

THEORETICAL PARTICLE PHYSICS AT DESY

Wilfried Buchmüller, DESY

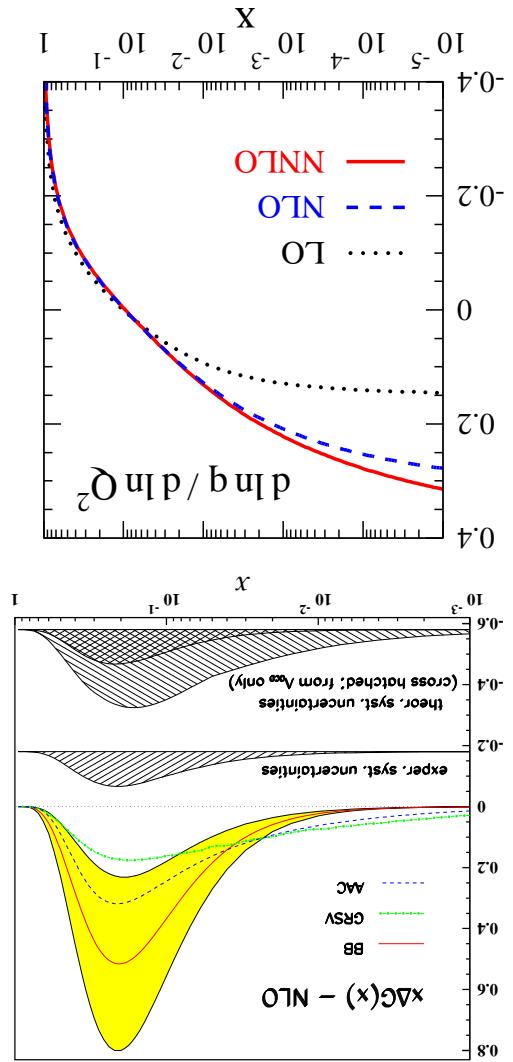
DESY PRC, November 2005, Open Session

Some guidance is provided by the classification of the ‘archive’:

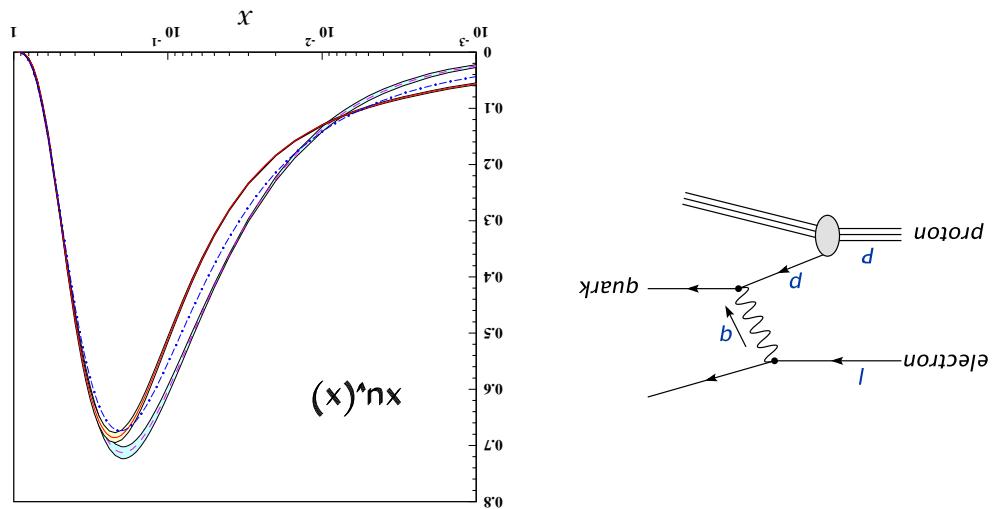
The DESY theory group always tried to cover the range from ‘phenomenology’ to ‘theory’, in its activities; in the following the hep-ph activities will be illustrated with some representative examples, with emphasis on staff members; there are also important publications by postdocs and students.

- hep-ph (‘phenomenology’) [HH: Schomerus, Teschner]
- hep-th (‘theory’) [Z: Simma, Sommer, NIC: Janssen, Schierholz]
- hep-Lat (Lattice gauge theories) [HH: Buchmiller, Covi, Hamaguchi, Ringwald]
- B. Beyond the SM, Particle Cosmology and Astroparticle Physics: neutrino physics, grand unification, proton decay, high energy cosmic rays, model building, extra dimensions, leptogenesis, ... [HH: Ali, Dierl, Kilian, Schrempp, Zerwas; Z: Blumlein, Moch, Riemann]
- A. Gauge theories and collider physics: processes at HERA, Tevatron, LHC, ILC; radiative corrections (electroweak, SUSY), precision tests, QCD, B-physics, ... [HH: Buchmiller, Covi, Hamaguchi, Ringwald]
- B. Beyond the SM, Particle Cosmology and Astroparticle Physics: neutrino physics, model building, extra dimensions, leptogenesis, ... [HH: Buchmiller, Covi, Hamaguchi, Ringwald]
- hep-th (‘theory’) [Z: Simma, Sommer, NIC: Janssen, Schierholz]

What is Theoretical High Energy Physics ?



- Parton distributions with correlated errors (at 2 and 3 loops)
- Precision calculations of anomalous dimensions and Wilson coefficients to 3 loops
- Precision determination of α_s (1% uncertainty) in close collaboration with H1, ZEUS, HERMES
- Detailed comparisons with Lattice results



Deep-Inelastic Scattering at HERA: unpolarized and polarized

$$\text{Instanton-driven} \left\{ \begin{array}{l} \text{Color Glass Condensate} \\ \Leftrightarrow \\ \text{QCD Sphaleron state} \end{array} \right\} \Leftrightarrow \left\{ \begin{array}{l} \text{Saturation scale} \\ \Leftrightarrow \\ \langle p \rangle \approx 0.5 \text{ fm (lattice)} \end{array} \right\}$$

Strong classical fields $\propto \frac{1}{\alpha_s} \Leftarrow \text{Instantons!? [Schrempp \& Petermann]}$
Expect: many non-perturbative gluons, multiplicity $\langle n_g \rangle \propto \frac{1}{\alpha_s}$:

- Gluon saturation at small x_B [$\alpha_s(Q^2)$ still small]

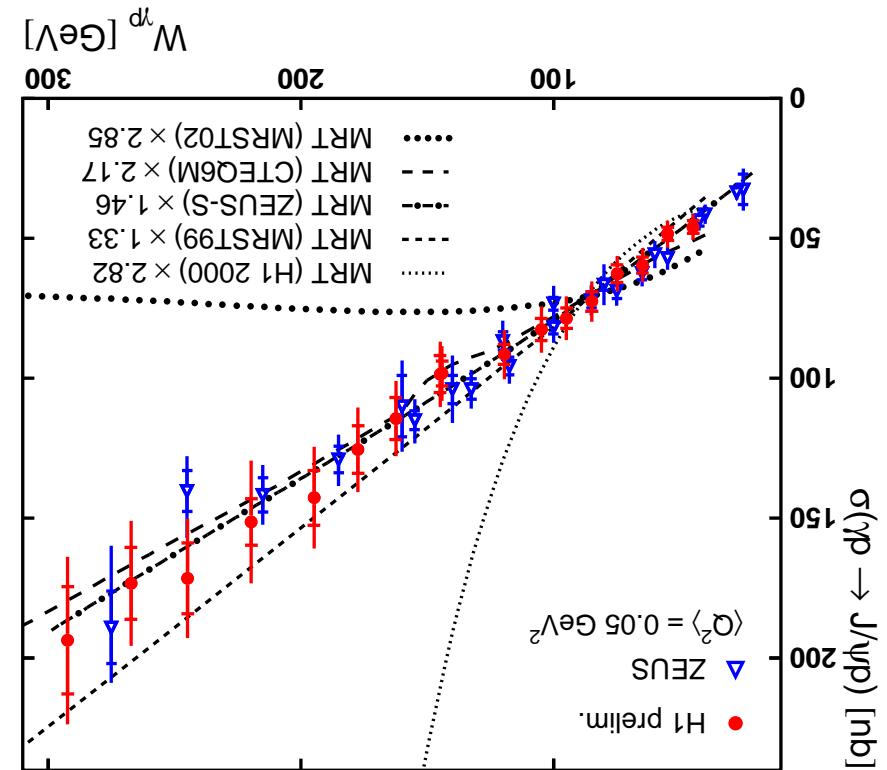
$$\text{HERA} \left\{ \begin{array}{l} \Leftrightarrow \text{2 dedicated } I\text{-search exps. (H1, ZEUS)} \\ \Leftrightarrow [\text{Moch, Ringerwald \& Schrempp}] \\ \Leftrightarrow \text{study of } I\text{-discovery potential} \\ [\text{Petermann \& Schrempp}], \text{in progress} \end{array} \right\} \text{LHC}$$

Characteristic, I -induced processes, calculable within I -perturbation theory.
Basic non-perturbative aspect of QCD, yet direct exp'tl evidence lacking.

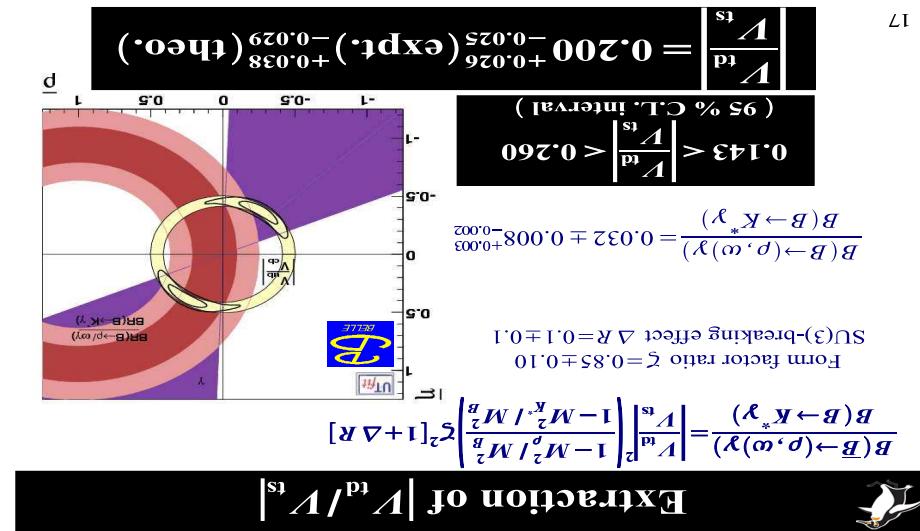
HERA and LHC: Instantons (I)

H1 and ZEUS data on J/ψ production (theory: T. Teubner, DIS'05)

- accessible in vector meson production, virtual Compton scattering, hard diffraction
- theoretical framework for analysis: Generalized Parton Distributions
- extend QCD factorization to exclusive final states
- information on proton structure
- spatial distribution of partons
- orbital angular momentum carried by quarks
- strong sensitivity to gluon distribution
- various measurements by H1, ZEUS, HERMES



Exclusive Processes at HERA



BELLE $\bar{B}_{\text{exp}}[B \rightarrow (p, \omega) \gamma] = (1.34^{+0.34}_{-0.34} \text{ (stat)}^{+0.14}_{-0.10} \text{ (sys)}) \times 10^{-6}$

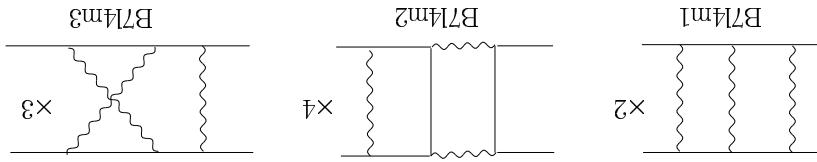
$$R(p, \omega/K_*) \equiv \frac{\bar{B}[B \rightarrow K^* \gamma]}{\bar{B}[B \rightarrow (p, \omega) \gamma]} = 0.033 \pm 0.010$$

$$\bar{B}[B \rightarrow (p, \omega) \gamma] = (1.38 \pm 0.42) \times 10^{-6}$$

Theoretical Predictions:

Experiment vs. SM ($b \rightarrow d \gamma$): Test of CKM unitarity, determination of $|V_{td}/V_{ts}|$

Flavour Physics: Rare decays and CP violation



representations

- Use of n -dim. complex Mellin-Barnes
- Solve systems of differential equations
- Non-trivial techniques:
massive 2-loop box diagrams

First systematic evaluation of master integrals for

Carls

In order to complement the Bhabha Motte

massive 2-loop corrs. to Bhabha scattering

← need for this:

scattering

is possible with small angle Bhabha

$\bullet 10^{-4}$ luminosity measurement at ILC

Electroweak Precision Predictions for LHC and ILC

<http://www-zethen.desy.de/theory/research>

ZFITTER Support Group

in close collaboration with experimentalists:

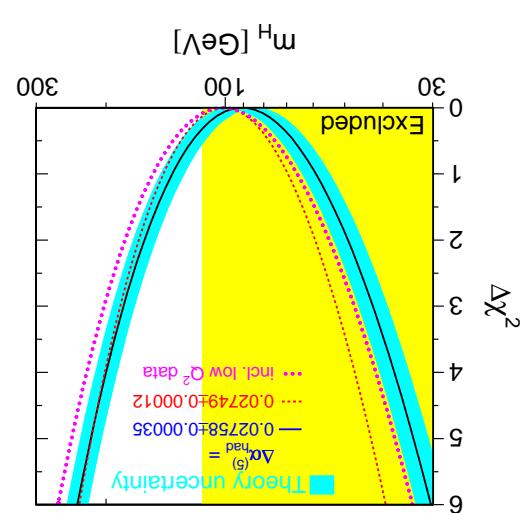
7/2005: now with latest weak 2-loops

made with **ZFITTER** (1985 – 2005)

See the “blue-band-plot”

 The m_{Higgs} prediction for the LHC

 The m_{top} prediction for Tevatron



for precision physics

⇒ serious limiting factor

| | | | | |
|---|--------|-------------------------------|--------|-------------------------------|
| $\frac{\alpha(M_Z)}{\delta\alpha(M_Z)}$ | \sim | 3.6×10^{-9} | \sim | 3.6×10^{-9} |
| $\frac{G_\mu}{\delta G_\mu}$ | \sim | 8.6×10^{-6} | \sim | 8.6×10^{-6} |
| $\frac{M_Z}{\delta M_Z}$ | \sim | 2.4×10^{-5} | \sim | 2.4×10^{-5} |
| $\frac{\alpha}{\delta\alpha}$ | \sim | $1.6 \div 6.8 \times 10^{-4}$ | \sim | $1.6 \div 6.8 \times 10^{-4}$ |
| | | | | (present) |
| | | | | (TESLA requirement) |

for VB physics (Z, W) etc.

$a(M_Z), G_\mu, M_Z$ best effective input parameters

relationship $\sin^2 \Theta_f, v_f, a_f, M_W, T_Z, T_W, \dots$

non-perturbative precision predictions

partially

a, G_μ, M_Z most precise input parameters

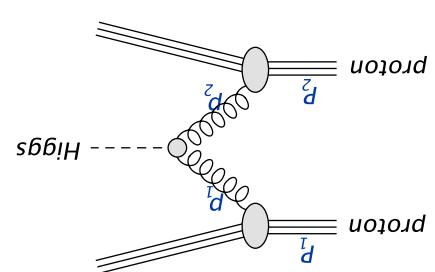
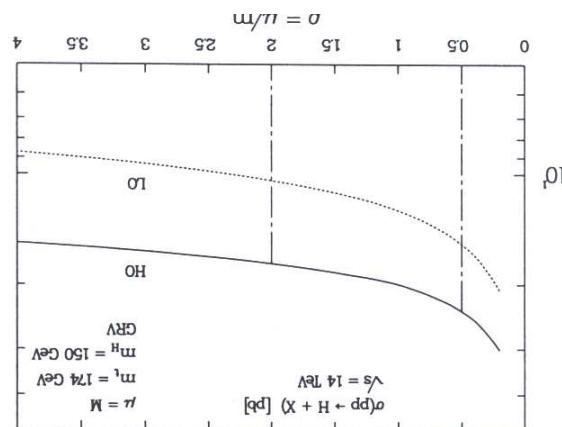
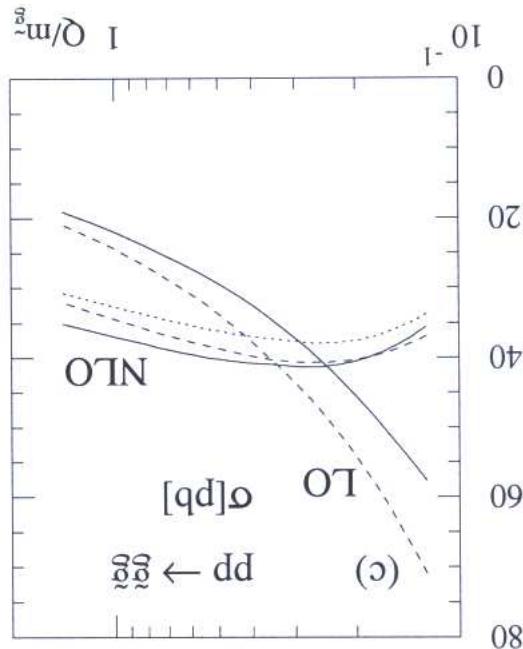
Precision physics limitations:

electroweak precision physics:

Uncertainties of hadronic contributions to effective a are a problem for

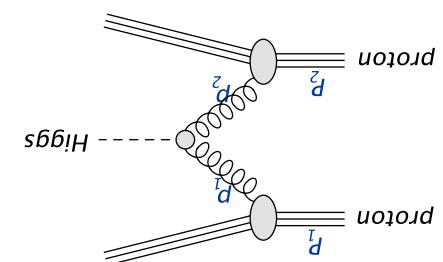
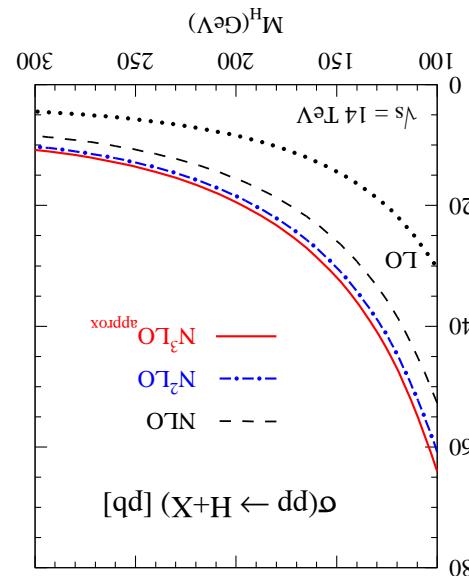
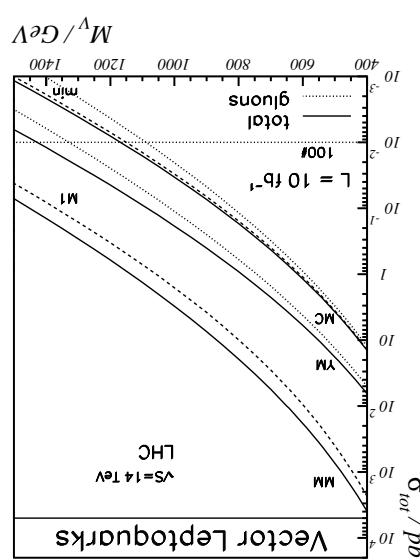
② $\alpha_{em}(s)$ in Precision Physics

- SUSY particles: $pp \rightarrow q\bar{q}, g\bar{g}$ etc; cross section nearly doubled, prediction stabilized
- Higgs bosons: SM and SUSY; $pp \rightarrow H$ and $t\bar{t}H$, etc; NLO/LO: cross section nearly doubled
- Many production channels of Higgs bosons and SUSY particles based on gluon-gluon fusion \rightarrow large QCD radiative corrections



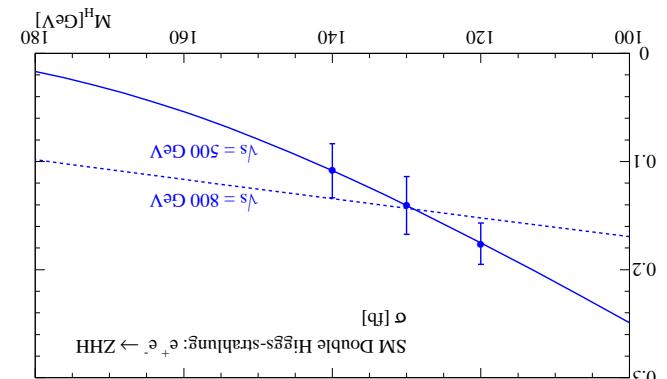
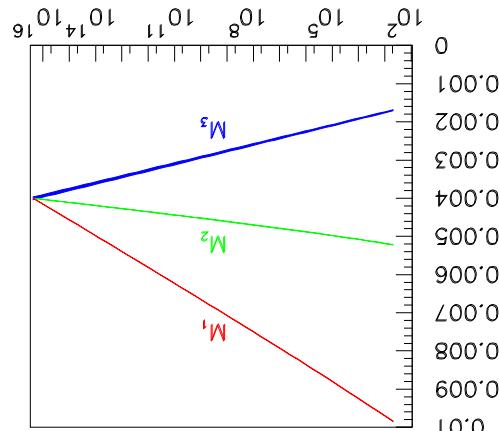
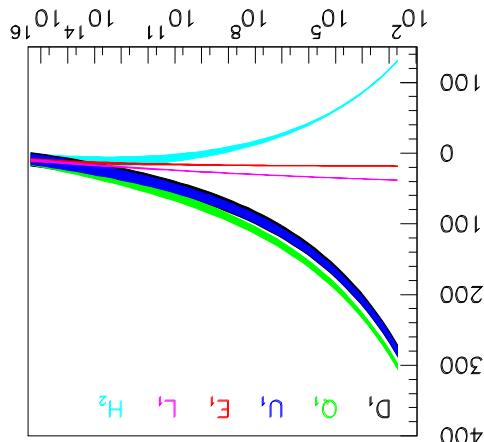
Higgs and SUSY at LHC

- Extension to physics beyond the Standard Model
- Calculation of time-like processes with different mass scales
- Transfer of DIS higher order technology to LHC physics
- Fast Mellin-based evolution codes for QCD matrix elements for inclusive LHC processes
- Higgs Production: QCD corrections to total cross section at 3 loops



Predictions for the LHC: in the Standard Model and Beyond

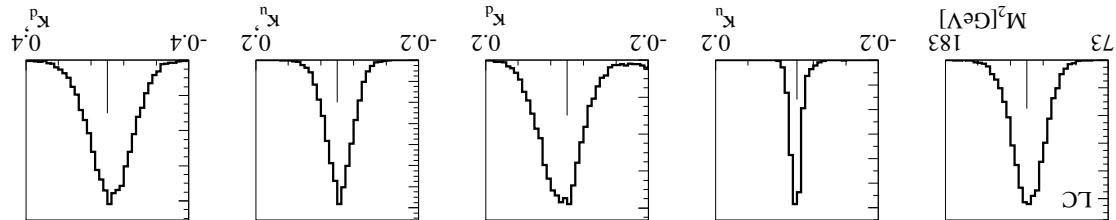
- Higgs at ILC: spin-parity and **trilinear coupling**, important for reconstruction of Higgses
- SUSY at LHC/ILC: particle masses/couplings, **extrapolation to GUT/Planck scale potential**
- Reconstitution of fundamental SUSY theory; identification of mechanism for supersymmetry breaking



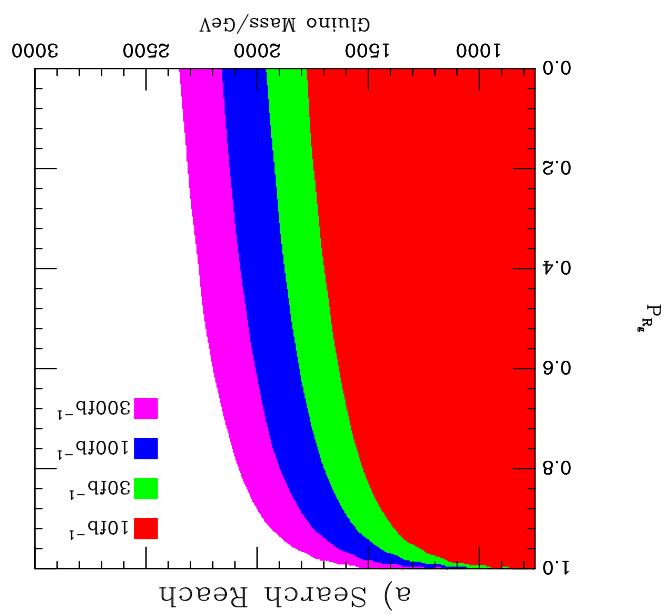
Higgs and SUSY at LHC/ILC

[Kilian, Plehn, Richardson, Schmidt]

⇒ ILC: Anomalous Yukawa couplings can be measured with sufficient precision ($< 10\%$)

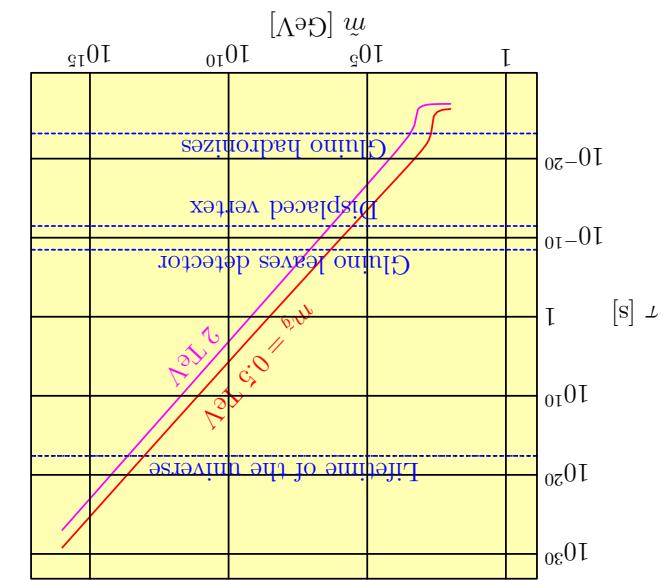


Establish high-scale supersymmetry: measure chargino/neutralino parameters at ILC



Phenomenology:

- Gluino is long-lived
- Gluino pairs at LHC:
- Heavy stable hadrons
- Charginos/neutralinos
- Higgs boson
- With $m > 130$ GeV

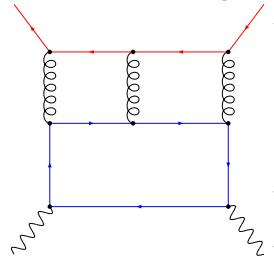


Fine-tuning resolved by string-theory vacuum multiplicity?

All scalars (except Higgs) superheavy = no flavor problem, but fine-tuning

Split Supersymmetry at LHC/ILC

Mathematical Challenges and Computer Algebra



The solution of multi-loop problems as needed by LHC, GiGAsZ, ILC faces us with serious technology challenges.

Feynman diagrams are no longer solved by conventional technology: Bytes of integrals

$$-24H_{1,-} \zeta_3 - 16H_{1,-} - 2H_{1,0} + \frac{67}{6} H_{1,0} - 2H_{1,0} \zeta_2 + \frac{3}{31} H_{1,0,0} + 11H_{1,0,0,0} + 8H_{1,1,0,0} - 8H_{1,3} + H_4$$

$$+ \frac{9}{67} H_{2,-} - 2H_{2,0} + \frac{11}{3} H_{2,0} \zeta_2 + \frac{3}{31} H_{2,0,0} + H_{3,0} + p_{q4}(x) \left[\frac{1}{6} \zeta_2^2 - \frac{9}{67} \zeta_3 + \frac{31}{3} H_{3,0} - 3H_{3,0} \zeta_2 - 3H_{2,-} + 2H_{2,0} - 2H_{1,-} \zeta_2 + \frac{10}{3} H_{1,0,0} - \right.$$

$$- 3H_{2,-} \zeta_2 - 4H_{2,-} - 4H_{2,0} - 2H_{2,0} \zeta_2 - \frac{31}{3} H_{2,0,0} - 3H_{2,0,0} \zeta_2 - 42H_{1,-} \zeta_2 - H_{3,0} +$$

$$- 4H_{1,-} - 2,0 + 56H_{1,-} - 1\zeta_2 - 36H_{1,-} - 1,0,0 - 56H_{1,-} - 1,2 - \frac{9}{134} H_{1,-} - 42H_{1,-} \zeta_2 - H_{3,0}$$

Need of:

$$(722731935998670312187894999 + 6419601) q_{F,2}^{127884509146146400000}$$

$$(77502662711876824509176089050514652421741 + 2849482004138921491531) q_{F,2}^{221491531}$$

$$+ (5987886538667 + 6419601) q_{F,2}^{222202799333557529} + (1531530 + 6419601) q_{F,2}^{453}$$

$$+ (5984963680800 + 6419601) q_{F,2}^{453} + (-4021001384400 + 6419601) q_{F,2}^{453}$$

$$+ (493290512768148 + 6419601) q_{F,2}^{453} + (-3984963680999 + 6419601) q_{F,2}^{453}$$

$$+ (52977469944553728278848870400000 + 6419601) q_{F,2}^{453} + (1531530 + 6419601) q_{F,2}^{453}$$

$$+ (5987886538667 + 6419601) q_{F,2}^{453}$$

Algorithm in FORM

$$+ 13H_{0,0} \zeta_2 + \frac{89}{12} H_{0,0,0} - 5H_{0,0,0,0} - 7H_{2,0} - \frac{31}{6} H_3 - 10H_4 + (1-x) \left[\frac{133}{36} + 4H_{0,0,0,0} \right]$$

$$+ 4H_{1,0,0} + \frac{3}{4} H_{1,0} \left[1 + (1+x) \left(\frac{43}{4} \zeta_2^2 - 3 \zeta_2^2 + \frac{25}{2} H_{2,0} - 3H_{1,-} \zeta_2 - 14H_{1,-} - 1,0 - \frac{3}{13} H_{1,-} \right) \right]$$

$$+ 24H_{1,2} + 23H_{1,0,0} + \frac{55}{2} H_{0,0,2} + 5H_{0,0,2,2} + \frac{1457}{36} H_0 - 1025 H_0 - \frac{155}{6} H_2 + H_2 \zeta_2 - 15H_3$$

$$+ 2H_{2,0,0} - 3H_4 \left[- 5 \zeta_2^2 - \frac{1}{2} \zeta_2^2 + 50 \zeta_3 - 2H_{-3,0} - 7H_{-2,0} - H_0 \zeta_3 - \frac{37}{2} H_0 \zeta_2 - \frac{24}{9} H_0 \right]$$

$$+ (185 + 16CA_Cp(x) \left[\frac{245}{67} \zeta_2^2 - \frac{67}{18} \zeta_2^2 + \frac{5}{12} \zeta_2^2 - \frac{1}{12} \zeta_3 + \frac{64}{216} H_0 \right] + 16CA_Cp(x) \left[\frac{245}{67} \zeta_2^2 - \frac{67}{18} \zeta_2^2 + \frac{5}{12} \zeta_2^2 - \frac{1}{12} \zeta_3 + \frac{64}{216} H_0 \right])$$

$$+ (60 + 12CA_Cp(x) \left[\frac{247}{67} \zeta_2^2 + \frac{211}{12} \zeta_3 + \frac{15}{2} \zeta_3 \right] + 16CA_Cp(x) \left[\frac{245}{67} \zeta_2^2 - \frac{67}{18} \zeta_2^2 + \frac{5}{12} \zeta_2^2 - \frac{1}{12} \zeta_3 + \frac{64}{216} H_0 \right])$$

$$- H_{-3,0} + 4H_{-2,0} - H_{-2,0} + 2H_{-2,0} - H_{-2,0} \left(\frac{31}{12} H_{0,0,2} + 4H_{0,0,2,2} + \frac{389}{12} H_{0,0} - 2H_{0,0} \right)$$

$$\leftarrow \text{in cooperation with mathematicians}$$

$$+ H_{0,0,0} + 9H_{1,0,0,0} + 6H_{1,0,0,2} - \frac{4}{11} H_{1,0,0,0} - 3H_{1,0,0,0} - 4H_{1,0,0,2} + 4H_{1,0,0,3} + \frac{31}{12} H_{0,0,0} - 2H_{0,0,0}$$

$$+ H_{-3,0} + 4H_{-2,0} - H_{-2,0} - H_{-2,0} + 2H_{-2,0} - H_{-2,0} \left(\frac{31}{12} H_{0,0,2} + 4H_{0,0,2,2} + \frac{389}{12} H_{0,0,0} - 2H_{0,0,0} \right)$$

$$- 3H_{-1,0,0,0} + \frac{3}{11} H_{-1,0,0,2} + 12H_{-1,0,0,2,2} - 16H_{-1,-} \zeta_3 - 8H_{-1,-} - 1,0,0 + 16H_{-1,-} - 1,2 + \frac{9}{67} H_{-1,0,0}$$

$$+ \frac{11}{12} H_3 + H_4 \left[+ p_{q4}(-x) \left(\frac{67}{18} \zeta_2^2 - \zeta_2^2 - \frac{4}{11} \zeta_3 - H_{-3,0} + 8H_{-2,0} \zeta_2 + \frac{6}{11} H_{-2,0} - 4H_{-2,0,0} \right) \right]$$

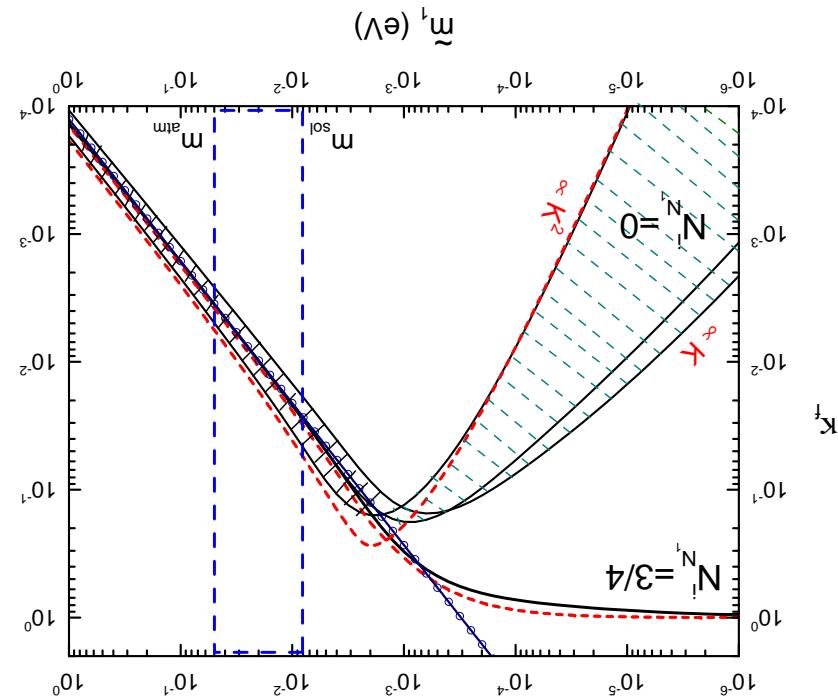
$$- Blumlein, Moch, Riemann]$$

Implications: light neutrino mass window, $10^{-3} \text{ eV} < m_\nu < 0.1 \text{ eV}$, can be tested in laboratory experiments and cosmology; lower bound on the heavy neutrino mass and maximal temperature of the early universe, $T^B \sim M^i > 10^9 \text{ GeV} \rightarrow$ important implications for dark matter.

$$\eta_B \approx 0.01 k_f.$$

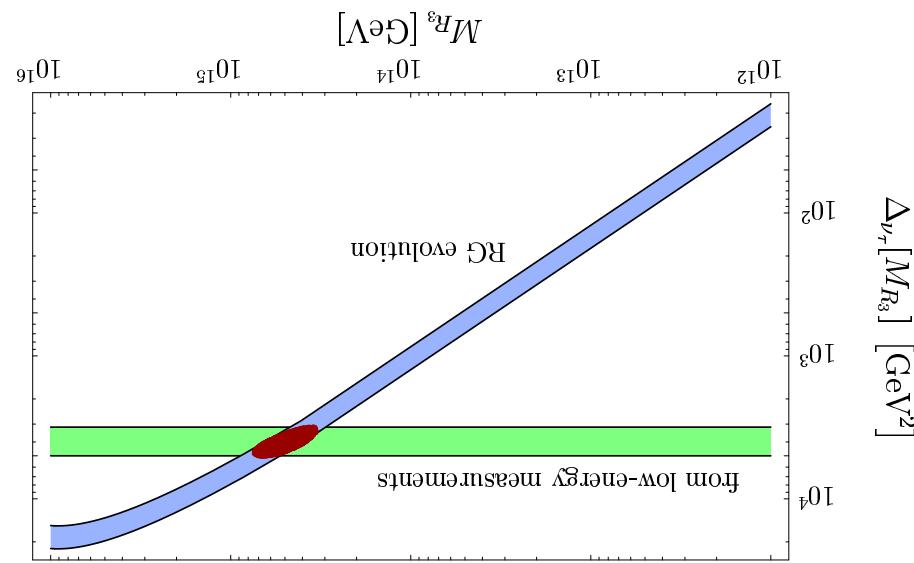
- Non-equilibrium process: quantitative relation between neutrino masses and baryon asymmetry (CP asymmetry ϵ_1 , efficiency factor k_f):

- Leptogenesis: leading (?) theory for origin of matter (Fukugita, Yanagida; DESY (86))



Neutrino Masses and Baryon Asymmetry

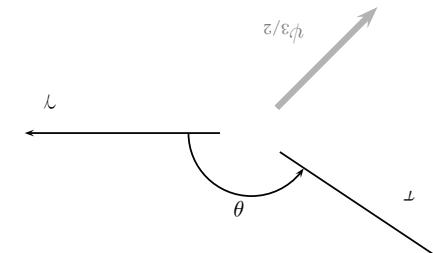
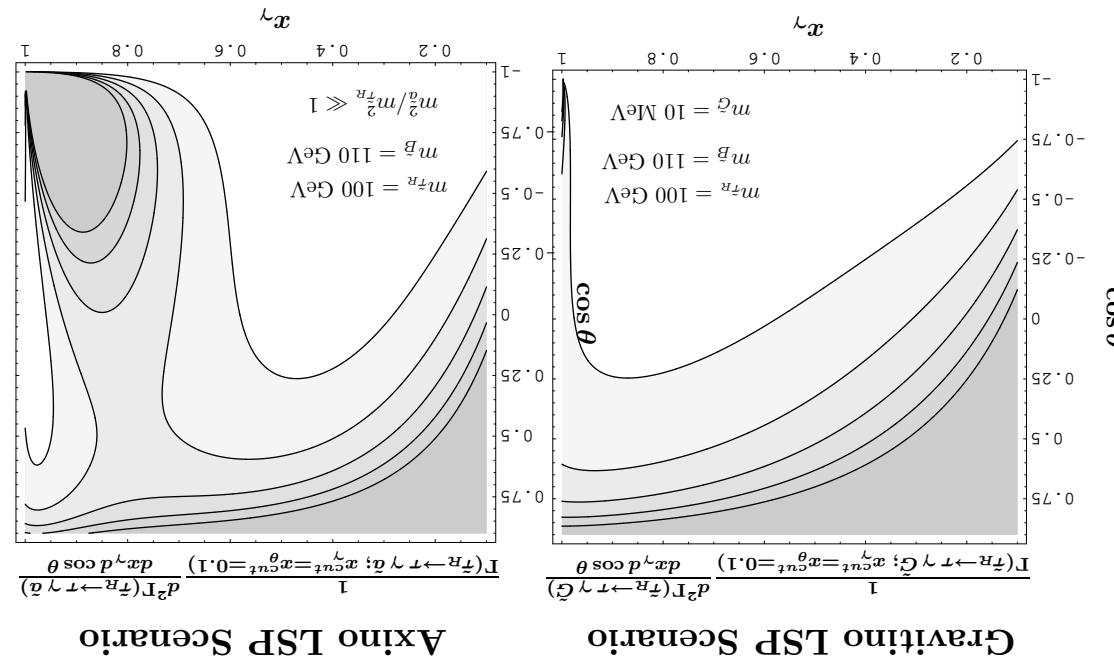
The mass differences between scalar tau-leptons and selectrons depend on the seesaw scale in supergravity $SO(10)$ models. For light SUSY mass patterns, the mass differences can be accurately determined at ILC.



The reconstruction of the heavy neutrino mass scale is a theoretical/experimental challenge! For supersymmetric theories this may be possible in some regions of parameter space.

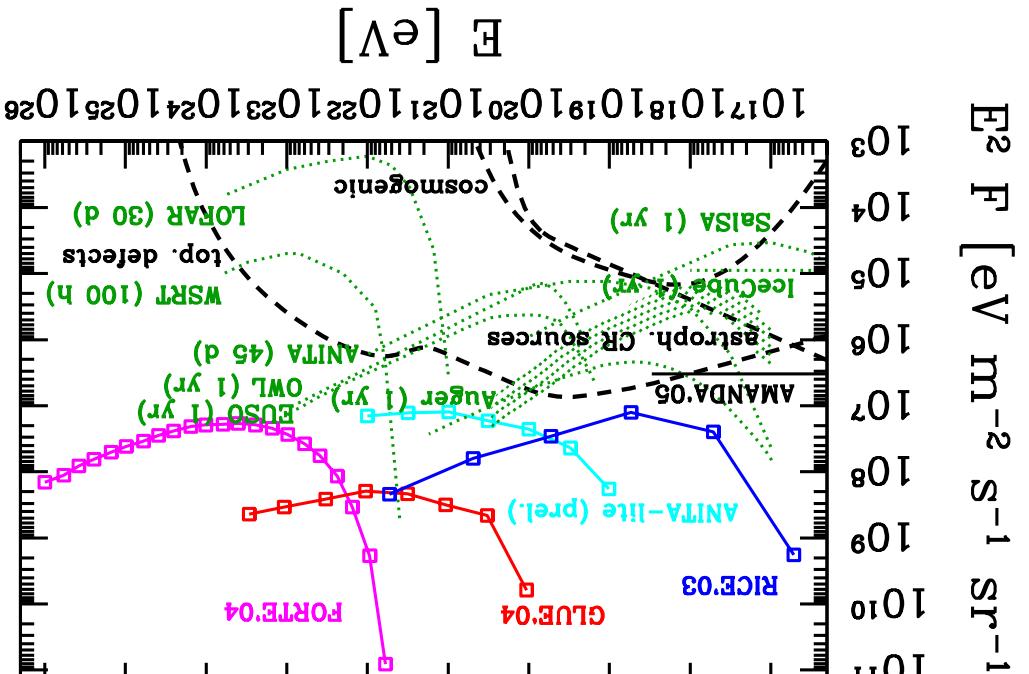
Reconstruction of Heavy ν_R Mass Scale at the ILC

- From the 3-body decay $\tau \rightarrow \tau + G + \gamma$ one can determine the gravitino spin, $s=3/2$, (‘Gedankenexperiment’).
- For a large range of parameters a quasi-stable scalar-tau is the **NLSP**, its lifetime yields a microscopic determination Newton’s constant.
- The large temperature required by leptogenesis favours a **gravitino LSP** which can be the dominant component of dark matter.



Discovering the Gravitino at the LHC/ILC

- Many observatories running (AMANDA,...) or under construction (IceCube,...)
- Already now stringent constraints on high energy neutrino flux with significant discovery potential in the next decade; strong impact on astrophysics, particle physics and cosmology
- astrophysics: processes associated with acceleration of cosmic rays
- particle physics: Ω_N far beyond the reach of HERA and LHC; study of SM contributions beyond SM
- cosmology: big bang relics (topological defects, relic neutrinos,...)



Extremely Energetic Cosmic Neutrinos

- Broad research activities related to the colliders HERA, LHC/ILC, QCD, electroweak theory and B-physics, Higgs and SUSY with implications for cosmology and astroparticle physics will be crucial for the physics analysis at LHC/ILC.
- Sophisticated ‘tools’ have been/are being developed which and SUSY with implications for cosmology and astroparticle physics will be crucial for the physics analysis at LHC/ILC.
- Research in Astroparticle Physics is closely related to AMANDA/IceCube
- Physics beyond the SM (neutrinos, supersymmetry, GUTs...) tries to bridge the gap between the SM and string theory
- Particle Cosmology is focussed on applications of particle physics to the physics of the early universe, with implications for collider physics

Summary of the hep-ph activities:

HGF Evaluation in 2003: "The building-up/strengthening of groups in String Theory and Particle Cosmology is strongly supported",

activity in particle cosmology (Virtual Institute)

- Hamburg: lattice group ends, string group starts; more for strong QCD lattice group (NIC)
- Zeuthen: expansion of astroparticle physics (exp), support

changes at (German) universities:

major changes in recent past reflect development of particle physics and

- center for theoretical particle physics in Germany and beyond

LHC/ILC, IcCube

- close connection with experimental program: HERA,

General Strategy

- HGF-Nachwuchsgruppe VH-NG-004: QCD in exklusiven Prozessen bei HERA
- HGF-Nachwuchsgruppe VH-NG-005: Universelle Ereignisgeneratoren für zukünftige Lepton-Collider
- HGF-Nachwuchsgruppe VH-NG-006: Partikel physik und kosmologie: beyond the two standard models
- HGF-Nachwuchsgruppe VH-NG-105: Computer algebra and higher orders in partikel theory
- HGF-Virtuelles Institut VH-VI-106: Partikel Cosmology (ViPAC, coordination DESY)
- HGF-Virtuelles Institut VH-VI-032: Hochenergiestrahllungen aus dem Kosmos (VIHKOS)
- HELEN: High Energy Physics Latinamerican-European Network (Coordination DESY: F. Schrempp)
- European Network for Theoretical Astroparticle Physics ENTAPP
- SFB Transregio-9: Computational Particle Physics (Karlsruhe, Berlin, Aachen, Zeuthen/NIC)
- SFB Particles, Strings and the Early Universe (positive pre-evaluation; evaluation of full application January 2006; coordination Hamburg University)

Additional funding from HGF, DFG, EU, ...

corner stone of the scientific life at DESY, includes:

Collaboration with (German) Universities

the continuous flow of excellent graduate students from (German) Universities is decisive for the scientific life and the success of the DESY theory group...

- the close collaboration with the Universities in Berlin, Potsdam and Hamburg
- frequent joint workshops on collider physics (HERA, LHC and ILC)
- the new Center for Mathematical Physics with the II. Institute of Theoretical Physics and Department of Mathematics (Hamburg)
- Virtual Institute on Particle Cosmology with Univ. Bonn, Heidelberg, Munich
- HGF-Nachwuchsguppen with Univ. Karlsruhe, Padua, Regensburg, Würzburg
- the annual Theory Workshop in Hamburg, organized jointly with German universities
- the conference series Loops and Legs in QFT, organized by the Zeuthen group
- the Computer-Algebra School, organized by Zeuthen/German Universities and MPG
- all DESY staff members supervise graduate students. During the past 4 years, 10 permanent positions in particle theory were filled in Europe with former graduate students from Hamburg, including 6 professorships at German Universities