

m_w FROM POLARISED THRESHOLD SCAN

G.W. WILSON

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UPDATE

- July slide

- 1) Polarisation systematics..
- 2) Reasonable background spin model
- 3) Control of bkgd. spin structure

Will produce a nice plot - but presently mostly concerned with demonstrating that this measurement is indeed feasible!

(N.B. L , σ_b , ϵ , E_b^* , beamstrahlung $\rightarrow \checkmark$)

* If $\langle \sqrt{s} \rangle$ from Z^0 works.

M_w FROM POLARISED THRESHOLD SCAN

STATUS REPORT/PLANS

G.W. WILSON
14/7/00

Available documentation.

- 1) Sitges talk (on web)
- 2) Sitges proceedings

BASIC OPEN ISSUES (ie. potential show stoppers)

- 1) E_{beam} error
- 2) Beam-strahlung / energy spread of $\frac{dL}{dIs}$ (Is current TESLA design in CIRC inc. ΔE_b ?)
- * 3) Polarisation systematics. (Use in situ $Z(\gamma)$ events to measure P ?)
- 4) Theoretical errors. (presumably scan helps....)
- 5) Control of bkgd. spin structure.

REFINEMENTS

- 1) Reasonable bkgd. spin model. *
- 2) Upgrade event selection performance wrt LEP2.
- 3) SIMDET studies? May be useful for evaluating whether beamstrahlung can be an event selection issue. (Is SIMDET updated now?)

POLARISATION SYSTEMATICS

GWW, Sitges proceedings : $\Delta m_W = 4.9 \text{ MeV}$ $\frac{\Delta P}{P} = 0.25\%$
 $\Delta m_W = 4.6 \text{ MeV}$ $\frac{\Delta P}{P} = 0\%$
 $\Delta m_W = ?$ $\frac{\Delta P}{P}$ larger
 convergence problems in the fits.

The solution

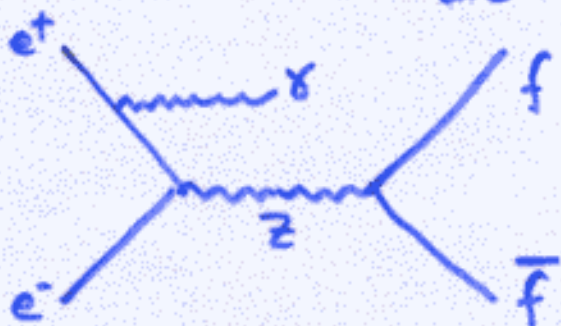
use $e^+e^- \rightarrow Z(\gamma)$ at $\sqrt{s} = 161 \text{ GeV}$
 to measure $P_{e^+}, P_{e^-}, \sigma_{RR}, A_{LR}^{Z(\gamma)}$
 with "Blondel scheme".

discussed previously by K. Mönig.

At $\sqrt{s} = 161 \text{ GeV}$,

$$\sigma_{\sqrt{s} \lesssim 120 \text{ GeV}} \approx 100 \text{ pb}$$

and Zee, WW, ZZ backgrounds are small



$$f = e, \mu$$

$$m_{f\bar{f}} \approx m_Z$$

10^7 radiative return events
in 100 fb^{-1} !

POLARISATION SYSTEMATICS

e^+ : L, R, 0

e^- : L, R, 0

so 6 ABSOLUTE POLARISATIONS TO MEASURE

P_L^+ , P_R^+ , P_0^+

P_L^- , P_R^- , P_0^-

PROPOSED HELICITY COMBINATION SHARING

$e_R^+ e_L^-$: 70% WW enhancement

$e_L^+ e_R^-$: 23% WW suppression.

L L
R R
0 0
L 0
R 0
0 L
0 R

} 1% each.

⇒ CAN MEASURE ALL SIX POLARISATIONS FROM PP

MINUIT TASK: FIT W MASS TO NUMBER OF EVENTS IN EACH CHANNEL

FCN= 211.4589 FROM MINOS STATUS=SUCCESSFUL 1984 CALLS 3004 TOTAL
 EDM= .70E-06 STRATEGY= 1 ERROR MATRIX ACCURATE

EXT NO.	PARAMETER NAME	VALUE	PARABOLIC		MINOS ERRORS	
			ERROR	NEGATIVE	POSITIVE	
1	WMASS	80.383	.52990E-02	-.53041E-02	.52938E-02	
2	BKGLL	.19168E-01	.10968E-02	-.10973E-02	.10963E-02	
3	BKGLQ	.80181E-01	.30944E-02	-.30967E-02	.30919E-02	
4	BKGQQ	.39732	.47314E-02	-.47324E-02	.47302E-02	
5	FLUMI	1.0000	.23151E-02	-.23138E-02	.23161E-02	
6	REFFLL	1.0023	.21894E-02	-.21892E-02	.21897E-02	
7	REFFLQ	.99853	.19873E-02	-.19867E-02	.19878E-02	
8	REFFQQ	.99922	.20900E-02	-.20896E-02	.20903E-02	
9	ALPHAS	.12000	constant			
10	ALRLL	.15000	constant			
11	ALRLQ	.33000	constant			
12	ALRQQ	.49000	constant			
13	ALRMZ	.15940	.20791E-02	-.20803E-02	.20776E-02	
14	PELL	.78225	.50929E-02	-.50870E-02	.50950E-02	
15	PELR	.82069	.28556E-02	-.28522E-02	.28577E-02	
16	PELZ	.60436E-02	.33653E-02	-.33673E-02	.33629E-02	
17	PPOSL	.61810	.29707E-02	-.29685E-02	.29717E-02	
18	PPOSR	.58001	.37921E-02	-.37892E-02	.37927E-02	
19	PPOSZ	-.11807E-01	.28543E-02	-.28538E-02	.28547E-02	
20	XSR	99.948	.85487E-01	-.85206E-01	.85759E-01	

$\Rightarrow \Delta M_W = 5.3 \text{ MeV}$ for $\int L dt = 100 \text{ fb}^{-1}$
 assuming $\Gamma_W = f_{sm}(M_W)$, A_{LR} ^{has} known.

MINUIT TASK: FIT W MASS TO NUMBER OF EVENTS IN EACH CHANNEL

FCN= 199.8177 FROM MINOS STATUS=SUCCESSFUL 1104 CALLS 1749 TOTAL
 EDM= .74E 05 STRATEGY= 1 ERROR MATRIX ACCURATE

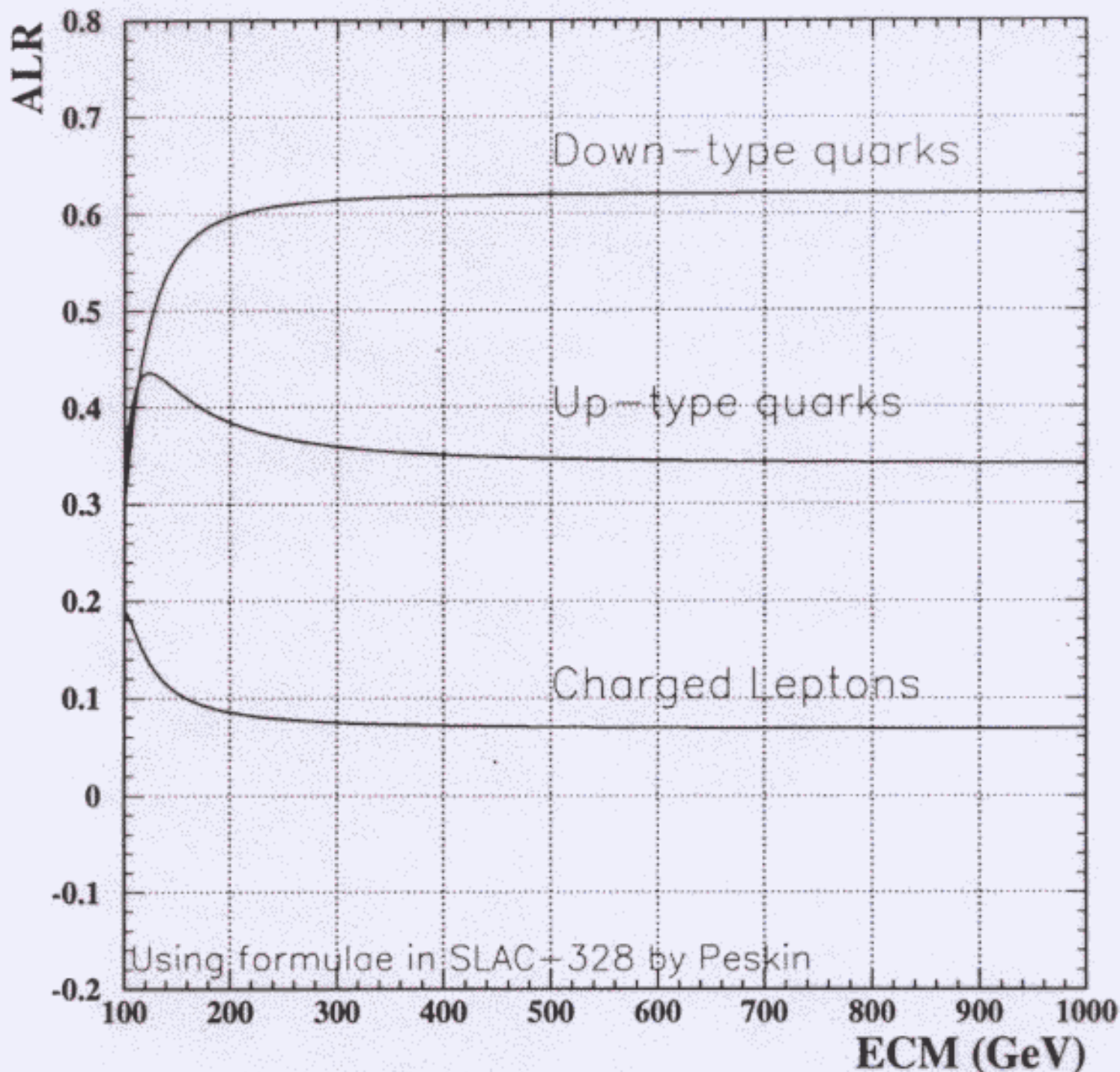
EXT PARAMETER NO.	NAME	VALUE	PARABOLIC ERROR	MINOS ERRORS NEGATIVE	POSITIVE
1	WMASS	80.382	.51967E-02	-.52023E-02	.51913E-02
2	BKGLL	.19312E-01	.11177E-02	-.11171E-02	.11182E-02
3	BKGLQ	.79986E-01	.30280E-02	-.30269E-02	.30290E-02
4	BKQQ	.39700	.46713E-02	-.46694E-02	.46731E-02
5	FLUMI	1.0007	.17910E-02	-.17901E-02	.17919E-02
6	REFFLL	1.0023	.21697E-02	-.21708E-02	.21685E-02
7	REFFLQ	.99879	.19045E-02	-.19042E-02	.19047E-02
8	REFFQQ	.99955	.20057E-02	-.20060E-02	.20053E-02
9	ALPHAS	.12000	constant		
10	ALRLL	.15000	constant		
11	ALRLQ	.33000	constant		
12	ALRQQ	.49000	constant		
13	ALRMZ	.16029	.33911E-03	-.33917E-03	.33904E-03
14	PELL	.80001	.22214E-02	-.22208E-02	.22215E-02
15	PELR	.80000	constant		
16	PELZ	.00000E+00	constant		
17	PPOSL	.59999	.18000E-02	-.17949E-02	.18048E-02
18	PPOSR	.60000	constant		
19	PPOSZ	.00000E+00	constant		
20	XSR	99.976	.53469E-01	-.53474E-01	.53464E-01

Also assume $P_L^+ = P_R^+$, $P_L^- = P_R^-$, $P_0^+ = P_0^- = 0$
 $\Rightarrow \Delta M_W = 5.2 \text{ MeV}$
 (so don't NEED these assumptions)

REASONABLE BKGD. SPIN MODEL?

HAD NEGLECTED A_{LR}^{bkgd} .

$$e^+e^- \rightarrow f\bar{f}$$



Thanks to Arnd Brandenburg for pointing out the fallacies in the Phys. Rep. polarization discussion!

Sensitivity to A_{LR}^{blyd}

<u>A_{LR}^{blyd}</u>	ΔM_W (MeV)
0	4.6
0.45	5.8
0.50	6.2
0.55	6.9

EVEN if we KNOW A_{LR}^{blyd} well
sensitivity degrades!

Assuming i)
$$\left. \begin{aligned} A_{LR}^{blyd}(\nu\nu) &= 0.15 \\ A_{LR}^{blyd}(\nu e\nu) &= 0.33 \\ A_{LR}^{blyd}(e e e) &= 0.49 \end{aligned} \right\} \Rightarrow \Delta M_W = 5.2 \text{ MeV}$$

If $A_{LR}^{blyd}(\nu\nu) = A_{LR}^{blyd}(\nu e\nu) = A_{LR}^{blyd}(e e e) = 0.5$

then $A_{LR} = 0.50 \pm 0.01 \rightarrow \pm 4 \text{ MeV}$
on M_W .

CONCLUSIONS

- RADIATIVE RETURNS SOLVE POLARISATION SYSTEMATIC ISSUES
- $A_{LR}^{bhad} \neq 0 \Rightarrow$ DEGRADES SENSITIVITY A BIT
- A_{LR}^{bhad} CONTROL IS AN OPEN ISSUE WHICH MAY ARGUE FOR A HIGHER PURITY SELECTION.
ie. $l\nu l\nu$
 $q\bar{q}e\nu$
 $q\bar{q}\mu\nu$
WITH LITTLE $e^+e^- \rightarrow q\bar{q}$ CONTAMINATION.