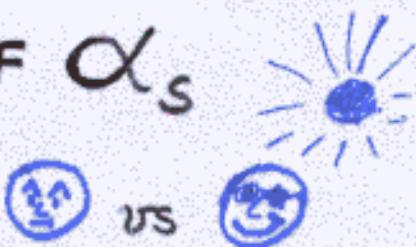


DETERMINATION OF α_s

AT GIGA-Z



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Outline

- Z observables vs α_s
- ADLO performances
- Improvement forecasts :
 - $q\bar{q}$ selection
 - $\mu^+\mu^- \neq e^+e^-$ selections
 - luminosity determination
- TESLA potential for $\alpha_s(M_Z^2)$
- Summary - Conclusion

? how much may ΔR_l be reduced ?

? which improvement from Γ_Z and σ_0^h ?

Motivations for α_s determinations

- Fundamental parameter
- Essential to test gauge structure of strong interactions
- Allows to test hypotheses on nature well above \exp^{lat} energy range (GUT)
- Enters theoretical prediction of many observables measured/used in future \exp^{ts}

Sensitivity of Z Observables to α_s

- $\Gamma_h = \Gamma_h^0 (1 + \delta_{QCD})$ where $\delta_{QCD} \sim \frac{\alpha_s}{\pi} \lesssim 4\%$

$$\hookrightarrow R_\ell = \frac{\Gamma_h}{\Gamma_\ell}$$

- $\Gamma_Z = \Gamma_L + \Gamma_h = \Gamma_Z^0 + \Gamma_h^0 \delta_{QCD} \simeq \Gamma_Z^0 (1 + \underbrace{0.7 \delta_{QCD}}_{\lesssim 3\%})$

$$\bullet \sigma_0^h = \frac{12\pi \Gamma_\ell}{M_Z^2} \cdot \frac{\Gamma_h^0 (1 + \delta_{QCD})}{[\Gamma_L + \Gamma_h^0 (1 + \delta_{QCD})]^2} = \frac{12\pi \Gamma_\ell \Gamma_h^0}{M_Z^2 (\Gamma_Z^0)^2} \cdot \frac{1 + \delta_{QCD}}{1 + 2 \left(\frac{\Gamma_L \cdot \Gamma_h^0}{(\Gamma_Z^0)^2} + \left(\frac{\Gamma_h^0}{\Gamma_Z^0} \right)^2 \right) \delta_{QCD}}$$

$$\simeq \sigma_0^{h0} \cdot \frac{1 + \delta_{QCD}}{1 + 1.4 \delta_{QCD}} \simeq \sigma_0^{h0} \underbrace{(1 - 0.4 \delta_{QCD})}_{\sim 1.5\%}$$

goals :	R_ℓ	$\Gamma_Z [\text{MeV}]$	$\sigma_0^h [\text{nb}]$
S.M.	20.740	2495.7	41.479
ADLO	20.767 ± 0.025	2495.2 ± 2.3	41.541 ± 0.037
QCD corr.	~ 0.760	~ 65	~ 0.640
$\delta \alpha_s = 0.001 \Rightarrow$	~ 0.006	~ 0.55	~ 0.005

$$\overset{\text{ADLO}}{\alpha_s(M_Z^2)} = 0.1200 \pm 0.0024 + \frac{0.0029 - 0.000}{M_H = 100 \overset{+900}{-0} \text{ GeV}}$$

ADLO performances

$$\bullet \frac{\Delta E_h}{E_h} \oplus \frac{\Delta bg_h}{bg_h} = 0.04 - 0.10\%$$

$$\bullet \frac{\Delta E_\mu}{E_\mu} \oplus \frac{\Delta bg_\mu}{bg_\mu} = 0.09 - 0.31\%$$

$$\bullet \frac{\Delta E_\tau}{E_\tau} \oplus \frac{\Delta bg_\tau}{bg_\tau} = 0.18 - 0.65\%$$

$$\bullet \frac{\Delta L_{\text{syst}}^{\text{exp}}}{L} = 0.033 - 0.09\%$$

$$\bullet \frac{\Delta L_{\text{syst}}^{\text{th}}}{L} = 0.054\%$$

$\mu^+\mu^-$ selection of ALEPH

$$\epsilon_{\text{sel}} \simeq 84\%$$

$$\text{present } \Delta_{\text{syst}}^A \sim \Delta_{\text{stat}}^{\text{TESLA}} \times 10$$

→ reduction is crucial

Source	$\Delta\sigma/\sigma (\%)$
Acceptance	0.05
Momentum calibration	0.006 (0.009)
Momentum resolution	0.005
Photon energy	0.05
Radiative events	0.05
Muon identification	$\simeq 0.001$ (0.02)
Monte Carlo statistics	0.06
Total	0.10 (0.11)

track. eff.
→ $\cos\theta^*$ cut
ISR/FSR simul.

} $\mu^+\mu^- j\gamma$

- tracking eff. can be better controlled
 - ISR/FSR simul. will improve (mat. budget!)
 - acceptance knowledge ↑ with stat.
 - $\mu^+\mu^- j\gamma$ MC " " " " & time
 - ECAL response for $j\gamma$ will be \gg LEP calo.
- Δ_{syst} can be reduced by 3-5(?)

$\tau^+\tau^-$ selection of ALEPH

$$\epsilon_{sel} \approx 94\%$$

"price" of
 ϵ_{sel} det. on

Source	A	B	$A \oplus B$
Acceptance	0.03	0.04	0.05
Efficiency			
Preselection cuts	0.05	0.04	0.07
$q\bar{q}$ cuts	0.06	0.09	0.11
Bhabha cuts	0.03	0.04	0.05
Dimuon cuts	0.03	0.03	0.05
Total Efficiency	0.08	0.11	0.14
Background			
$\gamma\gamma$	0.04	-	0.04
$q\bar{q}$	0.04	-	0.04
Bhabha	0.05	0.01	0.05
Dimuon	0.02	0.01	0.02
Total background	0.07	0.01	0.08
MC statistics	-	0.07	0.07
Total	0.12	0.13	0.18

scales $\sim \frac{1}{\sqrt{N}}$

many uncertainty components will follow
 the stat. accuracy of their estimate ...
 → may improve by $>> 3$ (?)

Multi-Hadron selection of L3

$$\epsilon_{\text{sel}} \approx 99.5\% \text{ for } \sqrt{s}' > 0.1\sqrt{s}$$

Source		1993 – 1995
Monte Carlo statistics	[‰]	0.04 – 0.10
Acceptance = fragm. ^{on}	[‰]	<u>0.21</u>
Selection cuts → move cut values	[‰]	<u>0.30</u>
Trigger	[‰]	<u>0.12</u>
Total scale	[‰]	0.39 – 0.40
Non-resonant background [pb]		<u>3</u>
Detector noise [pb]		<u>1</u>
Total absolute [pb]		3.2

reduced by
higher stat.
→ better
tuning

better constrai
from higher
stat.

Increase of stat by factor 200

⇒ all these uncertainties should go down
by at least a factor 2-3 (possibly 5!)

Luminosity determination of OPAL

Rad. Metrology	$1.4 \cdot 10^{-4}$	
Inner Anchor	$(1.4 \oplus 0.2) \cdot 10^{-4}$	
Background	$\underline{(0.8 \oplus 0.8) \cdot 10^{-4}}$	*
Total External	$(2.2 \oplus 0.8) \cdot 10^{-4}$	
Energy	$(1.8 \oplus 0.1) \cdot 10^{-4}$	*
Beam param.	$(0.6 \oplus 0.6) \cdot 10^{-4}$	*
clustering	$1.0 \cdot 10^{-4}$	
MC stat.	$\underline{(0.8 \oplus 0.3) \cdot 10^{-4}}$	
Total Simul. ^{on}	$(2.3 \oplus 0.6) \cdot 10^{-4}$	
ALL	$(3.2 \oplus 1.0) \cdot 10^{-4}$	

- "geometrical" systematics will →
- background (beam related) and ΔE_{LEP} may be difficult to squeeze
→ $(2 \oplus 1) \cdot 10^{-4}$?

Precision expected from TESLA (1)

Δ_{stat}

- $4 \cdot 10^7 \mu\bar{\mu} + 4 \cdot 10^7 c\bar{c} \Rightarrow \frac{1}{\sqrt{N_{\mu\bar{\mu}+c\bar{c}}}} \simeq 1.1 \cdot 10^{-4}$ (add e^+e^- in barrel?)
- $\sim 10^9 q\bar{q} \Rightarrow \frac{1}{\sqrt{N_{q\bar{q}}}} \simeq 3 \cdot 10^{-5}$ (negligible)
- $2-3 \cdot 10^9$ small angle Bhabha $\rightarrow \Delta L_{\text{stat}}$ neglected

Δ_{syst}

- $\Delta_{\text{syst}}^{q\bar{q}} = \pm 3.9 \cdot 10^{-4}$ from L3
- $\Delta_{\text{syst}}^{l^+l^-} = \pm 9 \cdot 10^{-4}$ from ALEPH
- $\Delta L_{\text{syst}}^{\text{exp}} = \pm 3.3 \cdot 10^{-4}$ from OPAL
- $\Delta L_{\text{syst}}^{\text{th}} = \pm 5.4 \cdot 10^{-4}$

→ consider improvements of:

- $\Delta_{\text{syst}}^{q\bar{q}}$, $\Delta_{\text{syst}}^{l^+l^-}$, $\Delta L_{\text{syst}}^{\text{exp}}$ by factor 3 and 5
- $\Delta L_{\text{syst}}^{\text{th}}$ by factor 2 and 3

Precision expected from TESLA (2)

R_L	$\Delta_{\text{syst}}^{99\bar{9}}$	$\Delta_{\text{syst}}^{l^+l^-}$	$\Delta_{\text{stat}}^{l^+l^-}$	ΔR_L	$\Delta \alpha_s$
$[10^{-4}]$	3.9	9	1.1	9.9	0.003
$\frac{1}{3}$	$\frac{1}{3}$	-	-	3.4	0.0011
$\frac{1}{5}$	$\frac{1}{5}$	-	-	2.2	0.0007

σ_0^h	$\Delta_{\text{syst}}^{99\bar{9}}$	$\Delta L_{\text{syst}}^{\text{exp}}$	$\Delta L_{\text{syst}}^{\text{th}}$	$\Delta \sigma_0^h$	$\Delta \alpha_s$
$[10^{-4}]$	3.9	3.3	5.6	7.5	0.0058
$\frac{1}{3}$	$\frac{1}{3}$	-	-	5.7	0.0045
$\frac{1}{3}$	$\frac{1}{3}$	$\frac{1}{3}$	$\frac{1}{3}$	2.5	0.0019
$\frac{1}{5}$	$\frac{1}{5}$	$\frac{1}{5}$	$\frac{1}{3}$	2.1	0.0016
$\frac{1}{5}$	-	-	$\frac{1}{3}$	3.9	0.0030

• Γ_Z $\Delta \Gamma_Z = \pm 1 \text{ MeV} \rightarrow \Delta \alpha_s = \pm 0.002$ (if no effect from $\Delta \alpha_s$)

• combination (ignoring correlations)

$\frac{\Delta R_L}{R_L} [10^{-4}]$	9.9	3.4	3.4	2.2	3.4
$\frac{\Delta \sigma_0^h}{\sigma_0^h} [10^{-4}]$	7.5	5.7	2.5	2.1	3.9
$\Delta \alpha_s (M_Z^2)$	± 0.0016	0.0009_5	0.0009_8	0.0006	0.0009_2
best component	0.002	0.0011	0.0011	0.0007	0.0011

Summary - Conclusion

- $\alpha_s(M_Z^2)$ likely to be determined at Giga-Z with $\sim 0.0015 - 0.0005$ accuracy
- the factor 2-5 improvement w.r.t. to LEP relies on:
 - $\Delta_{\text{stat}} < \frac{\Delta_{\text{stat}}^{\text{LEP}}}{10}$
 - better detector performance (tracking, calo.)
 - better detector understanding (data stat., mat. budget, redundancy, align t , ...)
 - more accurate simulations (gener., detector)
 - more precise S.M. predictions: $m_t, M_H \rightarrow \Delta g \sim 0$ (Δg_{em}^h , H.O.?)

⇒ will need substantial effort!

- R_ℓ will be the most sensitive observable;
 Γ_2 and Γ_0^h will reduce Δg by $\lesssim 20\%$
- present study needs refinement → end '00