Searching for axions and ALPs from string theory.

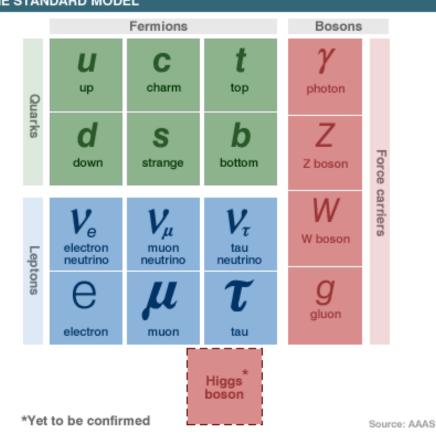
Andreas Ringwald (DESY)

PASCOS 2012, Merida, Mexico, 3-8 June 2012





Standard Model of Particle Physics: Fundamental description of known matter particles and gauge forces

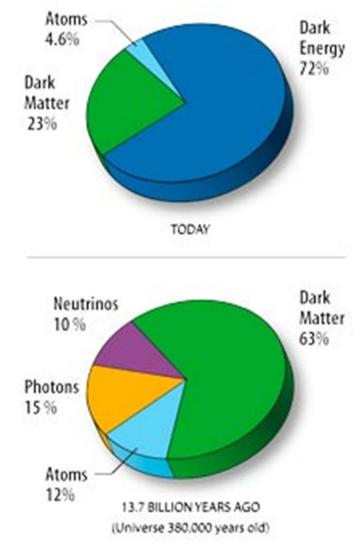






Particles beyond the Standard Model?!

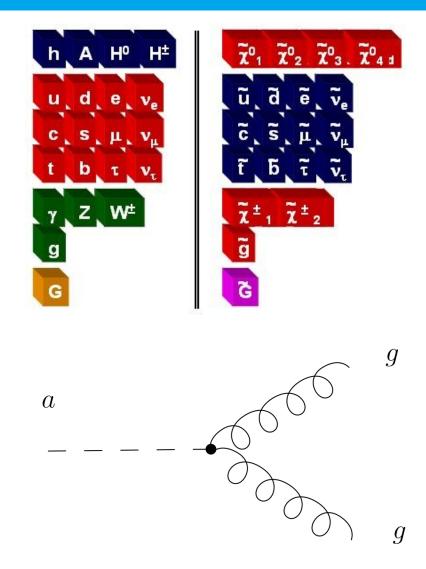
- Standard Model of Particle Physics: Fundamental description of known matter particles and gauge forces
- Standard Model of Cosmology: only about 5 % of energy content of present universe consists of known particles





Particles beyond the Standard Model?!

- Standard Model of Particle Physics: Fundamental description of known matter particles and gauge forces
- Standard Model of Cosmology: only about 5 % of energy content of present universe consists of known particles
- Extensions of the Standard Model of Particle Physics: several good motivated candidates for constituents of dark matter
 - SUSY: Neutralino, Gravitino
 - Peccei Quinn: Axion





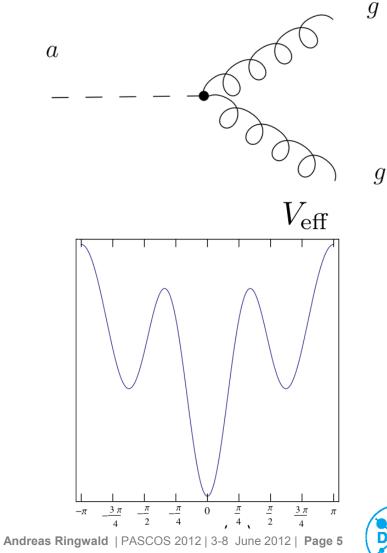
Motivation:

Explanation of unnatural smallness, $\theta < 10^{-10}$, of CP-violating topological term in QCD Lagrangian.

$$\mathcal{L}_{\rm CP-viol.} = \frac{\alpha_s}{4\pi} \theta \operatorname{tr} G_{\mu\nu} \tilde{G}^{\mu\nu}$$

> Axion field
$$\theta
ightarrow a(x)/f_a$$

•
$$\langle a \rangle = 0$$





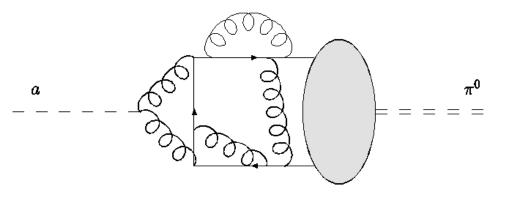
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- > Axion field $\theta \rightarrow a(x)/f_a$
 - $\langle a \rangle = 0$
 - Axion: ultralight particle, cf. Weinberg `78; Wilczek `78

$$m_a = \frac{m_\pi f_\pi}{f_a} \frac{\sqrt{m_u m_d}}{m_u + m_d} \simeq 6 \text{ meV} \times \left(\frac{10^9 \text{ GeV}}{f_a}\right)$$



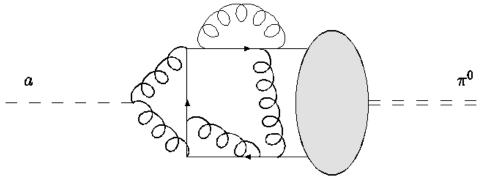


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 - extremely weak interactions with Standard Model particles

$$\mathcal{L}_{a\gamma\gamma} = -\frac{1}{4} g_{a\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu} = g_{a\gamma} a \vec{E} \cdot \vec{B},$$
$$g_{a\gamma} \simeq \frac{\alpha}{2\pi f_a} \sim 10^{-12} \text{ GeV}^{-1} \left(\frac{10^9 \text{ GeV}}{f_a}\right)$$



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 - extremely weak interactions with Standard Model particles
- > Welcome side effect:
 - Axion: candidate for dark matter: created non-thermally via misalignment mechanism in form of coherent oscillations of axion field

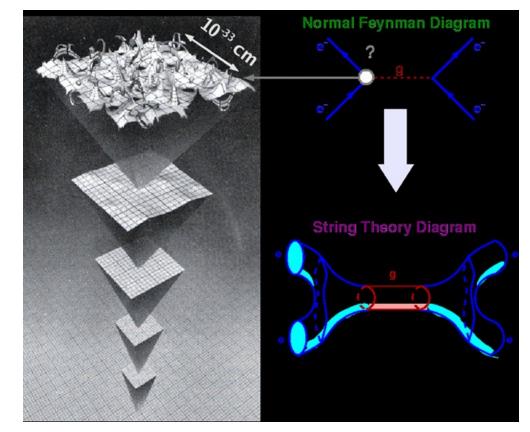
$$\Omega_{a}h^{2}\approx 0.7\left(\frac{f_{a}}{10^{12}~{\rm GeV}}\right)^{7/6}\left(\frac{\theta_{i}}{\pi}\right)^{2}$$



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Coexistence: SUSY and PQ extension in string theory

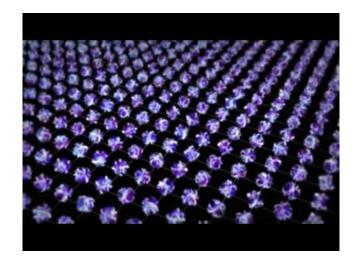
- Particularly strongly motivated extensions of Standard Model based on string theory:
 - Unification of all forces
 - Quantum gravity

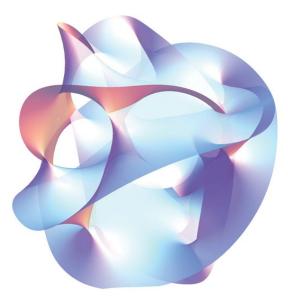




Coexistence: SUSY and PQ extension in string theory

- Particularly strongly motivated extensions of Standard Model based on string theory:
 - Unification of all forces
 - Quantum gravity
- Spectrum of low-energy effective theory in (3+1)-dimensions is supersymmetric and possibly contains several kinds of very weakly interacting slim particles (WISPs): Axion, ALPs (Axion-Like Particles
 - if the compact space comprised of the 6 extra dimensions has certain geometrical and topological properties (Calabi-Yau; several cycles)







String compactifications generically contain pseudo-scalar fields with axionic coupling to gauge fields and anomalous global shift symmetry

$$a_i F \tilde{F} \qquad a_i \to a_i + \epsilon$$

These axion and axion-like particle (ALP) candidates arise in string compactifications as KK zero modes of antisymmetric tensor fields: cf. Witten `84

heterotic string : B_2 IIB string : C_2, C_4

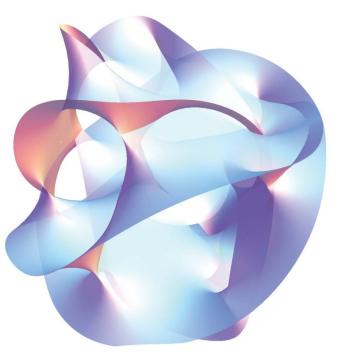


 Concentrate on IIB case (moduli stabilisation best understood): Realisation of brane-world scenarios in string theory

KK reduction (expansion in harmonic forms):

$$C_{2} = c^{a}(x)\omega_{a}, \ a = 1, ..., h_{-}^{1,1}$$

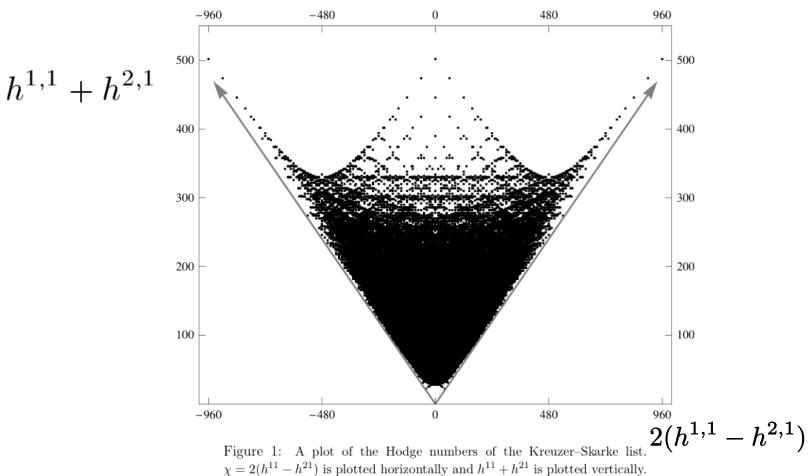
$$C_{4} = c_{\alpha}(x)\tilde{\omega}^{\alpha} + ..., \ \alpha = 1, ..., h_{+}^{1,1}$$



Number of axionic fields determined by topology of CY orientifold: number of topologically non-equivalent 2-cycles or 4-cycles



> Number of cycles generically O(100):

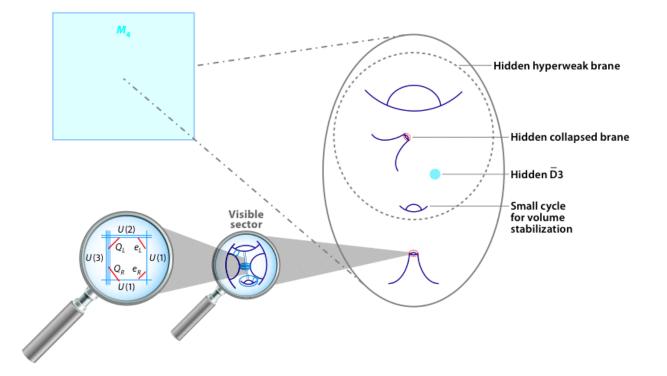


The oblique axes bound the region $h^{11} \ge 0, h^{21} \ge 0$.

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> Cycles can be wrapped by space-time filling D-branes



- Each of these branes gives rise to a gauge theory at low energy
 - Visible sector gauge theory realized by stacks of D7 branes wrapping small 4-cycles

 Hidden sectors, in particular hidden photons, realized by branes wrapping cycles not intersecting visible sector branes Each axionic field comes along with a real scalar field – saxion – which is real part of lowest component of chiral superfield,

$$T_{\alpha} = \tau_{\alpha} + i c_{\alpha}$$

> au_{lpha} ... Kähler modulus measuring the volume of 4-cycle lpha> 4D EFT from KK reduction of D-brane action

$$S_p = \frac{-2\pi}{(2\pi\sqrt{\alpha'})^{p+1}} \left(\int_{\Sigma} d^{p+1}\xi e^{-\phi}\sqrt{\det(g+B+2\pi\alpha' F)} + i \int_{\Sigma} e^{B+2\pi\alpha' F} \wedge \sum_q C_q \right)$$

> T_{α} is gauge kinetic function for theory on D7-brane:

- volume measures gauge coupling, $au_{lpha} \sim g^{-2}$
- c_{lpha} has axionic coupling, $\ \sim c_{lpha} F \wedge F$

L

> 4D effective field theory, cf. Jockers, Louis `05

$$S \supset -dc_{\alpha} \frac{\mathcal{K}_{\alpha\beta}}{8} \wedge \star dc_{\beta} - \frac{r^{i\alpha}\tau_{\alpha}}{4\pi M_{P}} (F_{i} \wedge \star F_{i}) + \frac{r^{i\alpha}c_{\alpha}}{4\pi M_{P}} (F \wedge F),$$

with $\mathcal{K}_{\alpha\beta} \equiv \frac{\partial^{2}K}{\partial\tau_{\alpha}\partial\tau_{\beta}}, K = -2\ln\mathcal{V}, r^{i\alpha} \equiv \ell_{s}^{-4} \int_{D_{i}} \tilde{\omega}^{\alpha}$

Decay constants and coupling to gauge bosons via canonical normalization of axion and gauge kinetic terms and matching to:

$$\begin{split} \mathcal{C} &\supset \frac{1}{2} \,\partial_{\mu} a_{i} \,\partial^{\mu} a_{i} - \frac{g_{3}^{2}}{32\pi^{2}} \left(\theta_{0} + C_{i33} \frac{a_{i}}{f_{a_{i}}}\right) F_{3,\mu\nu}^{b} \tilde{F}_{3}^{b,\mu\nu} \\ &- \frac{g_{2}^{2}}{32\pi^{2}} \,C_{iWW} \frac{a_{i}}{f_{a_{i}}} \,F_{W,\mu\nu}^{b} \tilde{F}_{W}^{b,\mu\nu} - \frac{g_{Y}^{2}}{32\pi^{2}} \,C_{iYY} \frac{a_{i}}{f_{a_{i}}} \,F_{Y,\mu\nu} \tilde{F}_{Y}^{\mu\nu} \end{split}$$



IIB axiverse

- An axiverse QCD axion plus possibly many ultra-light ALPs whose mass spectrum is logarithmically hierarchical – may naturally arise from strings, cf. Arvanitaki et al. `09
- > Challenges to obtain an axiverse:
 - Only axions which are not projected out by orientifold projection appear in LE EFT
 - Only axions which do not get too heavy by K\u00e4hler moduli stabilisation can be candidates for QCD axion and other light ALPs, cf. Conlon `06
 - Only axions which are not eaten by Stückelberg mechanism to give masses to branelocalized anomalous U(1) gauge bosons will appear in LE EFT
- > Acharya, Bobkov, Kumar `11: Moduli stabilisation via single non-perturbative correction to superpotential (cf. Bobkov,Braun, Kumar, Raby `10) fixes all Kähler moduli plus one axion combination: axiverse with $W_0 \ll 1$ and $f_a \sim 10^{16} \,\text{GeV}$
- > Cicoli, Goodsell, AR, 1206.0819: Moduli stabilisation of the so-called LARGE Volume Scenario (LVS) which exploits both perturbative and non-perturbative effects to fix Kähler moduli gives rise to an axiverse with $W_0 \sim 1$ and $f_a \sim 10^{10} \, {\rm GeV}$

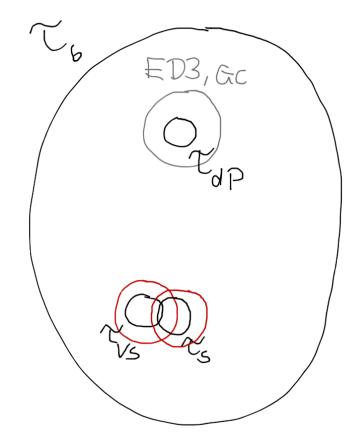
- Strategy to fix moduli in LVS: (cf. Cicoli, Mayrhofer, Valandro `11)
 - Exploit del Pezzo four-cycle supporting single nonperturbative effect; dP modulus fixed at small size, dP saxion and axion heavy; interplay with leading order alpha' correction yields exponentially large CY volume
 - Visible sector with chirality built by wrapping magnetised D7-branes around rigid but not del Pezzo four cycles; D-term conditions stabilise d combinations of Kähler moduli, leaving

 $n_{\rm ax}\equiv h^{1,1}-1-d\geq 1$ flat directions; latter fixed by pert. corrections

- LVS requires $n_{\rm ax} \geq 2$: one of the remaining cycles should be small to obtain correct value of $g_{\rm vs}^2 \sim 1/\tau_{\rm vs}$, while there should be at least one further which can be large; latter fixed at

$$\mathcal{V} \sim \tau_b^{3/2} \sim W_0 \, e^{2\pi \tau_{\rm dP}}$$

More on this: talk by Michele Cicoli

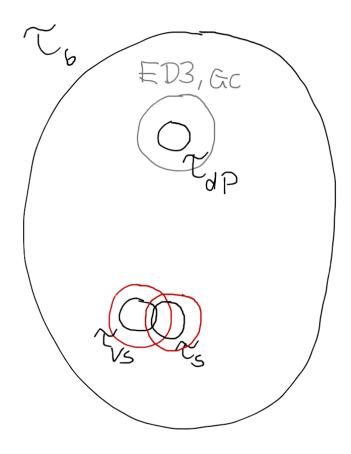




> Mass scales for $g_s \sim 0.1, W_0 \sim 1, \mathcal{V} \sim 10^{14}$:

$$\begin{split} M_s &\sim \frac{M_P}{\sqrt{4\pi\mathcal{V}}} \sim 10^{10} \,\mathrm{GeV} \\ m_{\tau_s} &\sim \frac{M_P}{\mathcal{V}^{1/2}} \sim 10^{10} \,\mathrm{GeV} \\ m_{\tau_{\mathrm{dP}}} &\sim \frac{M_P}{\mathcal{V}} \ln \mathcal{V} \sim 30 \,\mathrm{TeV} \\ m_{3/2} &\sim \sqrt{g_s/(4\pi)} W_0 \frac{M_P}{\mathcal{V}} \sim 1 \,\mathrm{TeV} \\ m_{\tau_{\mathrm{vs}}} &\sim \alpha_{\mathrm{vs}} m_{3/2} \sim 40 \,\mathrm{GeV} \\ m_{\tau_b} &\sim \frac{M_P}{\mathcal{V}^{3/2}} \sim 0.1 \,\mathrm{MeV} \end{split}$$

- No cosmological moduli problem since τ_b diluted by entropy production due to decay of τ_{vs} reheating universe to O(GeV)





> Scaling of axion decay constants and couplings:

$$\begin{split} f_{a_b} &= \frac{\sqrt{3}}{4\pi} \frac{M_P}{\tau_b} \simeq \frac{M_P}{4\pi \mathcal{V}^{2/3}} \simeq \frac{M_{\rm KK}^{10\rm D}}{4\pi} \,, \qquad f_{a_s} = \frac{1}{\sqrt{6} \left(2\tau_s\right)^{1/4}} \frac{M_P}{4\pi \sqrt{\mathcal{V}}} \simeq \frac{M_s}{\sqrt{4\pi}\tau_s^{1/4}} \,, \\ C_{bbb} &\simeq g_b^{-2} \frac{f_{a_b}}{M_P} \simeq \mathcal{O}\left(1\right) \,, \qquad C_{sbb} \simeq g_b^{-2} \frac{f_{a_s}\tau_s^{3/4}}{\mathcal{V}^{1/2}M_P} \simeq \mathcal{O}\left(\epsilon\right) \simeq \mathcal{O}\left(\mathcal{V}^{-1/3}\right) \,, \\ C_{bss} &\simeq g_s^{-2} \frac{f_{a_b}}{M_P} \simeq \mathcal{O}\left(\epsilon^2\right) \simeq \mathcal{O}\left(\mathcal{V}^{-2/3}\right) \,, \qquad C_{sss} \simeq g_s^{-2} \frac{f_{a_s}}{\tau_s^{3/4}M_s} \simeq \mathcal{O}\left(1\right) \,. \\ &= a_{\rm VS} : \text{QCD axion with } f_{a_{\rm VS}} \sim f_{a_s} \sim 10^{10} \, {\rm GeV} \\ &= a_b \,: \text{essentially massless ALP with } f_{a_b} \sim 10^8 \, {\rm GeV} \,, \\ &\text{but nearly decoupled} \,, \, C_{bss} \sim 10^{-10} \, {\rm GeV} \,. \end{split}$$

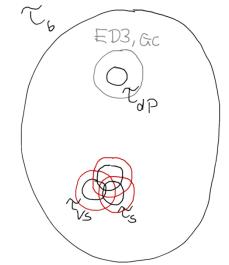


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- $a_{
 m vs}$: QCD axion with $~f_{a_{
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- a_b : essentially massless ALP with $f_{a_b}\sim 10^8\,{
 m GeV}$, but nearly decoupled, $C_{bss}\sim 10^{-10}$
- Possibly more ultralight, $m_{a_b} \ll m_a$, ALPs with

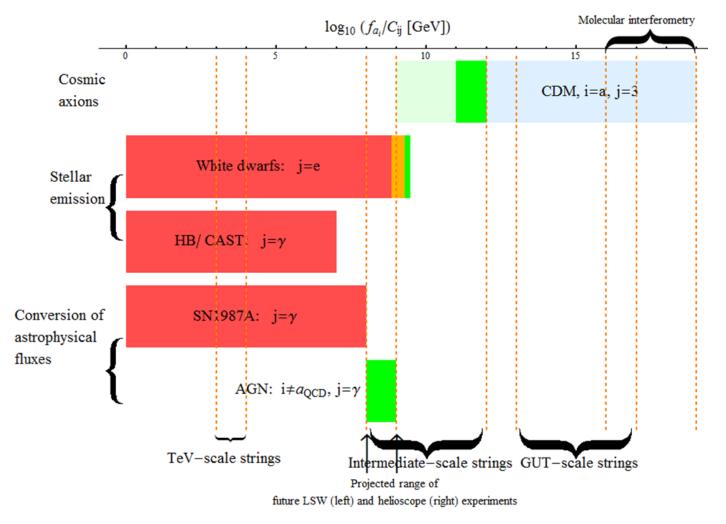
$$f_{a_s} \sim 10^{10} \,\mathrm{GeV}, C_{sss} \sim 1$$





Axion and ALPs with intermediate scale decay constant?

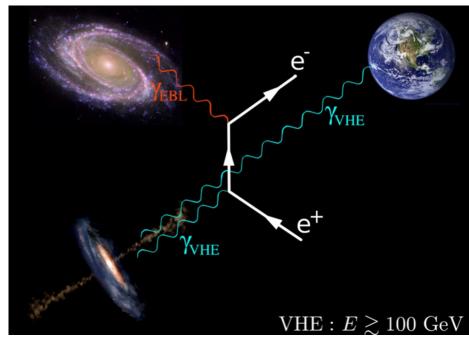
Current limits and possible hints from astrophysics and cosmology:







TeV photons of distant Active Galactic Nuclei (AGN) should feature absorption breaks due to electronpositron pairproduction on the Extragalactic Background Light (EBL)





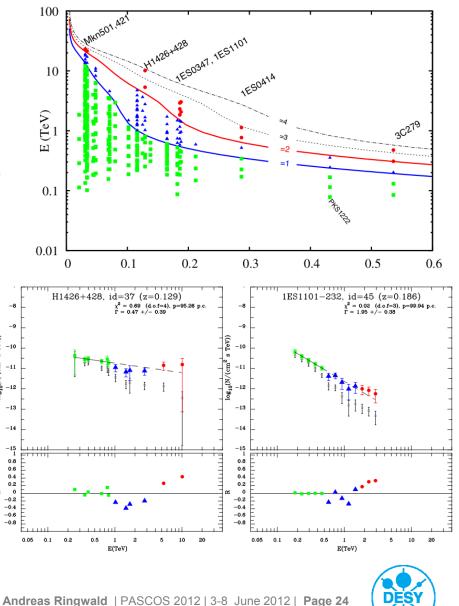
TeV))

 $\log_{10}(N/(cm^2$

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> Not the case: (e.g. Horns, Meyer `12)

 50 spectra (HESS, MAGIC, Veritas), assumption: minimal EBL; absorption ruled out by more than 4 sigma

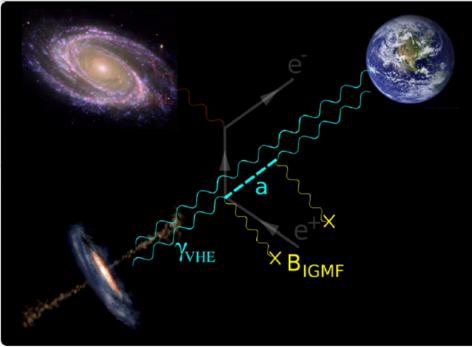


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Possible explanation: Photon <-> ALP conversion in astrophysical magnetic fields (Roncadelli et al.,Sanchez-Conde et al.)



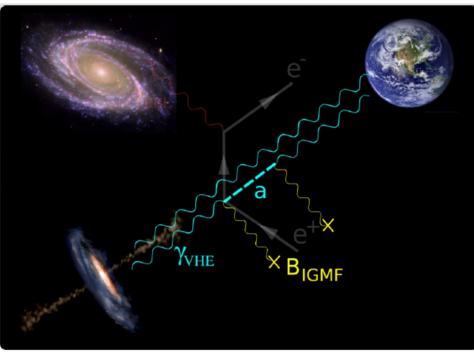
 $P(a \leftrightarrow \gamma) = 4 \frac{(g_{a\gamma} \omega B)^2}{m^4} \sin^4$



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- Possible explanation: Photon <-> ALP conversion in astrophysical magnetic fields (Roncadelli et al.,Sanchez-Conde et al.)
- Has to be an ALP: too light for a QCD axion with such a decay constant



$$g_{i\gamma} \equiv \frac{\alpha C_{i\gamma}}{2\pi f_{a_i}} = 10^{-12} \div 10^{-11} \,\text{GeV}^{-1}$$
$$\rightarrow \frac{f_{a_i}}{C_{i\gamma}} \simeq 10^8 \div 10^9 \,\text{GeV}$$
for $m_{\gamma} \leq \text{neV}$

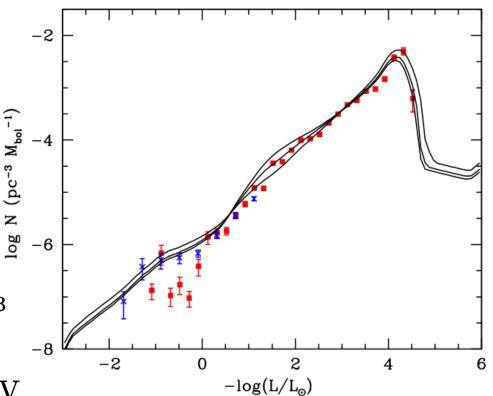


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Hint for axion/ALP-production in white dwarfs?

- Non-standard energie-loss mechanism in white dwarfs, cf. Isern et al.
- Compatible with axion or ALP production in electronbremsstrahlung

$$g_{ie} \equiv \frac{C_{ie}m_e}{f_{a_i}} = (2.0 \div 7.0) \times 10^{-13}$$
$$\rightarrow \frac{f_{a_i}}{C_{i,e}} \simeq (0.7 \div 2.6) \times 10^9 \text{ GeV}$$
for $m_{a_i} \lesssim \text{keV}$





Can we explain these hints within IIB axiverse?

Anomalous transparency of universe and anomalous energy loss of white dwarfs could be explained by

 $C_{i\gamma}/C_{ie} \simeq 10, \qquad f_{a_i}/C_{i\gamma} \simeq 10^8 \text{ GeV}, \qquad m_{\text{ALP}} \lesssim 10^{-9} \div 10^{-10} \text{ eV}.$

Model where visible sector build from intersecting branes in geometric regime

$$\frac{C_{i\gamma}}{C_{ie}} \sim \frac{8\pi\tau_*}{3}, \qquad \frac{f_{a_i}}{C_{i\gamma}} = \frac{1}{8\pi N_{i\gamma}\tau_*^{1/4}} \frac{M_P}{\sqrt{\mathcal{V}}} = \frac{1}{8\pi N_{i\gamma}\tau_*^{1/4}} \sqrt{\frac{g_s M_P m_{3/2}}{W_0}}.$$

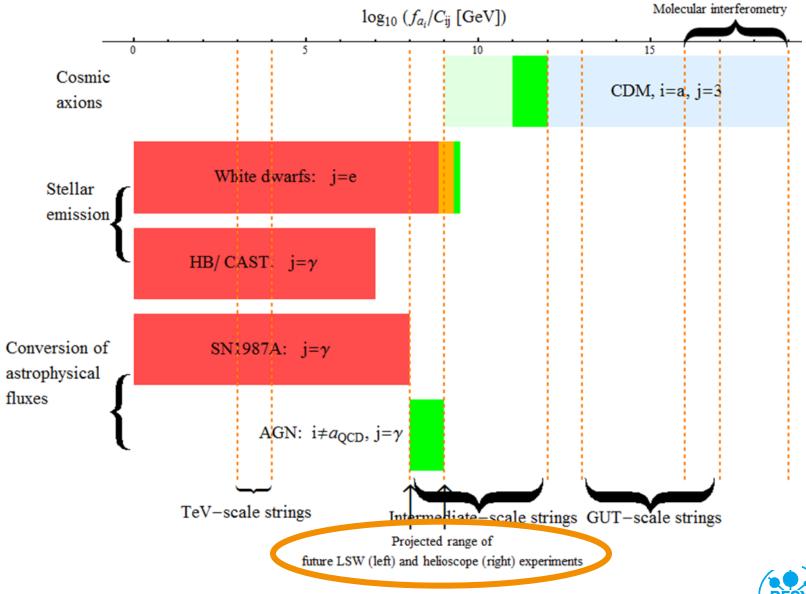
- > Yields required values for $m_{3/2} = 10$ TeV, $g_s \simeq 0.1$ and $W_0 \sim 10^{-1}$
- > ALP mass could be generated by single Kähler potential instanton, e.g. $m_{\rm ALP} \sim m_{3/2} e^{-\pi \tau_*} \sim 10^{-10} \, {\rm eV}$ requires

$$\tau_* \sim \frac{1}{\pi} \ln \left(\frac{g_s m_{3/2}}{m_{\rm ALP}} \right) \sim 16.$$

> Astrophyical hints compatible with intermediate string-scale scenario

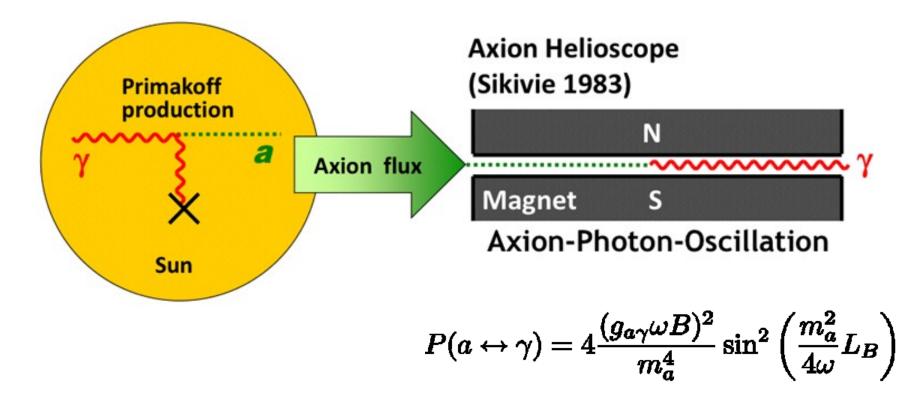
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Laboratory probes of axion and ALPs with intermediate scale decay constant?



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- Sun strong source of axions and ALPs
- > Helioscope searches for axions/ALPs and HPs





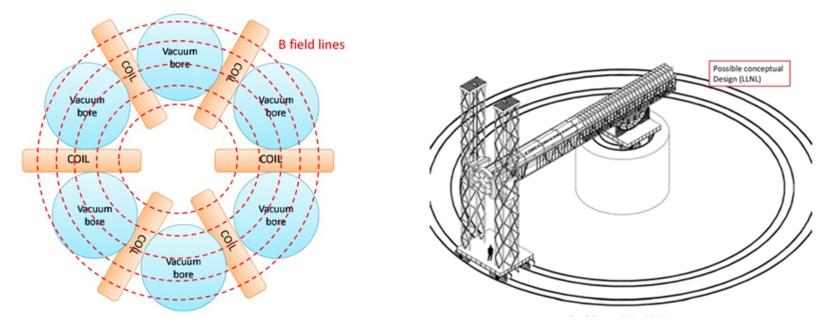
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CAST ... CERN Axion Solar Telescope

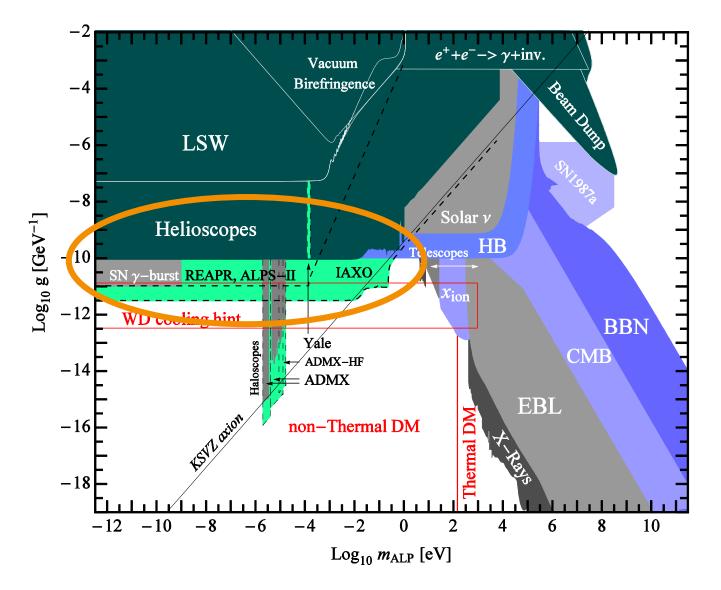


- Sun strong source of axions and ALPs
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- CAST ... CERN Axion Solar Telescope
- IAXO ... International Axion Observatory (under investigation)







> ALPs can pass walls

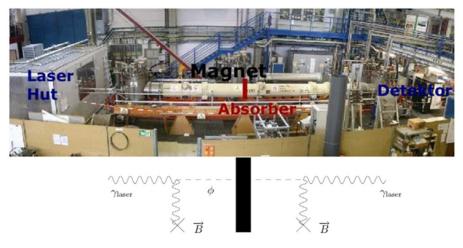
Light-shining-through-walls experiments: (here ALPS (@DESY)):

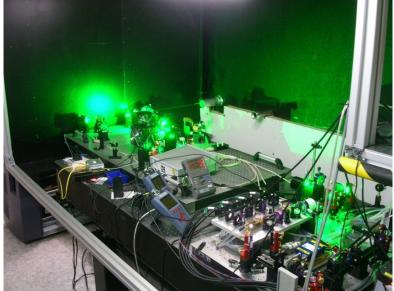


$$P(a\leftrightarrow\gamma)=4rac{(g_{a\gamma}\omega B)^2}{m_a^4}\sin^2\left(rac{m_a^2}{4\omega}L_B
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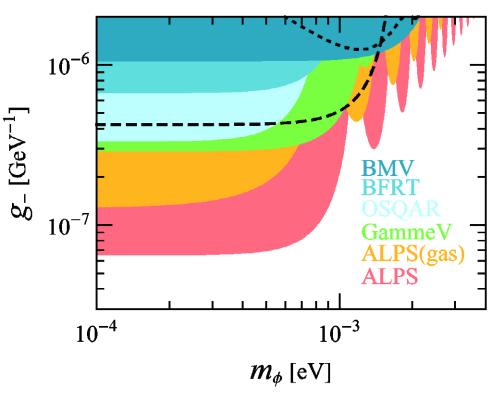
- HERA dipole (8.4 m, 5 T)
- Primary laser: enhanced LIGO laser (1064 nm, 35 W)
- Frequency doubled: 523 nm
- 300-fold power build-up in cavity





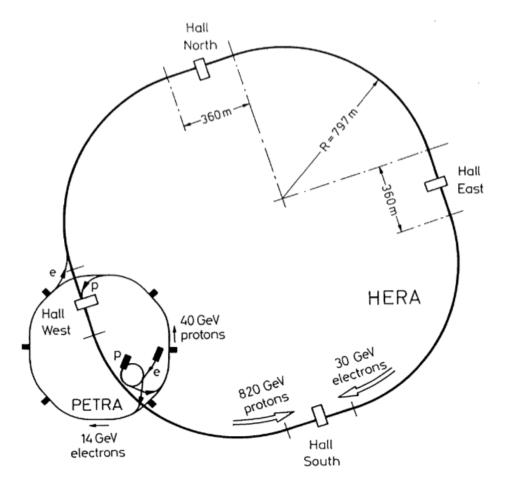


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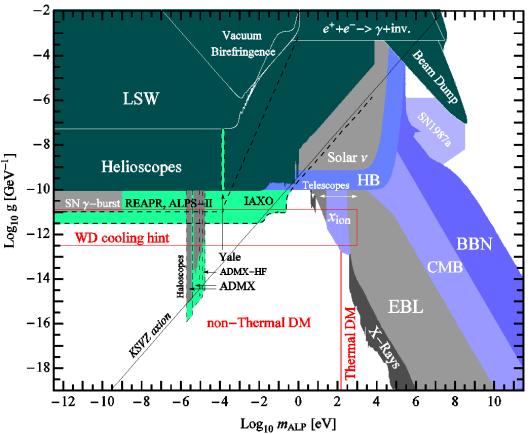


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- > ALPS-II plans (2016+):
 - 12 + 12 HERA dipoles
 - Increased power build-up (~5000)
 - Cavity also on regeneration part





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- Similar plans also at Fermilab (REAPR)

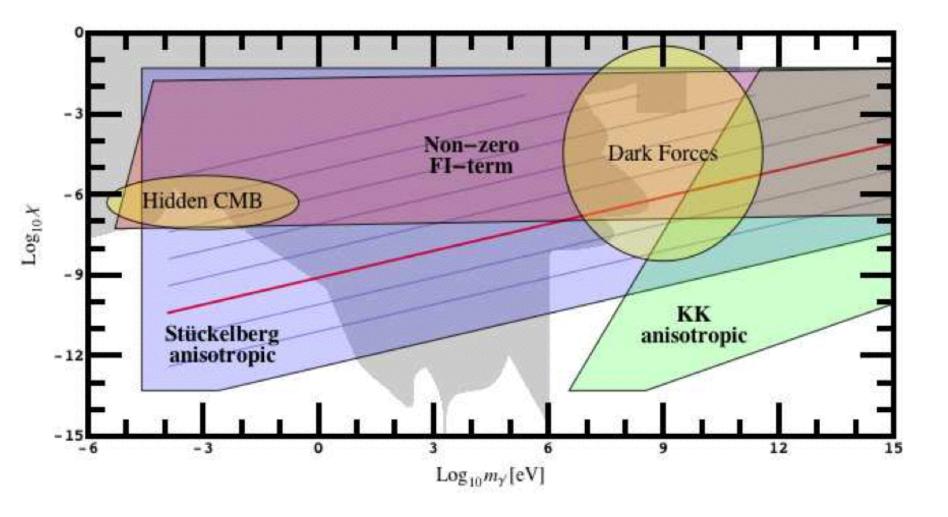




- String phenomenology holds the promise of an axiverse the QCD axion plus a (possibly large) number of further ultralight axion-like particles, possibly populating each decade of mass down to the Hubble scale
- Promise fulfilled in LARGE Volume Scenario of IIB string compactifications
- Models that exhibit a QCD axion with an intermediate-scale decay constant and additional even lighter axion-like particles having the same decay constant and coupling to the photon can explain astrophysical anomalies and be tested in the next generation of helioscopes and light-shining-through-walls experiments
- Cosmology of LVS axiverse still to be investigated in detail



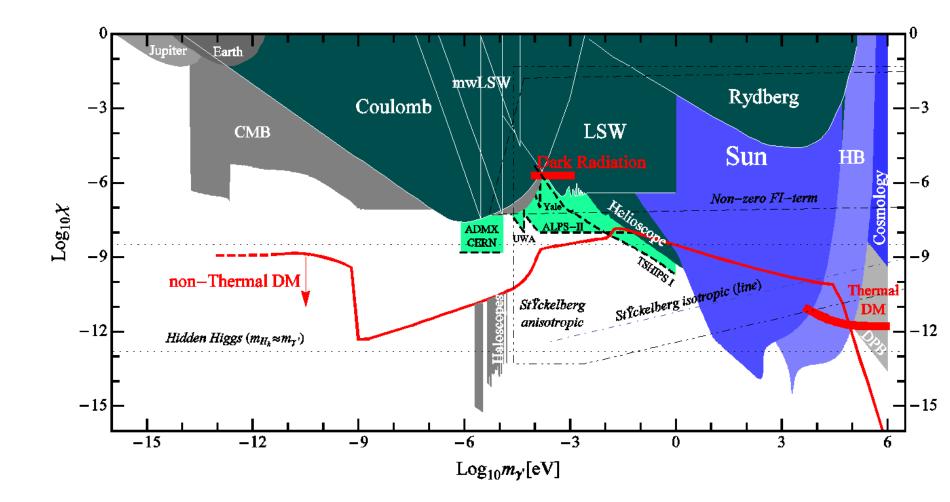
Backup: Hidden photons



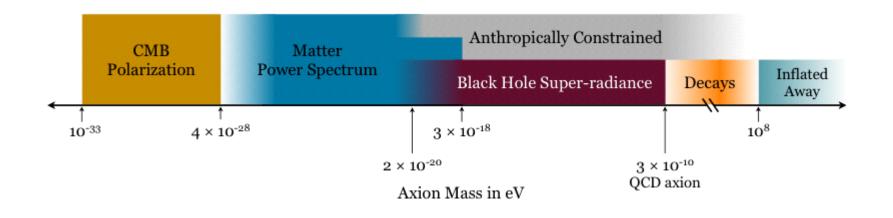


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Backup: Hidden photons







cf. Arvanitaki, Dimopoulos, Dubovsky, Kaloper, March-Russell `09

