

### Simulation of EUDET Pixel Beam Telescope using ILC Software

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**EUDET pixel beam telescope** 

• JRA1: Test beam infrastructure

Comprises large bore magnet (B<1.2 Tesla) and pixel beam telescope

- Purpose of telescope: precise track reconstruction used for pixel sensors, as well as for large volume tracking devices (e.g. TPC)
- Should have very high precision (<3  $\mu$ m)
- Suitable for different test beam environments:
  - DESY: electrons up to 6 GeV
  - CERN: pions 100-120 GeV
- For telescope planes use CMOS sensors developed by IPHC-Strasbourg
- DAQ development: Switzerland, Italy, France, Germany, UK
- Will be assembled at DESY

# Simulated Configuration I (symmetric geometry)



Simulation is important for detector design optimisation

**Electron energies:** 1, 2, 3, 4, 5, 6 GeV

Assumed intrinsic resolution of a telescope plane is 3  $\mu$ m (hit positions are smeared)

# Simulated Configuration II (asymmetric geometry)



# **Software Tools**

- Simulation: Mokka 06.00 (based on Geant 4)
  - New geometry driver EUTelescope has been created
  - Class TRKSD00 is used for telescope and DUT sensitive detectors
  - No hard-coded numbers
  - All parameters of the model are stored in MySQL database
  - Output: LCIO format files
  - Stored information: hit positions, deposited energy
- Analysis: Marlin 00.09.04 and Root
- Simulated 50000 events without magnetic field for different configurations

#### Validation of Multiple Scattering model

The width of the projected angular distribution is defined as

$$heta_0 = heta_{ ext{plane}}^{ ext{rms}} = rac{1}{\sqrt{2}} heta_{ ext{space}}^{ ext{rms}}$$

For small scattering angles Gaussian approximation is used:

$$heta_0 = rac{13.6 \ \mathrm{MeV}}{eta c p} z \sqrt{rac{x}{X_0}} \left[1 + 0.038 \ln \left(rac{x}{X_0}
ight)
ight]$$

 $p, \beta c, z$  are momentum, velocity and charge number of the incident particle  $x/X_0$  is the thickness of the scattering medium in radiation lengths

#### To check the validity of MS description:

- Simulate silicon wafer of 300  $\mu$ m thickness
- Shoot 1 GeV electrons (100000 events)
- Look at the projection of the scattering angles  $\theta$
- Compare the width of  $\theta$  distribution to theoretical prediction:  $\theta_0 = 0.602$  mrad

# **Projection of scattering angle**

Theory:  $heta_0=0.602$  mrad



# **Analysis Procedure**

- Fit a track (straight line model in the absence of magnetic field) through hits in telescope planes
- Find a position of the intersection of the track with the DUT  $(x_{
  m pred}, \, y_{
  m pred})$
- Find DUT residuals:

 $r_{x \; \mathrm{DUT}} = x_{\mathrm{pred}} - x_{\mathrm{DUT}}$ 

 $r_{y \text{ DUT}} = y_{\text{pred}} - y_{\text{DUT}}$ 

where  $x_{\text{DUT}}$  and  $y_{\text{DUT}}$  are hits in the DUT

• Fit Gaussian to residual distributions and find  $\sigma_x$  and  $\sigma_y$ 

## **Comparison of Geant 4 (Prague) and Mokka simulations**

#### The results look similar

Symmetric geometry

In symmetric geometry (1 AI box) DUT is shifted 5 mm left in comparison with picture on slide 4

In asymmetric geometry (3 AI boxes) DUT is shifted right from the center in comparison with picture on slide 5

**Asymmetric geometry** 

# **Track selection**

•  $\chi^2_{
m track}$  < 30 for 6 plane geom.,  $\chi^2_{
m track}$  < 10 for 4 and 2 plane geometries

• track slope < 2 mrad

$$ullet$$
 distance =  $\sqrt{(x_{DUT}-x_{pred})^2+(y_{DUT}-y_{pred})^2}$  < 200  $\mu$ m

#### **Efficiencies**

Energy	2-pl.symm	4-pl.symm.	6-pl.symm.	2-pl.asymm	4-pl.asymm.	6-pl.asymm.
1 GeV	81%	36%	27%	79%	28%	21%
2 GeV	97%	77%	74%	97%	71%	67%
3 GeV	99%	89%	90%	99%	86%	87%
4 GeV	99%	93%	95%	99%	92%	94%
5 GeV	100%	95%	97%	100%	94%	97%
6 GeV	100%	96%	98%	100%	96%	98%

### **Comparison of different geometries**

For 2- and 4-plane geometries the closest to DUT planes are considered



At low energies contribution of MS from telescope planes is big  $\implies$  2-plane geometries show better results

With increasing energy 4-plane geometry is an optimal variant

Asymmetric geometry gives worse results due to bigger MS from Aluminium

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#### **Comparison with close setup**

Standard setup: symmetric setup with closer planes 10 mm from DUT Close setup: symmetric setup with closer planes 5 mm from DUT



Telescope planes should be put as close as possible to the DUT

### **Comparison with high resolution setup**

Standard setup: all telescope planes have 3  $\mu$ m intrinsic resolution High resolution setup: two telescope planes closer to DUT have 1.5  $\mu$ m intrinsic resolution (high density sensors); all other planes - 3  $\mu$ m

#### Symmetric geometry

**Asymmetric geometry** 



#### In both cases the configurations with high density sensors for closest telescope planes give the best results

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### Simulation of pion beam 100 GeV

#### Assumed telescope plane resolution 3 $\mu$ m



#### Symmetric geometry

Asymmetric geometry

#### With increasing energy MS effects become negligible

 $\implies$  6-plane geometry is better

#### The results for symmetric and asymmetric geometries are similar

# Alignment package Millepede

- When detector is ready a proper software alignment will be an important issue for telescope precision
- $\bullet \Longrightarrow$  Test alignment procedures with simulated data
- Alignment package Millepede is developed by Volker Blobel (Uni Hamburg)
- Used in H1, ZEUS, CMS for tracker alignment
- Aligns all planes simultaneously
- Based on linear least squares fits
- Local parameters: track parameters (here track slopes and curvatures)
- Global parameters: alignment coefficients (here x and y shifts)
- Simulated 50000 events (6 GeV electron beam) for 6-plane symmetric telescope configuration without DUT

### First attempt to find x and y shifts using Millepede



#### Should investigate more and play around with constraints, etc.

## **Conclusions**

- Simulation of future pixel beam telescope is done using ILC software
- In general all geometries show good performance
- For high energy beams 6-plane geometry gives the best results
- For use in both high energy and low energy beams an optimal variant would be 4-plane telescope

#### Outlook

- Track fit taking into account multiple scattering (Kalman filter fit)
- Make simulation with magnetic field
- Implement alignment for plane rotations
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