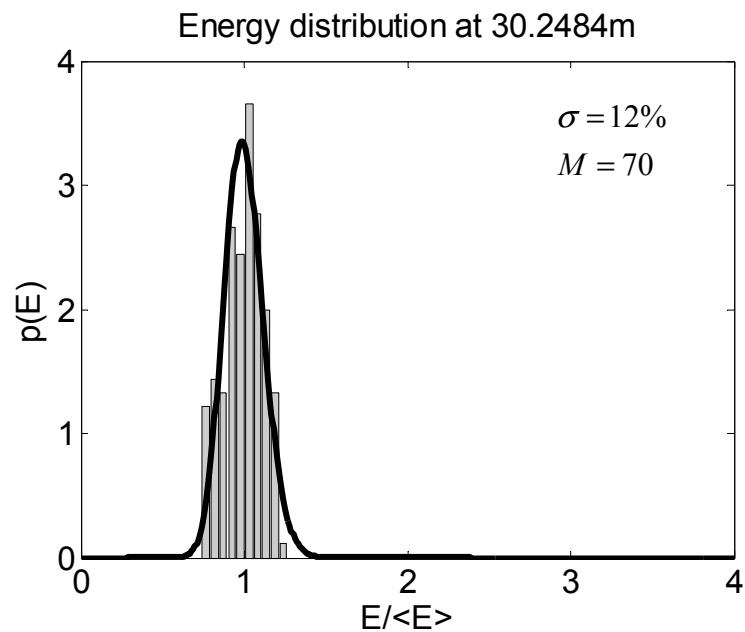
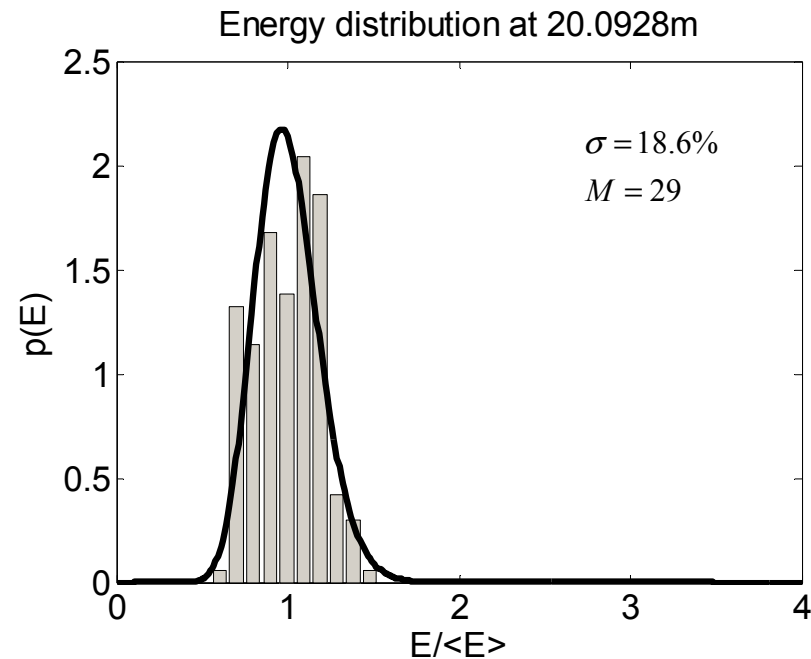
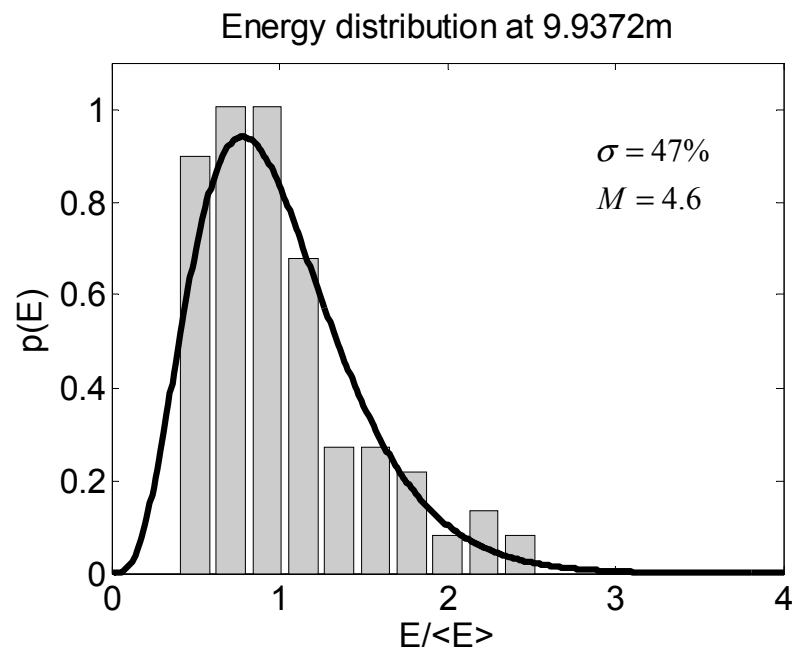
Genesis (SASE)



Genesis (SASE, N=170)

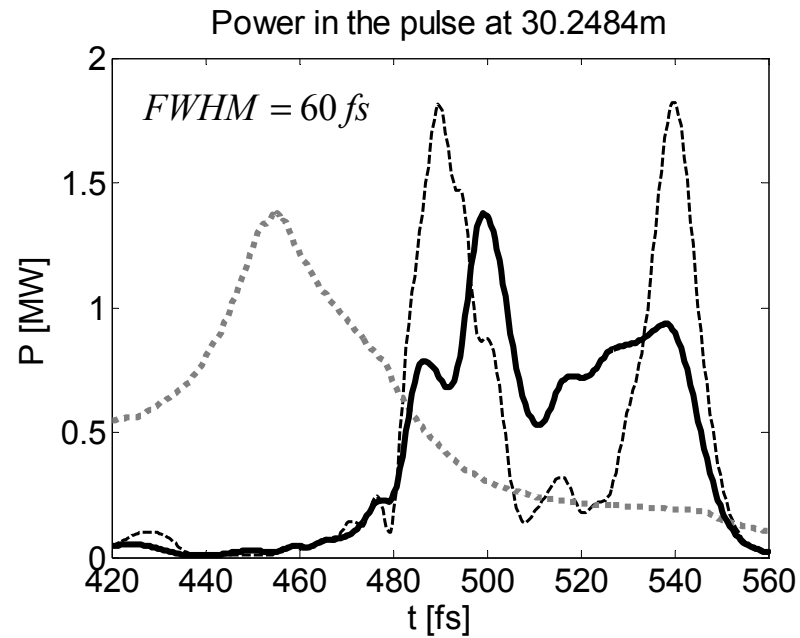
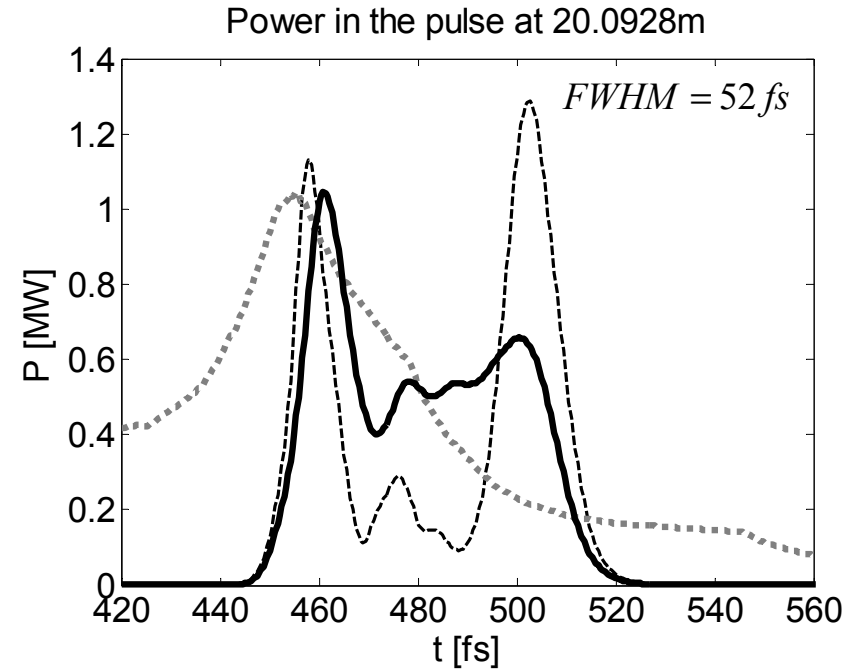
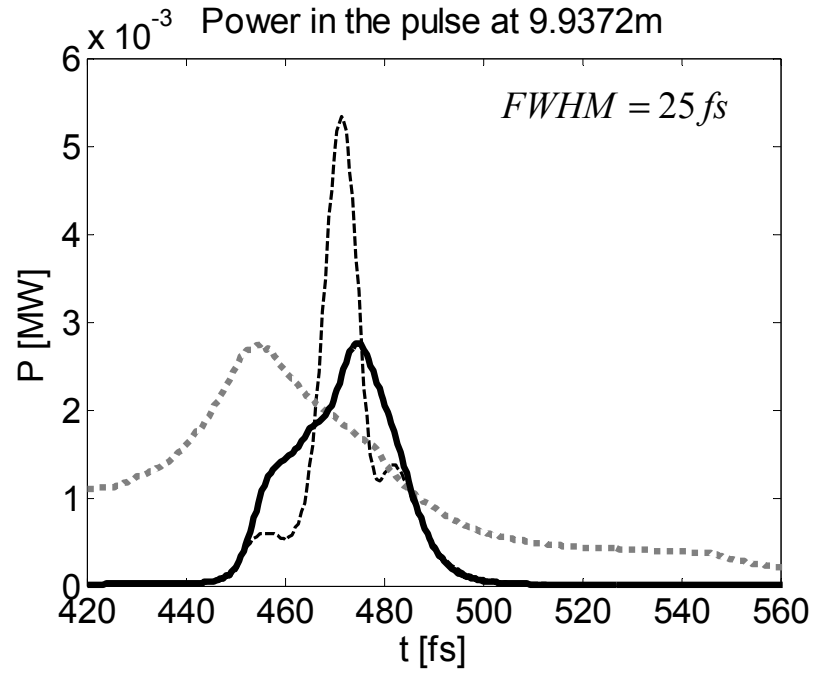


Probability distribution of energy in the radiation pulse (Genesis, N=170)

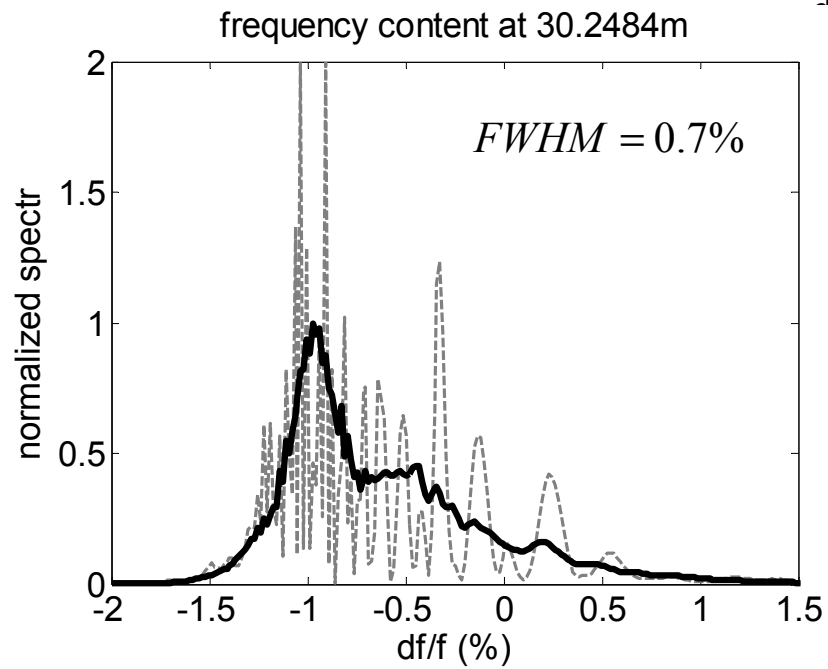
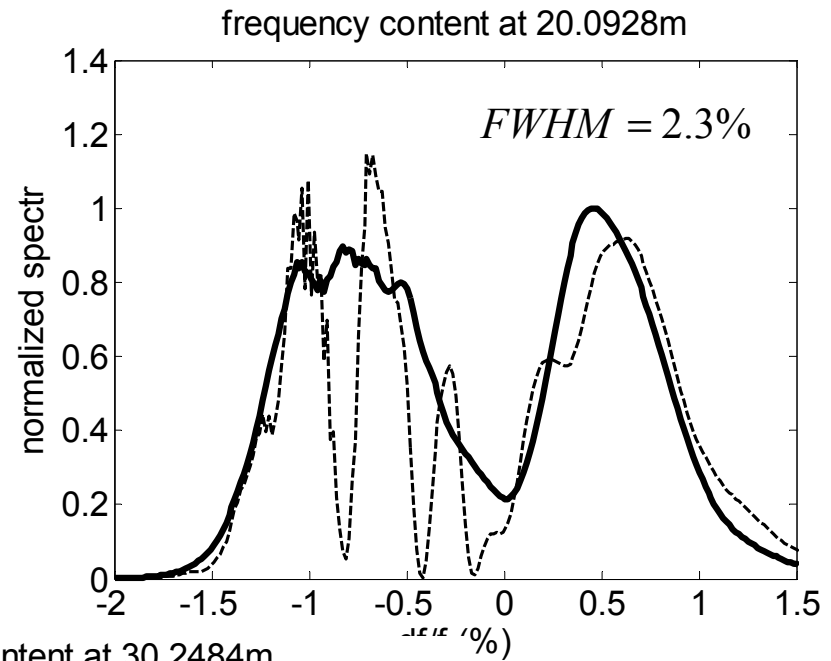
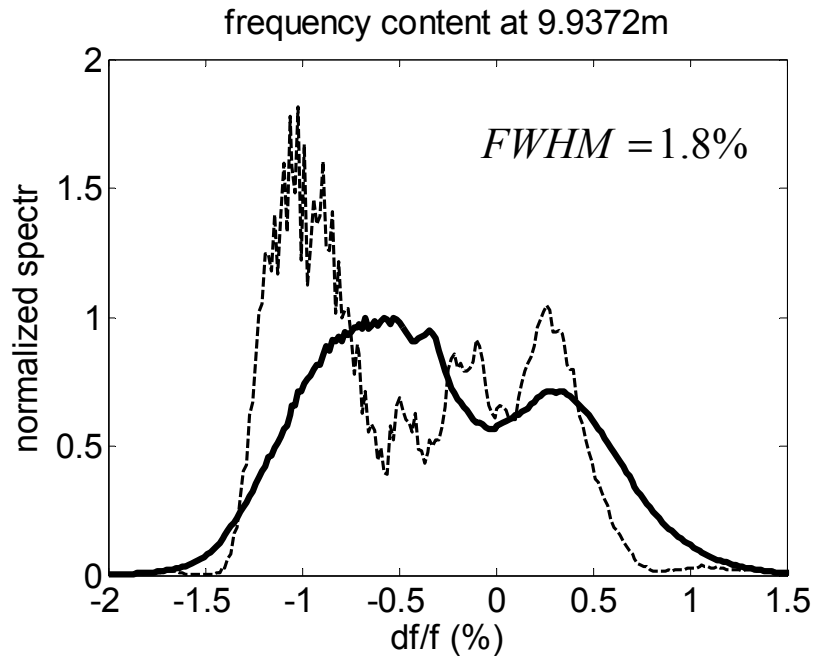


$$p(E) = \frac{M^M}{\Gamma(M)} \left(\frac{E}{\langle E \rangle} \right)^{M-1} \frac{1}{\langle E \rangle} \exp \left(-M \frac{E}{\langle E \rangle} \right)$$

Radiation power in the pulse and pulse duration (Genesis, N=170)



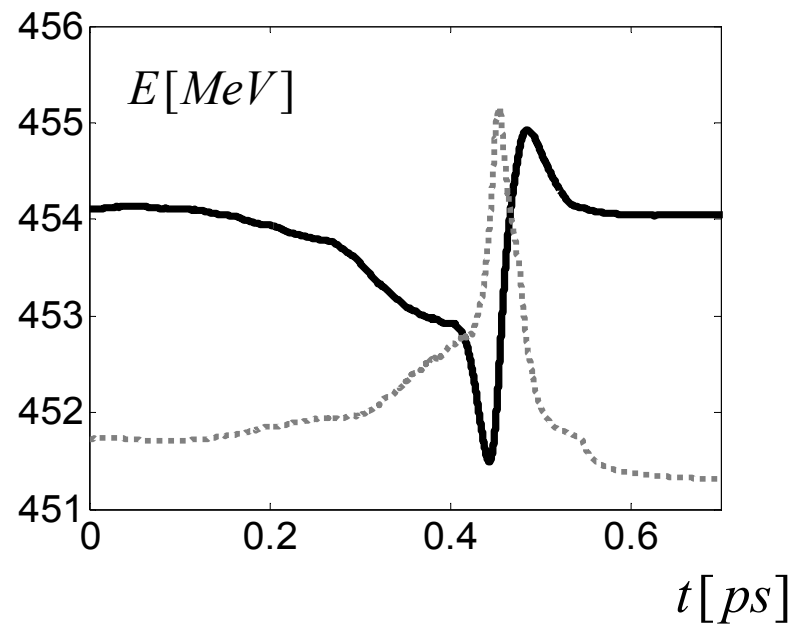
Spectrum of the radiation (Genesis, N=170)



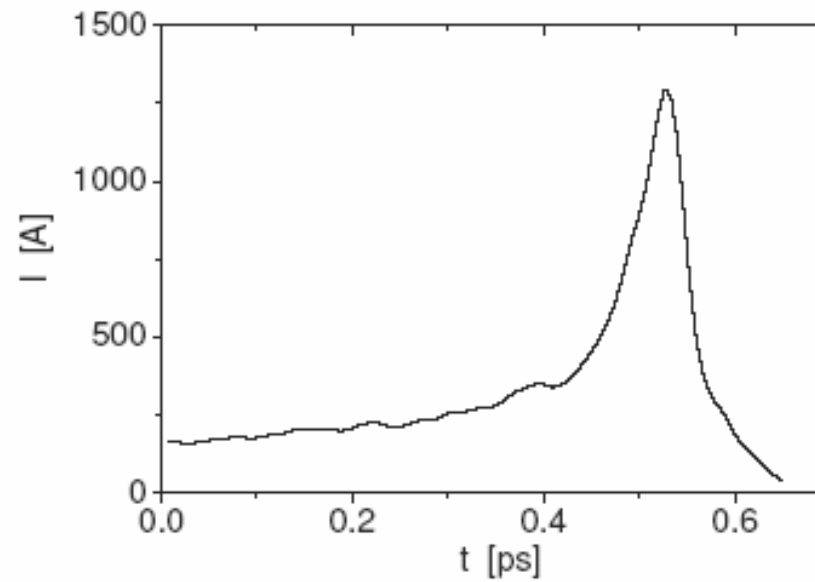
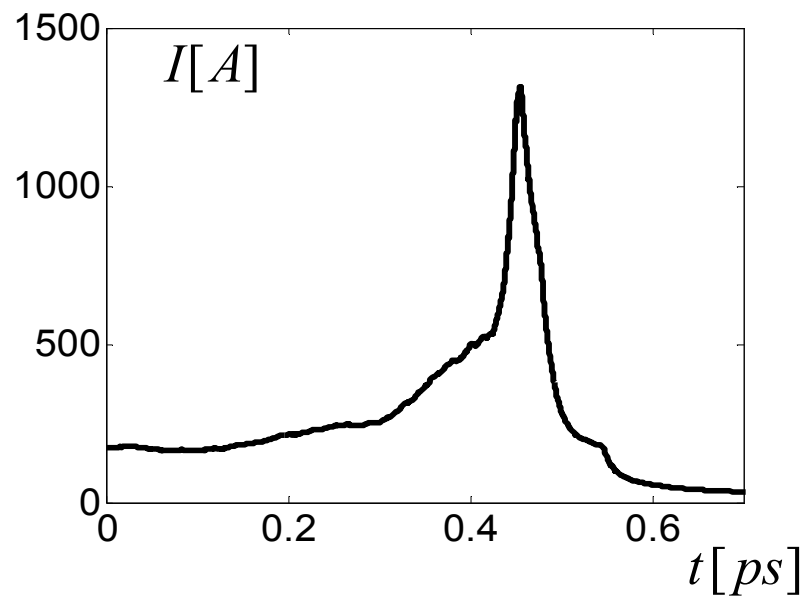
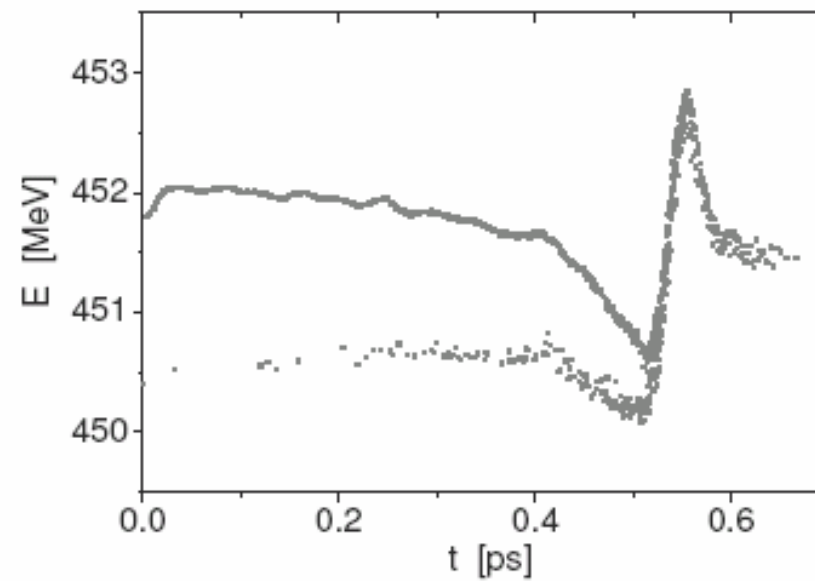
	SSY*	Genesis
Energy, MeV	450	454
Bunch charge, nC	0.5-1	1
Current, kA	1.3-2.2	1.3
Emittance, mm-mrad	1.5-3.5	2.4/1.4
Saturation length	18-22	~25
Energy in the rad. pulse	50-150	~60
Radiation pulse duration, fs	15-50	60
Radiation peak power, GW	2-4	3.8
Spectrum width, %	0.8	0.7

SSY*- E.L.Saldin et al, TESLA FEL 2004-06

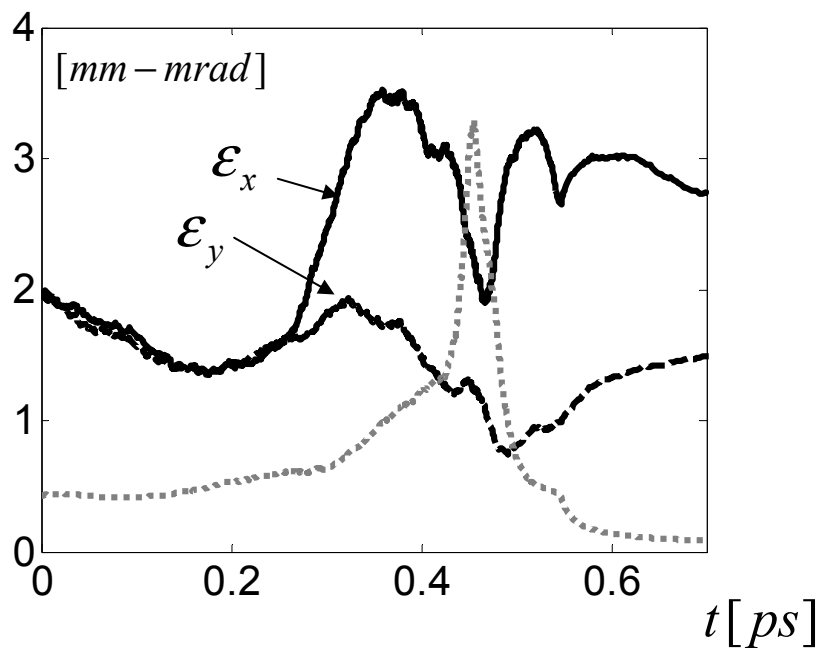
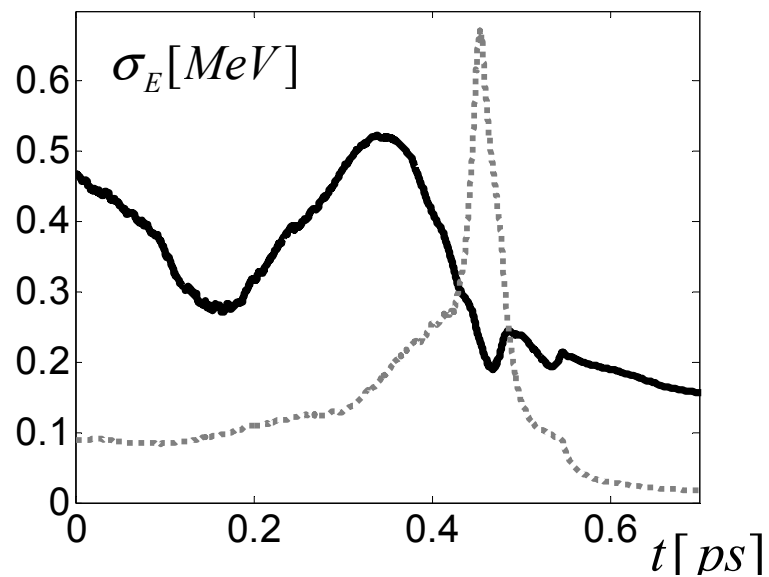
Genesis
30 m with intersections



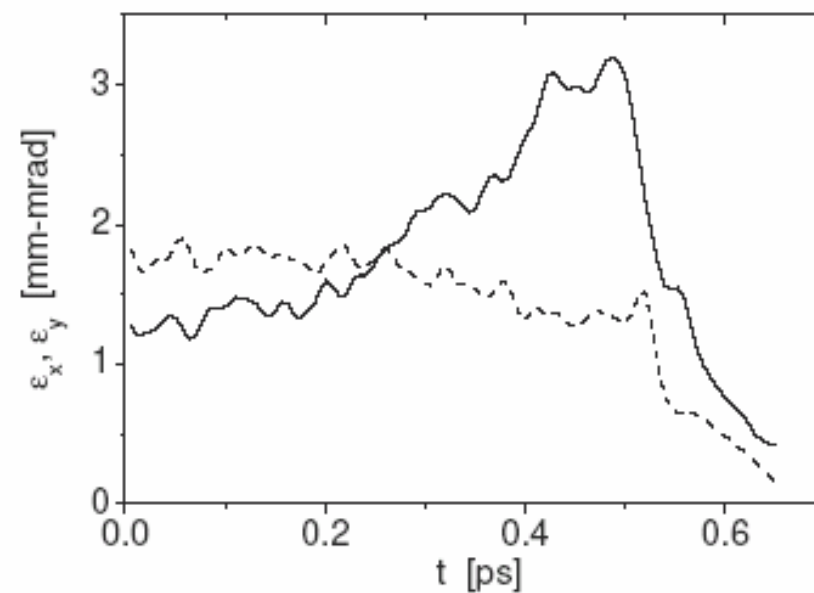
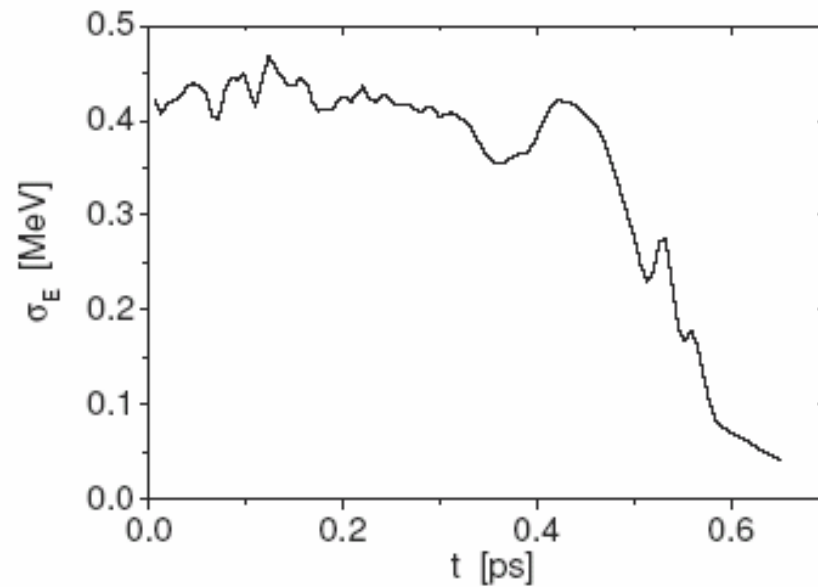
FAST (E.L.Saldin et al., DESY 05-239)
21 m without intersection



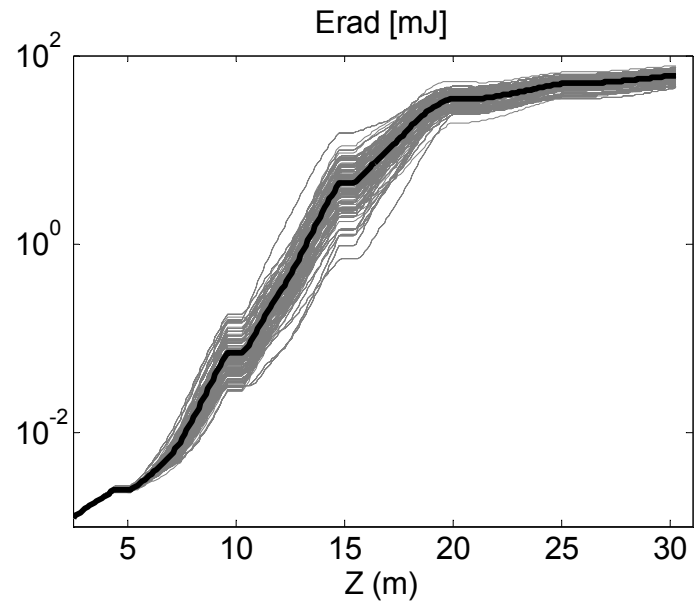
Genesis



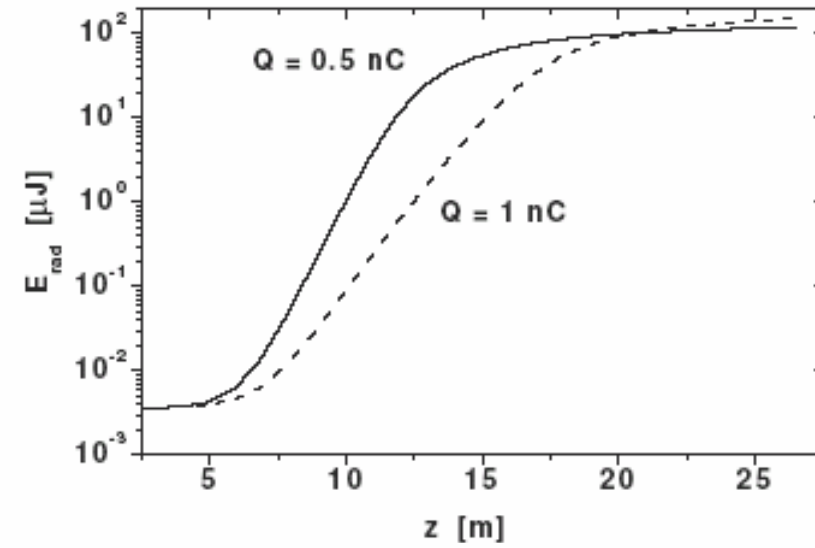
FAST (E.L.Saldin et al., DESY 05-239)



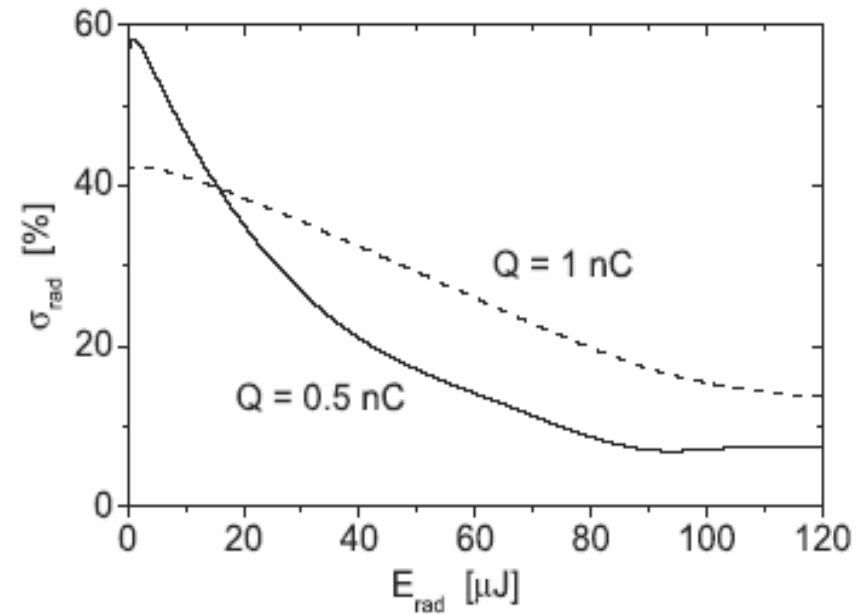
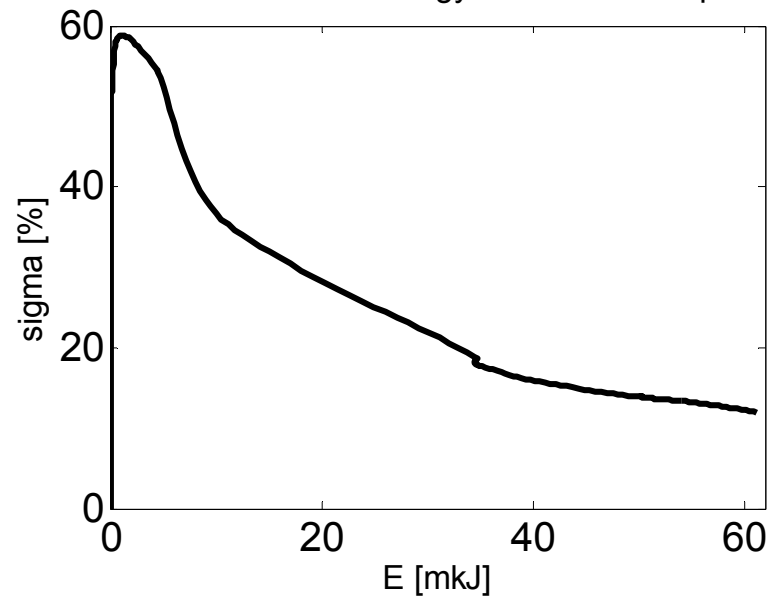
Genesis



FAST (E.L.Saldin et al., DESY 05-239)

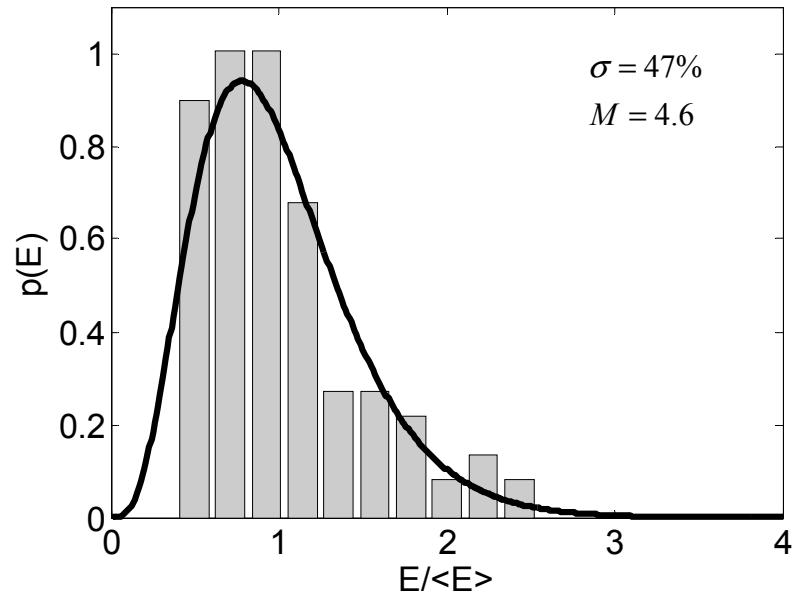


Fluctuations of the energy in the radiation pulse

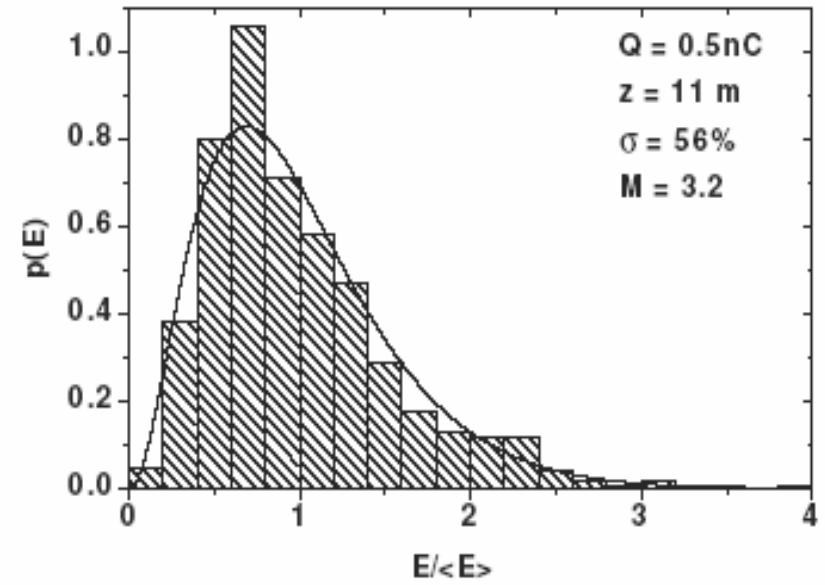


Genesis

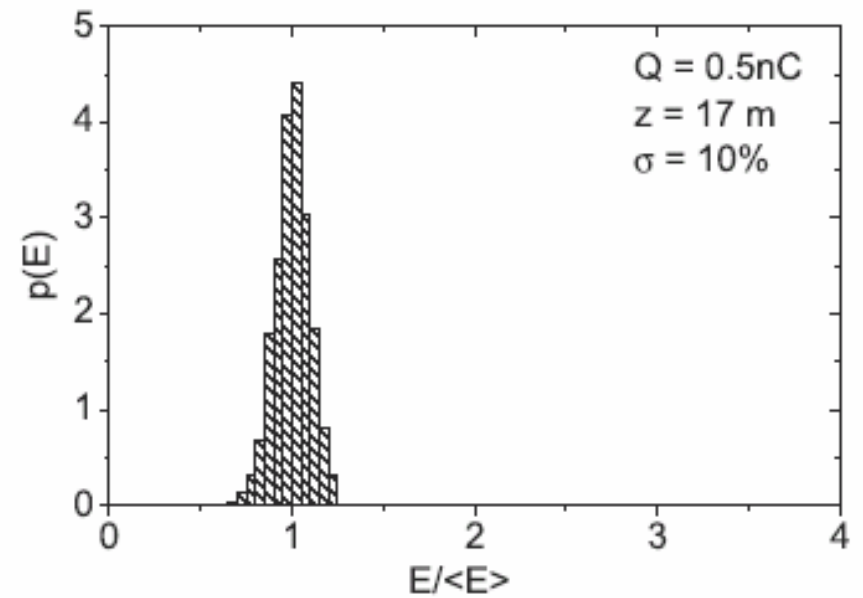
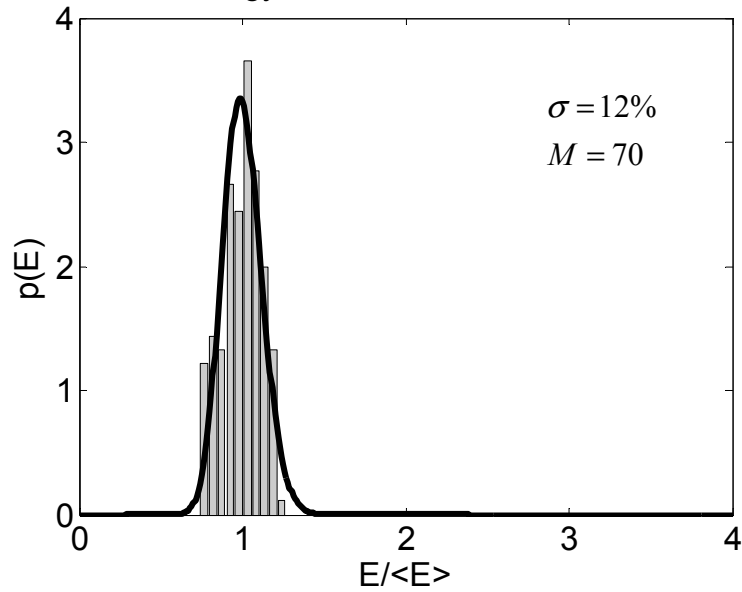
Energy distribution at 9.9372m



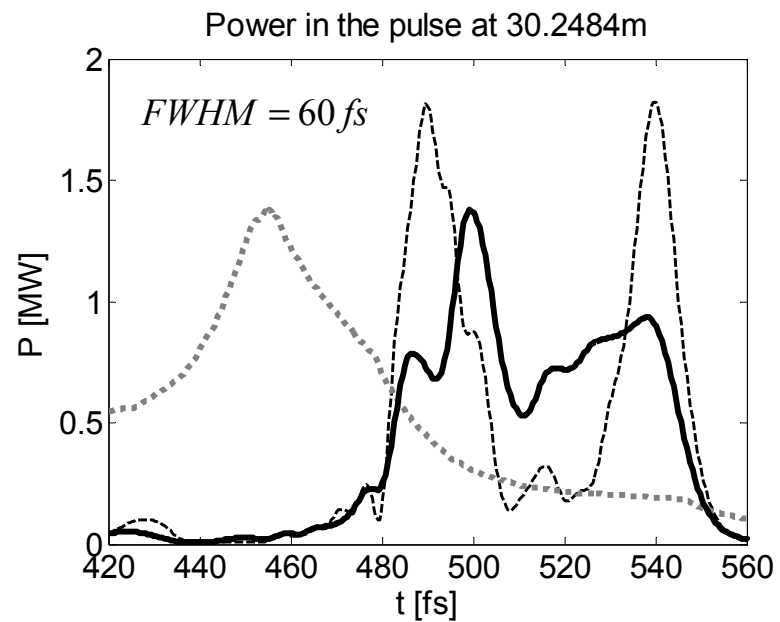
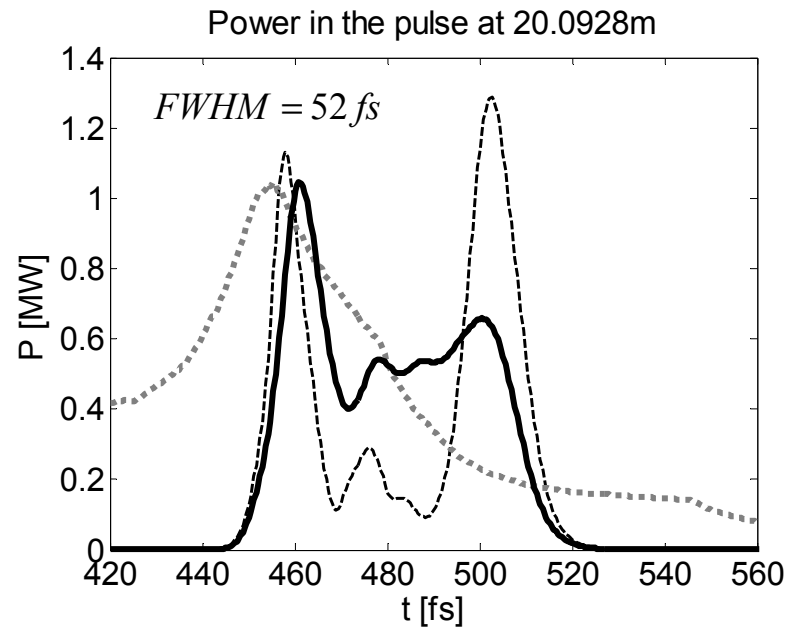
FAST (E.L.Saldin et al., DESY 05-239)



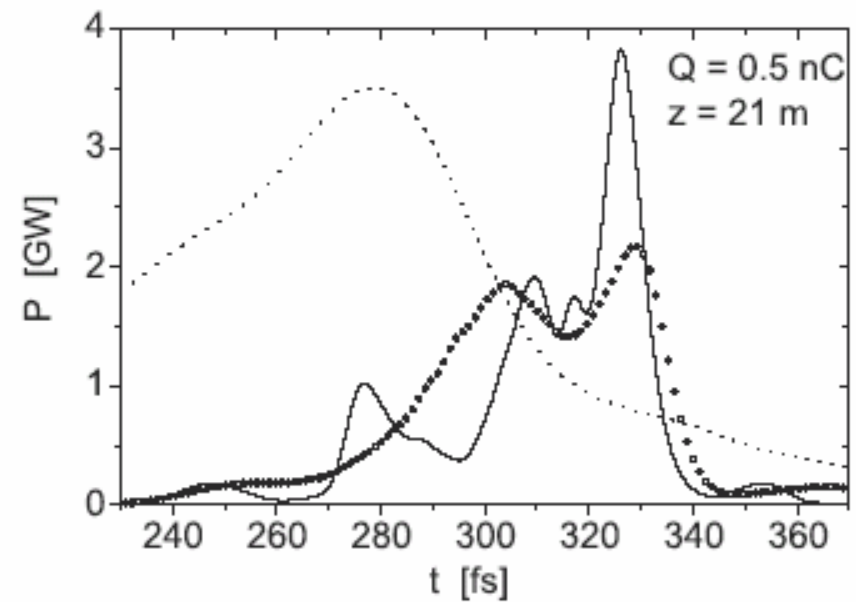
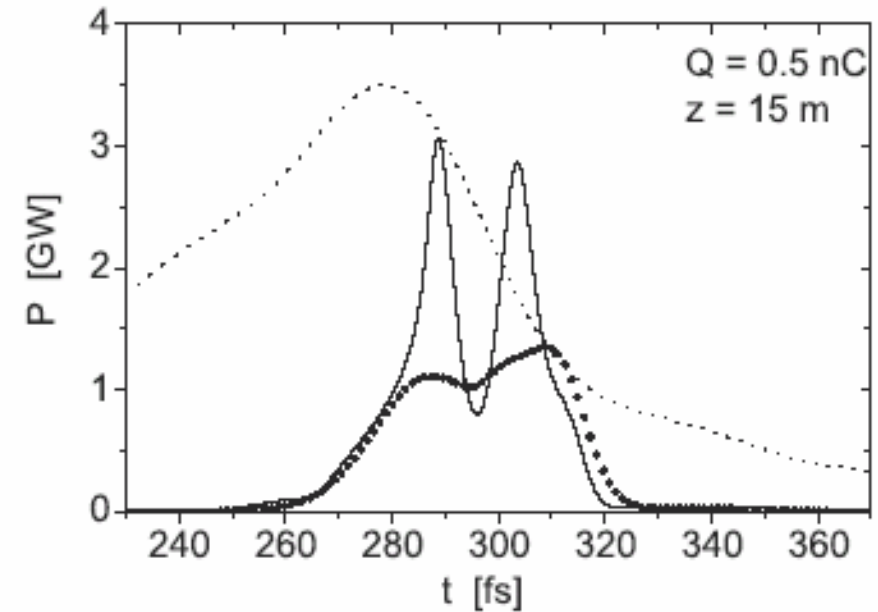
Energy distribution at 30.2484m



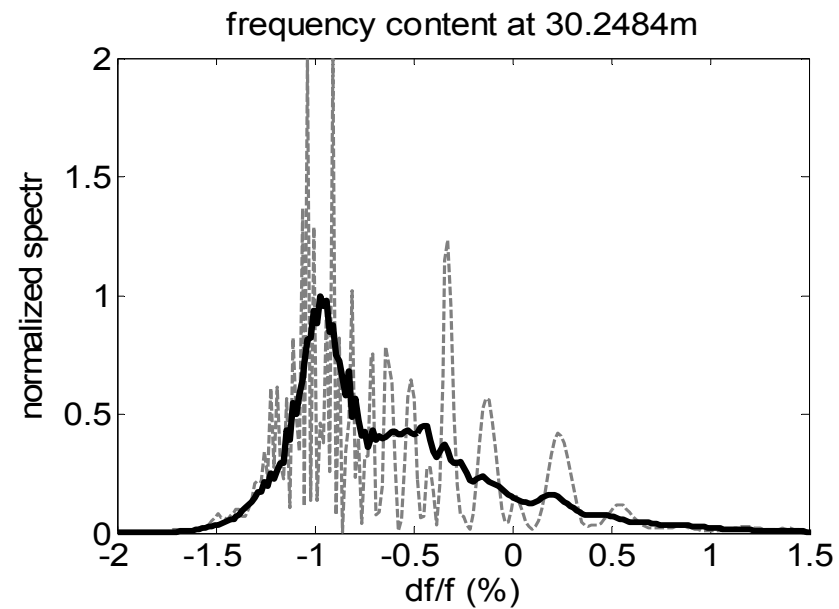
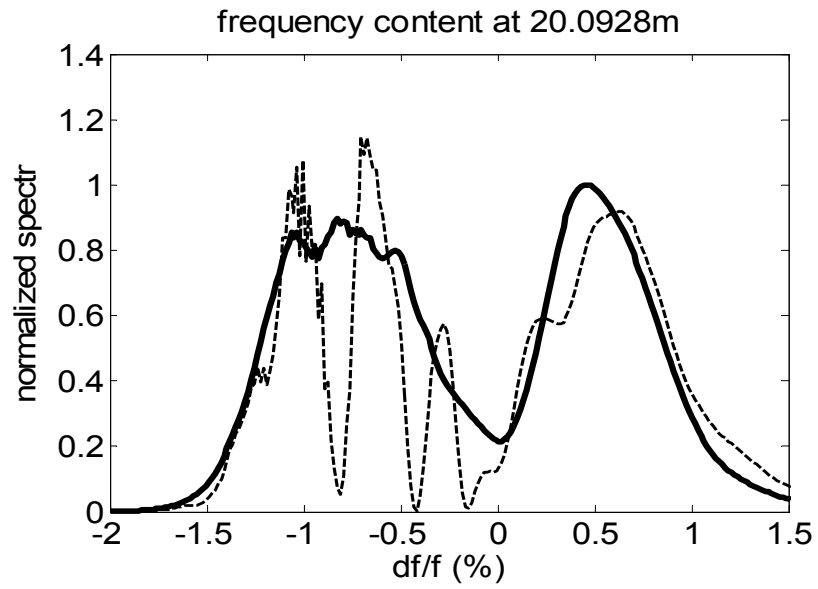
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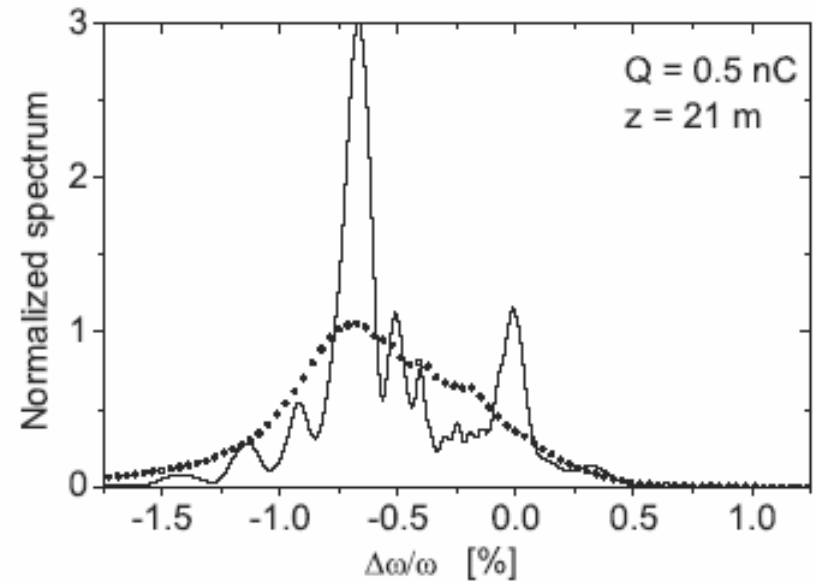
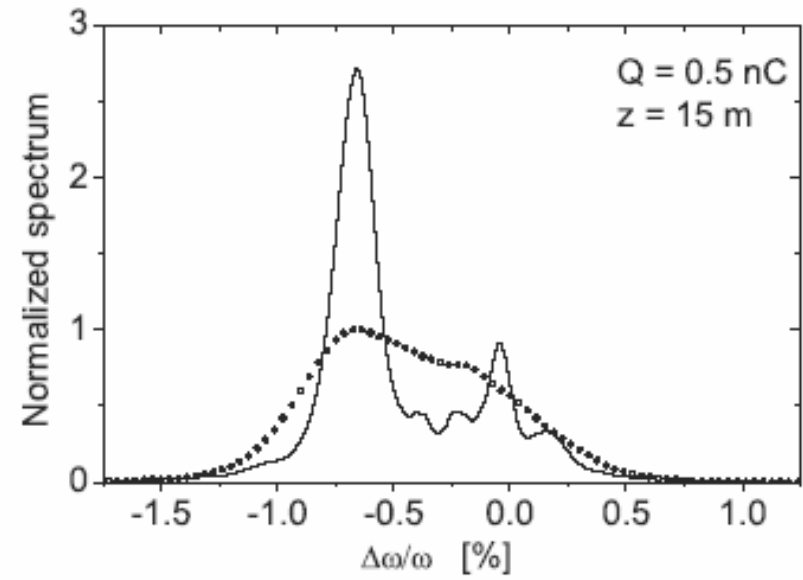
FAST (E.L.Saldin et al., DESY 05-239)



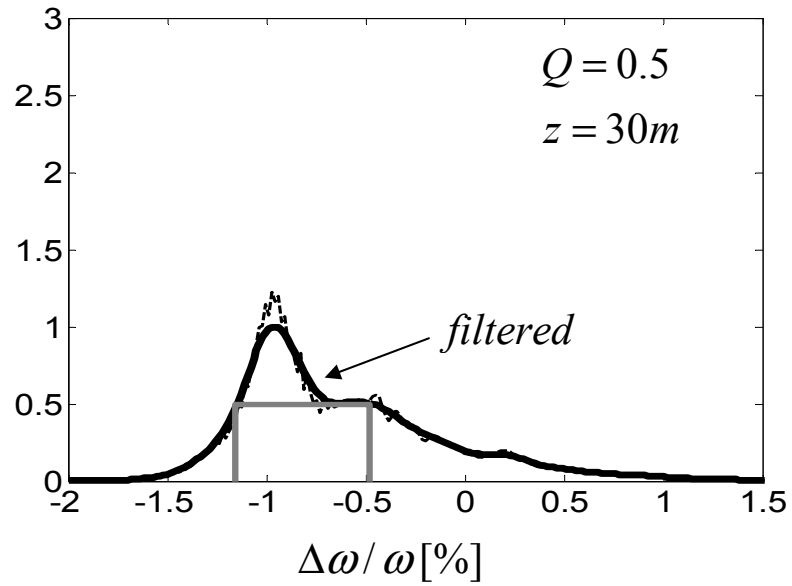
Genesis



FAST (E.L.Saldin et al., DESY 05-239)

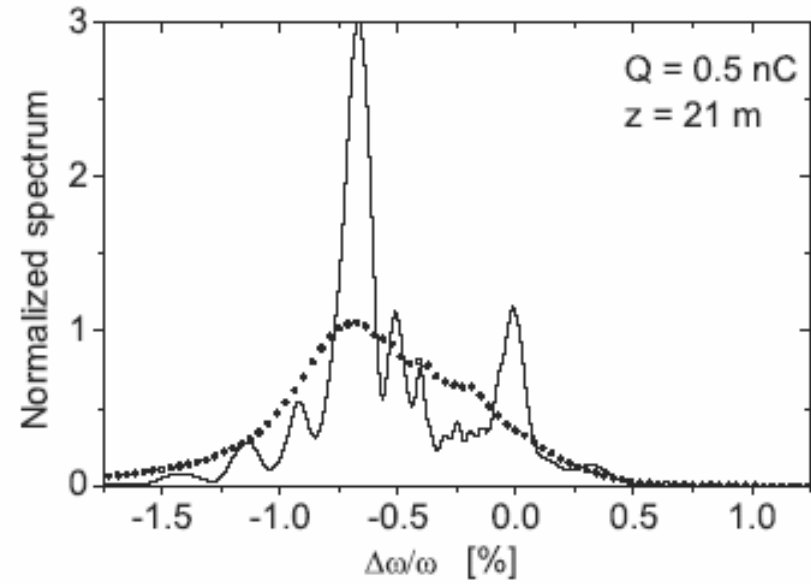


Genesis

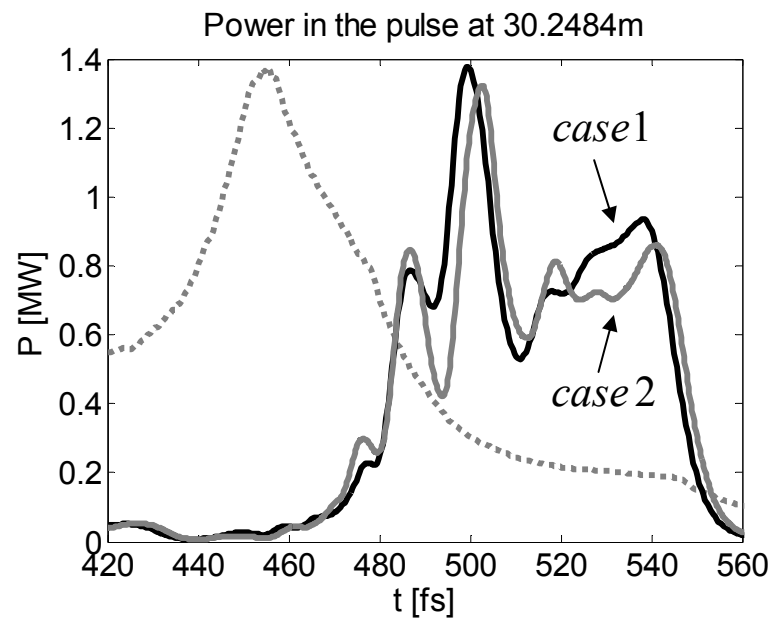
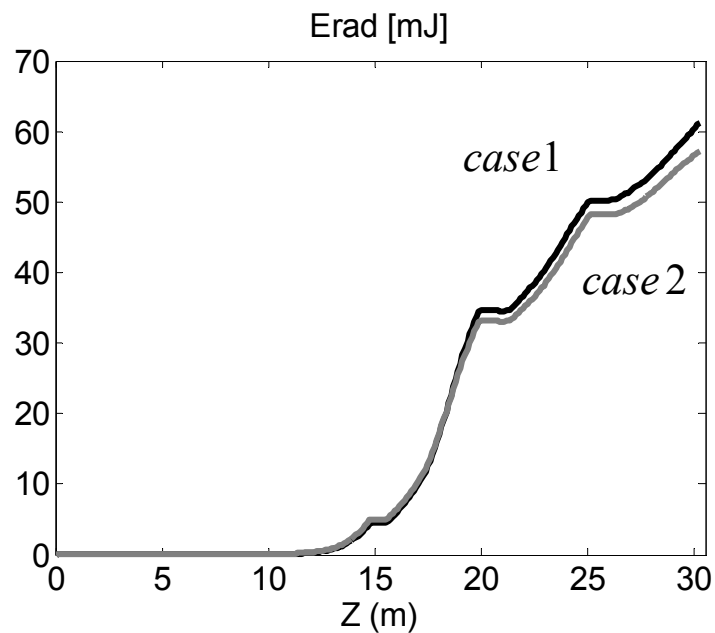


$FWHM = 0.7\%$

FAST (E.L.Saldin et al., DESY 05-239)



Genesis accuracy check



case1: slice width = $4\lambda_s$, N of particles / slice = 8192, N runs = 170

case2: slice width = $2\lambda_s$, N of particles / slice = 16384, N runs = 17

Conclusions

1. The pulse energy is smaller by factor 2 compared to FAST simulations
2. The possible reasons:
 - higher energy chirp,
 - smaller width of the current pulse
 - including of intersections in the lattice
3. Other statistical properties are in good agreement
4. The impact of wakefields?