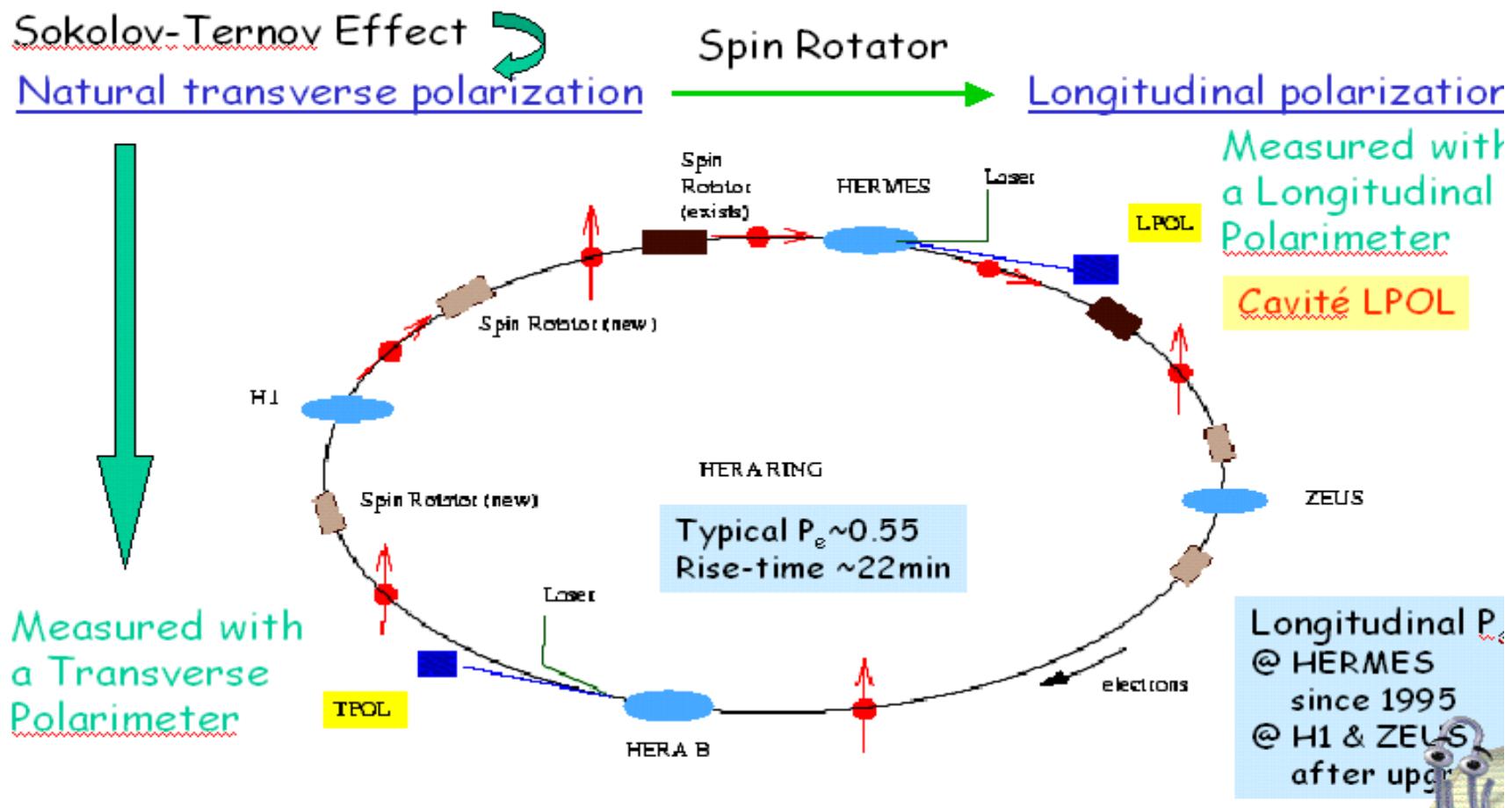


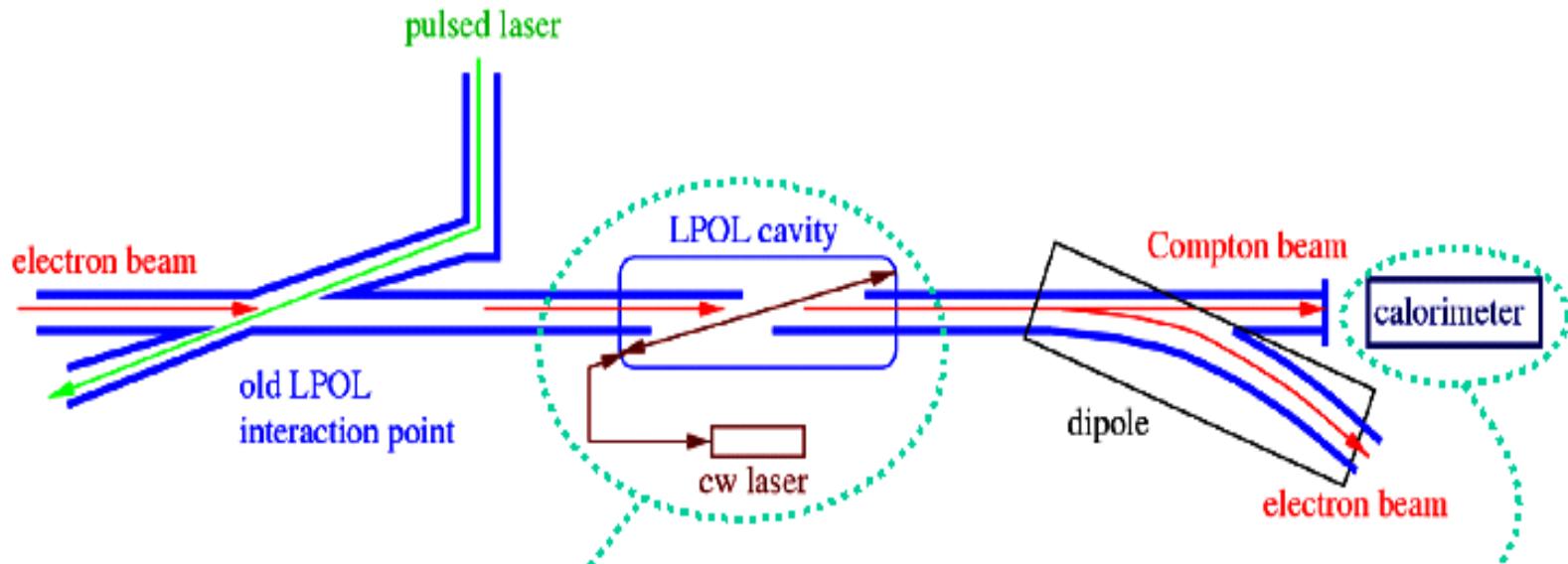
Polca

- Apparatus
- Principle
- Potential for comparison with Tpol Lpol
- Analysis
- Future prospect

Hera and the three polarimeters



Polca setup



→ Two key elements:

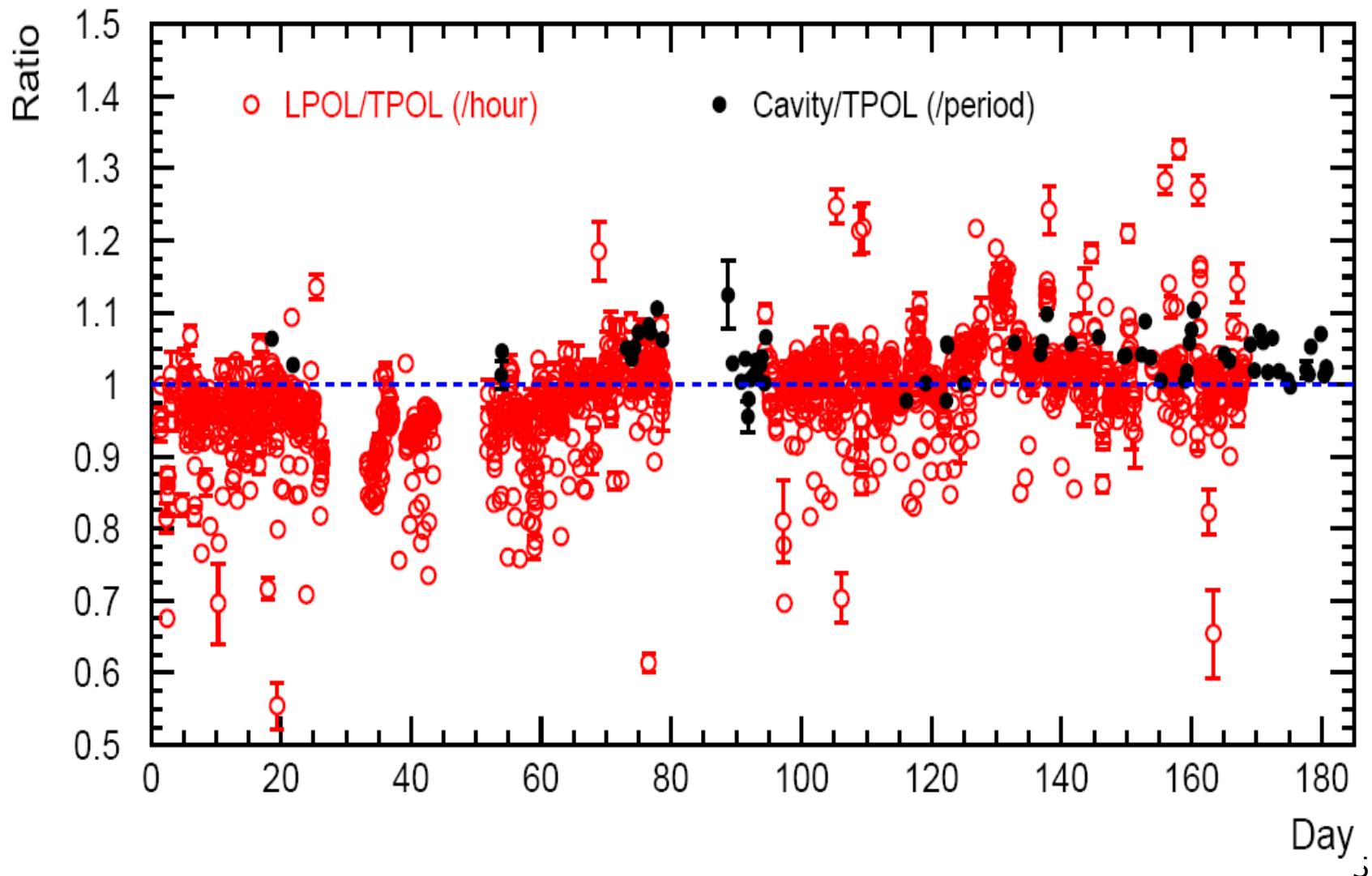
Fabry-Pérot cavity (to increase the laser power by $>10^3$)

Sandwich or fiber calorimeter (to measure E_{γ} of the scattered photon)

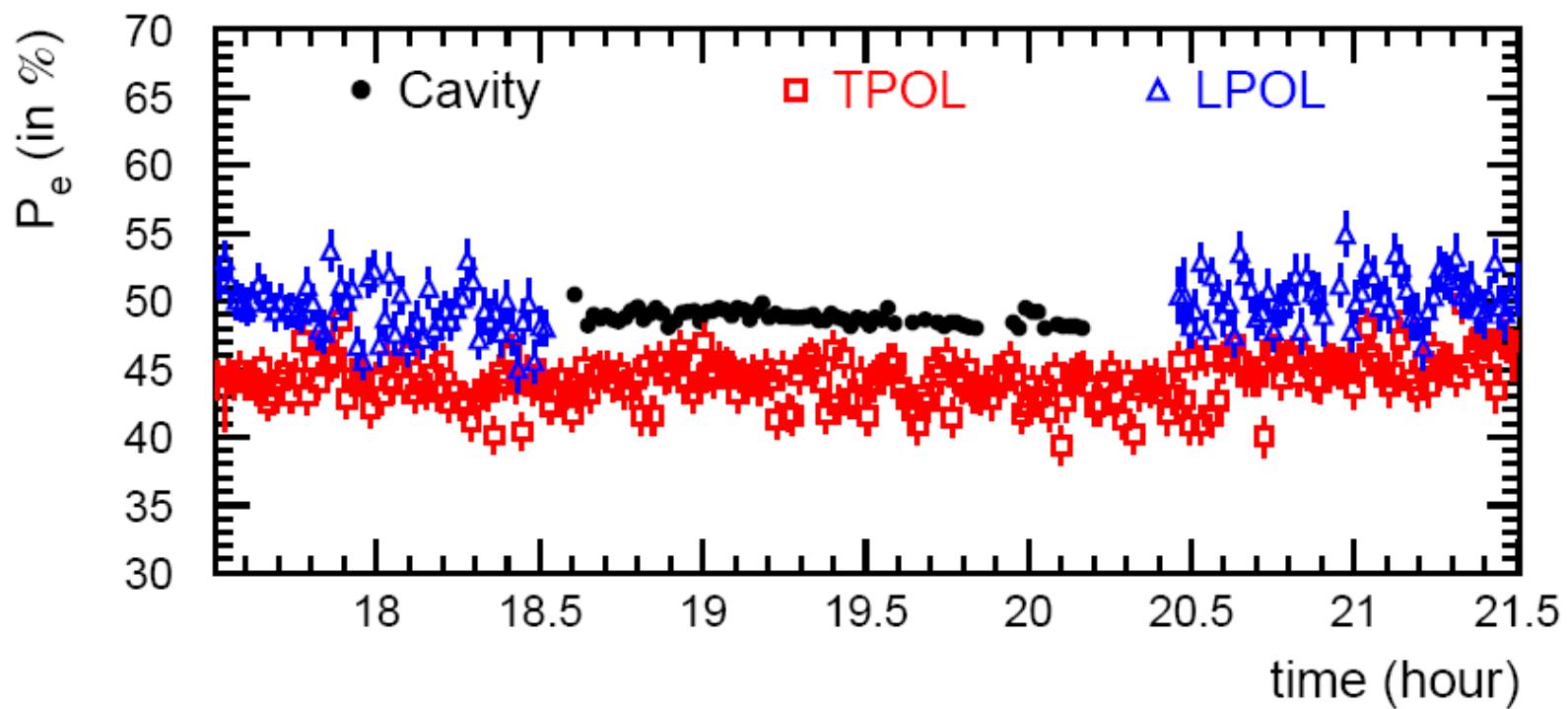
LPOL Cavity statistics

- Efficient running during about 300 hours (from 6 oct 2006 to end)
- Calo energy : 7 Tera measurements
- Brems : 2 Tera
- Comptons : 900 Giga (~ energy stored in one Hera bunch)
- Left and Right laser polarization differ by 60 Giga
- Systematics dominated

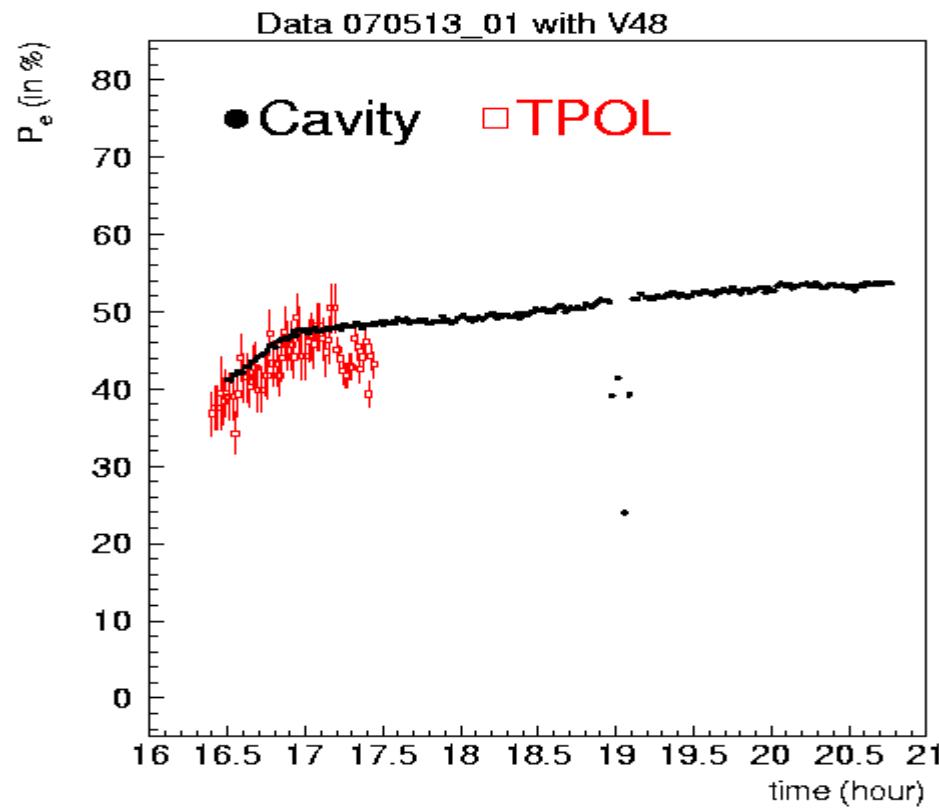
Comparison with Tpol Lpol

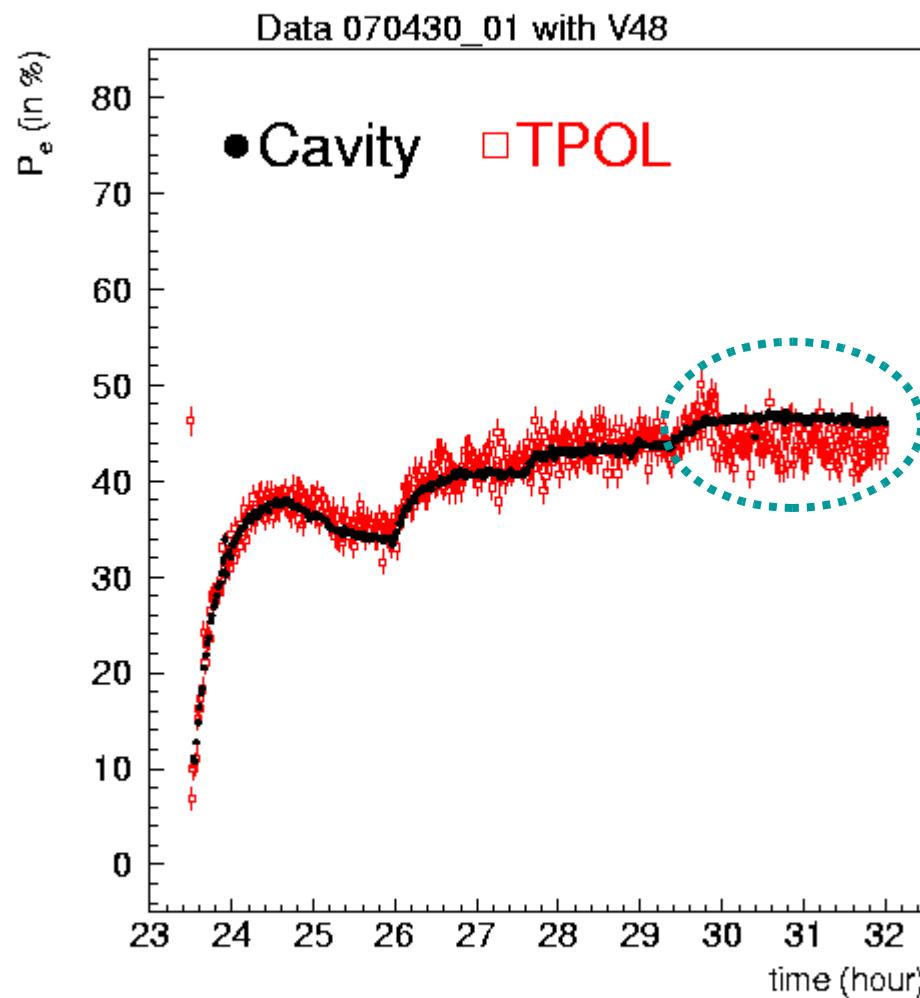


Data of 070518_02

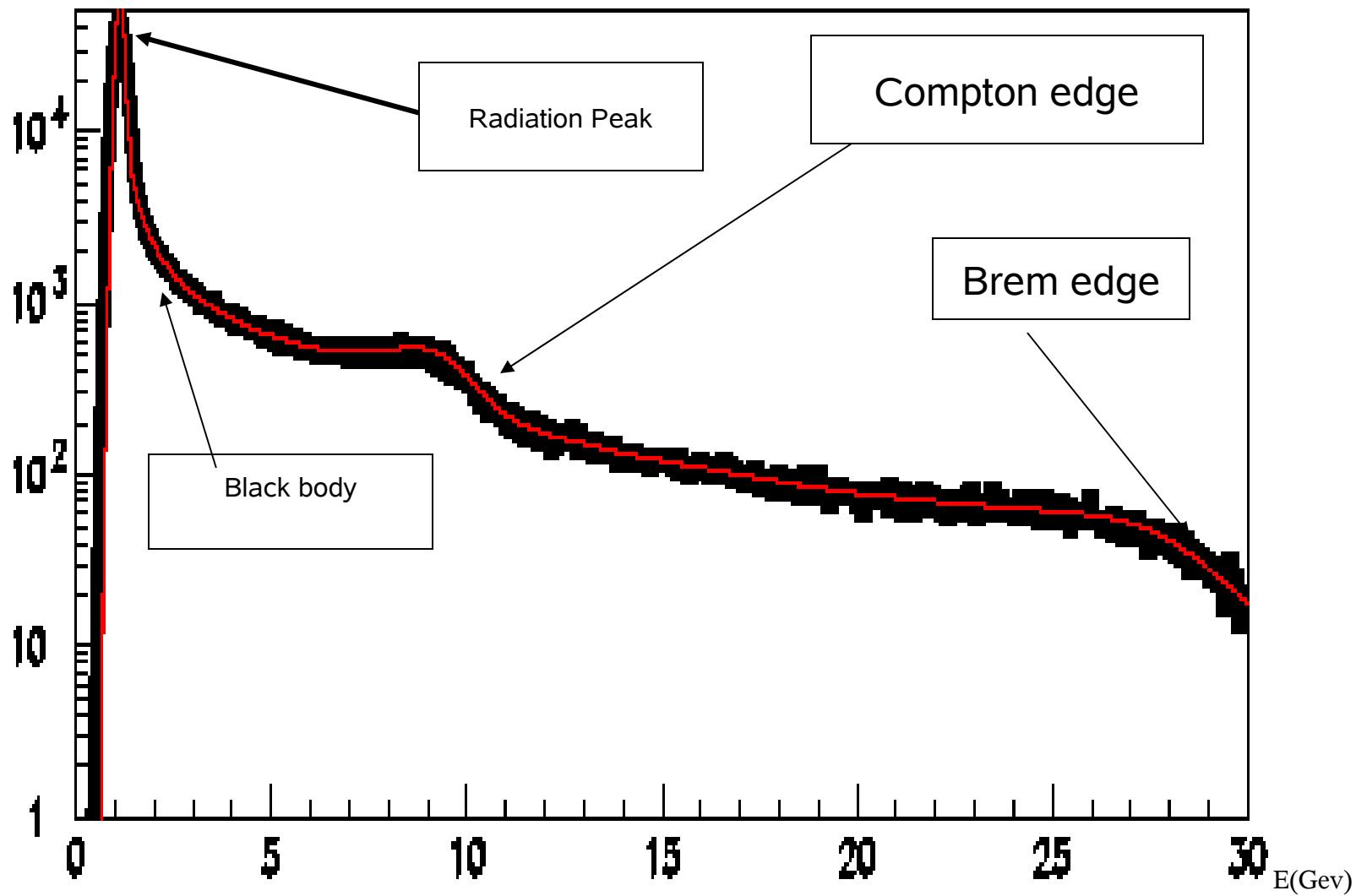


One good example to show only cavity measurements are available

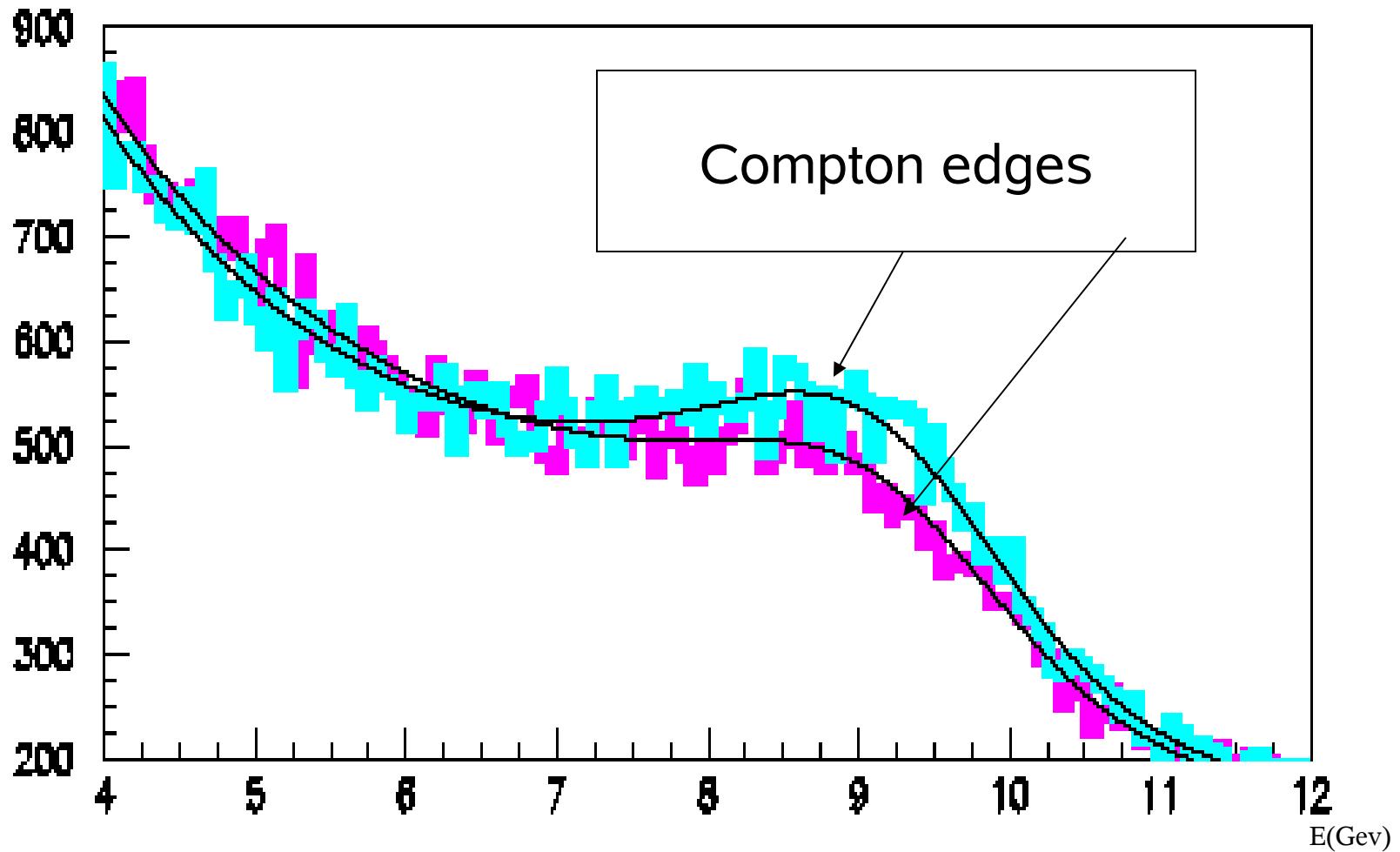




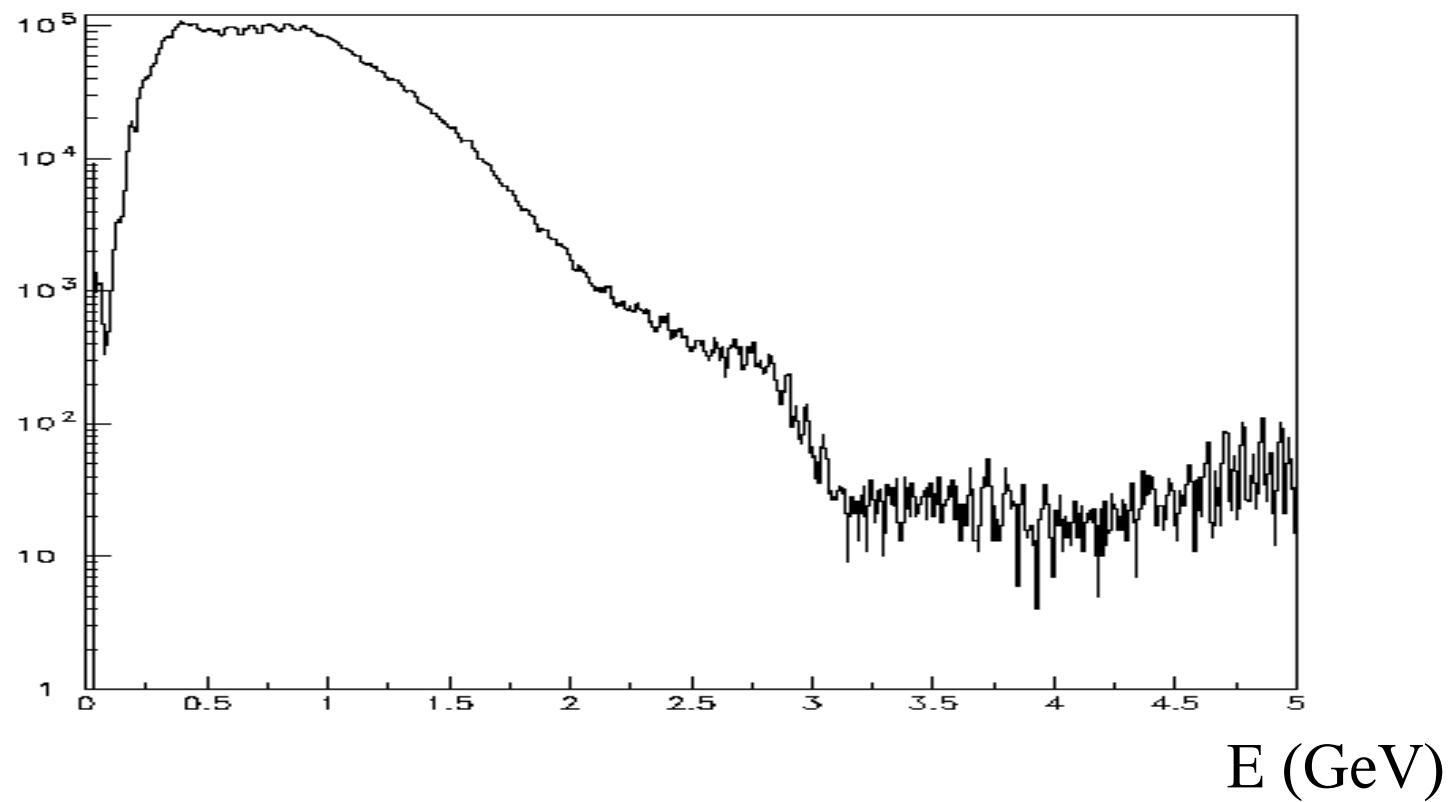
Measurement example



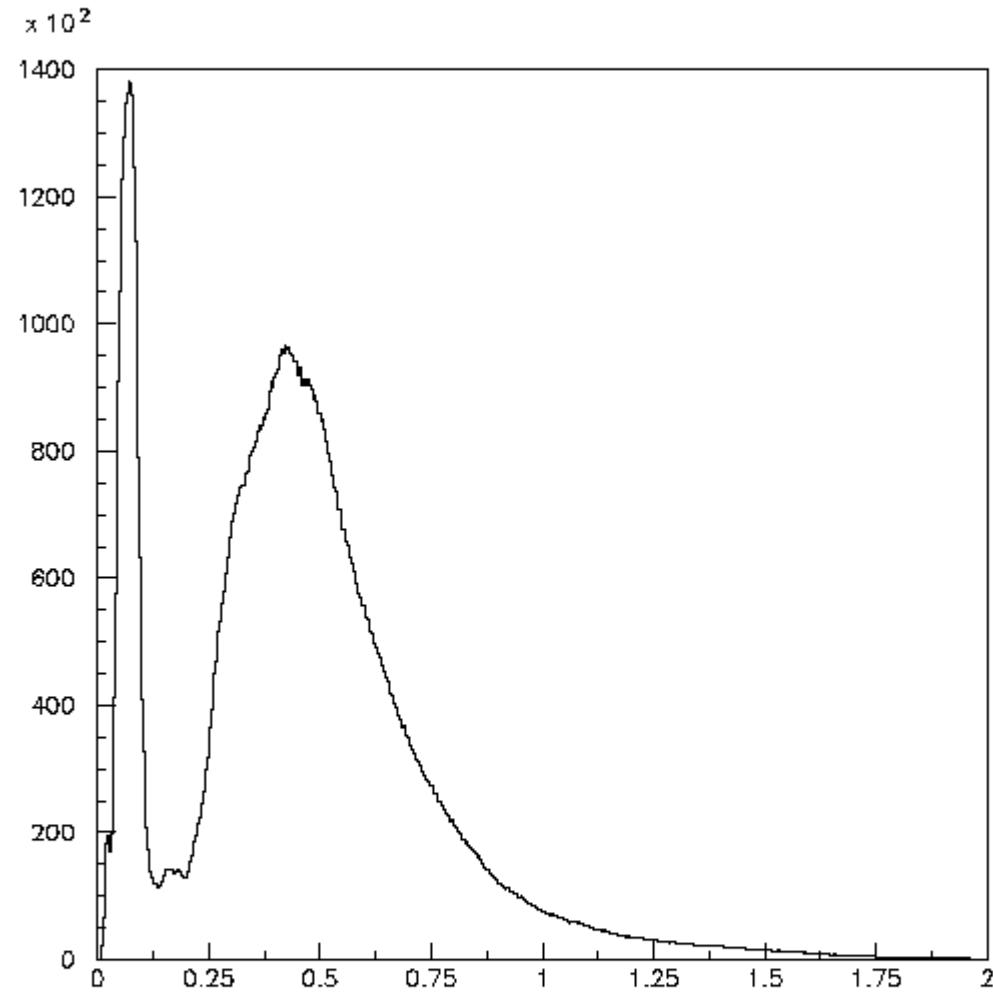
Measurement example



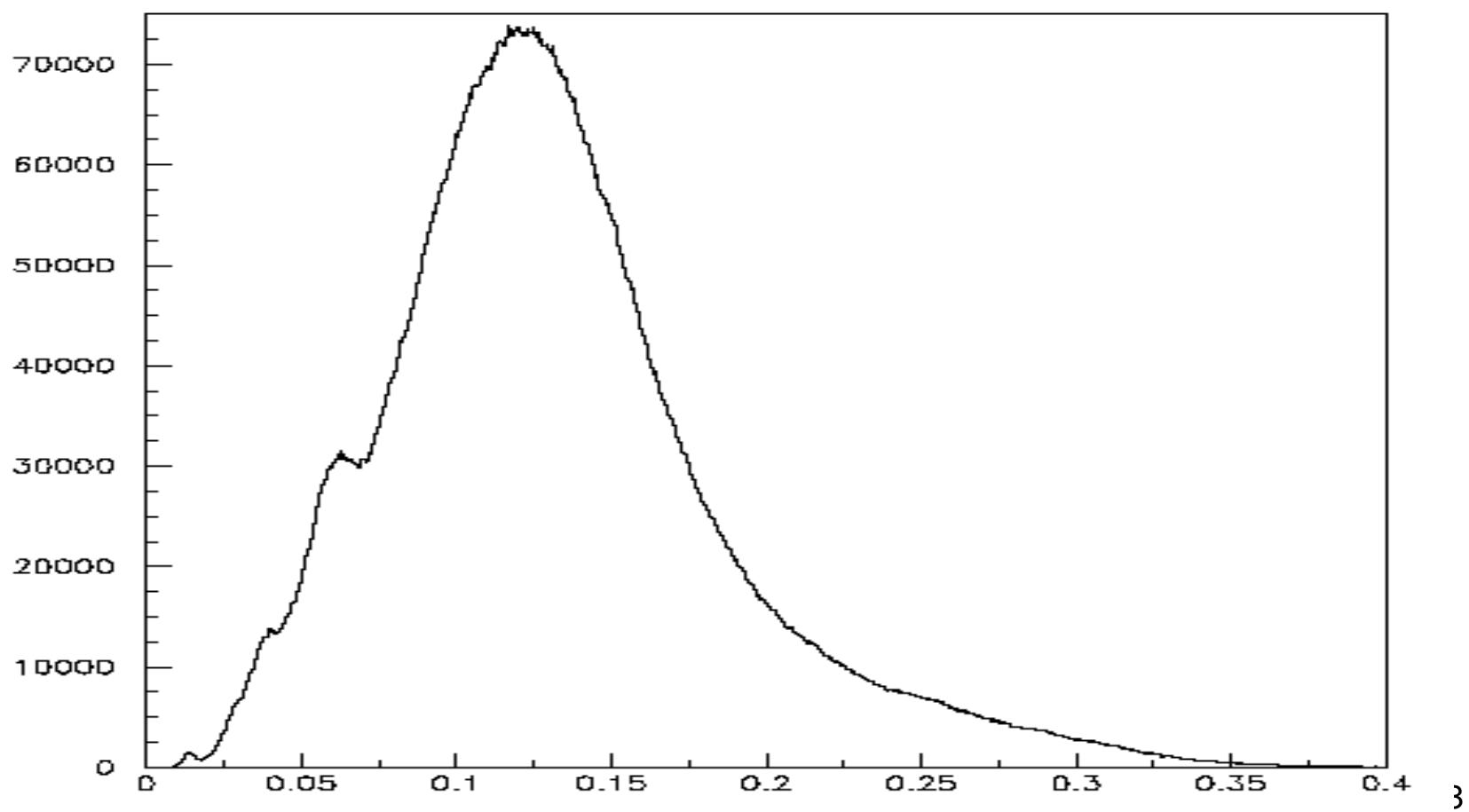
Rads peak energy



Distribution of number of brem



Distribution of number of Compton



Analysis method

- Simulation

Cross section: (Brem,Compton,Black) or Energy (synchrotron peak)

Calo simulation: Rad peak $\sigma_r^2 = aE$ photon $\sigma_\gamma^2 = aE+bE^2$

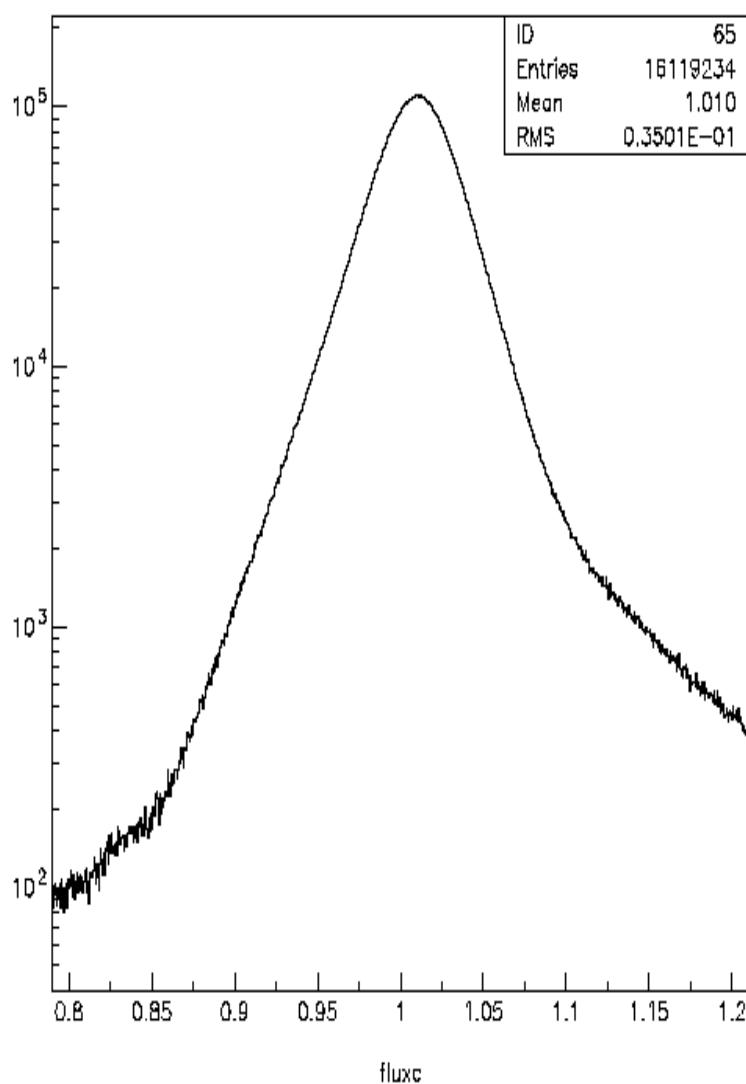
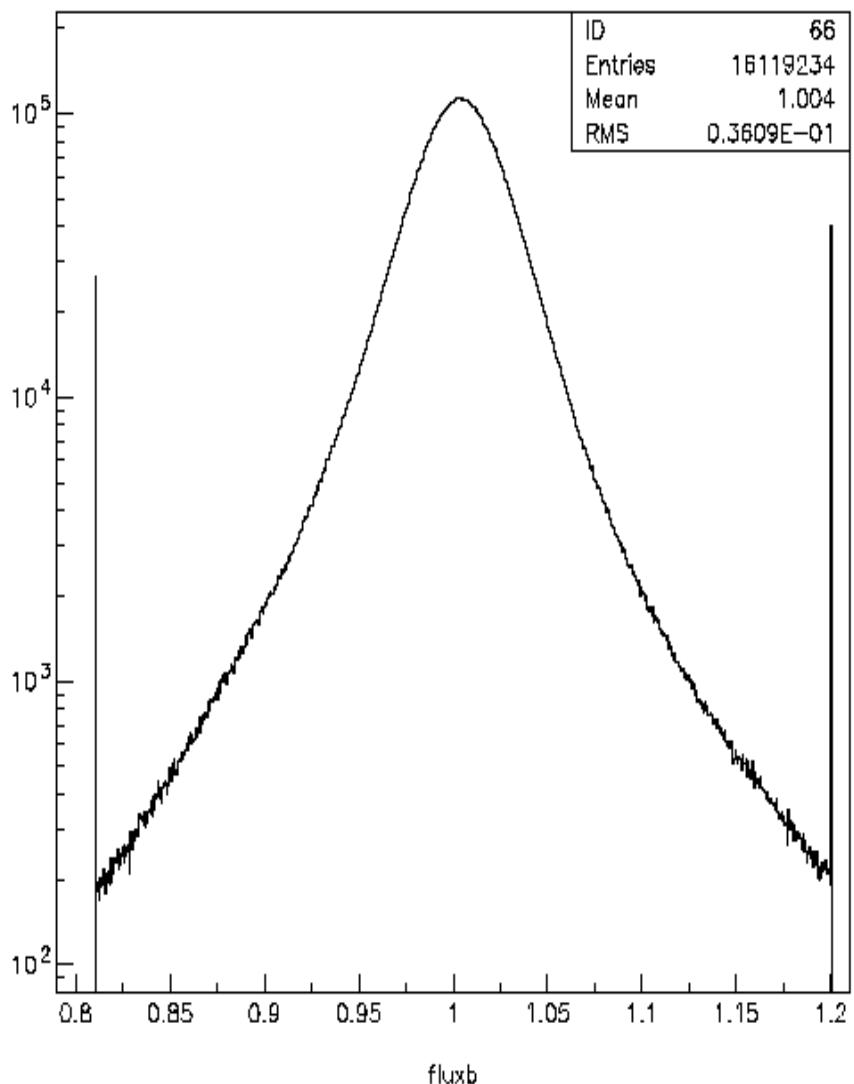
$E=\alpha x$ x following a χ^2 distribution

E to ADC conversion: five alternate models such as $A=cE(1+d/E+E*s)$

Electronic noise: Gaussian

- Likelihood to compare simulation and histogram
Obtain Erads Brem Black Compton polar (and fluxes)
- Calo optimization

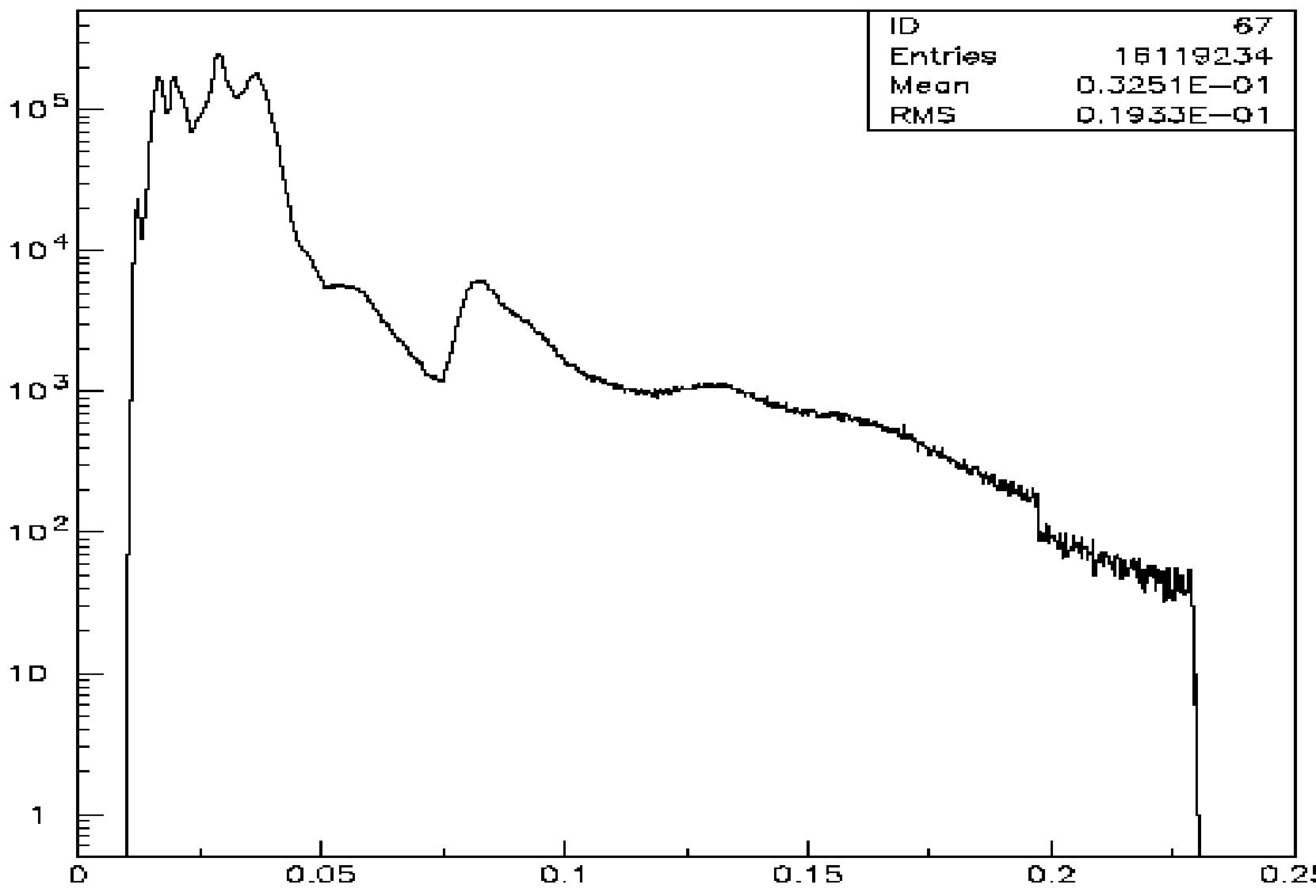
Fluxes for the second histoset



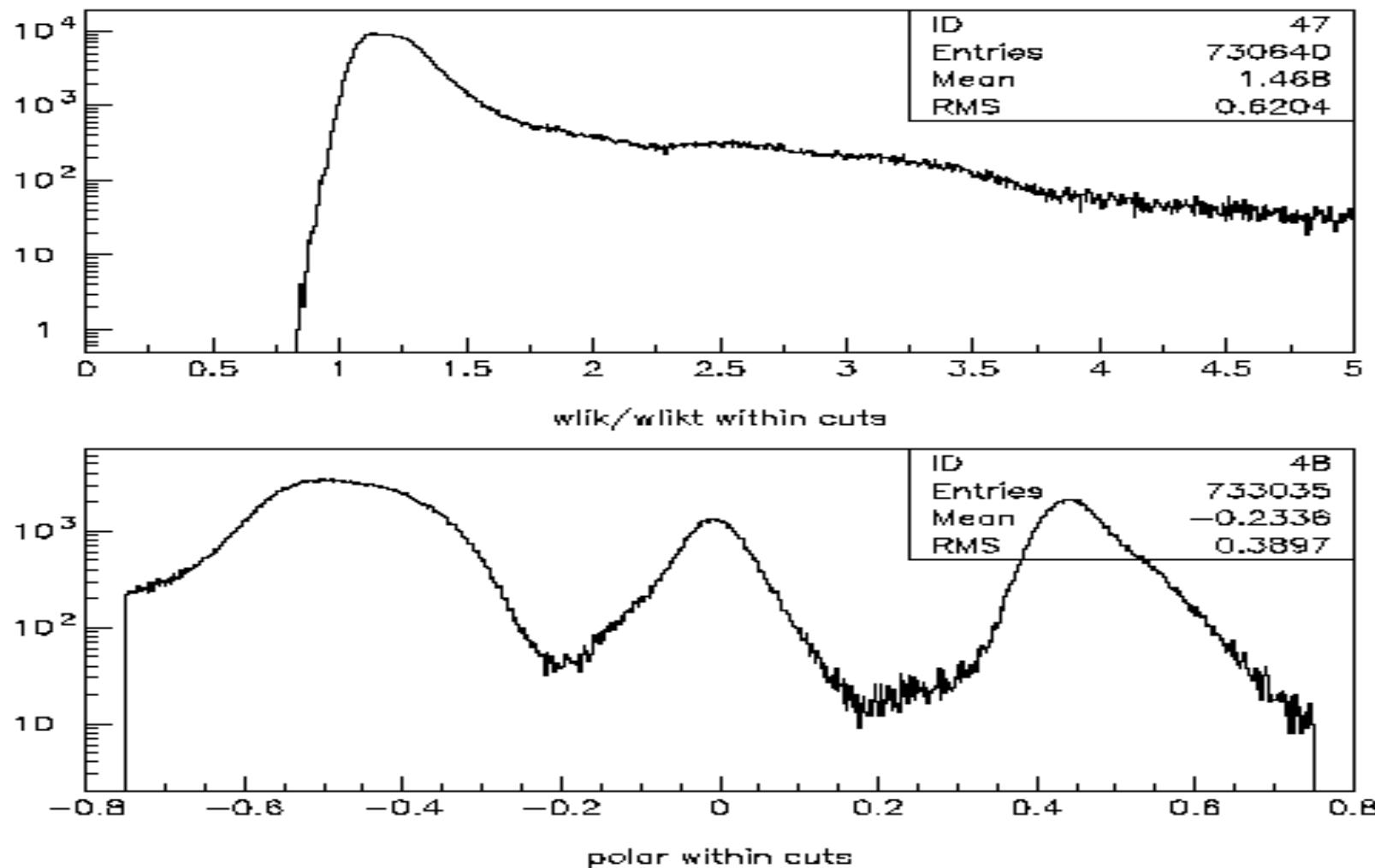
Errors determination

- Produce all the doublets
- Subdivide whole sample in ~80 sub periods
- Discriminate on rads brem compton and Tpol flatness
- For each error source choose a smaller sub sample (10,6,2) and vary conditions
- Run detector parameter measurements every 25 doublets
- Run polarization measurements
- Make plots comparison

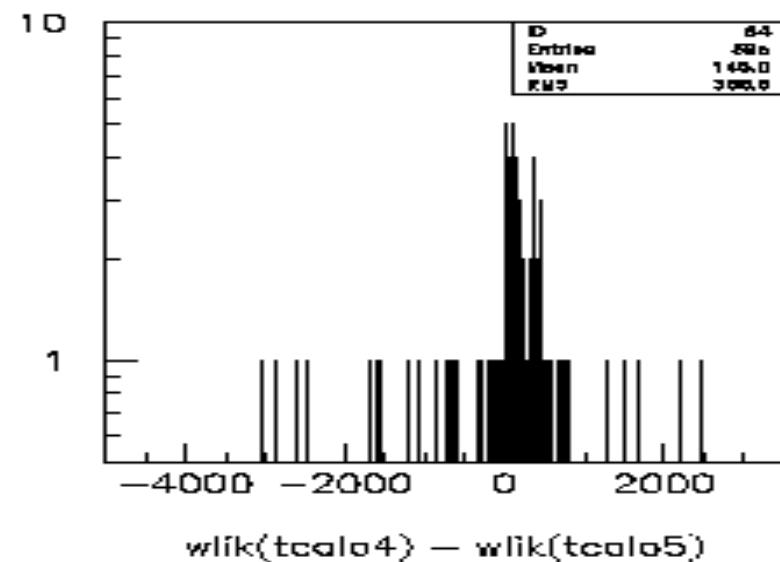
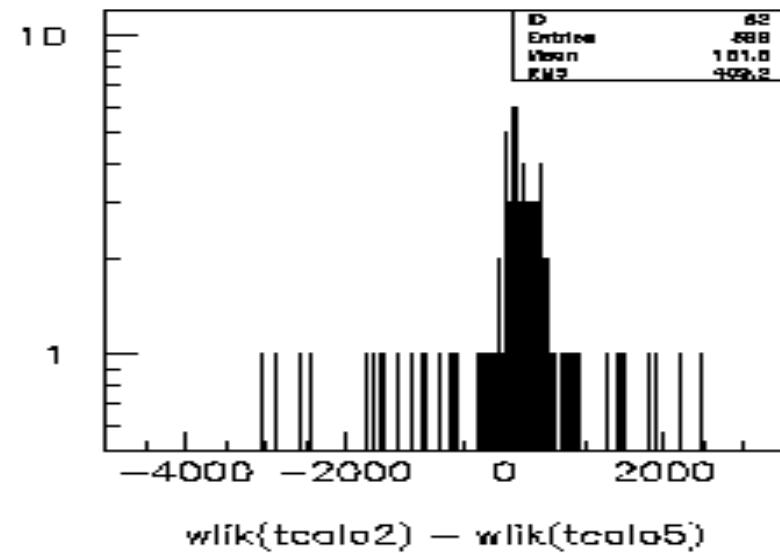
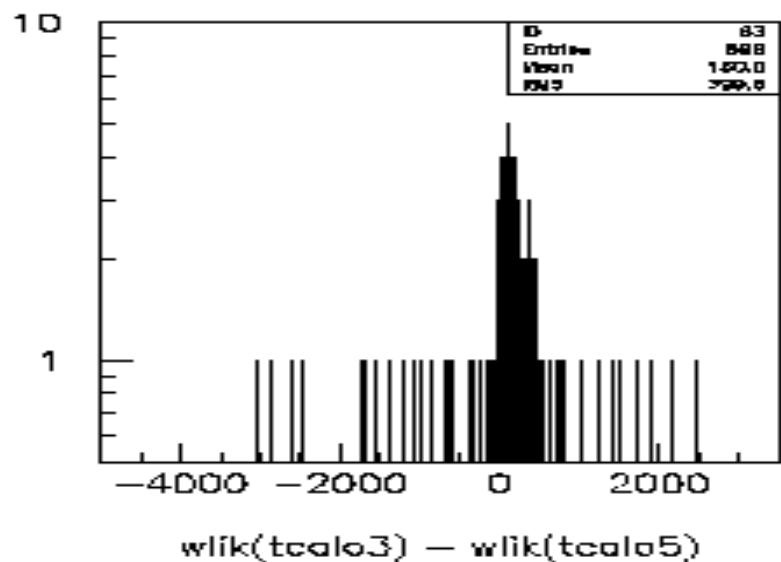
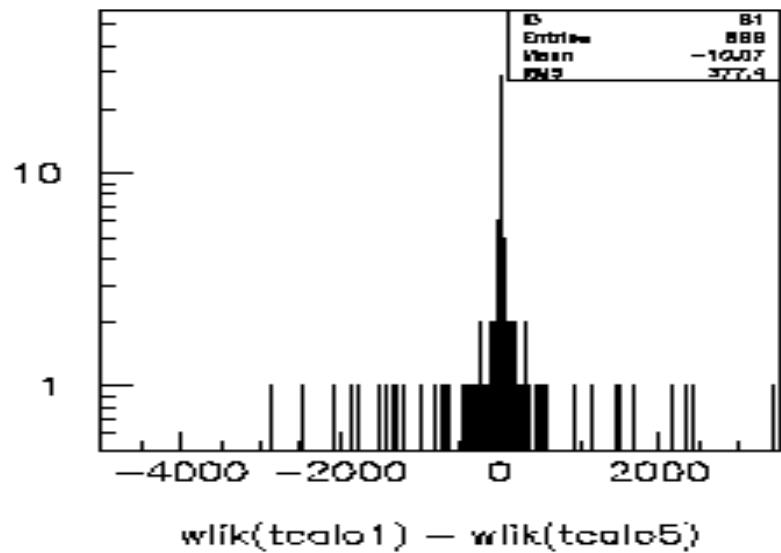
Error distribution for one bunch per 10 s



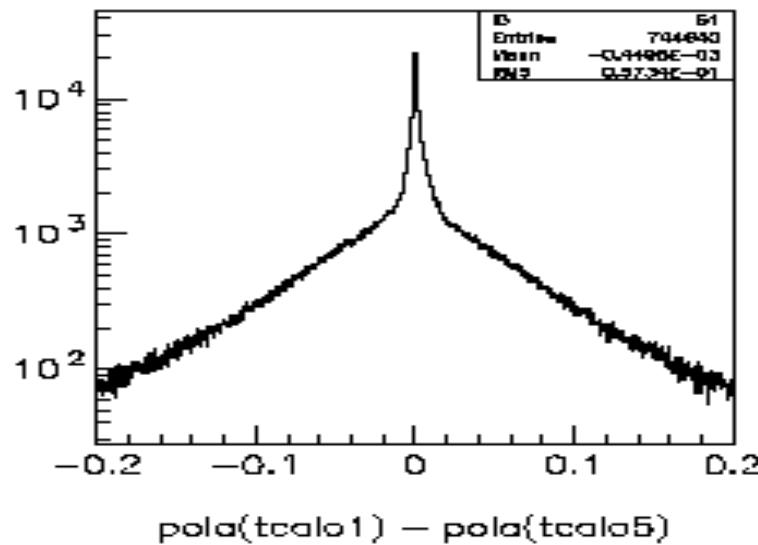
Detector analytical representation



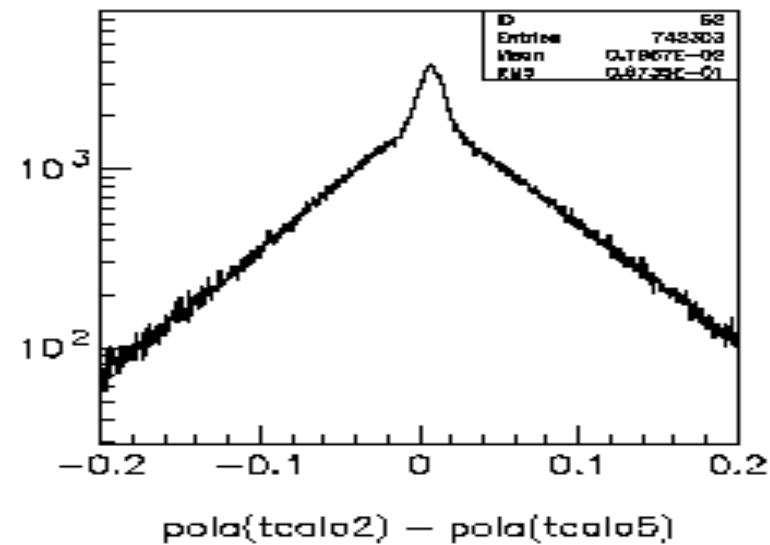
Detector analytical representation



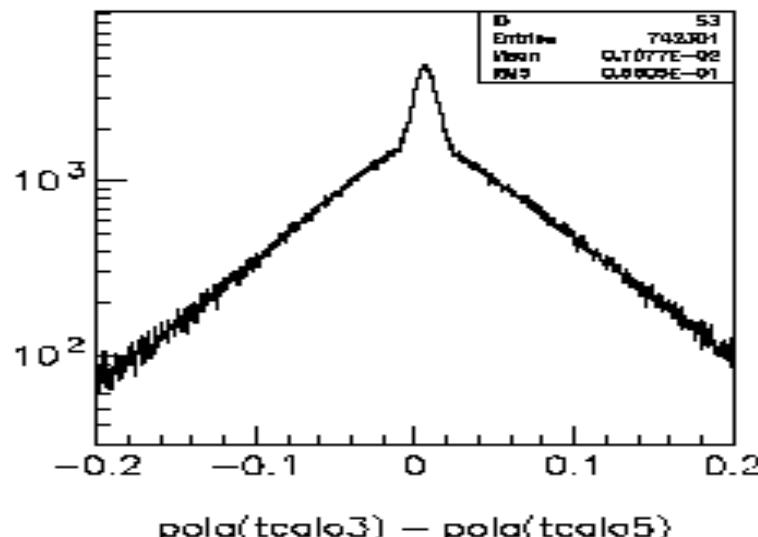
Detector analytical representation



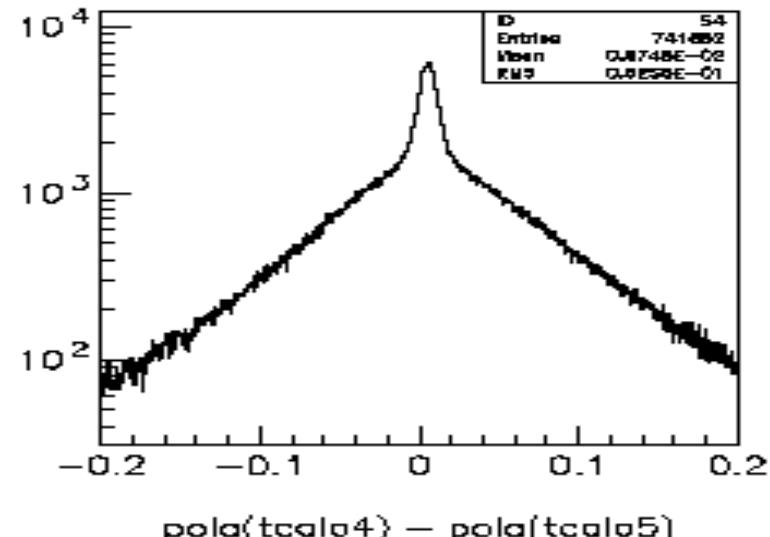
pola(tcalo1) – pola(tcalo5)



pola(tcalo2) – pola(tcalo5)

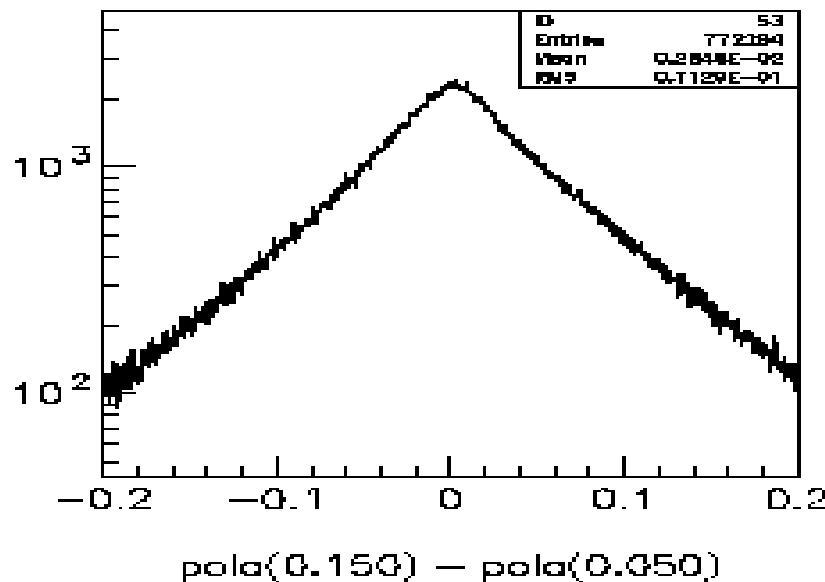
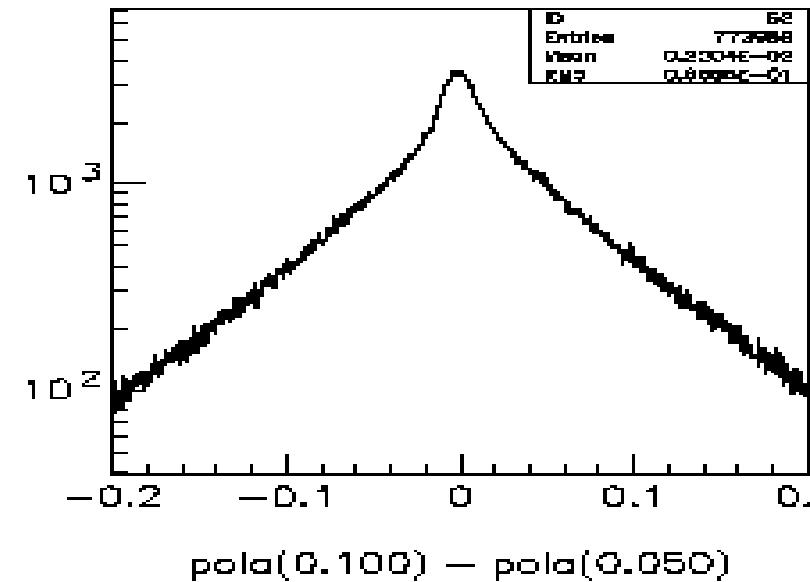
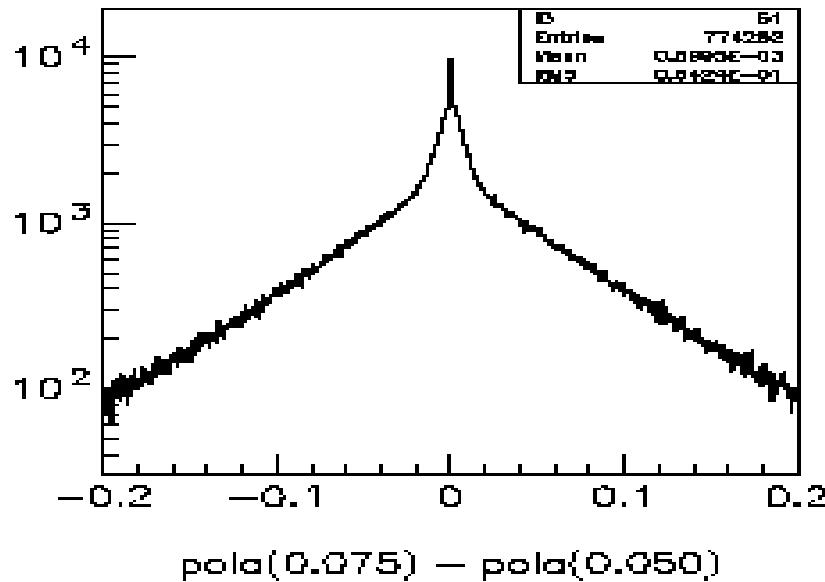


pola(tcalo3) – pola(tcalo5)

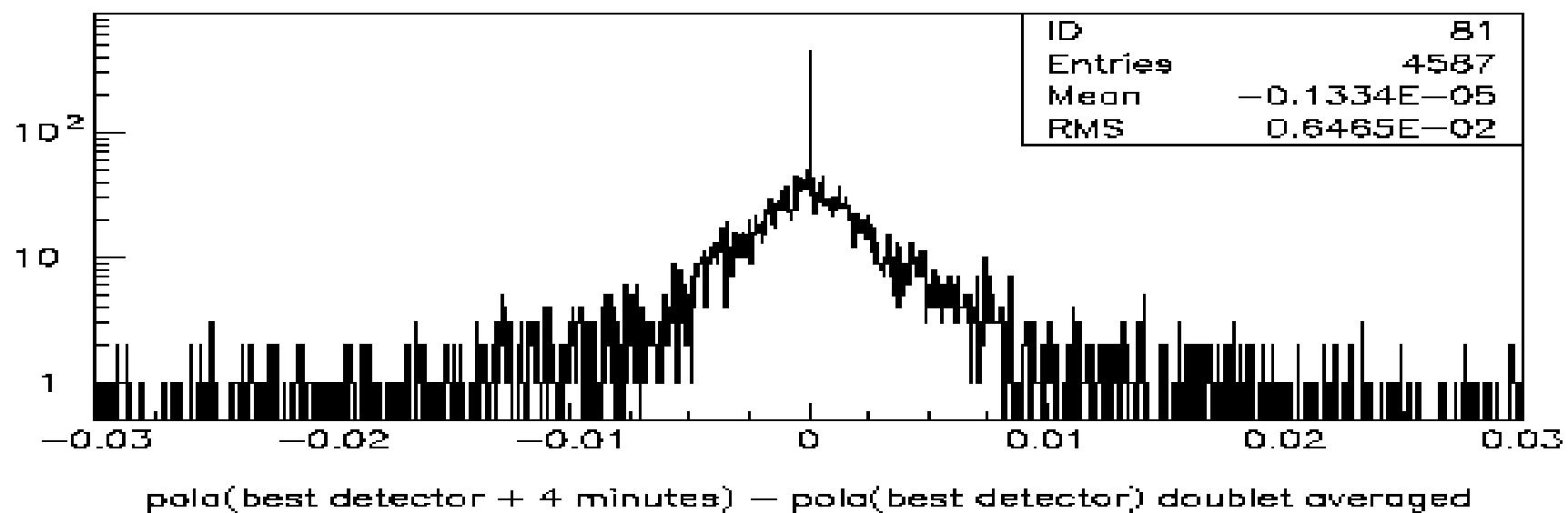
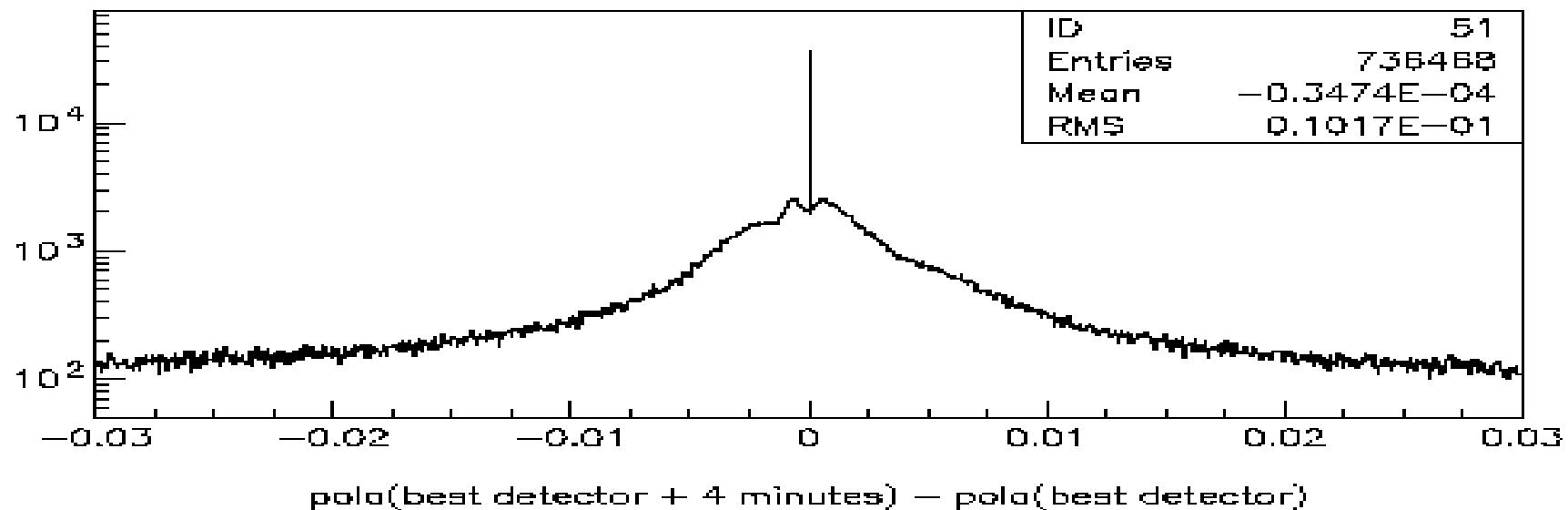


pola(tcalo4) – pola(tcalo5)

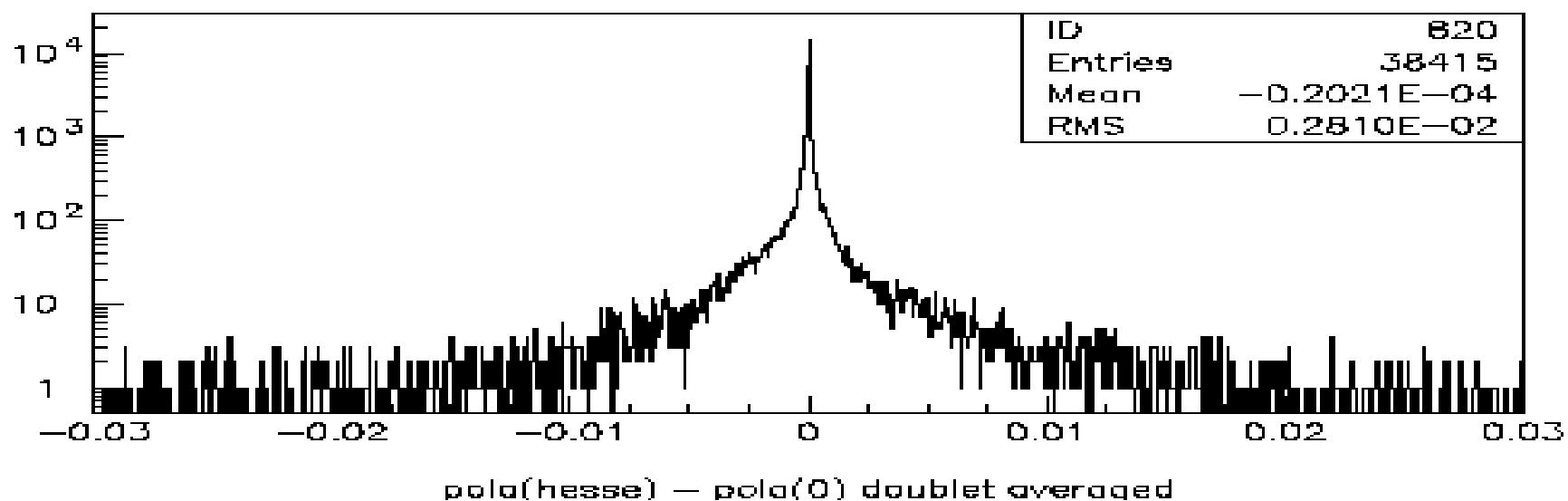
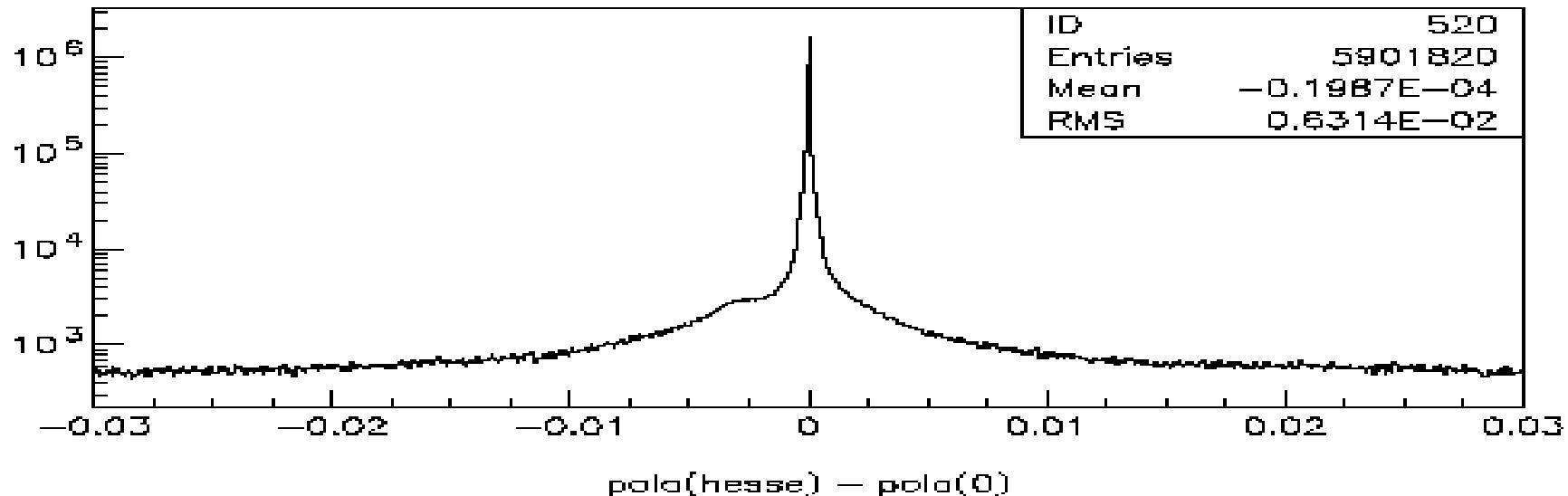
Radiation peak cut study



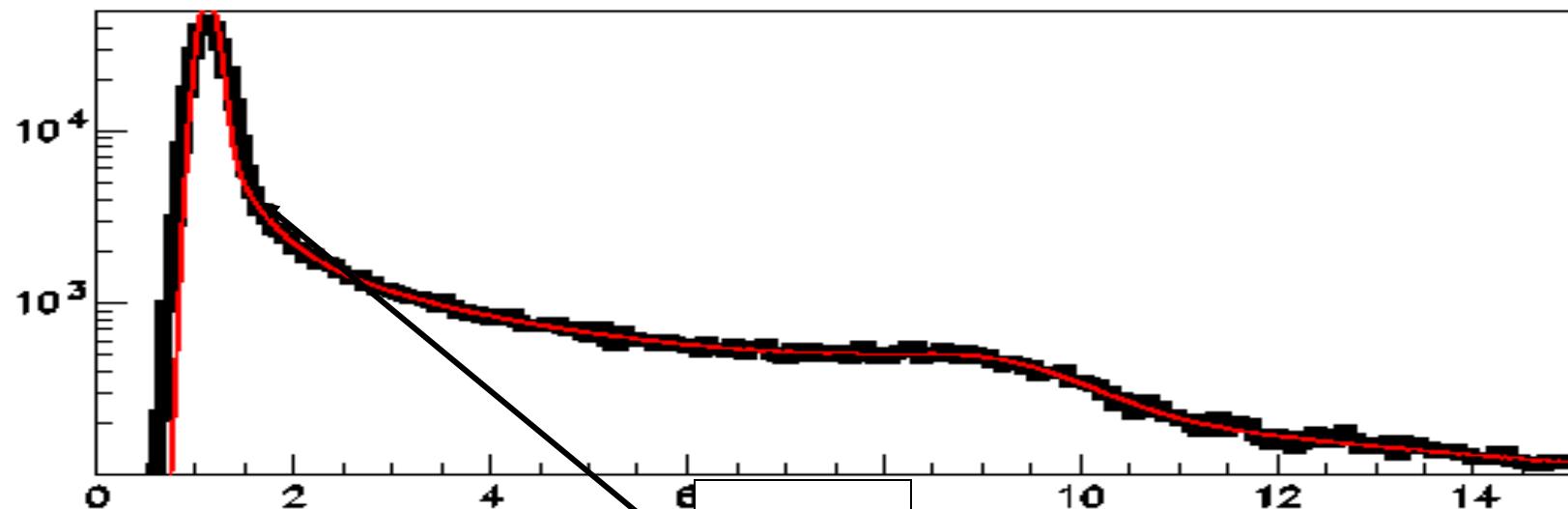
Detector fluctuation method 1



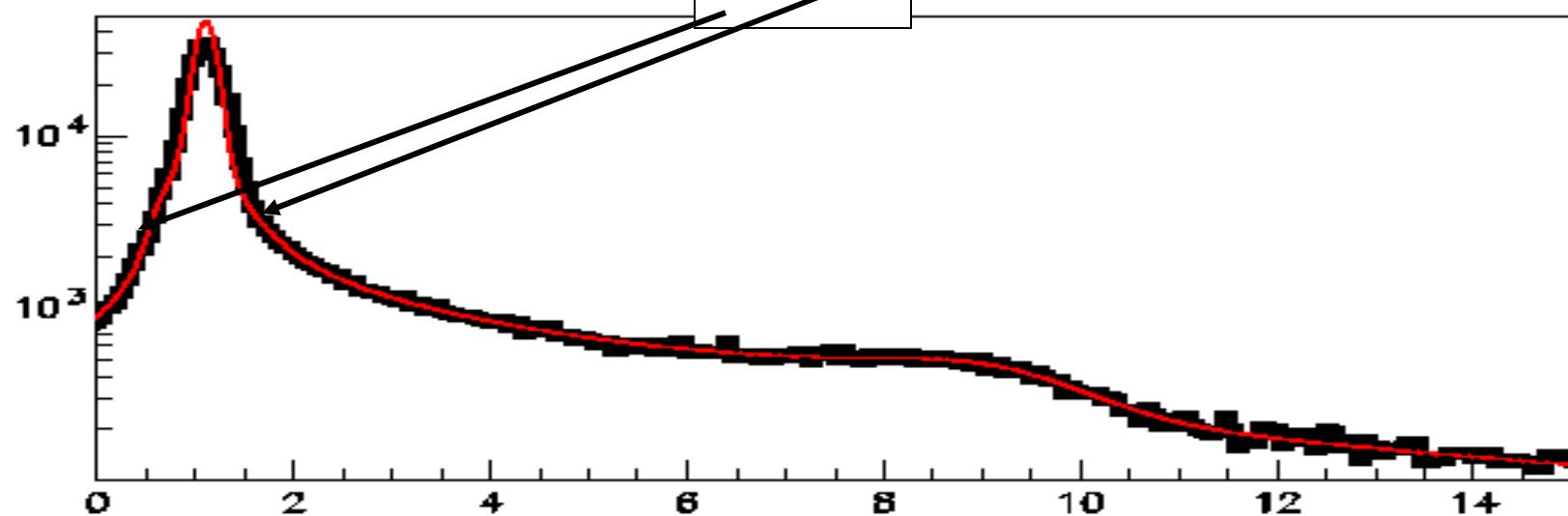
Detector fluctuation method 2



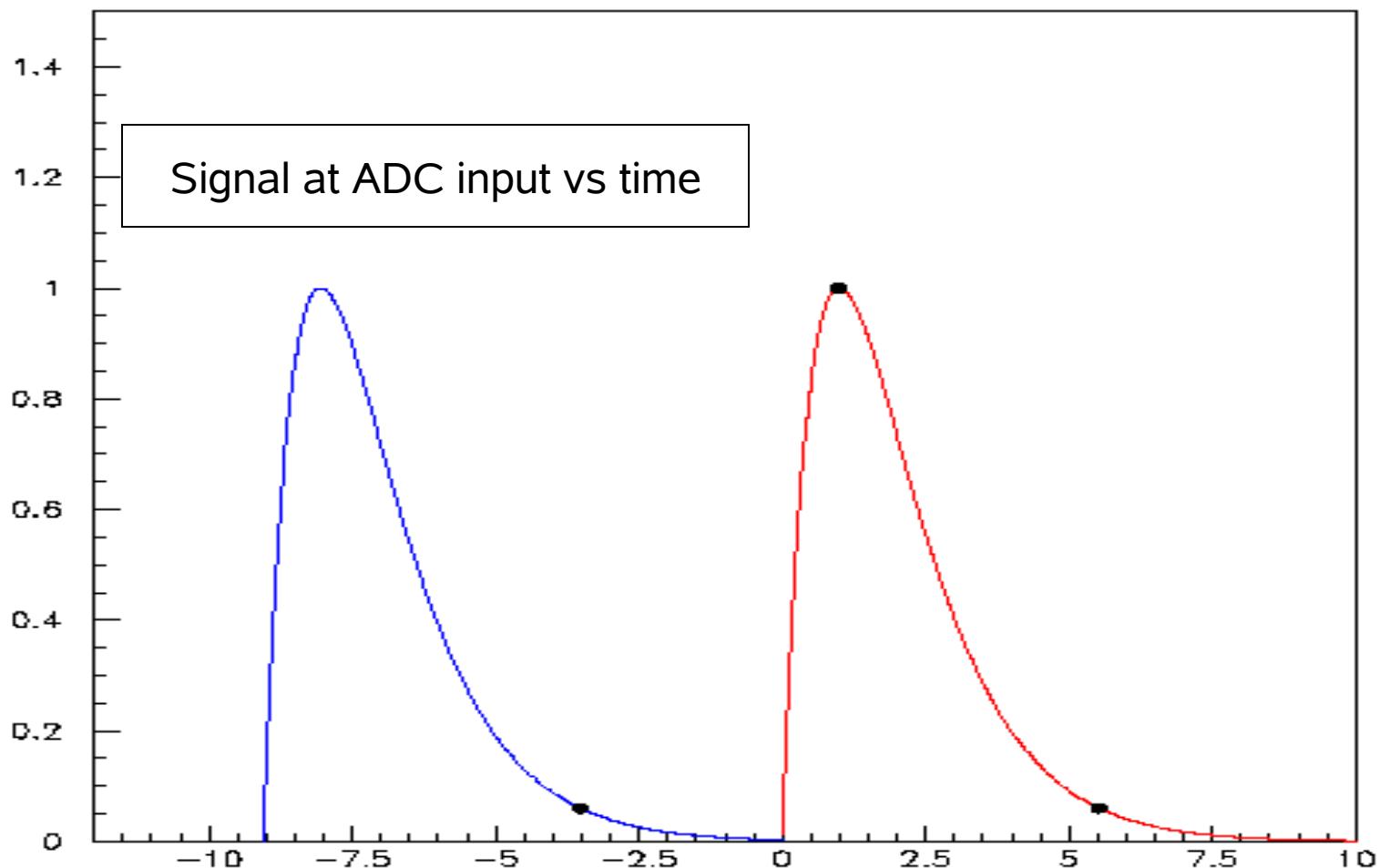
Leak effect



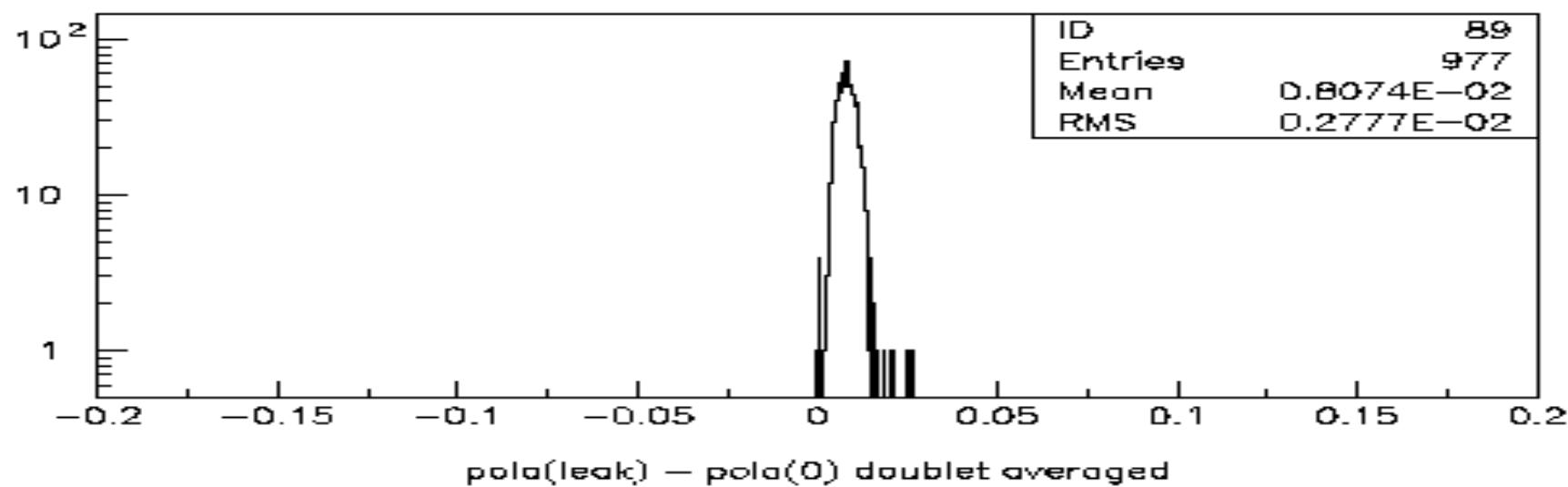
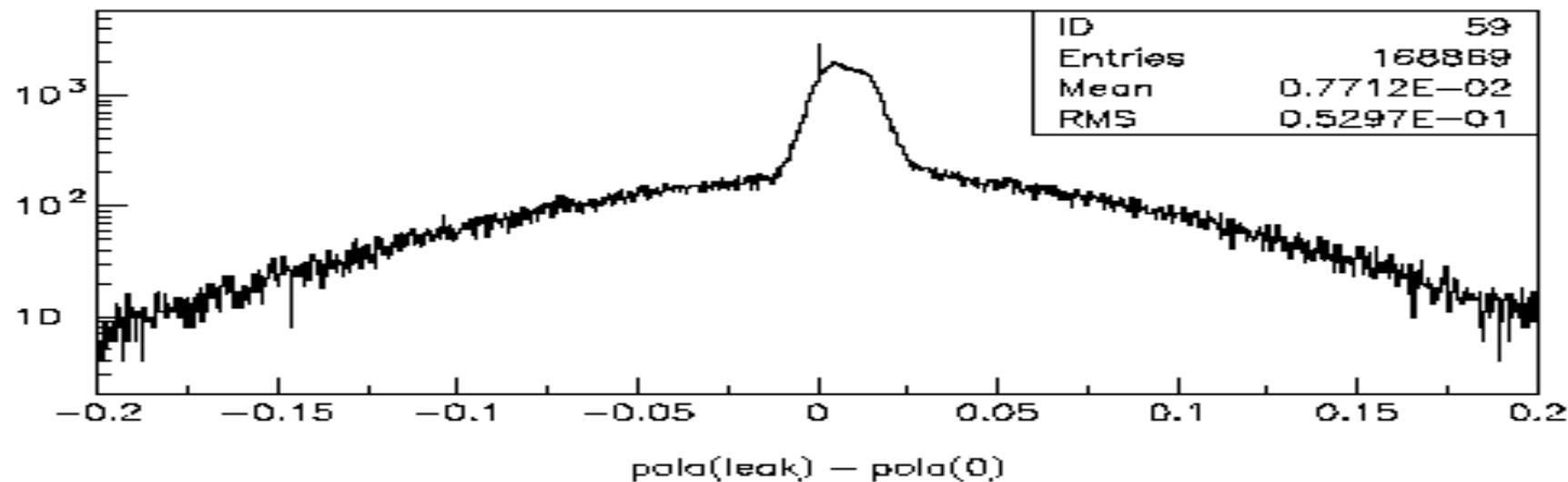
Rads cut



Leak effect and its systematic



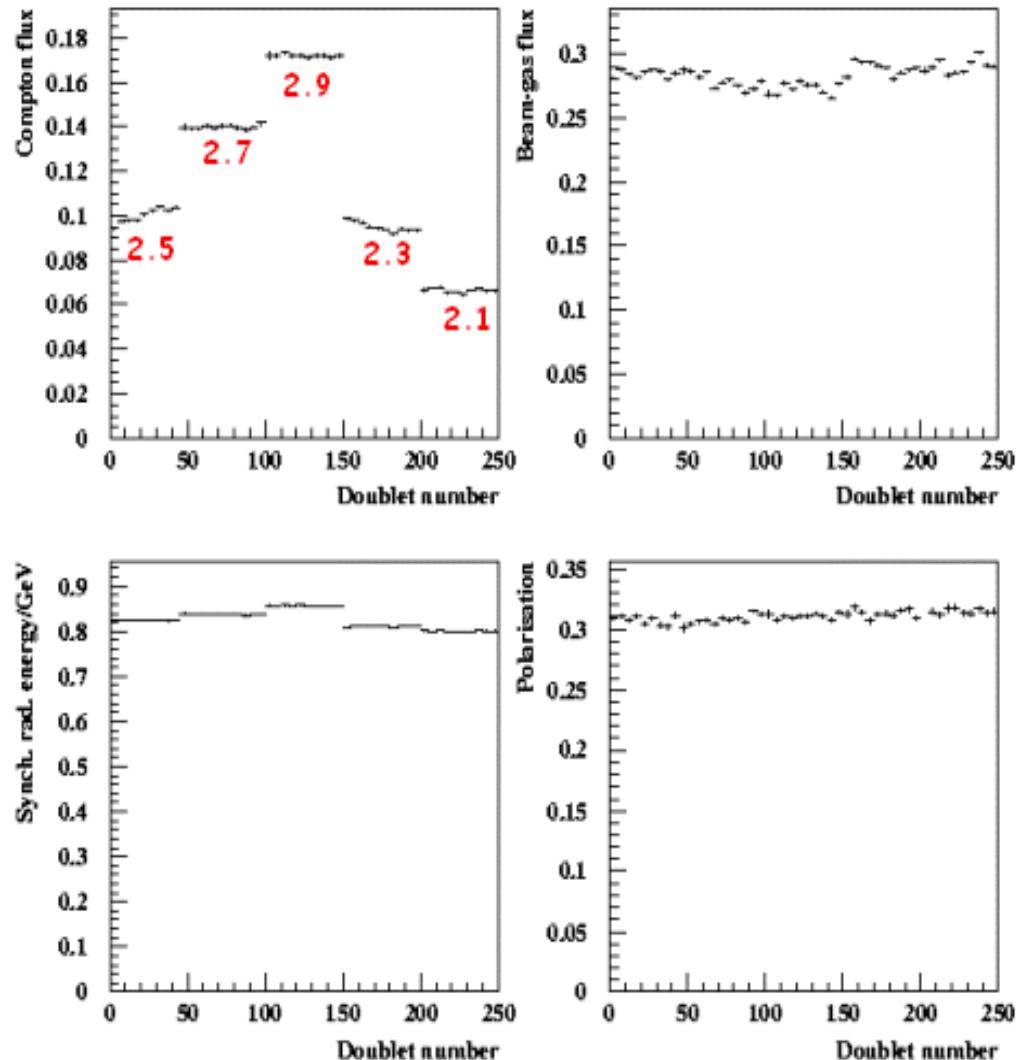
Leak



Beam position

Position (mm)	$\langle P_e \rangle$ (%)
2.5	30.98 ± 0.06
2.7	31.03 ± 0.05
2.9	31.20 ± 0.04
2.3	31.45 ± 0.07
2.1	31.77 ± 0.09

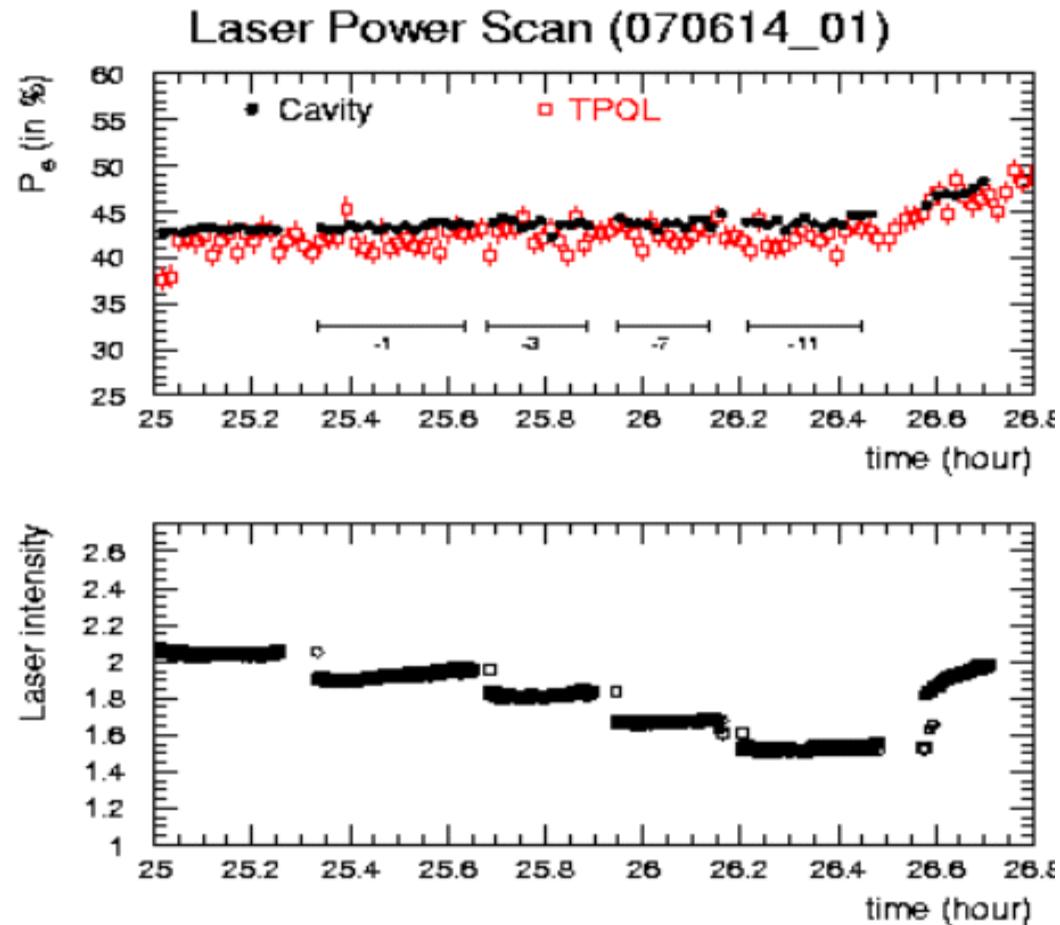
→ $\delta P_e \leq 0.32\%$



Laser power

Power	$\langle P_e \rangle$ (%)
Standard	43.44 ± 0.05
-1	43.70 ± 0.04
-3	44.07 ± 0.05
-7	44.09 ± 0.05
-11	44.21 ± 0.05

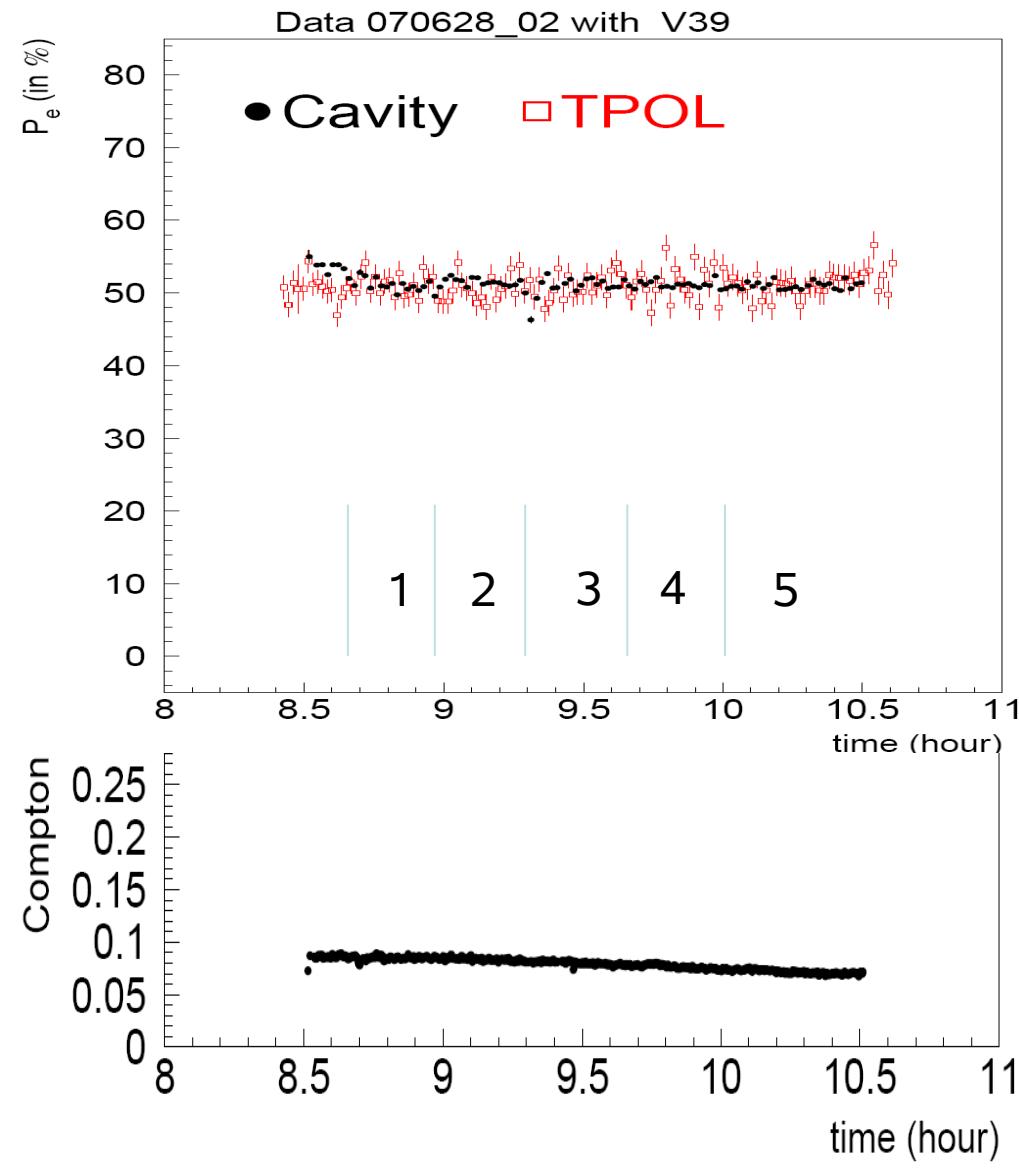
→ $\delta P_e \leq 0.37\%$



MOCO Position Scan

Position (S _{γ})	<P _e > (%)
1 (0.9936/-0.9834)	51.48 ± 0.09
2 (0.9936/-0.9909)	51.67 ± 0.08
3 (0.9973/-0.9968)	51.12 ± 0.08
4 (0.9842/-0.9968)	51.45 ± 0.08
5(default) (0.9957/-0.9950)	51.23 ± 0.06

→ $\delta P_e \leq 0.44\%$



Statistical uncertainties

- Pola error per bunch and for 10s : 3%
- From detector parameters (*) : 0.5 %
(fully bunches and doublet correlated
during 6mn)

Systematic uncertainties

- From Hera : 0.70%
- From Laser : 0.75%
- From detector: 0.10%
- -----
- Total 1. %

Future prospect

Everything done so far once but needs to be improved before final result

- Optimize detector parameters finding (done)
- Choose between tcalo1 and tcalo5 . Eventually run tcalo15 (one parameter more in widening technique $E=\alpha x^\mu$)
- Revisit systematic issues with test sub periods
- Run detector parameters finding on all periods
- Run polar extraction on all periods
- Extend to earlier 16 bits daq output and check quality (these go back to October 2006)

THE END

Polarimeter principle

Compton Scattering:

$$e + \gamma \rightarrow e + \gamma$$

Cross Section:

$$\frac{d\sigma}{dE_\gamma} = \sigma_0(E_\gamma) - P_e S_y \sigma_1(E_\gamma)$$

σ_0, σ_1 : known (QED)

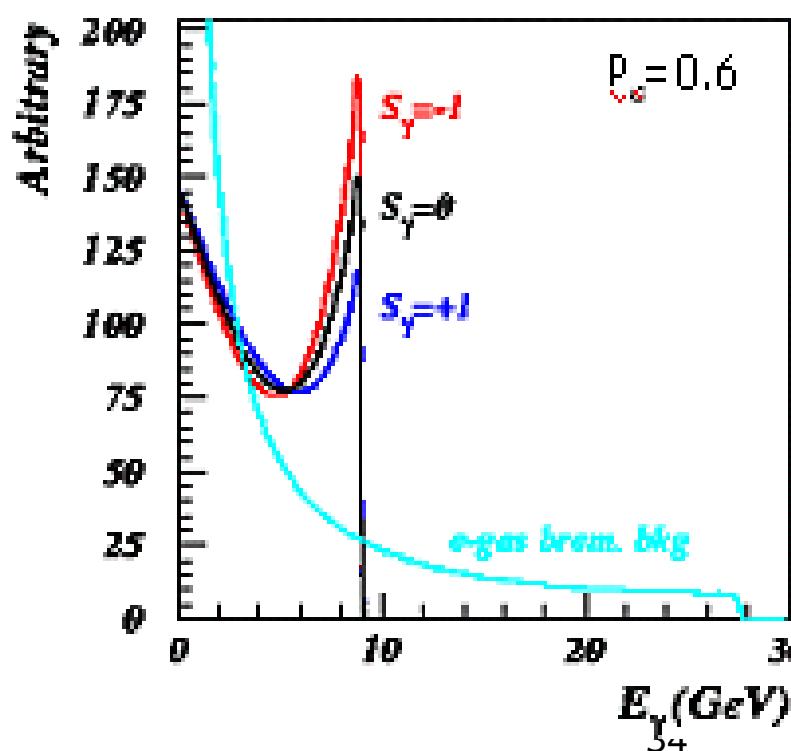
P_e : Polarization of the **e beam**
to be measured

S_y : Circular polarization (+-1)
of the **laser beam**

Luminosity (electron-laser):

$$L_{el} \propto \frac{P_L I_e}{k \alpha \sqrt{\sigma_{e,e}^2 + \sigma_{e,\gamma}^2}}$$

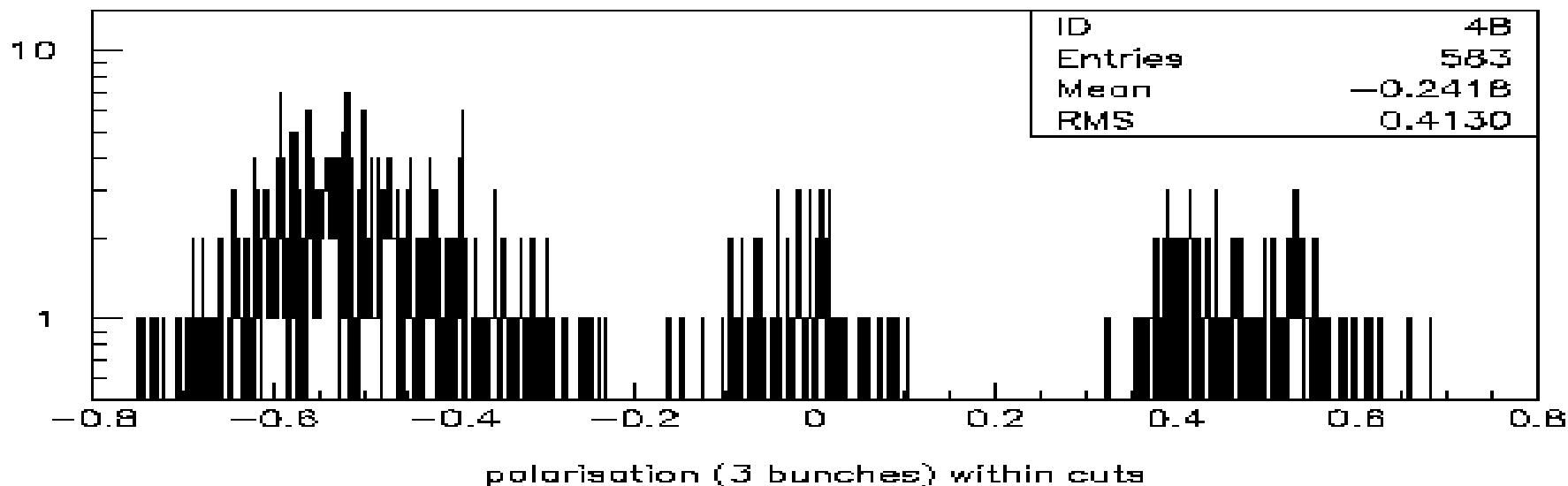
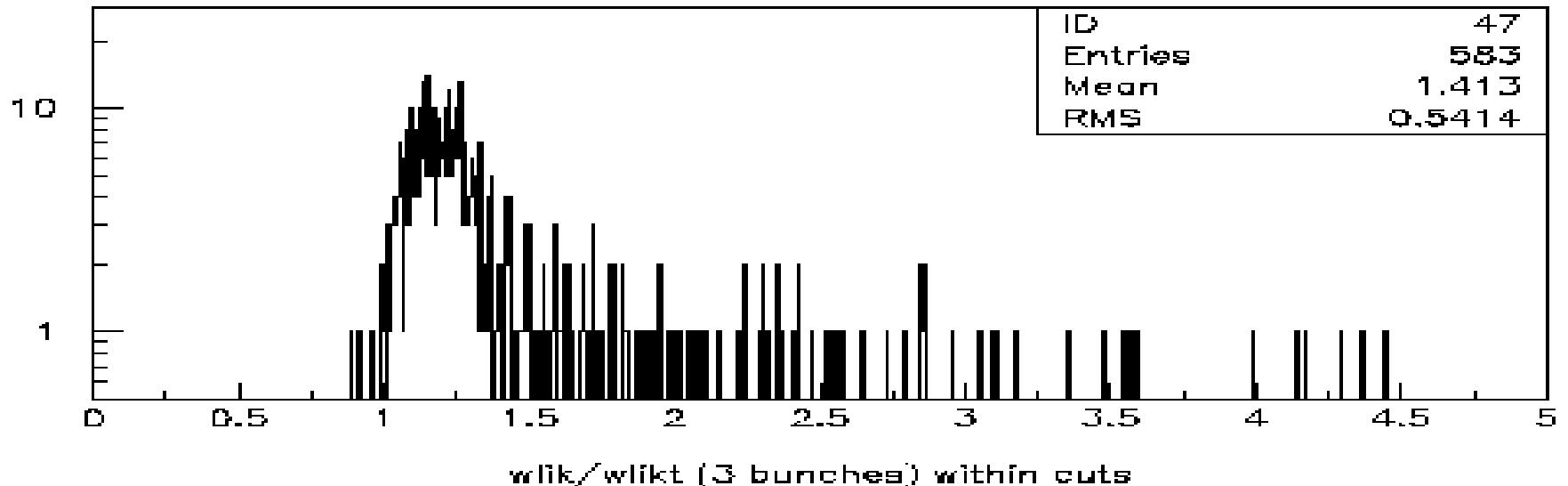
I_e : e beam intensity
 P_L : Laser beam power



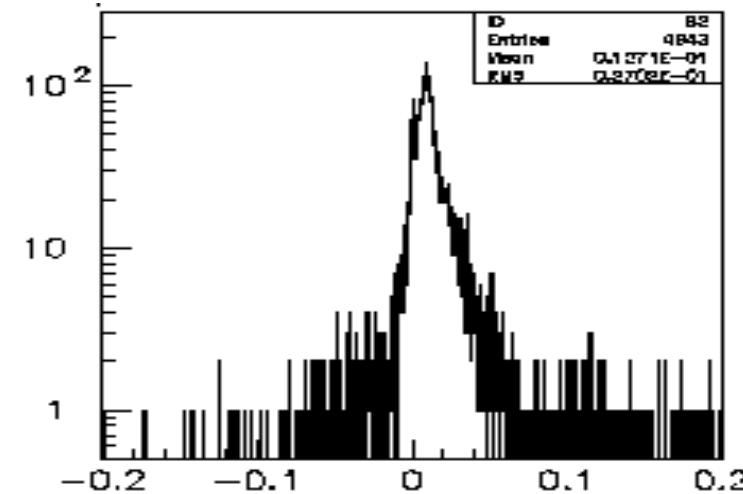
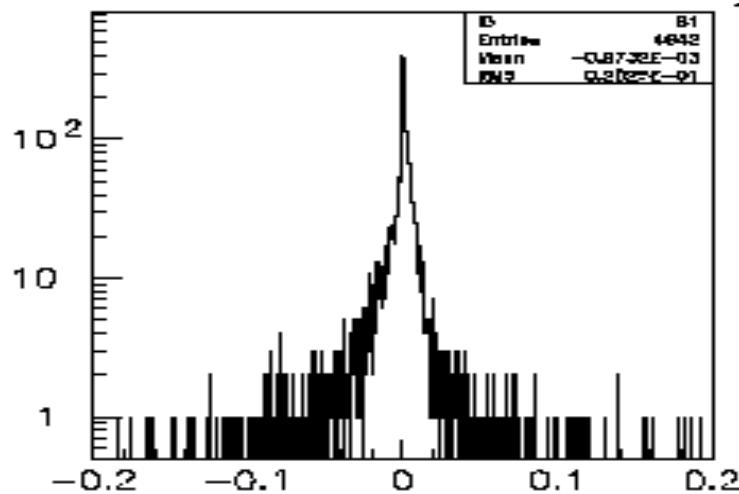
Previous systematics

Source	δP_e (%)	
Laser polarization uncertainty	≤ 0.50	Most of the errors are conservatively estimated.
Laser circularity [MOCO position scan]	≤ 0.44	
Laser power variation	≤ 0.37	
Electronic Noise [C_{empty} (1.43 - 1.35)]	0.20	
Detector parameters [$\pm 1\sigma$]	0.12	
Dead material in front of calorimeter?	0.17	
Calorimeter position scan in x & y	≤ 0.58	
Synchrotron radiation cut [0.05 → 0.01 - 0.1]	0.29	
Blackbody temperature [300K → 500K]	0.29	
Beam position scan	≤ 0.32	
e beam energy uncertainty (27.6GeV → 27.5GeV)	0.21	Error reduction is expected in the future with improved fit program.
Total	≤ 1.2	

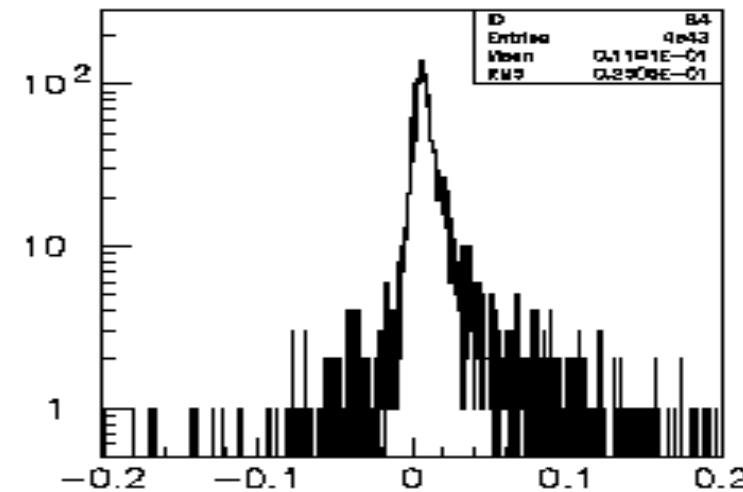
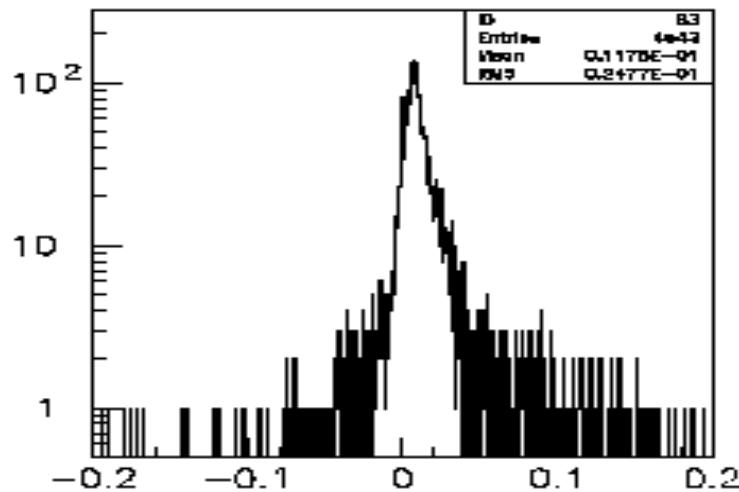
Detector analytical representation



Detector analytical representation

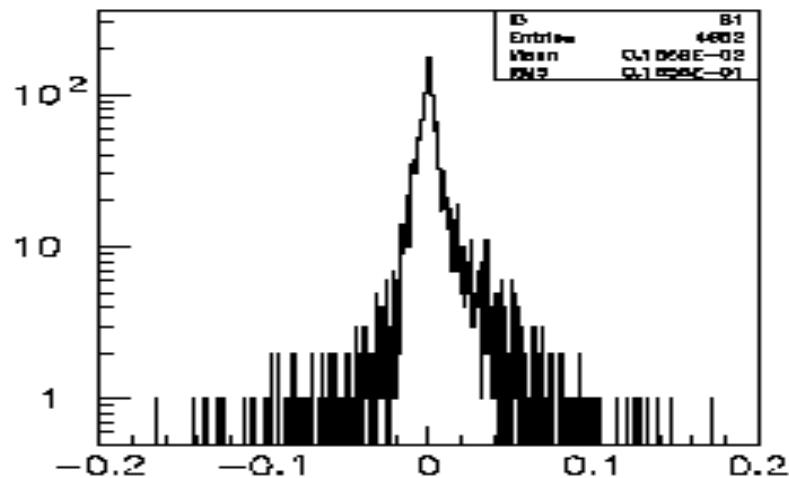


$\text{pola}(\text{tcalo1}) - \text{pola}(\text{tcalo5})$ doublet average $\text{pola}(\text{tcalo2}) - \text{pola}(\text{tcalo5})$ doublet averaged

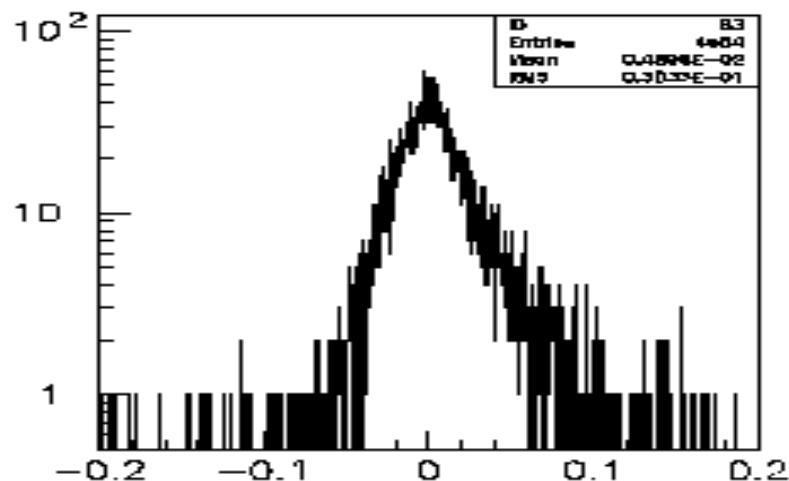


$\text{pola}(\text{tcalo3}) - \text{pola}(\text{tcalo5})$ doublet average $\text{pola}(\text{tcalo4}) - \text{pola}(\text{tcalo5})$ doublet averaged 17

Radiation peak cut study

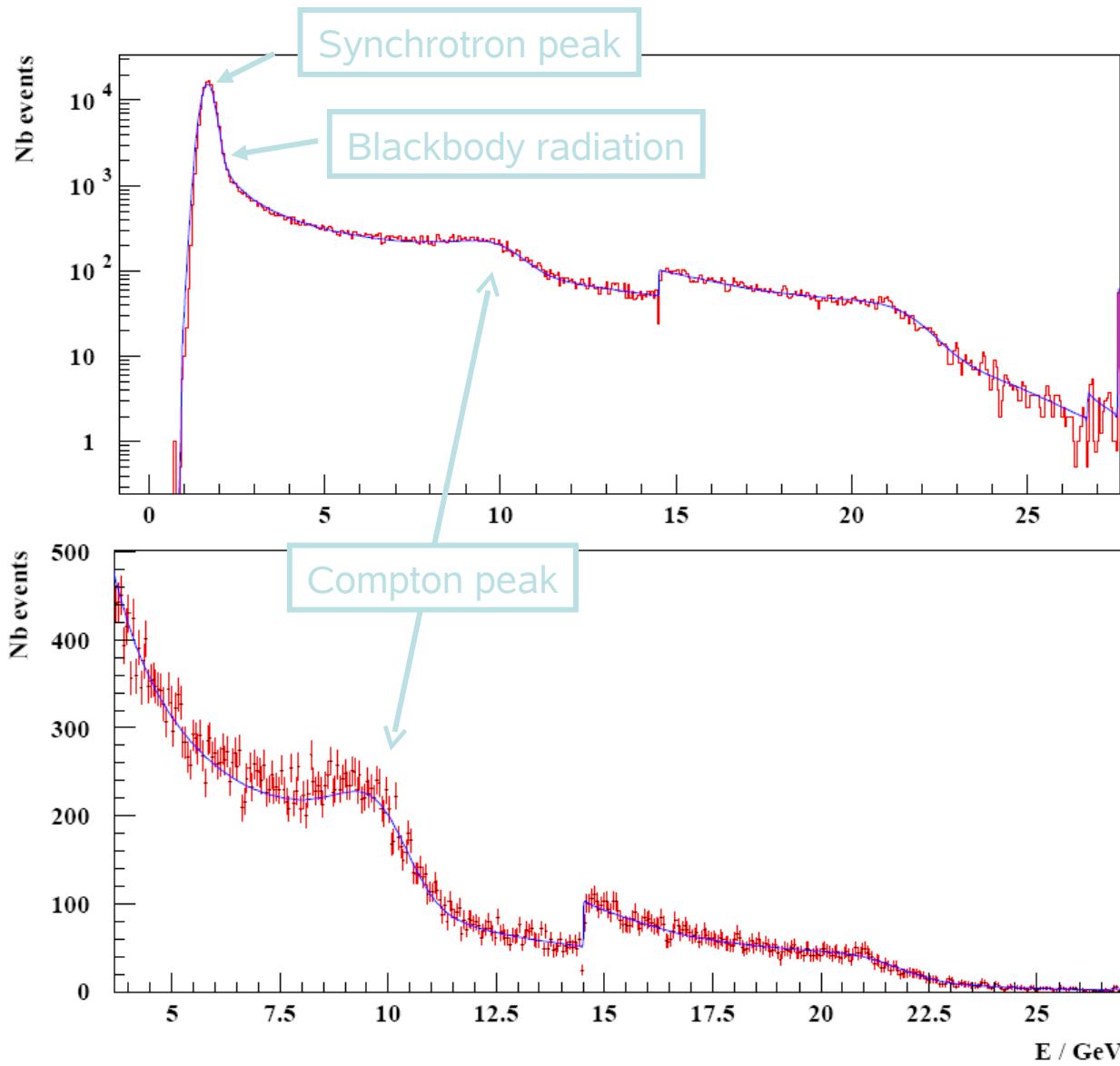


pola(0.075) — pola(0.050) doublet average pola(0.100) — pola(0.050) doublet averaged



pola(0.150) — pola(0.050) doublet averaged

Measurement (varying bin histogram) vs Fit



An arbitrary example of one single bunch (~4s)

→ good description of the data (in red) by the fit (in blue)

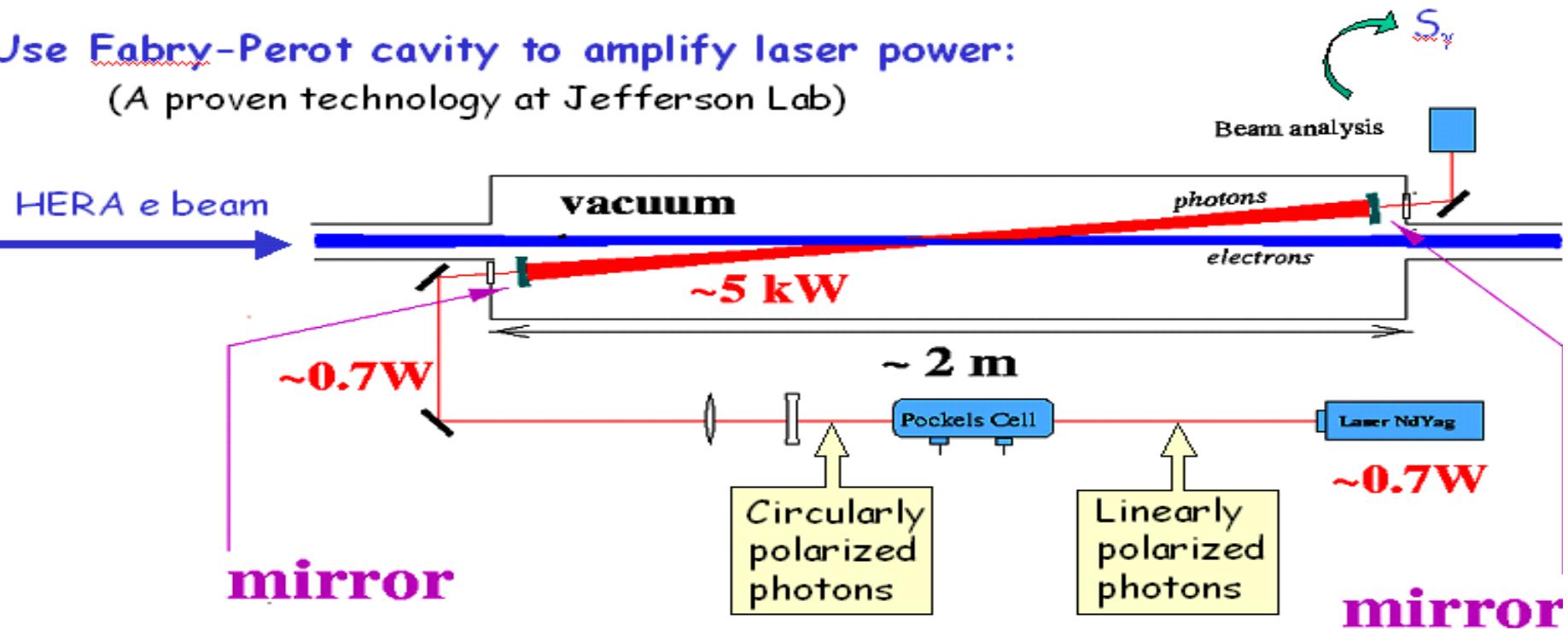
New feature: varying histogram bins [finer (coarse) bins @ low (high) energies]

→ to have optimal gain vs available ADC dynamical range

→ provide a complete low energy spectrum

Cavity setup

Use **Fabry-Perot cavity** to amplify laser power:
(A proven technology at Jefferson Lab)



All optical components are fixed rigidly on an optical table