

# COMPARISONS OF YFSWW AND KORALW

M. SKRZYPEK

CERN, GENEVA  
INP, CRACOW

W. PŁACZEK

S. JADACH

Z. WAS

B. FL. WARD

1) TECHNICAL TESTS FOR "COMMON"  
CC $\phi$ 3 RESULTS

2) PHYSICAL TESTS - SIZE OF  
4 fermion AND  $\mathcal{O}(\alpha)$  CORRECTIONS

3) UNCERTAINTY OF  
"KINEMATICS EXTRAPOLATION PROCEDURE"

ECFA-DESY

HAMBURG

20.09.2000



# MODIFICATIONS OF KORALW 1.42.3 TO BE INCLUDED IN FUTURE RELEASE FOR NOW AVAILABLE AS SEPARATE PATCH

1) NEW NORMALISATION OPTION (KEYBRA=3)  
TO ALLOW FOR CONVENTION AS IN  
RACONWW (W-BRANCHING RATIOS)

2) NEW "KINEMATIC EXTRAPOLATION PROCEDURE"

$$4f + N_f \leftrightarrow 4f$$

GENERATOR

BORN MATRIX EL.

UNDEFINED  
SUB-LEADING  
TERMS  $\equiv$

FREEDOM OF ROTATING  
THIS "EFFECTIVE" FRAME

NECESSARY TO ADJUST BREMSSTRAHLUNG  
TO THE ONE IN YFSWW

3) TECHNICALITIES - EASIER SIMULTANEOUS  
HANDLING OF CC $\phi$ 3 AND 4f WEIGHTS

$$WTSET(i) = 4f \quad i=1, \dots, 4$$

$$WTSET(10-i) = CC\phi 3$$



# CROSS-CHECKS:

YFSWW3  $\leftrightarrow$  KORALW

## CUTS:

(YR2K; 4F-SECTION; hep-ph/0005309)

### 1. FERMIONS:

- $\angle(f_{ch}, \text{beams}) > 10^\circ$ ,  $f_{ch}$  - CHARGED FERMION
- NO CUTS FOR  $\nu$ 'S

### 2. PHOTONS:

- $\angle(\gamma, \text{beams}) > 5^\circ$
- $\min\{M_{\gamma f_{ch}}, M_{\gamma \bar{f}_{ch}}\} < M_{rec}$  OR  $E_\gamma < 1 \text{ GeV}$   
 $\Rightarrow \gamma$  RECOMBINED WITH  $f_{ch}$   
(and discarded)

$$\rightarrow M_{rec} = \begin{cases} 5 \text{ GeV} : \text{BARE} \\ 25 \text{ GeV} : \text{CALO} \end{cases}$$

! PHOTON RADIATION FROM QUARKS INCLUDED

$\rightarrow$  SPECIAL VERSION OF PHOTOS (THANKS TO ZWAS)

$\rightarrow$  CC 11-TYPE CHANNELS

# TECHNICAL COMPARISONS

CCΦ3

YFSWW-KORALW

$$e^+e^- \rightarrow W^+W^- \rightarrow u\bar{d}\mu^-\bar{\nu}_\mu$$

161 GeV

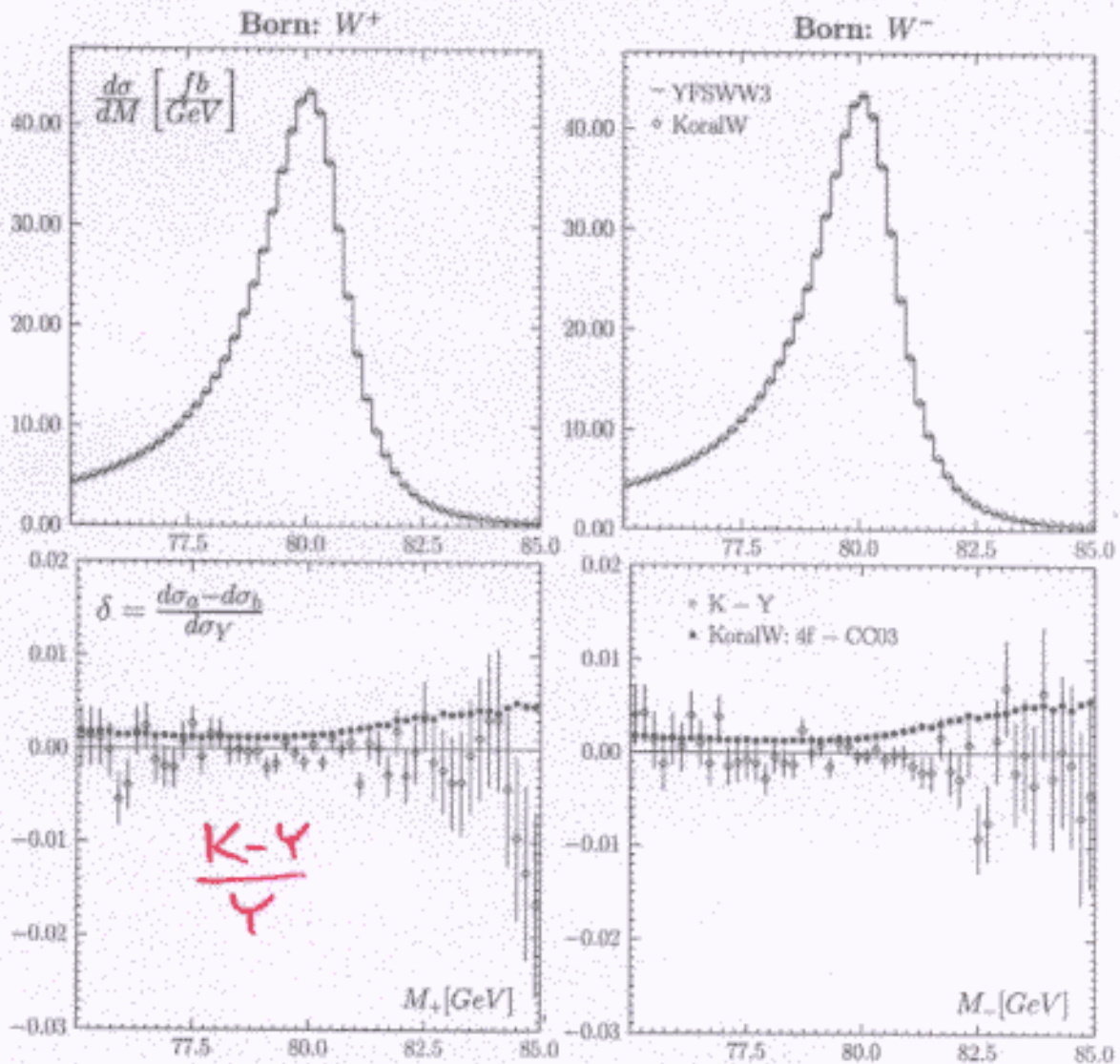


Figure 1: Distributions of the  $W^+$  and  $W^-$  invariant masses in the Born approximation at  $E_{CM} = 161$  GeV.



# TECHNICAL COMPARISONS

CC03 YFSWW - KORALW

$$e^+e^- \rightarrow W^+W^- \rightarrow u\bar{d}\mu^-\bar{\nu}_\mu$$

500 GeV

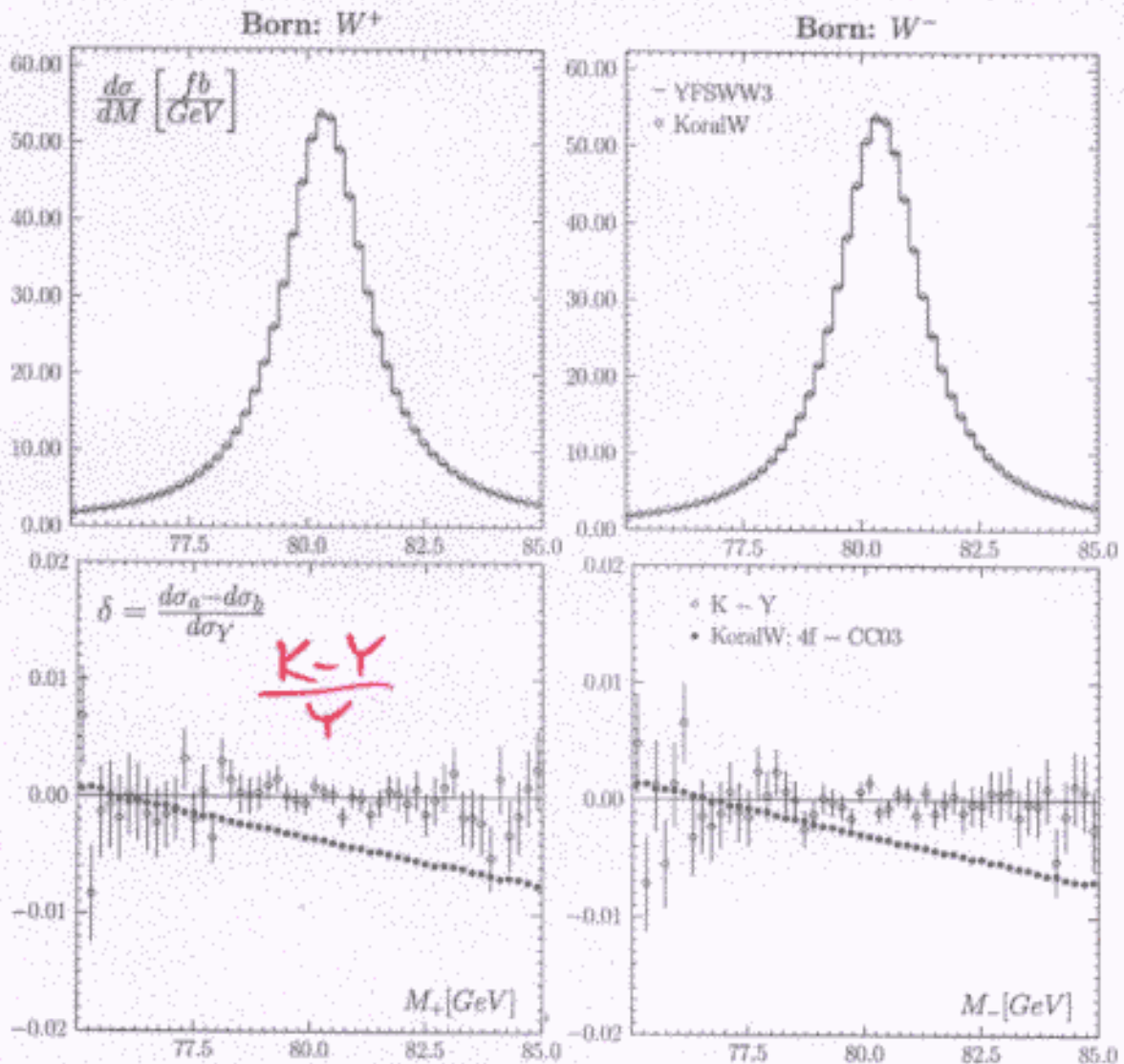


Figure 1: Distributions of the  $W^+$  and  $W^-$  invariant masses in the Born approximation at  $E_{CM} = 500 \text{ GeV}$ .

200 GeV

NO CUTS		$\sigma_{WW} [fb]$		$\delta_{4f} [\%]$		$\delta_{WW}^{NL} [\%]$
Final state	Program	Born	ISR	Born	ISR	
$\nu_{\mu}\mu^{+}\tau^{-}\bar{\nu}_{\tau}$	YFSWW3	219.793 (16)	204.198 (09)	—	—	-1.92 (4)
	KoralW	219.766 (26)	204.178 (21)	0.041	0.044	—
	(Y-K)/Y	0.01 (1)%	0.01 (1)%	—	—	—
$u\bar{d}\mu^{-}\bar{\nu}_{\mu}$	YFSWW3	659.69 (5)	635.81 (3)	—	—	-1.99 (4)
	KoralW	659.59 (8)	635.69 (7)	0.073	0.073	—
	(Y-K)/Y	0.02 (1)%	0.02 (1)%	—	—	—
$u\bar{d}s\bar{c}$	YFSWW3	1978.37 (14)	1978.00 (09)	—	—	-2.06 (4)
	KoralW	1977.89 (25)	1977.64 (21)	0.060	0.061	—
	(Y-K)/Y	0.02 (1)%	0.02 (1)%	—	—	—

Table 1: The total  $WW$  cross sections from YFSWW3 and KoralW at the Born and ISR level, the  $4f$  corrections from KoralW and  $\mathcal{O}(\alpha)$  NL correction from YFSWW3 at  $E_{CM} = 200 GeV$ , without cuts. The numbers in parentheses are statistical errors of the results corresponding to last digits.

WITH CUTS		$\sigma_{WW} [fb]$		$\delta_{4f} [\%]$		$\delta_{WW}^{NL} [\%]$
Final state	Program	Born	ISR	Born	ISR	
$\nu_{\mu}\mu^{+}\tau^{-}\bar{\nu}_{\tau}$	YFSWW3	210.938 (16)	196.205 (09)	—	—	-1.93 (4)
	KoralW	210.911 (26)	196.174 (21)	0.041	0.044	—
	(Y-K)/Y	0.01 (1)%	0.02 (1)%	—	—	—
$u\bar{d}\mu^{-}\bar{\nu}_{\mu}$	YFSWW3	627.22 (5)	605.18 (3)	—	—	-2.00 (4)
	KoralW	627.13 (8)	605.03 (7)	0.074	0.074	—
	(Y-K)/Y	0.01 (1)%	0.02 (1)%	—	—	—
$u\bar{d}s\bar{c}$	YFSWW3	1863.60 (15)	1865.00 (09)	—	—	-2.06 (4)
	KoralW	1863.07 (25)	1864.62 (21)	0.065	0.064	—
	(Y-K)/Y	0.03 (2)%	0.02 (1)%	—	—	—

Table 2: The total  $WW$  cross sections from YFSWW3 and KoralW at the Born and ISR level, the  $4f$  corrections from KoralW and  $\mathcal{O}(\alpha)$  NL correction from YFSWW3 at  $E_{CM} = 200 GeV$ , with cuts. The numbers in parentheses are statistical errors of the results corresponding to last digits.

4f  $\mathcal{O}(\alpha)$   
 $\mathcal{O}(\alpha) > 4f$

NO CUTS		$\sigma_{WW} [fb]$		$\delta_{4f} [\%]$		$\delta_{WW}^{NL} [\%]$
Final state	Program	Born	ISR	Born	ISR	
$u\bar{d}\mu^-\bar{\nu}_\mu$	YFSWW3	156.670 (16)	122.832 (08)	0.29	—	-1.41 (4)
	KoralW	156.601 (24)	122.836 (11)	—	0.25	—
	(Y-K)/Y	0.04 (2)%	0.00 (1)%	—	—	—

Table 1: The total  $WW$  cross sections from YFSWW3 and KoralW at the Born and ISR level, the  $4f$  corrections from KoralW and  $\mathcal{O}(\alpha)$  NL correction from YFSWW3 at  $E_{CM} = 161 GeV$ , without cuts. The numbers in parentheses are statistical errors of the results corresponding to last digits.

4f  $\mathcal{O}(\alpha)$

WITH CUTS		$\sigma_{WW} [fb]$		$\delta_{4f} [\%]$		$\delta_{WW}^{NL} [\%]$
Final state	Program	Born	ISR	Born	ISR	
$u\bar{d}\mu^-\bar{\nu}_\mu$	YFSWW3	151.158 (16)	118.482 (08)	0.29	—	-1.41 (4)
	KoralW	151.089 (24)	118.485 (11)	—	0.25	—
	(Y-K)/Y	0.05 (2)%	0.00 (1)%	—	—	—

Table 2: The total  $WW$  cross sections from YFSWW3 and KoralW at the Born and ISR level, the  $4f$  corrections from KoralW and  $\mathcal{O}(\alpha)$  NL correction from YFSWW3 at  $E_{CM} = 161 GeV$ , with cuts. The numbers in parentheses are statistical errors of the results corresponding to last digits.



TECHNICAL PREC. KORALW-YFSWW  $\sim 2 \times 10^{-4}$  (CC03)  
 $\mathcal{O}$  - "REDUCTION PROCEDURE" AMBIGUITY  $\leq 1 \times 10^{-3}$

4f  $\mathcal{O}(\alpha)$

Final state	NO CUTS Program	$\sigma_{WW}$ [fb]		$\delta_{4f}$ [%]		$\delta_{WW}^{NL}$ [%]
		Born	ISR	Born	ISR	
$u\bar{d}\mu^-\nu_\mu$	YFSWW3	261.368 (23)	292.029 (18)	---	---	-4.95 (4)
	KoralW	261.348 (17)	291.979 (19)	-0.51	-0.51	---
	(Y-K)/Y	0.01 (1)%	0.02 (1)%	---	---	---

$\mathcal{O}(\alpha) \gg 4f$

Table 1: The total  $WW$  cross sections from YFSWW3 and KoralW at the Born and ISR level, the 4f corrections from KoralW and  $\mathcal{O}(\alpha)$  NL correction from YFSWW3 at  $E_{CM} = 500$  GeV, without cuts. The numbers in parentheses are statistical errors of the results corresponding to last digits.

4f  $\mathcal{O}(\alpha)$

Final state	WITH CUTS Program	$\sigma_{WW}$ [fb]		$\delta_{4f}$ [%]		$\delta_{WW}^{NL}$ [%]
		Born	ISR	Born	ISR	
$u\bar{d}\mu^-\nu_\mu$	YFSWW3	181.505 (22)	209.449 (17)	---	---	-6.34 (4)
	KoralW	181.480 (17)	209.592 (18)	-0.69	-0.69	---
	(Y-K)/Y	0.01 (1)%	-0.07 (1)%	---	---	---

Table 2: The total  $WW$  cross sections from YFSWW3 and KoralW at the Born and ISR level, the 4f corrections from KoralW and  $\mathcal{O}(\alpha)$  NL correction from YFSWW3 at  $E_{CM} = 500$  GeV, with cuts. The numbers in parentheses are statistical errors of the results corresponding to last digits.



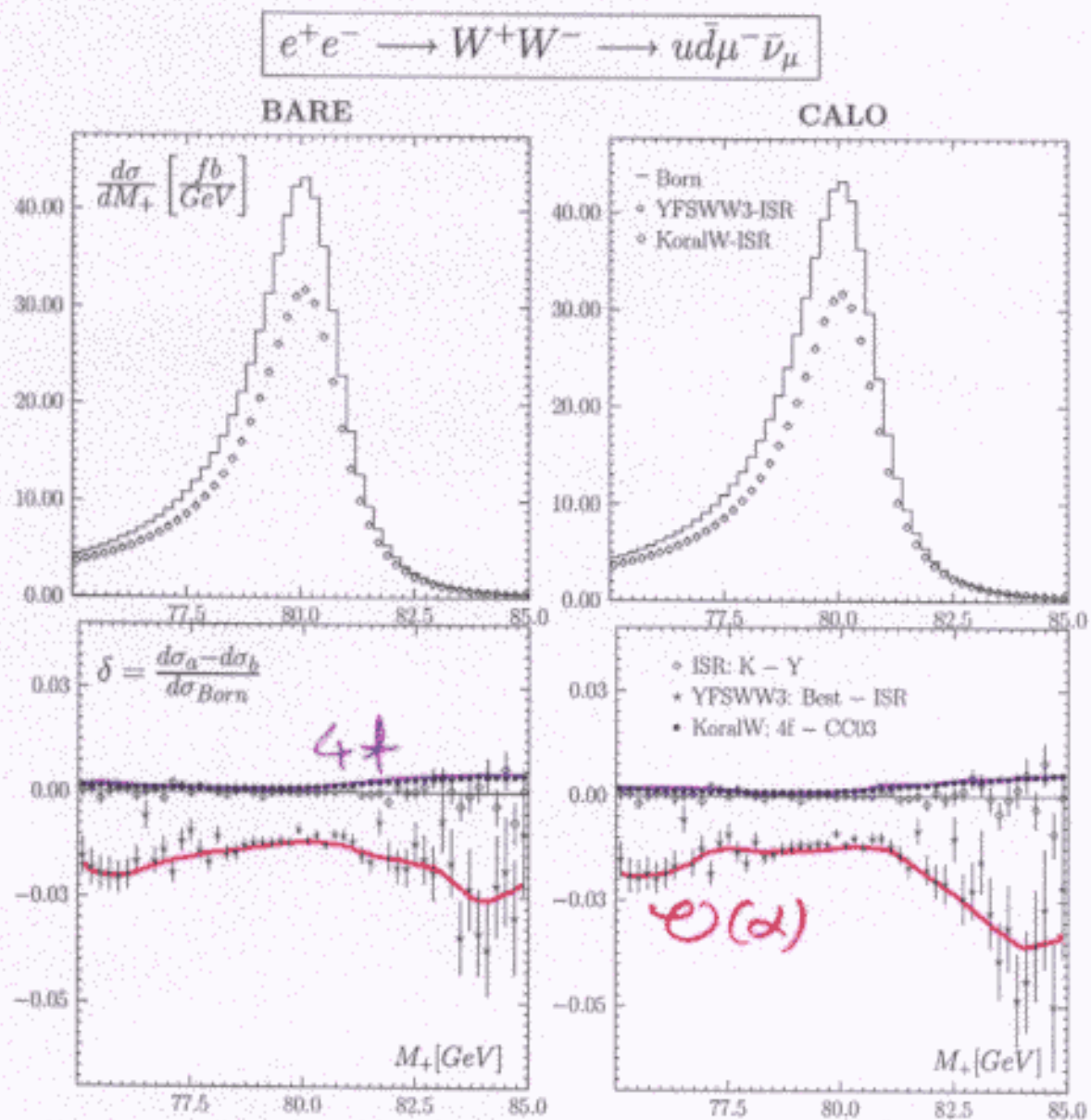


Figure 3: Distributions of the  $W^+$  invariant mass for BARE and CALO acceptances at  $E_{CM} = 161 GeV$ .

$O(\alpha) \gg 4^*$

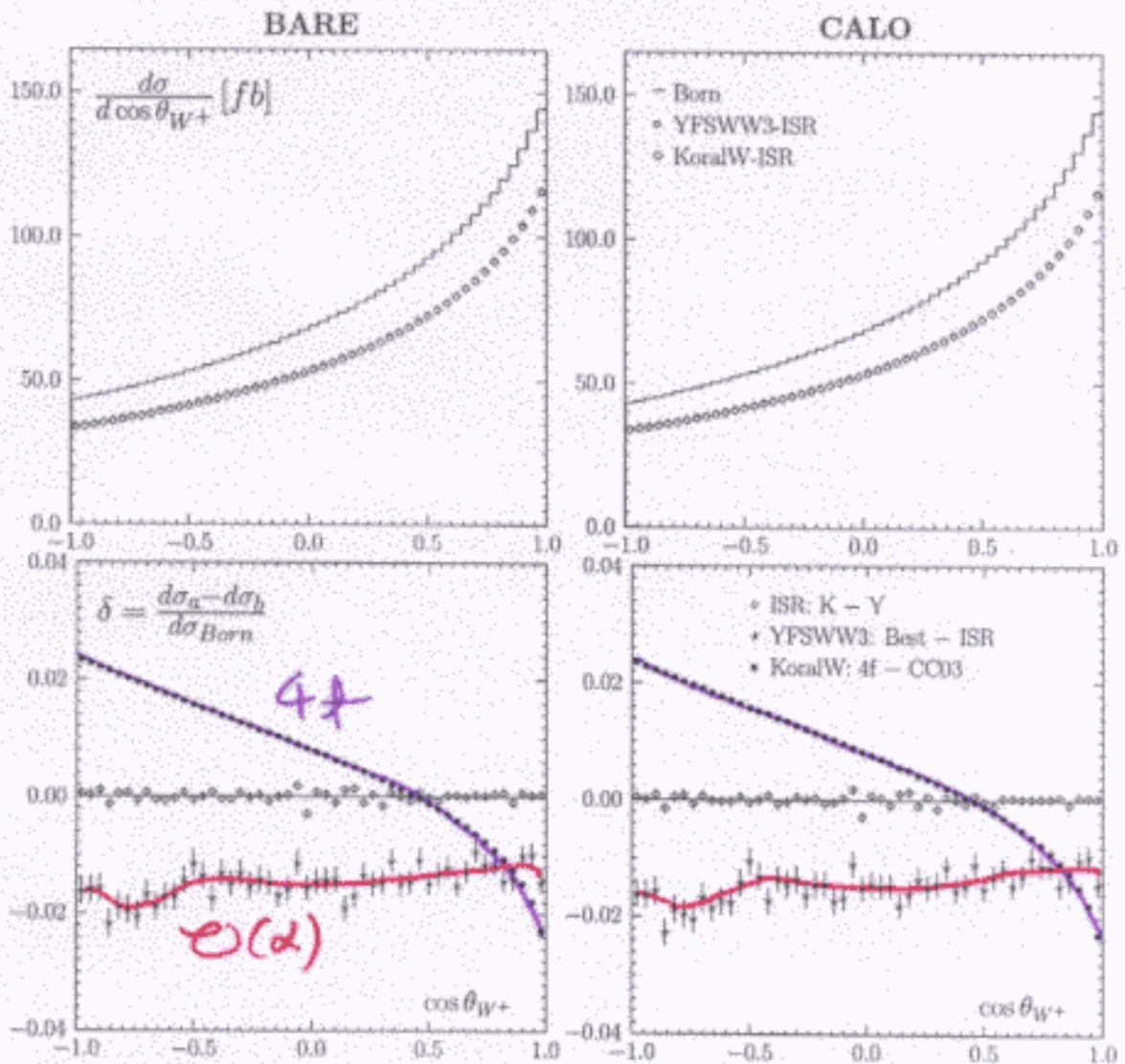
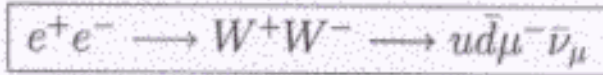


Figure 5: Distributions of cosine of the  $W^+$  polar angle w.r.t. the  $e^+$  beam for BARE and CALO acceptances at  $E_{CM} = 161$  GeV.

$$\psi(\alpha) \sim 4f$$

BUT! 4f CANCELS OUT!



200 GeV

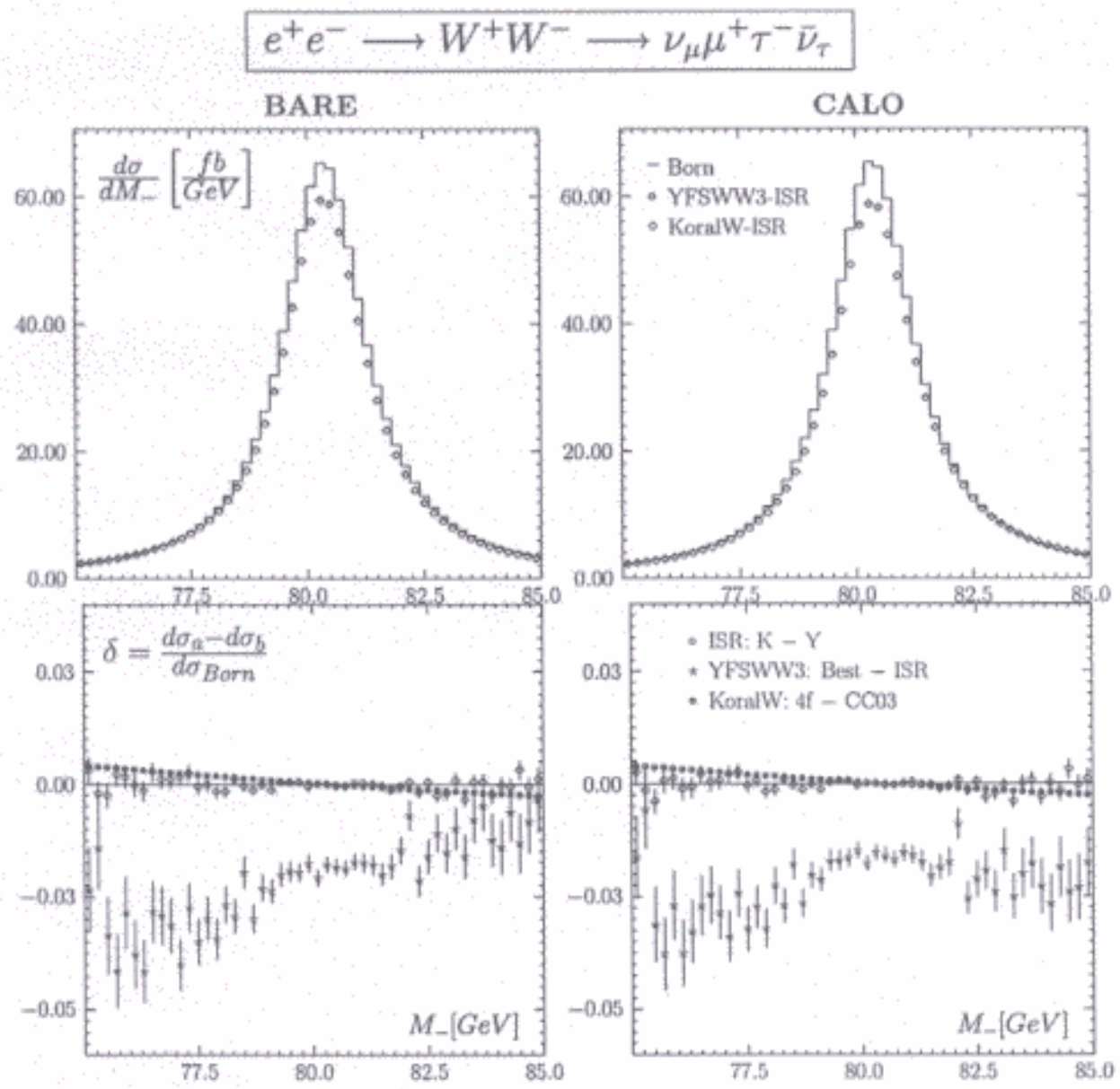


Figure 4: Distributions of the  $W^-$  invariant mass for BARE and CALO acceptances at  $E_{CM} = 200 \text{ GeV}$ .

200 GeV

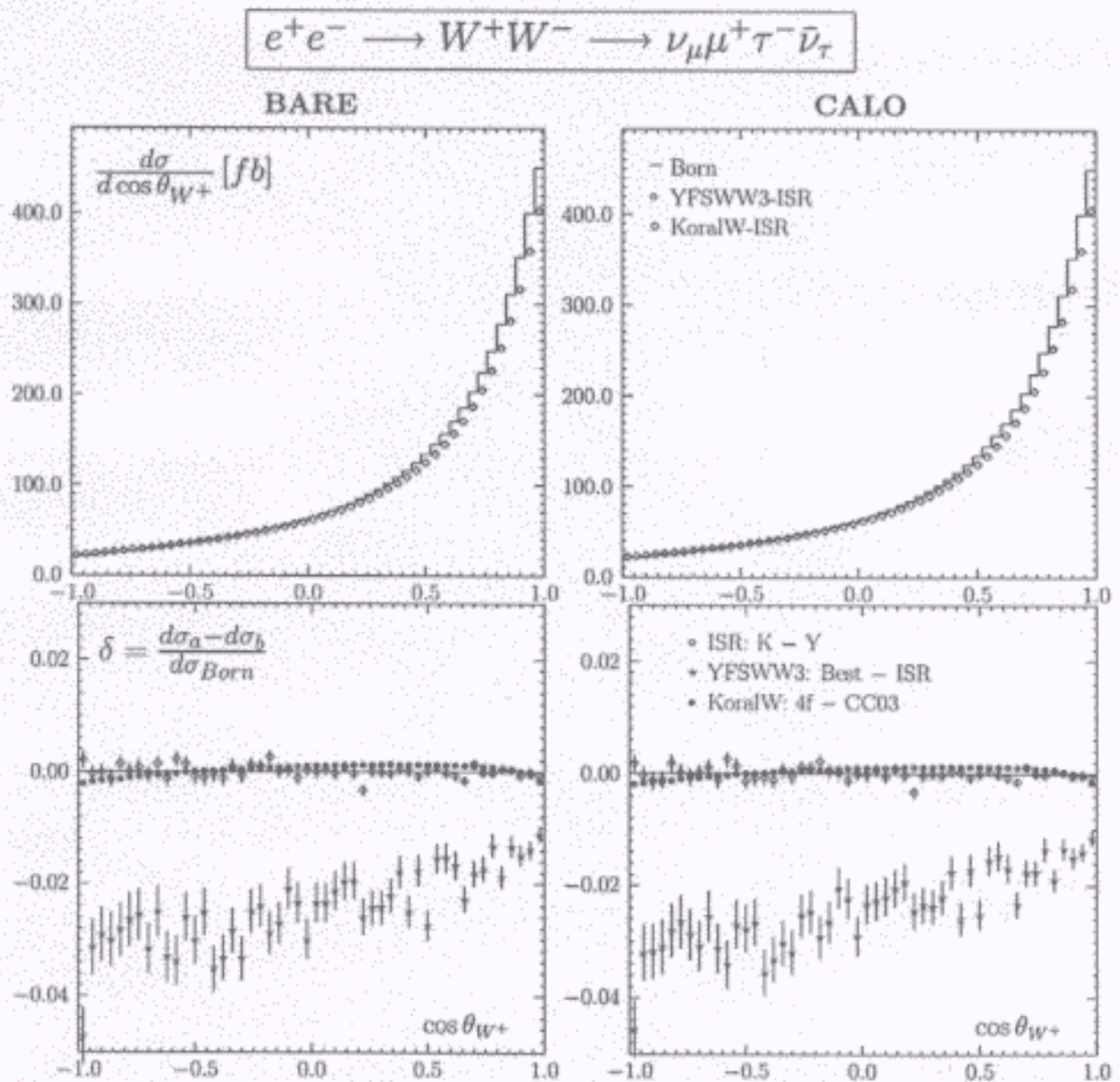


Figure 5: Distributions of cosine of the  $W^+$  polar angle w.r.t. the  $e^+$  beam for BARE and CALO acceptances at  $E_{CM} = 200\text{ GeV}$ .



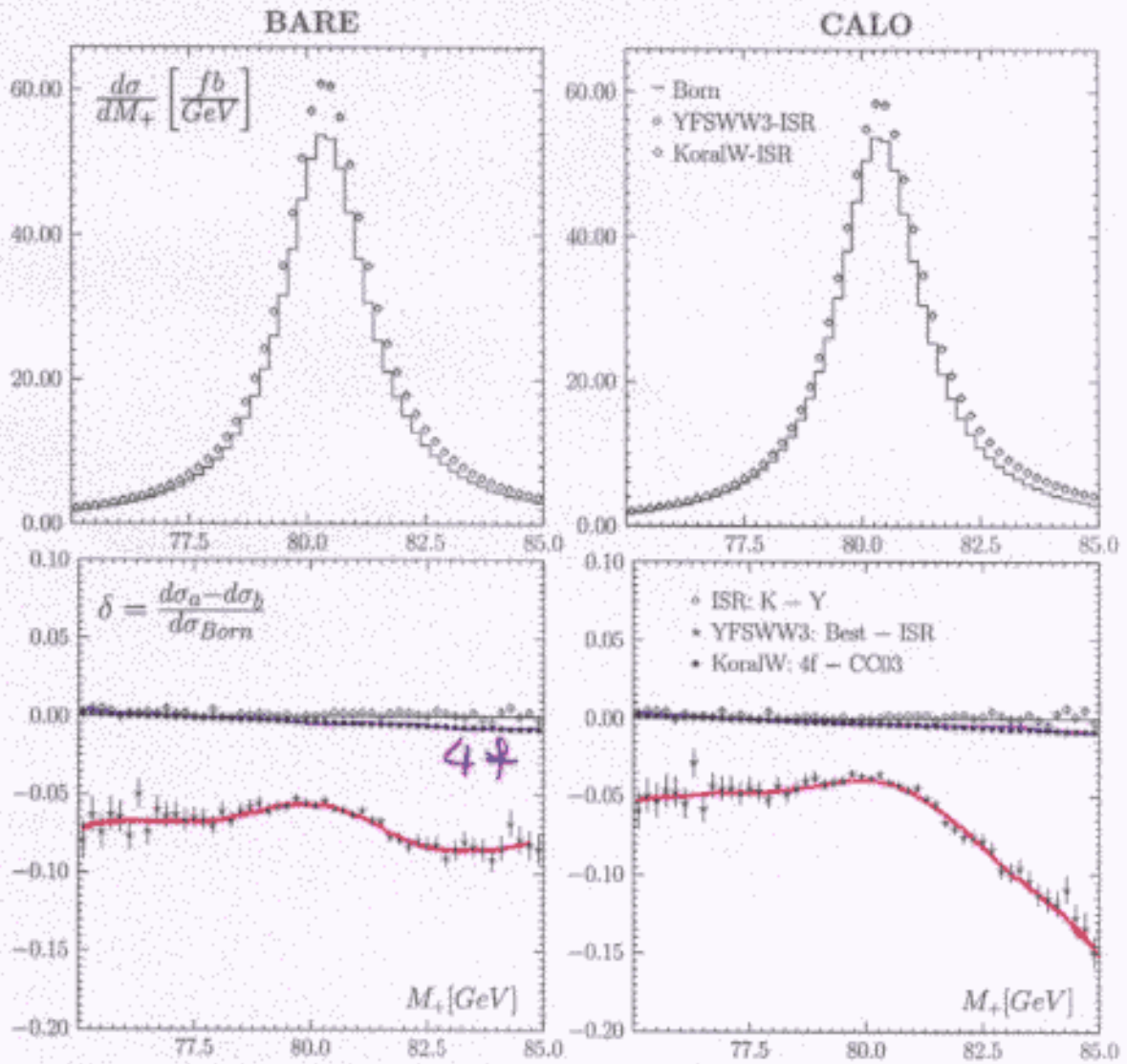
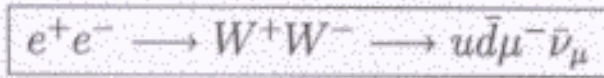


Figure 3: Distributions of the  $W^+$  invariant mass for BARE and CALO acceptances at  $E_{CM} = 500 \text{ GeV}$ .

$$\Theta(\alpha) \gg 4f$$

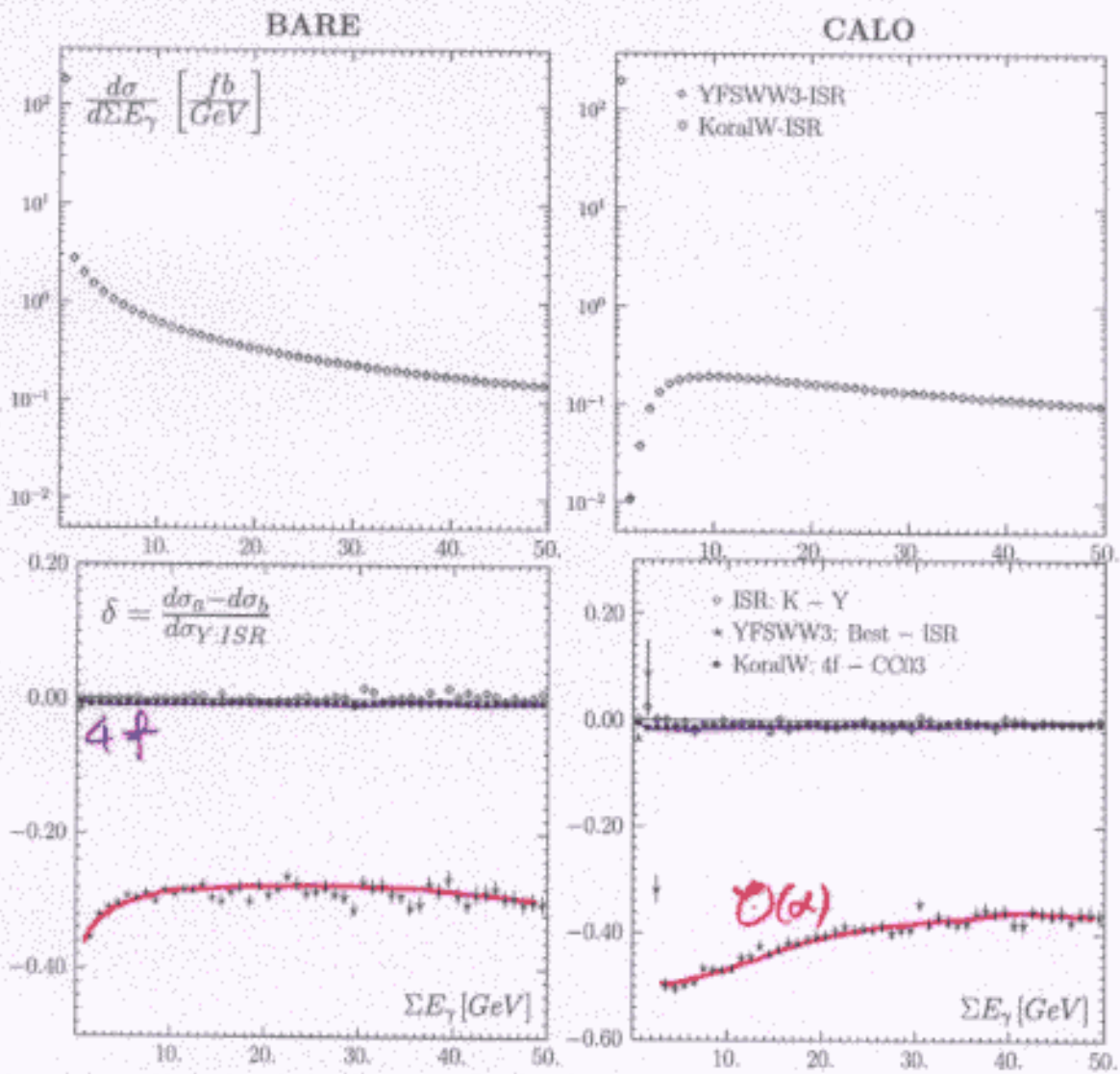
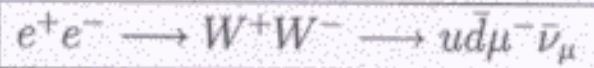


Figure 8: Distributions of the sum of the photon energy for BARE and CALO acceptances at  $E_{CM} = 500 \text{ GeV}$ .

$\mathcal{O}(\alpha) \gg 4f$



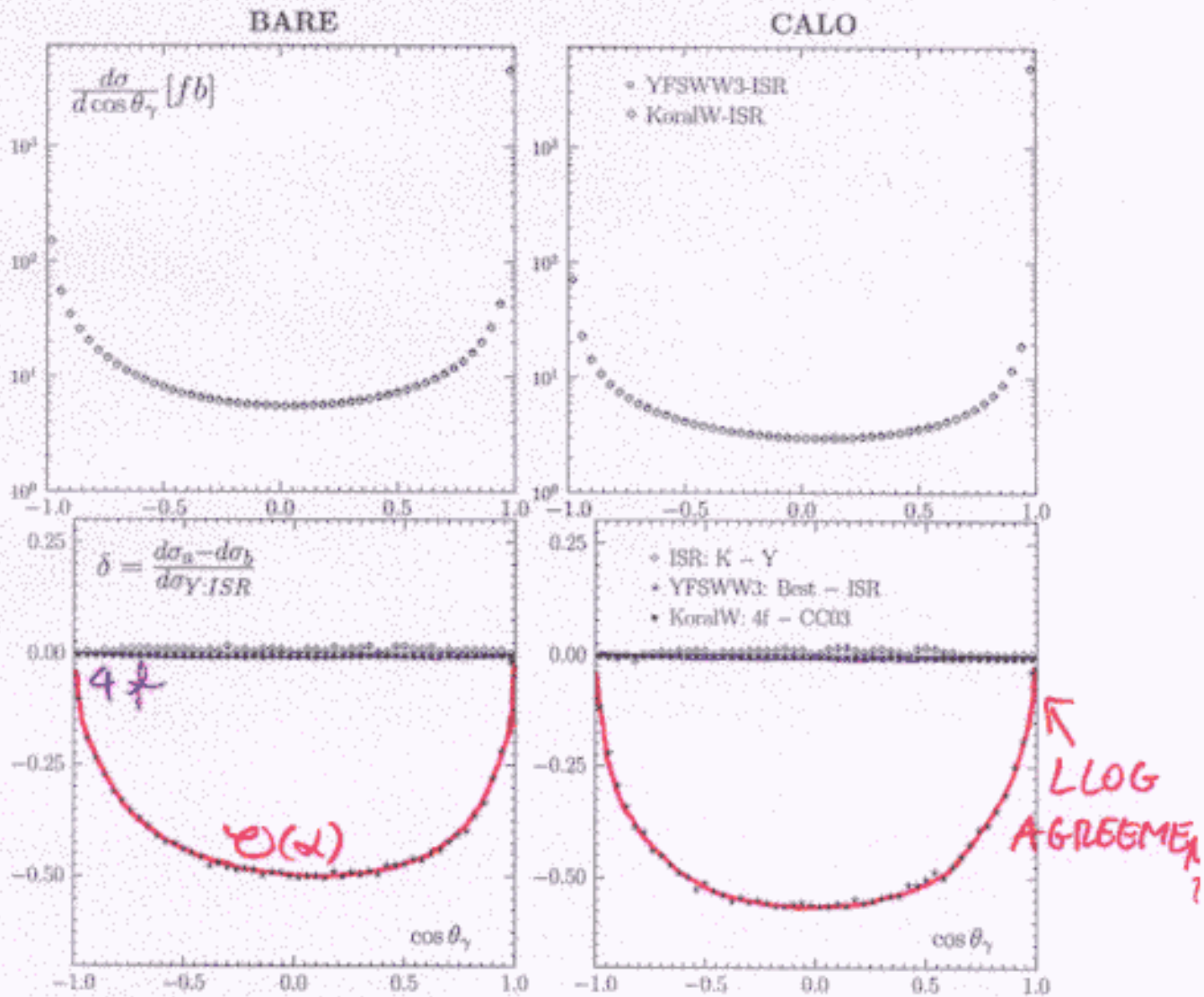
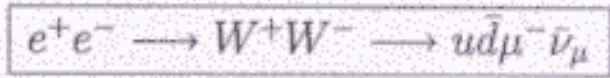


Figure 9: Distributions of cosine of the hardest photon polar angle w.r.t. the  $e^+$  beam for BARE and CALO acceptances at  $E_{CM} = 500 \text{ GeV}$ .

$$O(\alpha) \gg 4f$$

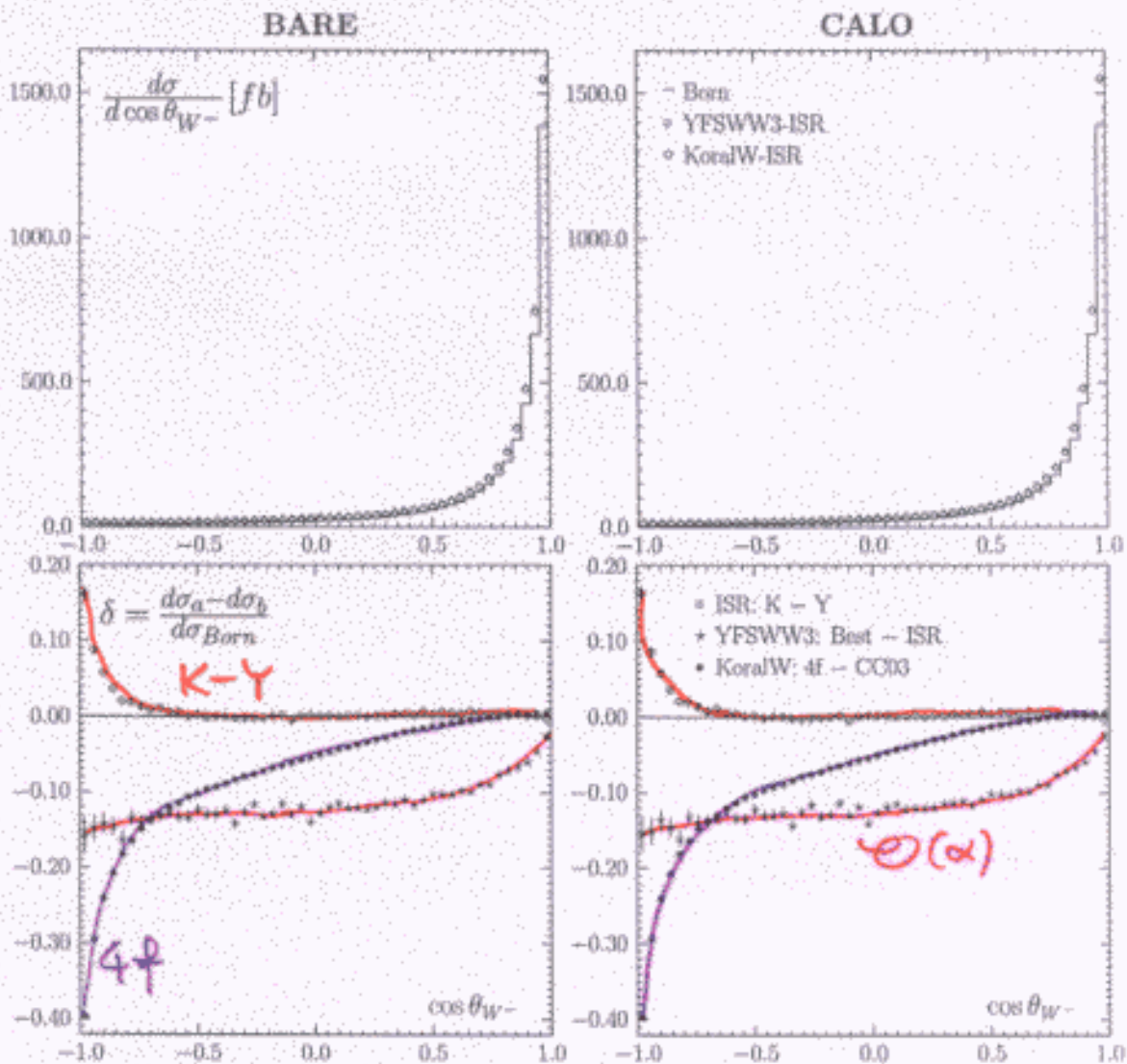
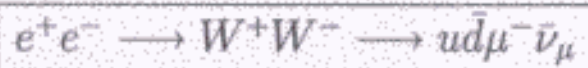


Figure 6: Distributions of cosine of the  $W^-$  polar angle w.r.t. the  $e^-$  beam for BARE and CALO acceptances at  $E_{CM} = 500 \text{ GeV}$ .

$$\Theta(\alpha) \sim 4f$$

EXTRAPOLATION  
 K-Y  $\rightarrow$  CC03 REDUCTION PROCEDURES  
 DIFFERENCE



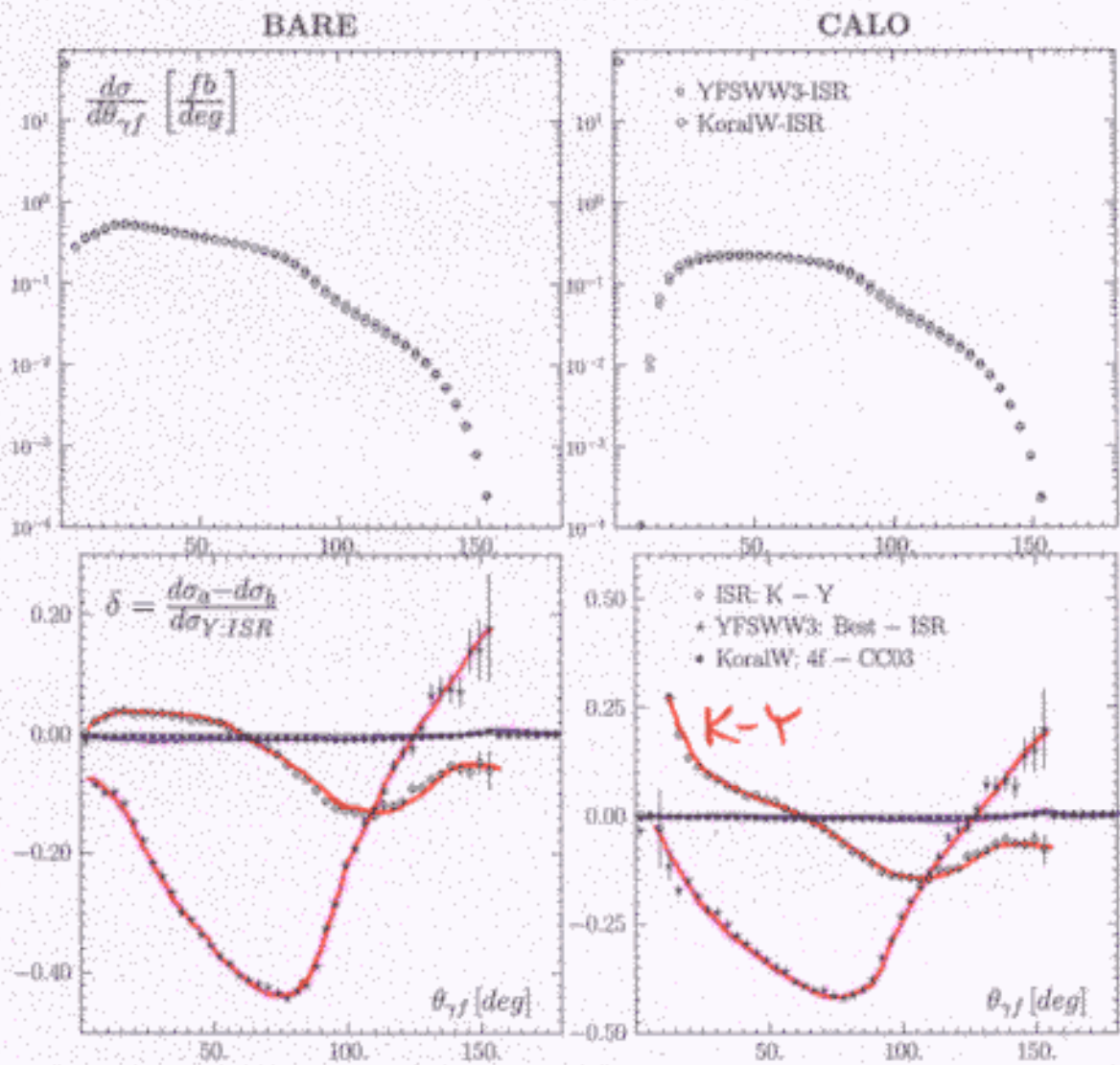
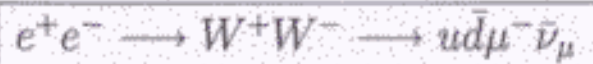


Figure 10: Distributions of the angle between the hardest photon and the nearest final state charged fermion for BARE and CALO acceptances at  $E_{CM} = 500 \text{ GeV}$ .

EXTRAPOLATION  
 K-Y  $\rightarrow$  CC $\phi$ 3 REDUCTION PROCEDURES  
 DIFFERENCE

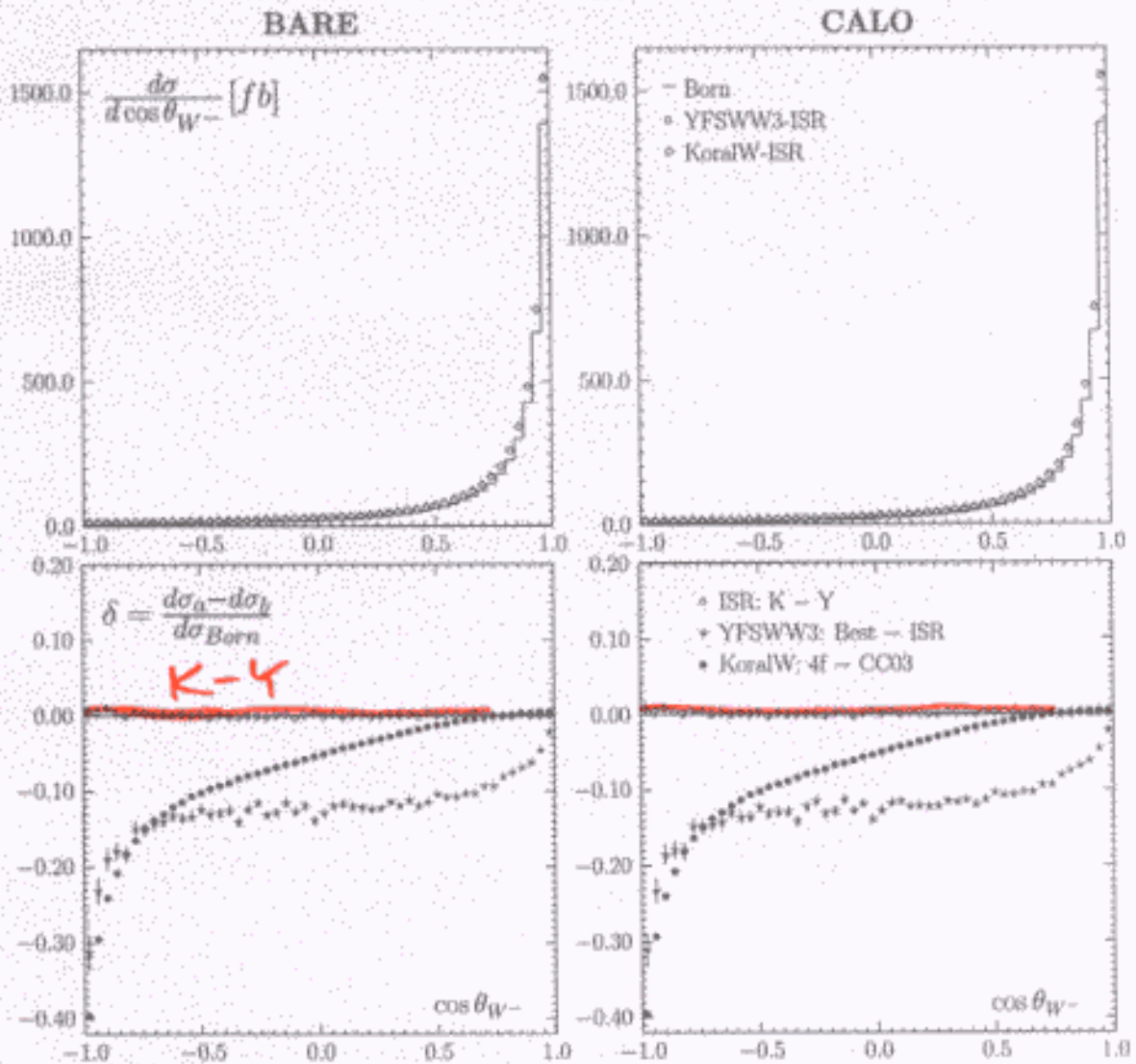
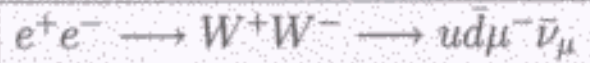


Figure 6: Distributions of cosine of the  $W^-$  polar angle w.r.t. the  $e^-$  beam for BARE and CALO acceptances at  $E_{CM} = 500 \text{ GeV}$ .

EXTRAPOLATION  
DIFFERENT REDUCTION PROCEDURE  
IN KORALW



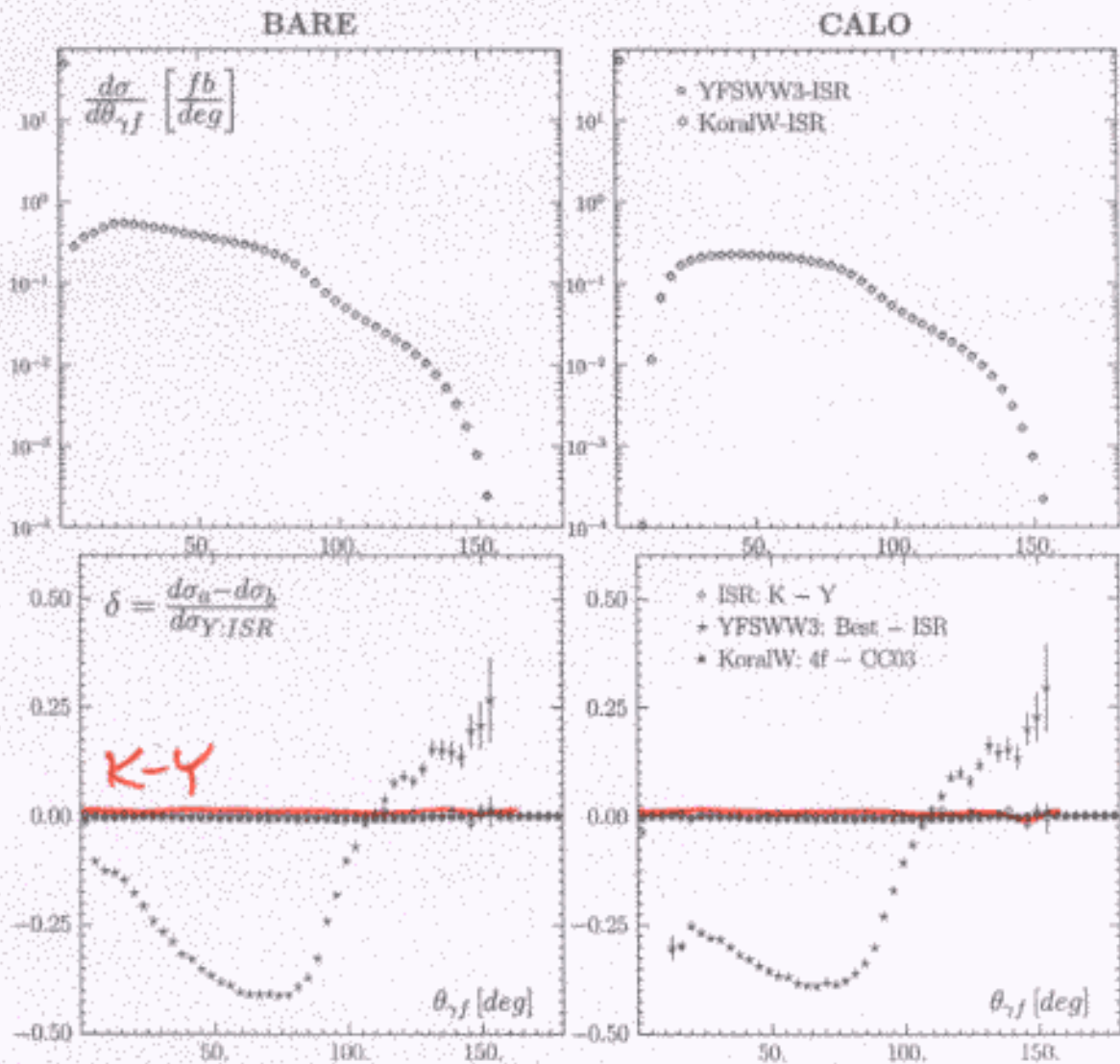
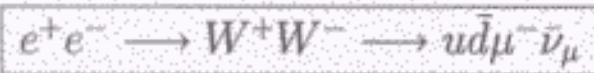


Figure 10: Distributions of the angle between the hardest photon and the nearest final state charged fermion for BARE and CALO acceptances at  $E_{CM} = 500 GeV$ .

EXTRAPOLATION  
DIFFERENT REDUCTION PROCEDURE  
IN KORALW

# CONCLUSIONS

- ① TECHNICAL PRECISION OF  
YFSWW - KORALW COMPARISONS

$$2 \times 10^{-4} \quad (\text{TOTAL X-SECT})$$

- ②  $\mathcal{O}(\alpha)$  CORR. DOMINATES  $4\text{f}$  CORR.  
**BUT** FOR SOME ANGULAR DISTRIBS.  
 $4\text{f}$  COMPARABLE TO  $\mathcal{O}(\alpha)$  BOTH  
AT 161 GeV AND 500 GeV.

**BUT** IN TOTAL X-SECT.  $4\text{f}$  CORR.  
SMALLER DUE TO ANGULAR  
CANCELLATIONS.

- ③ AN INTRINSIC FREEDOM/UNCERTAINTY  
PRESENT IN BREMSSTRAHLUNG  
SIMULATIONS DUE TO  
"EXTRAPOLATION PROCEDURE"  
(UNDEFINED RESIDUAL SUB-LEADING  
NON-IR TERMS)