

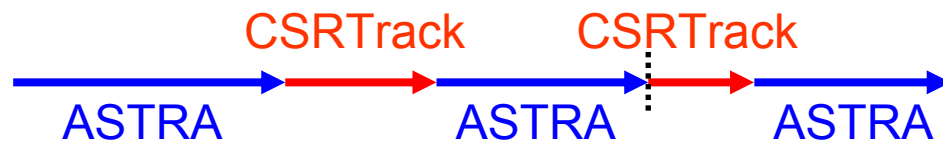
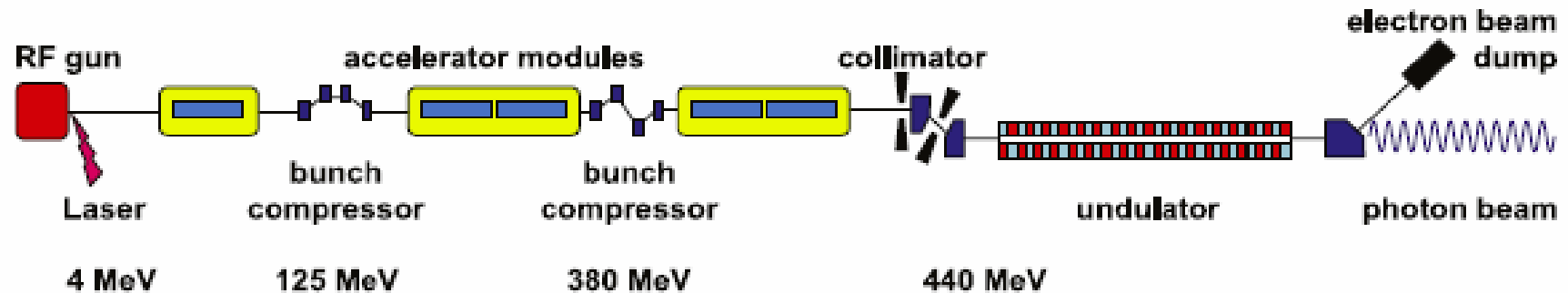
# New VUV-FEL Simulation results

27.3.2006 Bolko Beutner

# Motivation

- Goal : Observation of CSR Effects with LOLA
- Idea :
  - Use BC3 to achieve higher peak current and have a less complicated situation
  - Over-Compression in BC3 to reduce Space Charge Effects behind BC3

# Start to LOLA simulations



ACC1 : on crest

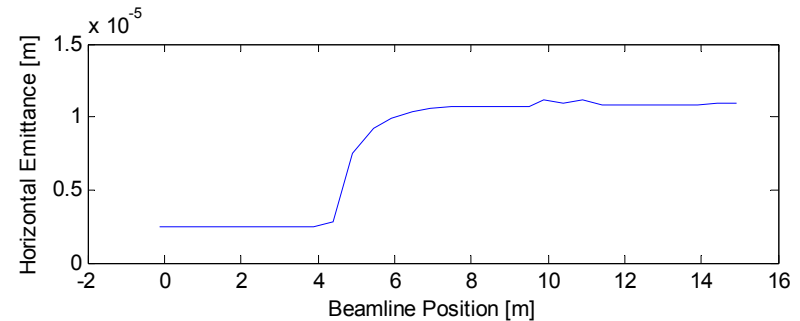
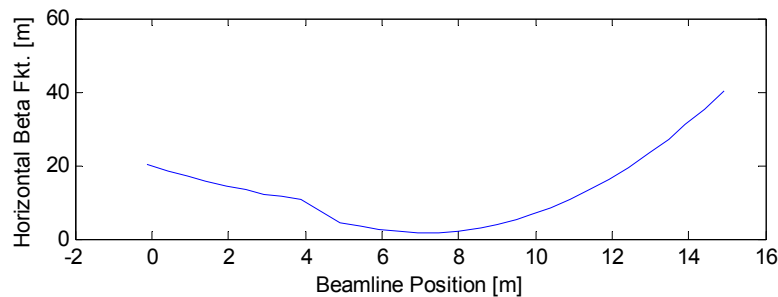
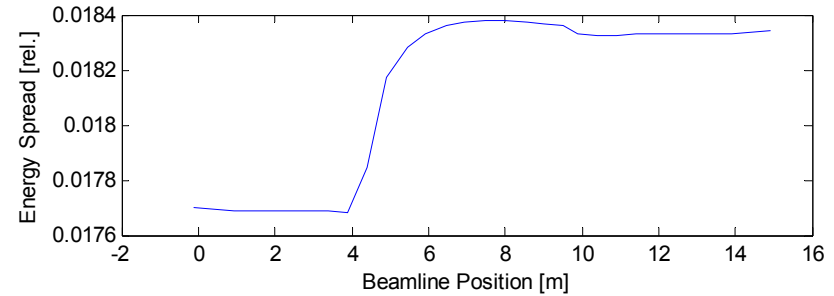
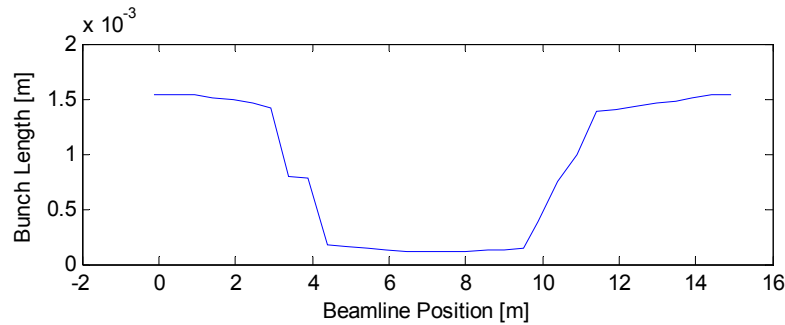
Bunch Charge : 2nC

ACC23 : maximum compression -28deg  
simulations with -40deg

Bunch Length : 1,5 mm

BC3: R = 5.3 m  
 $R_{56} = -0.099369$

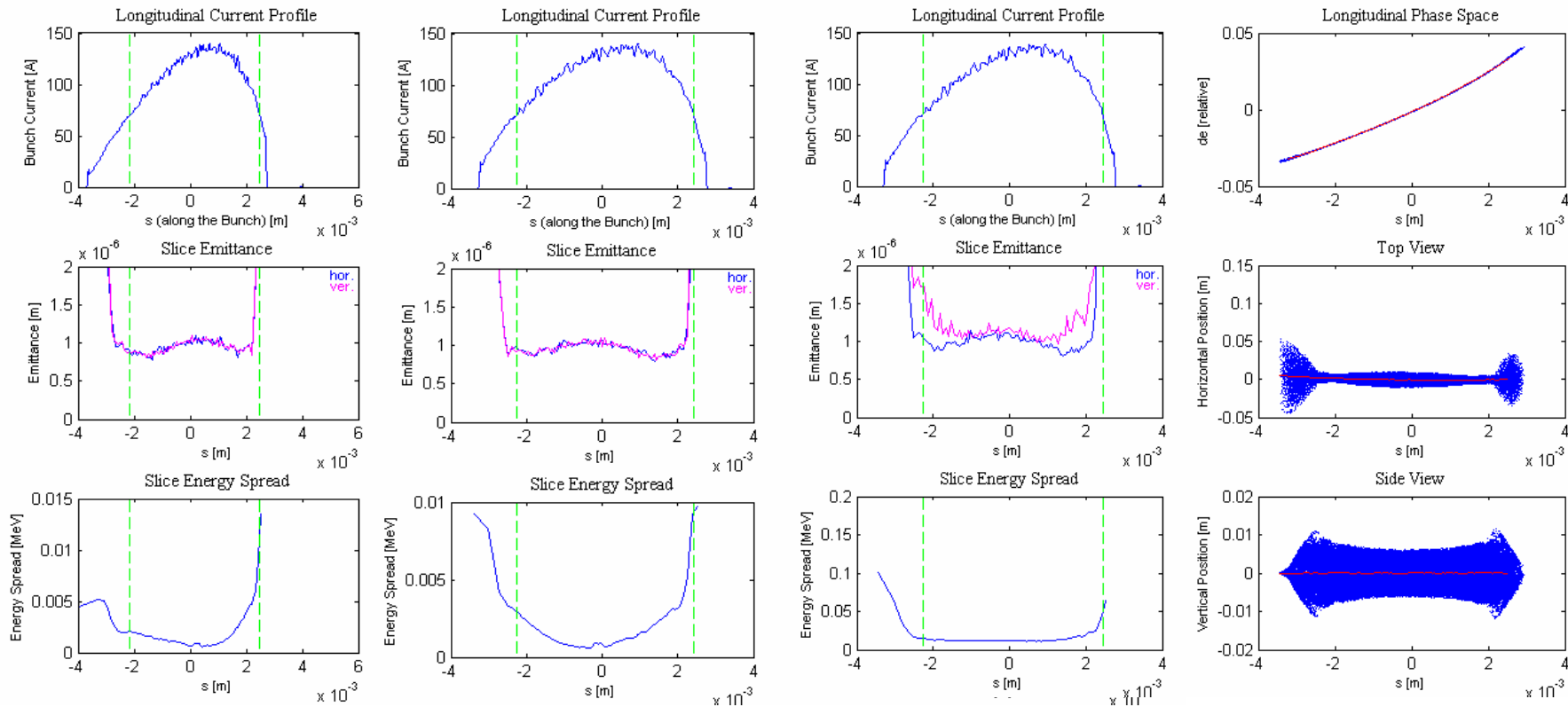
# Evolution along BC3 at -40deg off crest



At 40 deg off crest the bunch length after BC3 is the same as before.  
The Bunch is longitudinal flipped.

Energy spread and emittance growth indicates small interaction Region  
at the 3<sup>rd</sup> Dipole of BC3.

# s2e without CSR

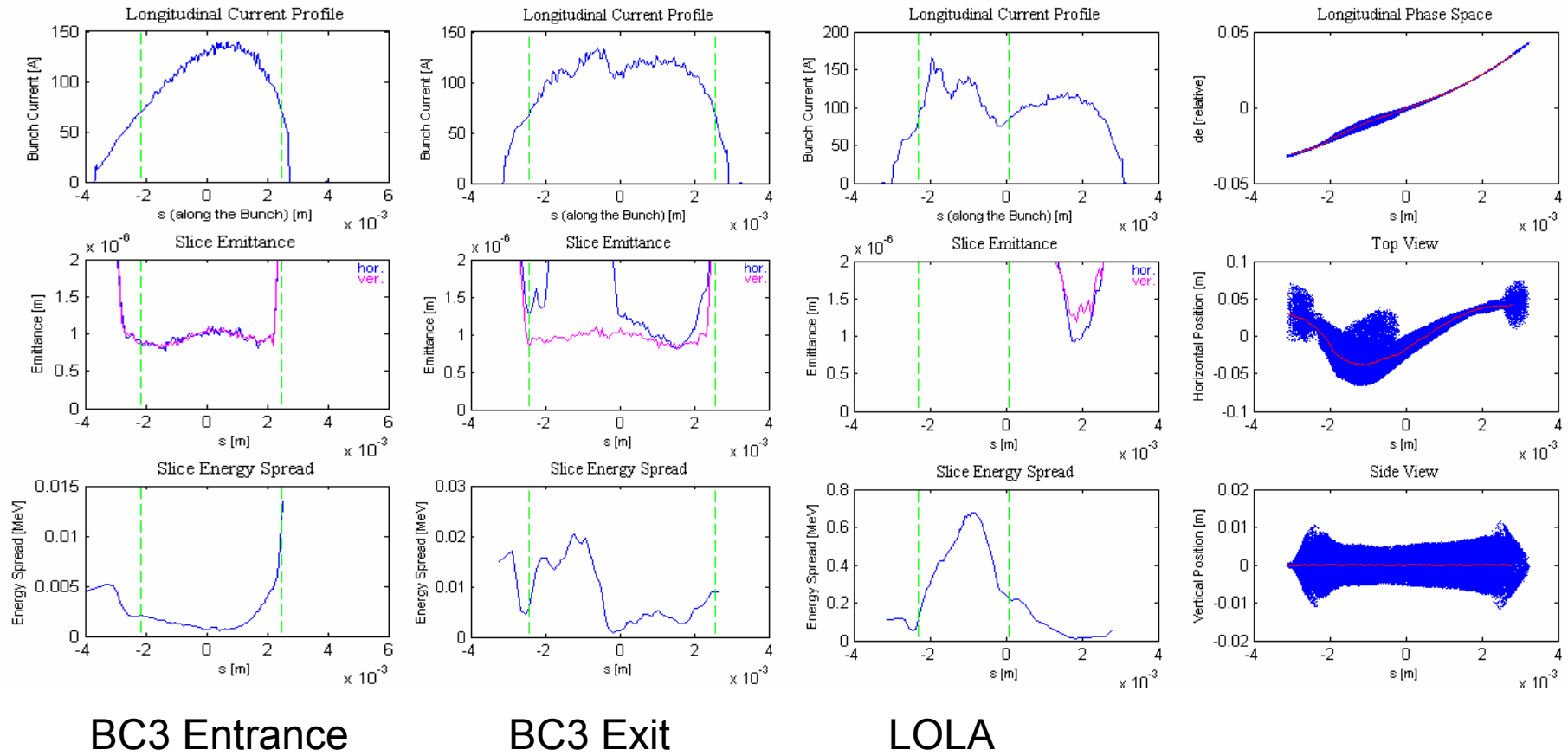


BC3 Entrance

BC3 Exit

LOLA

# s2e with CSR



# Interpretation

- Numerical Errors?
  - Results are stable under changes of tracking step width and smoothing parameters.
  - Always the same initial (BC3 entrance) Particle Distribution were used.
- Physics
  - CSR energy modulations at the Spike lead to longitudinal density changes. These longitudinal density modulations are transported to the tail of the bunch in BC3.
  - Due to longitudinal Space Charge these Modulations are amplified.

# Next Steps

- Use of different initial particle distributions
- Space Charge Effects within BC3 (Greens Function Method, ASTRA?)
- Scans of R, Phase, Charge.
- Complete ASTRA simulations (including CSR)
  
- Measurements at LOLA
  - Comparison of on crest and ~40 deg case