

A summary of experimental results at ICHEP04

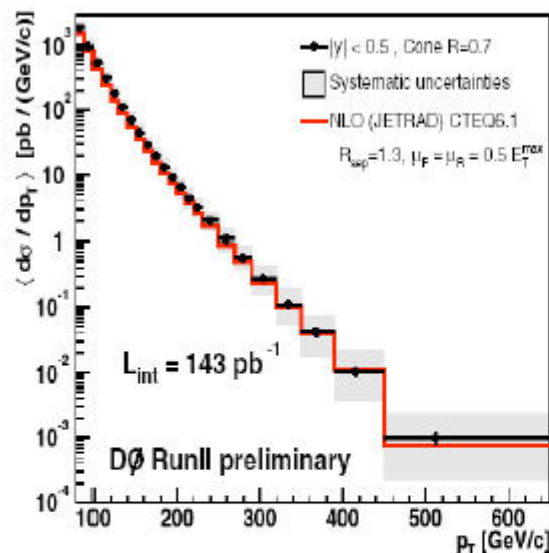
- QCD at Tevatron II
- Hadron spectroscopy
- Quark-Gluon plasma
- Quark mixing
- Neutrino mixing
- Electroweak (top mass)
- Higgs search
- Particle astrophysics and cosmology
- Final remarks
 - Not summarized: results from HERA, future colliders, detectors.. and many others.

Most slides
taken from
plenary talks:

My comments
are in red boxes

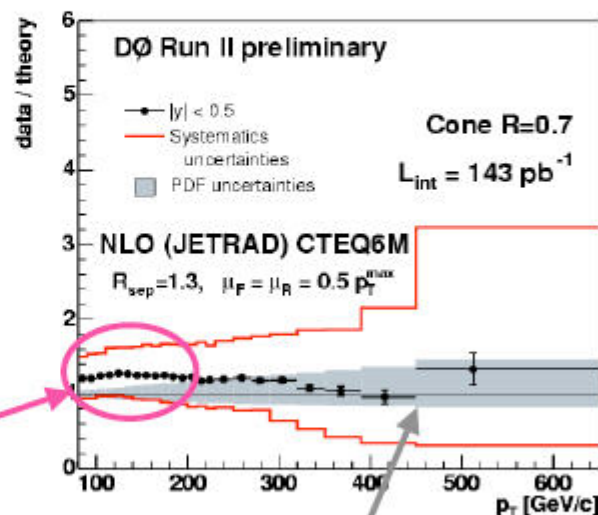
Apologies in advance for oversimplifications, omissions and mistakes.

Inclusive Jet P_T Cross section



Hadronization correction needed?

Agreement with theory within systematic uncertainties
(dominated by jet-energy scale)



NLO uncertainty due to gluon @ high x

August 21, 2004

Donatella Lucchesi

6

First jet cross-section results from the Tevatron II: Jet-energy scale unc. still large (should improve with more data)

Inclusive Jet Cross section: K_T algorithm

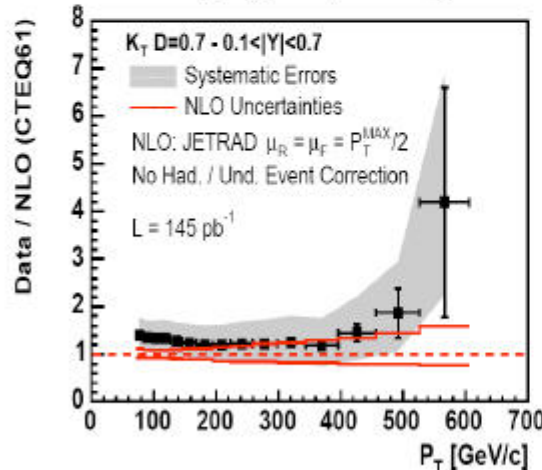
- Inclusive K_T algorithm

$$d_{ij} = \min(P_{T,i}^2, P_{T,j}^2) \frac{\Delta R^2}{D^2} \rightarrow (y_i - y_j)^2 + (\varphi_i - \varphi_j)^2$$

jet size

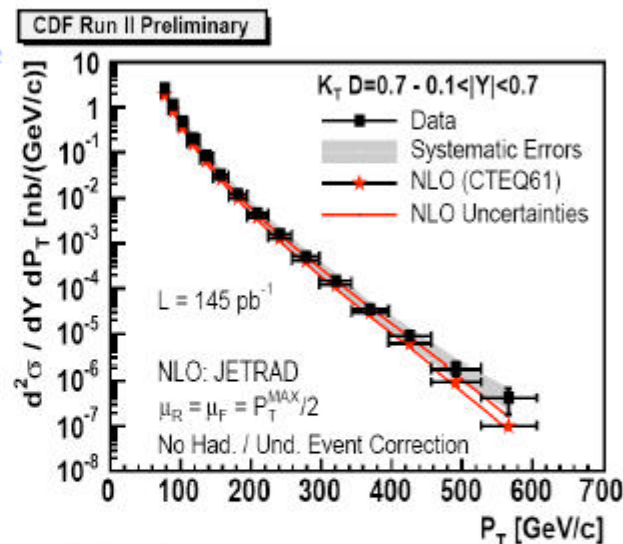
$$d_i = (P_{T,i})^2$$

- Infrared and collinear safe
- No merging / splitting



August 21, 2004

Donatella Lucchesi



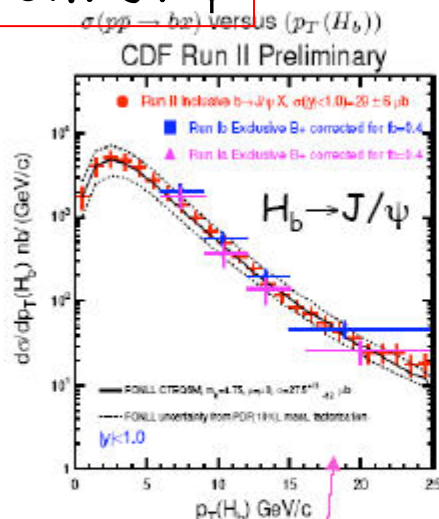
- Reasonable data-theory agreement
- NLO still needs to be corrected for Hadronization / Underlying Event

7

Tevatron is starting to use modern jet algorithms

B production derived from J/ψ

Beauty and Charm production at Tevatron

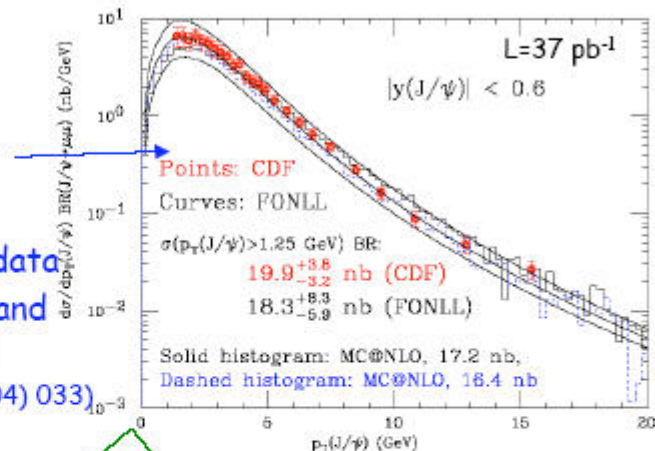


Data: $\sigma = 29 \pm 6 \mu\text{b}$,
 FONLL: $\sigma = 27.5$
 (CTEQ6M, $m_b = 4.75$, $\mu = \mu_0$)

See Paul De Jong talk

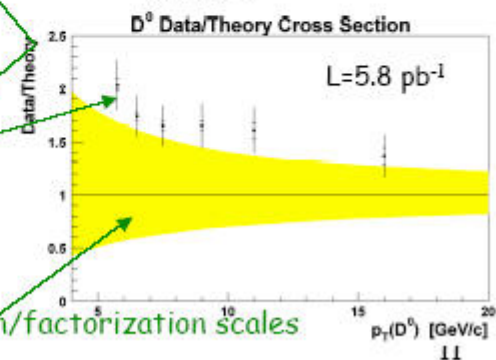
August 21, 2004

Cacciari,
 Frixione,
 Mangano,
 Nason,
 Ridolfi
 compared data
 to FONLL and
 MC@NLO:
 (JHEP07 (2004) 033)



$\sigma(pp \rightarrow D^0 X, |y| < 1.0, p_T > 5.5 \text{ GeV/c})$
 $= 13.3 \pm 0.2 (\text{stat.}) \pm 1.5 (\text{syst}) \mu\text{b}$

renormalization/factorization scales
 Donatella Lucchesi



Beauty and Charm production at Tevatron: agree with latest theory. Consistent picture with LEP and HERA?

Summary of Θ^+ pentaquark reports

Jin

Experiment	Mass(MeV)	Width(MeV)	Reaction	Production
SPRING-8	1540 ± 10	< 25	γn	nK^+
DIANA	1539 ± 2	< 9	$K^+ Xe$	$nK^+ \rightarrow pK_s^0$
CLAS-1	1542 ± 5	< 21	γd	nK^+
SAPHIR	$1540 \pm 4 \pm 2$	< 25	γp	nK^+
ITEP	1533 ± 5	< 20	$\nu CC, \bar{\nu} CC$	pK_s^0
CLAS-2	1555 ± 10	< 26	γp	nK^+
ALICE	$1532 \pm -$	$< -$	CC	pK_s^0
HERMES	$1528 \pm 2.6 \pm 2.1$	$17 \pm 9 \pm 3$	γd	pK_s^0
COSY-TOF	1530 ± 5	$< 18 \pm 4$	pp	$\Sigma^+ pK_s^0$
SVD-2	$1526 \pm 3 \pm 3$	< 24	pN	pK_s^0
JINR-1	1545.1 ± 12.0	16.3 ± 3.6	pC_3H_3	pK_s^0
ZEUS	$1521.5 \pm 1.5^{+2.8}_{-1.7}$	$6.1 \pm 1.6^{+2.0}_{-1.4}$	ep	$pK_s^0, \bar{p}K_s^0$
JINR-2	1541 ± 4	8 ± 4	np	nK^+
NA49	$1535 \pm -$	$-$	pp	pK_s^0

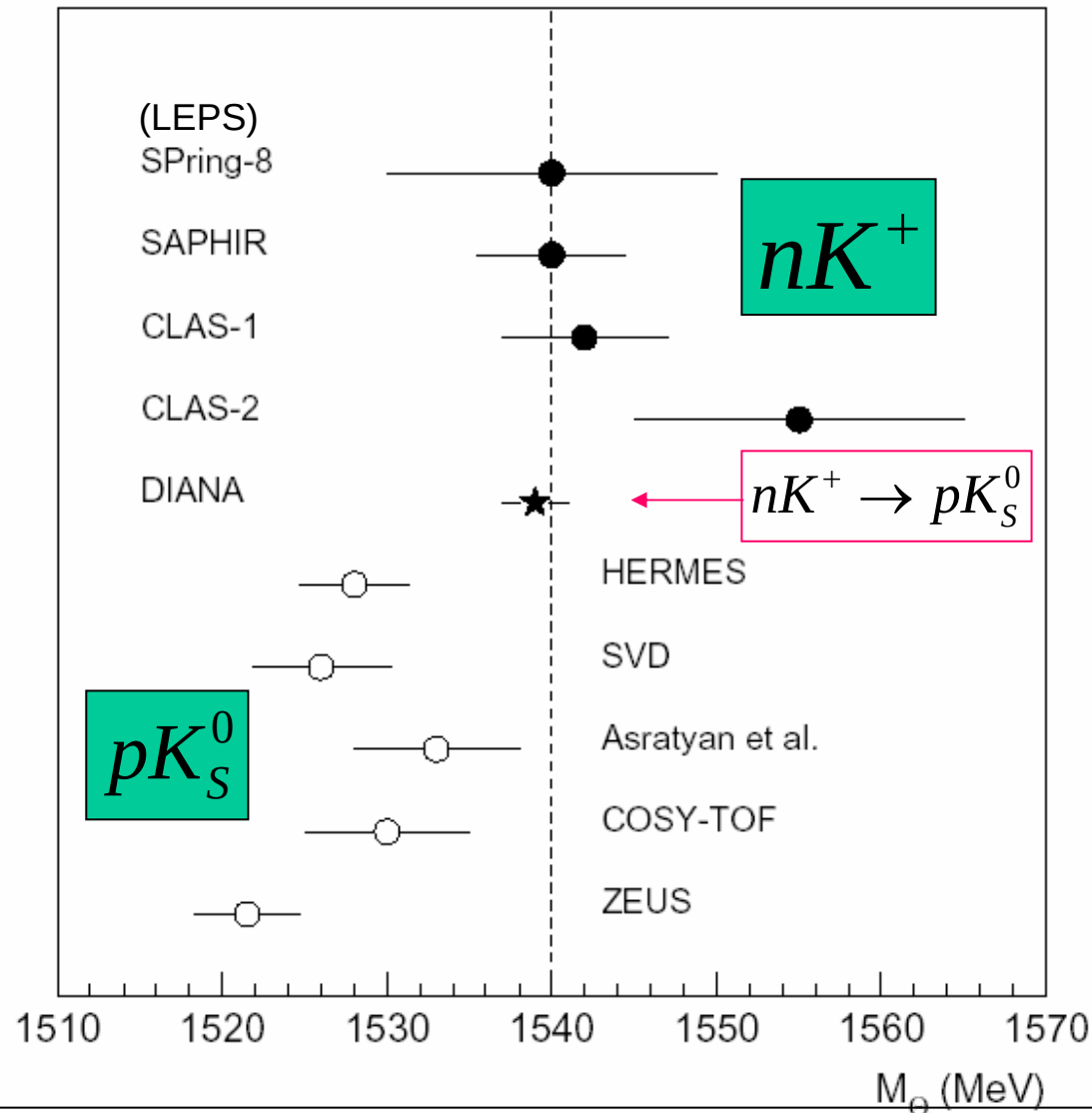
8 ± 4 MeV

Some speakers commented on a "width" inconsistency of Θ^+ (< 1 MeV from partial wave analysis of nK results. See review in PDG by Trilling.)

Hermes and ZEUS see significantly wider width than other positive expts. (?)

Close and Zhao

M_{Θ}



Masses of “two decay modes” appear to be inconsistent

Summary of negative pentaquark reports

Jin

Experiment	$\Theta^+(1540)$ ($uudd\bar{s}$)	$\Xi^{--}(1862)$ ($ddss\bar{s}$)	$D^{*-}p(3100)$ ($uudd\bar{c}$)	Reaction
→ HERA-B	NO	NO		$pA \rightarrow \Theta^+ X, \Xi^{--} X$
E690	NO	NO		$pp \rightarrow \Theta^+ X, \Xi^{--} X$
CDF	NO	NO	NO	$p\bar{p} \rightarrow \Theta^+ X, \Xi^{--} X, \Theta^c X$
HyperCP	NO			$\pi, K, p \rightarrow \Theta^+ X$
BaBar	NO	NO		$e^+e^- \rightarrow \Theta^+ X, \Xi^{--} X$
ZEUS	yes	NO	NO	$ep \rightarrow \Theta^+ X, \Xi^{--} X, \Theta^c X$
ALEPH	NO	NO	NO	$e^+e^- \rightarrow \Theta^+ X$
DELPHI	NO			$e^+e^- \rightarrow \Sigma^+ K^0 p$
→ PHENIX	NO			$AuAu \rightarrow \Theta^+ X$
FOCUS			NO	$\gamma A \rightarrow \Theta^c X$
→ BES	NO			$e^+e^- \rightarrow J/\Psi \rightarrow \Theta^+ \bar{\Theta}^-$

- Production rate (e.g. for $\Theta^+(1540)$)

- “Positive” experiments:

SAPHIR: $R_{\Lambda^*(1520)} \sim 0.3$

HERMES: $R_{\Lambda^*(1520)} \sim 1.6 \sim 3.5$

ZEUS: $R_{\Lambda^*(1520)} \sim 0.2$ (I estimated from $R_{\Lambda} \sim 0.04$)

SVD-2: $R_{\Lambda^*(1520)} > 0.2$ (estimated by SPHINX,
hep-ex/0407026)

- “Negative” experiments:

ALEPH: $R_{\Lambda^*(1520)} < 0.1$

BaBar: $R_{\Lambda^*(1520)} < \sim 0.01$ Belle: < 0.02

HERA-B: $R_{\Lambda^*(1520)} < 0.027 \sim 0.16$

SPHINX: $R_{\Lambda^*(1520)} < 0.02$

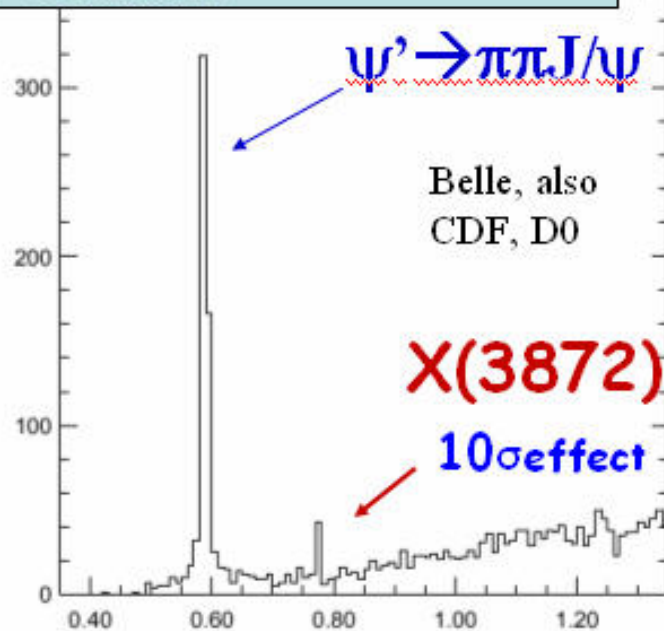
Quantitative comparisons difficult.
(ZEUS-H1 should be one of the few possible)

Jin

X(3872): Charmonium or $D^0\bar{D}^{*0}$ “Molecular State”?

$$M = 3872.0 \pm 0.6 \pm 0.5 \text{ MeV}$$

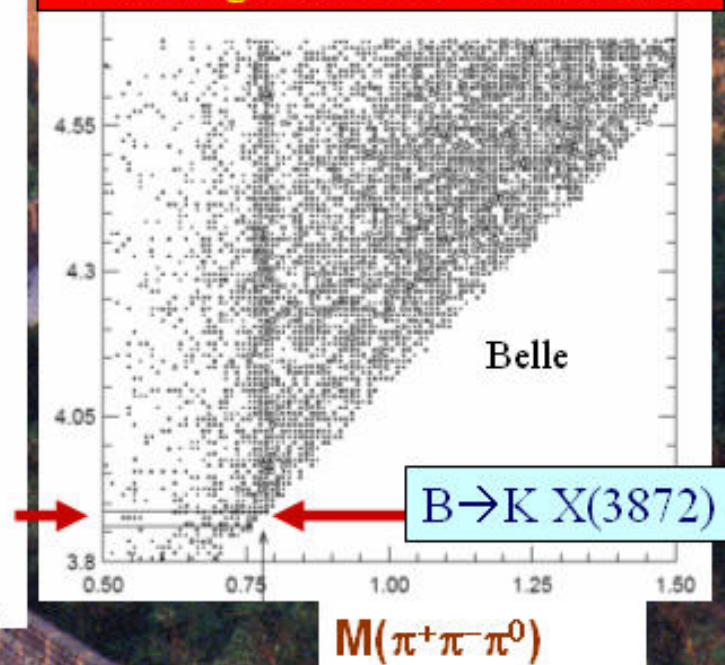
$$\Gamma < 2.3 \text{ MeV at } 90\% \text{ C.L.}$$



No $D^0\bar{D}^0 \rightarrow$ unnatural spin-parity

Also reported in $\omega J/\psi$:
mixed isospin would
→ mixing with $D^0\bar{D}^{*0}$ molecule

Close



$M(\pi^+\pi^-\pi^0)$

Mass, width, angular dist, etc. disfavor charmonium:
On the other hand DD^* molecule models also have difficulties

Quark-gluon plasma?

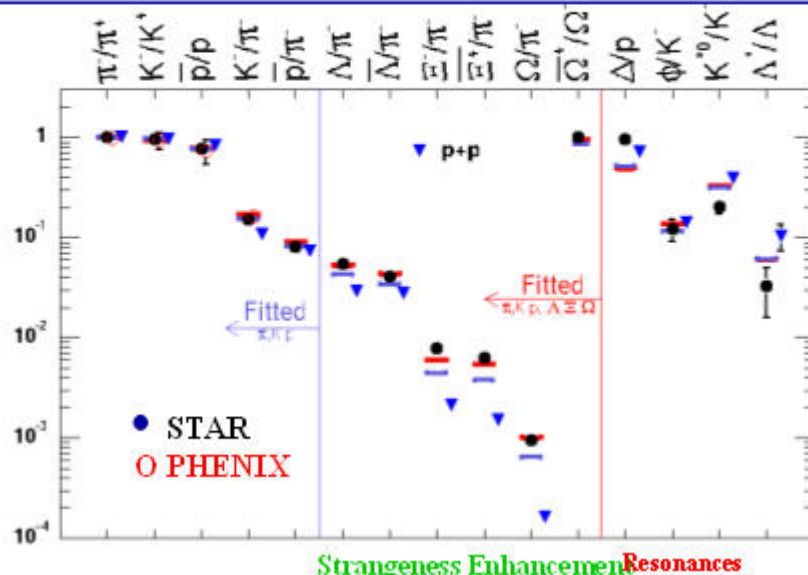
Dunlop

RHIC Implementation

- Flexibility is key to understanding complicated systems
 - Polarized protons, $\sqrt{s} = 50\text{-}500\text{ GeV}$
 - Nuclei from d to Au, $\sqrt{s_{NN}} = 20\text{-}200\text{ GeV}$
- Physics runs to date
 - Au+Au @ 20, 62, 130, 200 GeV
 - Polarized p+p @ 200 GeV
 - d+Au @ 200 GeV



Chemical Equilibration? Hadron Yield Ratios

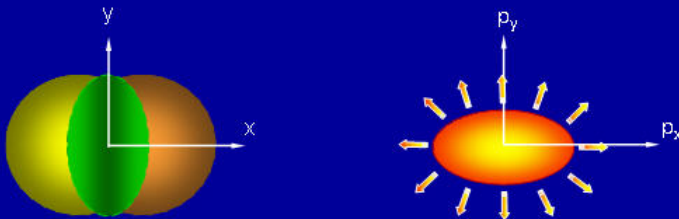


Some quark-gluon plasma models give good descriptions of data...

Dunlop

Collective Behavior: Azimuthal Anisotropy v_2

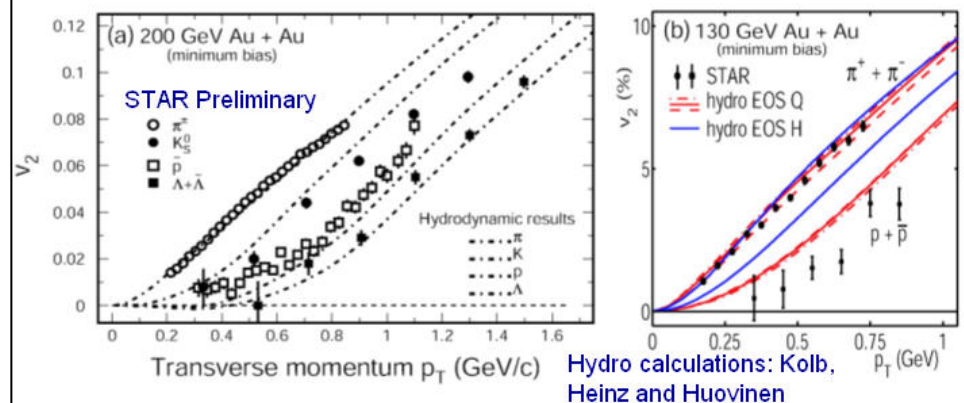
coordinate-space-anisotropy $\Leftrightarrow$ momentum-space-anisotropy

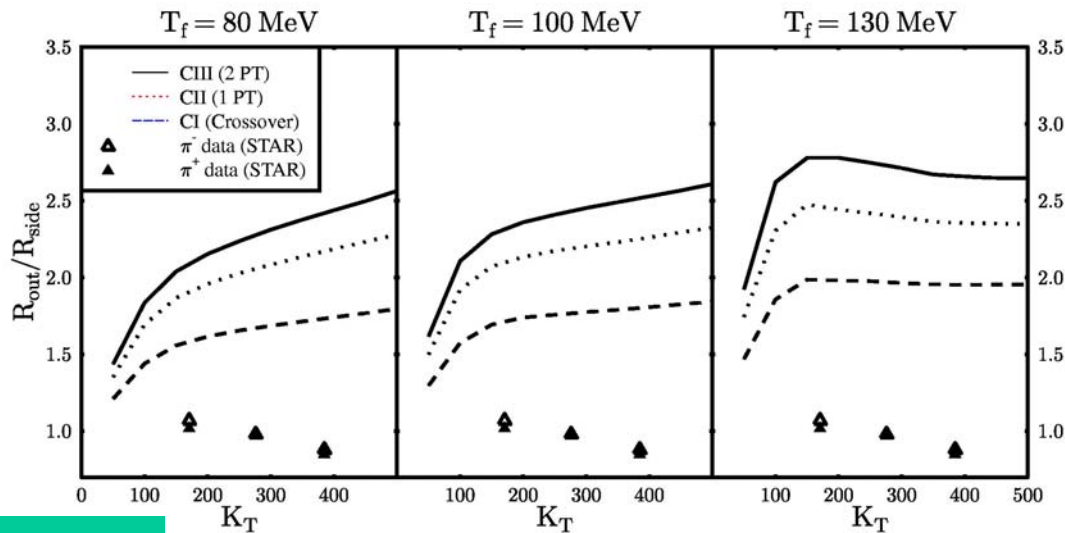


Pressure converts initial coordinate-space Anisotropy into final momentum-space anisotropy

$$\varepsilon = \frac{\langle y^2 - x^2 \rangle}{\langle y^2 + x^2 \rangle} \quad v_2 = \langle \cos 2\varphi \rangle, \quad \varphi = \tan^{-1}\left(\frac{p_y}{p_x}\right)$$

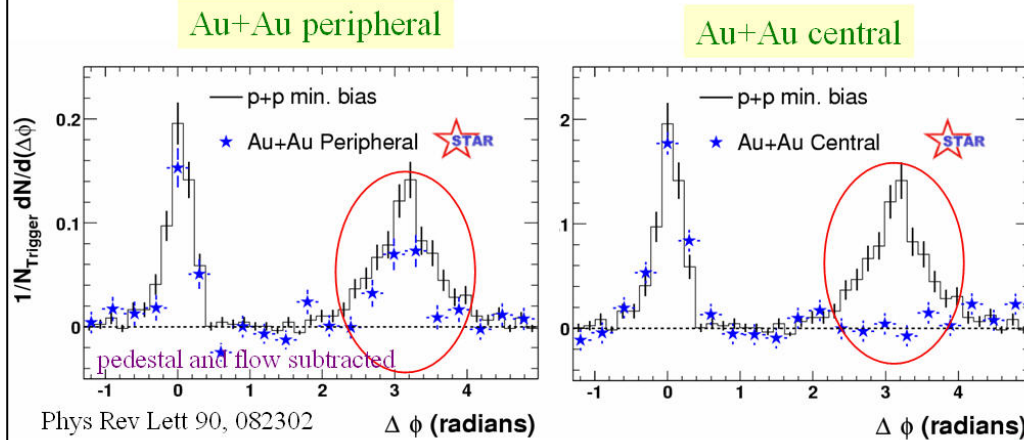
v_2 vs. Ideal Hydrodynamics





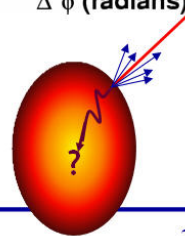
Dunlop

Azimuthal distributions in Au+Au



Near-side: peripheral and central Au+Au similar to p+p

Strong suppression of back-to-back correlations in central Au+Au



But doesn't work sometimes:

← source size compared to hydrodynamic model

Jet-quenching measurements look impressive: Quantitative predictions?

Summary

RHIC has made major advances in runs 1-3, leading to an appealing picture of bulk, dense, highly interacting matter.

- 1) *Extended reach in energy density appears to reach simplifying conditions in central collisions* -- ~ideal fluid expansion; approx. local thermal equilibrium.
- 2) *Extended reach in p_T gives probes for behavior difficult to access at lower energies* – jet quenching; ~constituent quark scaling.

However: *In the absence of a direct “smoking gun” signal of deconfinement revealed by experiment alone, a QGP discovery claim must rest on the comparison with a promising, but still not yet mature, theoretical framework. In this circumstance, clear predictive power with quantitative assessments of theoretical uncertainties are necessary for the present appealing picture to survive as a lasting one.*

J/ ψ suppression measurements yet to come

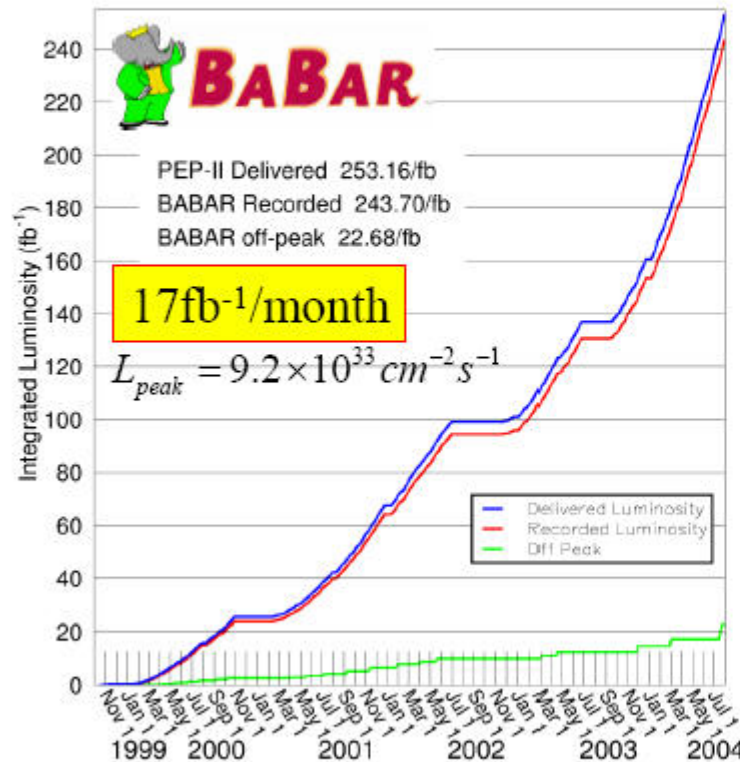
Quark mixing

Current luminosities and data samples

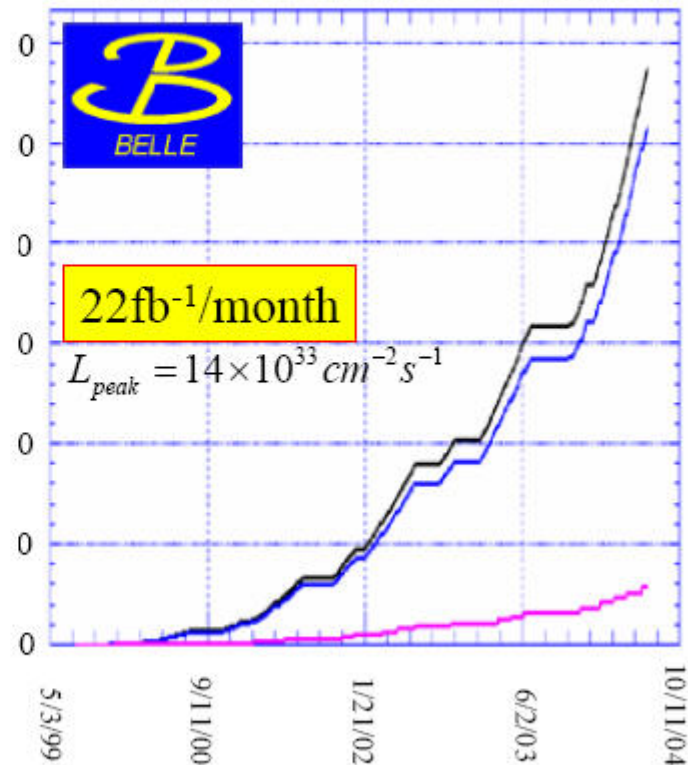
Total 244

+

$286 \text{ fb}^{-1} = 0.530 \text{ ab}^{-1}!!$



Delivered Recorded Continuum



ICHEP04-北京

August 20, 2004

Marcello A. Giorgi

2

B-factories are doing spectacularly well.
project 1.5 ab^{-1} (each !) end 2009.

CPV in decay amplitudes

First observation of Direct CPV in B decays

$$B^0 \rightarrow K^+ \pi^-$$

BABAR

hep-ex/0408057,
submitted to PRL

$$A_{CP} = -0.133 \pm 0.030 \pm 0.009$$

4.2 σ

Belle

Confirmation at ICHEP04

Signal (274M $B\bar{B}$ pairs): 2140 ± 53

$$A_{CP} = -0.101 \pm 0.025 \pm 0.005$$

3.9 σ

Average

$$A_{CP} = -0.114 \pm 0.020$$

$$B^0 \rightarrow K^+ \pi^0$$

$$A_{CP} = +0.06 \pm 0.06 \pm 0.01 \text{ BABAR}$$

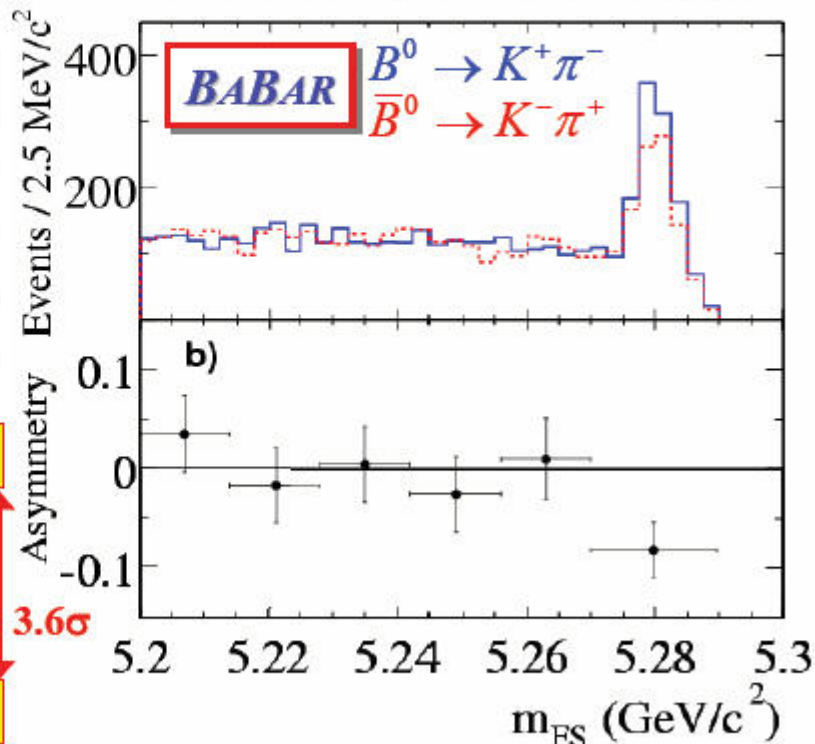
$$A_{CP} = +0.04 \pm 0.05 \pm 0.02 \text{ Belle}$$

Average

$$A_{CP} = +0.049 \pm 0.040$$

B^+

Signal (227M $B\bar{B}$ pairs): 1606 ± 51



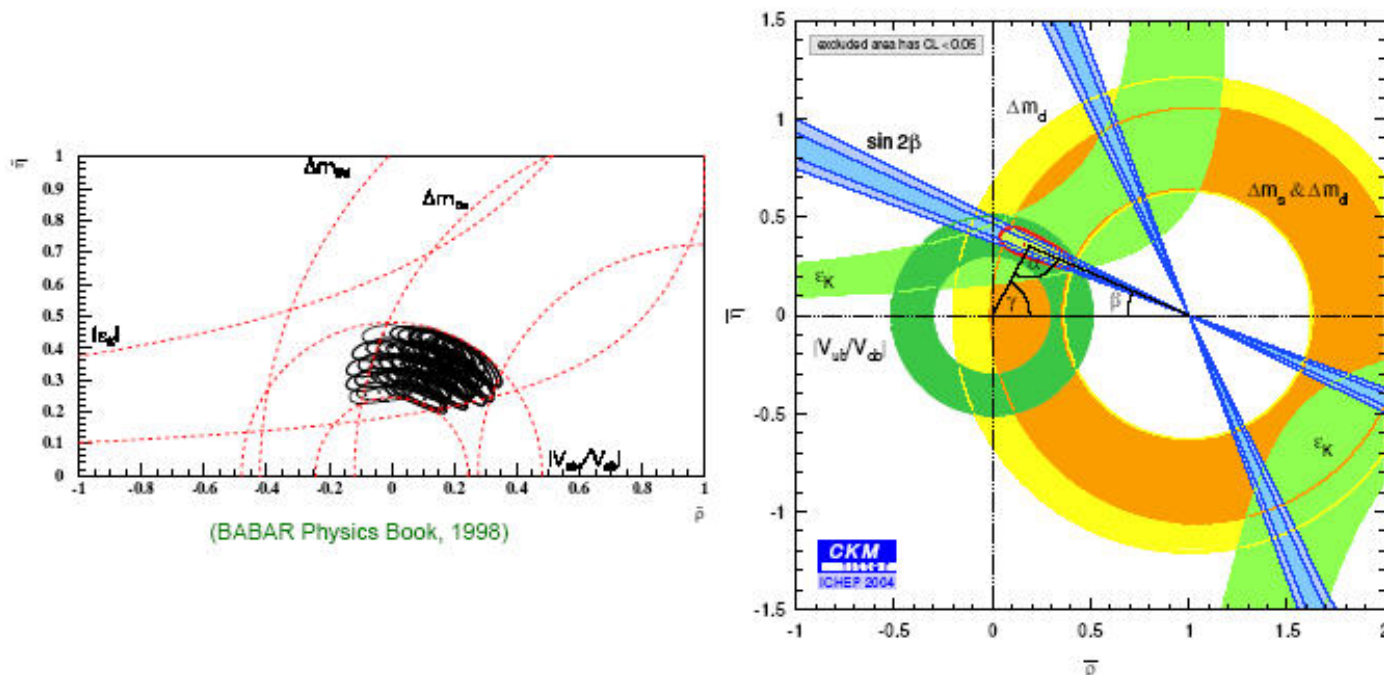
ICHEP04-北京
August 20, 2004

Marcello A. Giorgi

J. Wu, Y. Chao, CP-6

Testing the flavor sector

- For 35 years, until 1999, the only unambiguous measurement of CPV was ϵ_K



$\sin 2\beta = 0.726 \pm 0.037$, order of magnitude smaller error than first measurements

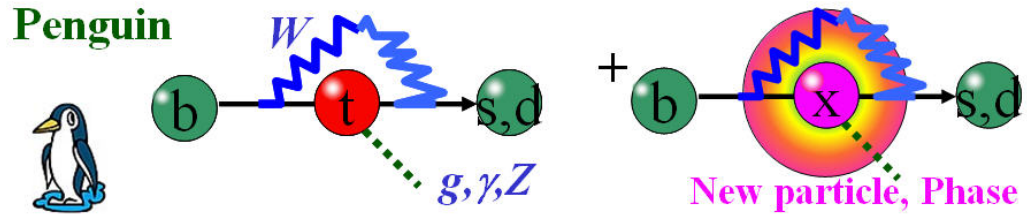


Z. Ligeti — p. 2



From $B \rightarrow J/\psi K_s$; other modes are (mostly) consistent

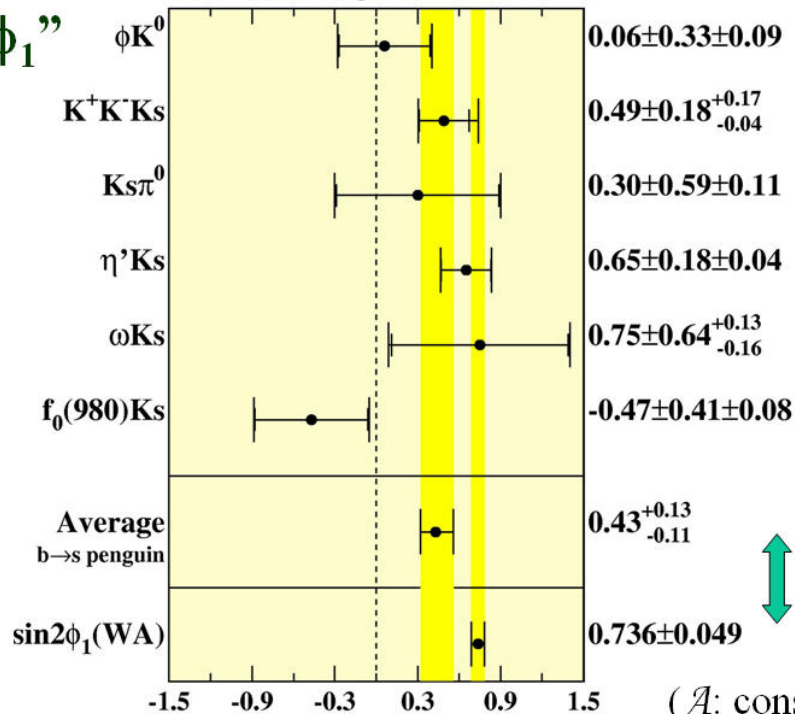
Penguins are sensitive to new physics



Summary of $b \rightarrow s\bar{q}q$ CPV



“ $\sin 2\phi_1$ ”



274M $B\bar{B}$

Slight deviation from expectations: similar results from Babar (2.7σ)

NP?
Hadronic corrections?

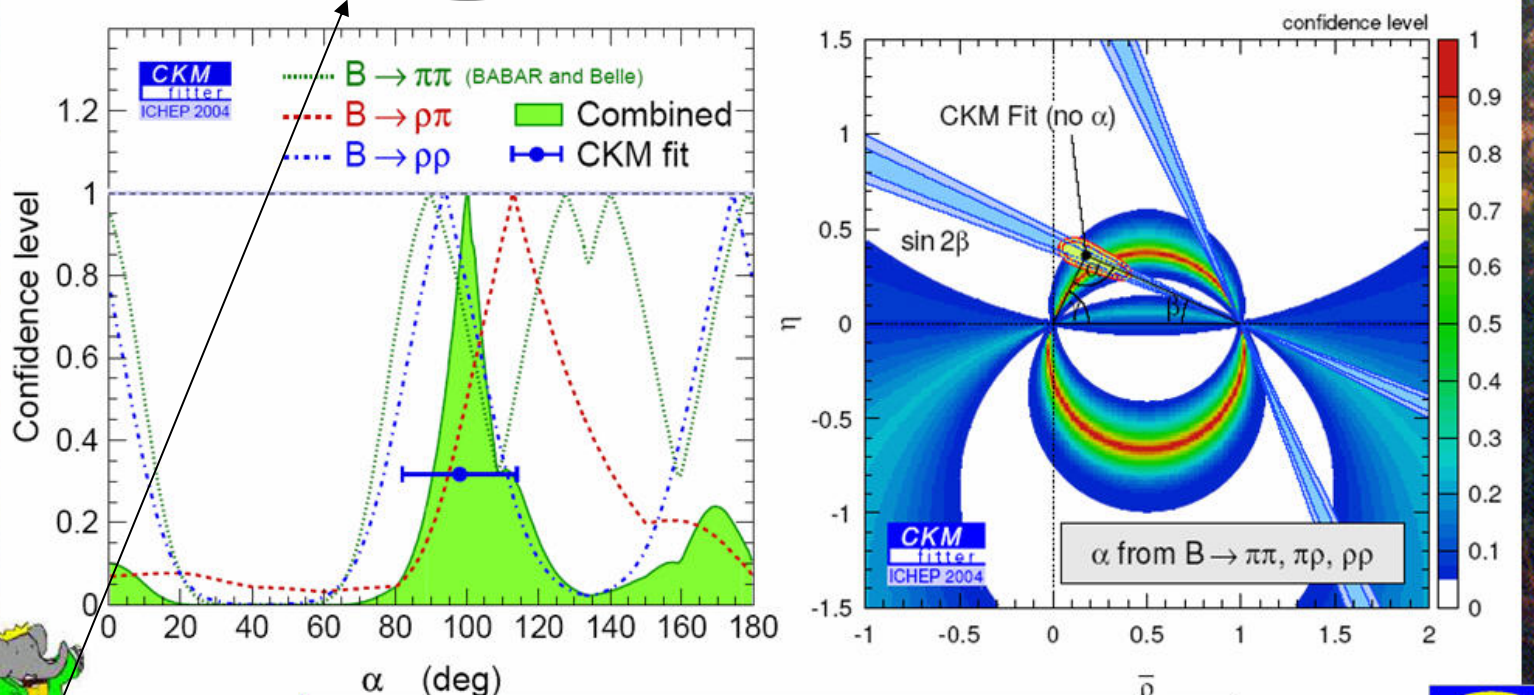
$\updownarrow 2.4\sigma$

Inconsistent result
from Belle and BaBar

Measuring α with $B \rightarrow \pi\pi$, $\rho\pi$ and $\rho\rho$ Decays

Sensitivity mainly from $S_{\rho^+\rho^-}$ and $\rho\pi$ Dalitz, $\pi\pi$ has small effect at present

Combined result $\alpha = 100^{+12}_{-10}^\circ$ ($103 \pm 11^\circ$ w/o $\pi\pi$); better than indirect fit $98 \pm 16^\circ$



Excellent agreement with global CKM fit



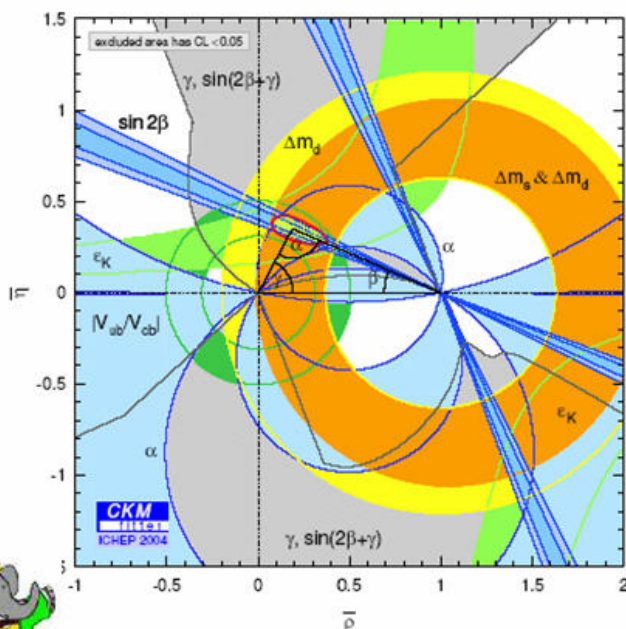
I only saw this in Ellis' summary.

Fit allowing for new physics in $B\bar{B}$ mixing amplitude

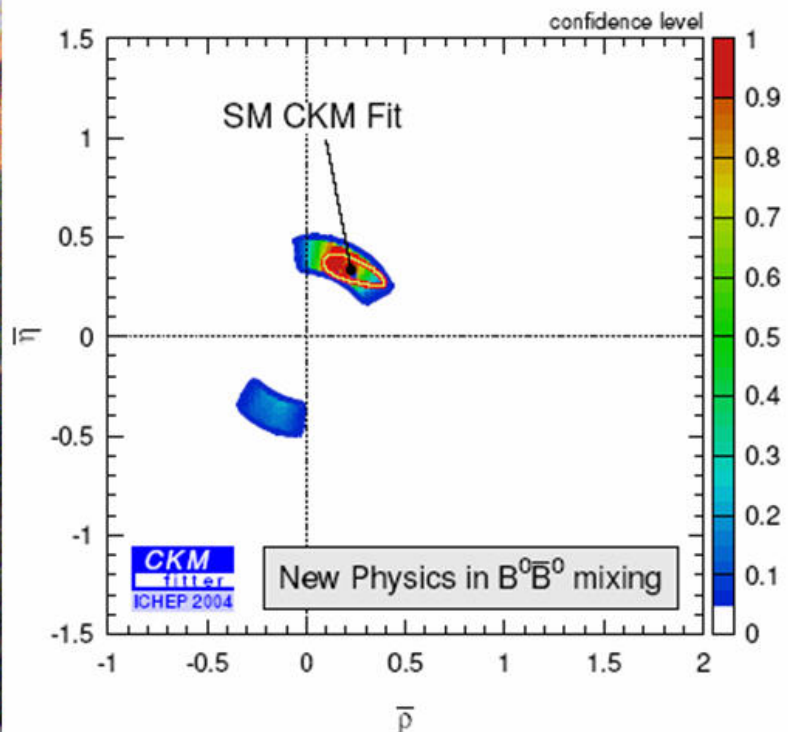
Putting all the Information Together

- α from $B \rightarrow \rho\rho$ and $\rho\pi$ Dalitz
- $2\beta + \gamma$ from $B \rightarrow D^{(*)}\pm\pi^\mp$
- γ from $B \rightarrow DK$ (with D Dalitz)
- $\cos 2\beta$ from ψK^* and A_{SL} (for NP)

Fit with above included



Is there room for NP?



Sakai, Ali, Giorgi, Ligeti

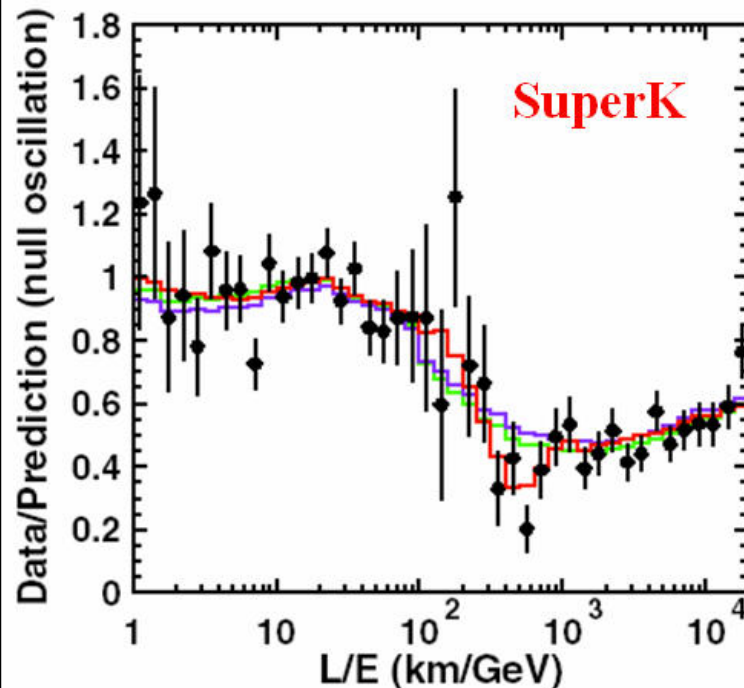
Not much room for NP

Neutrino sector

Neutrino oscillations established

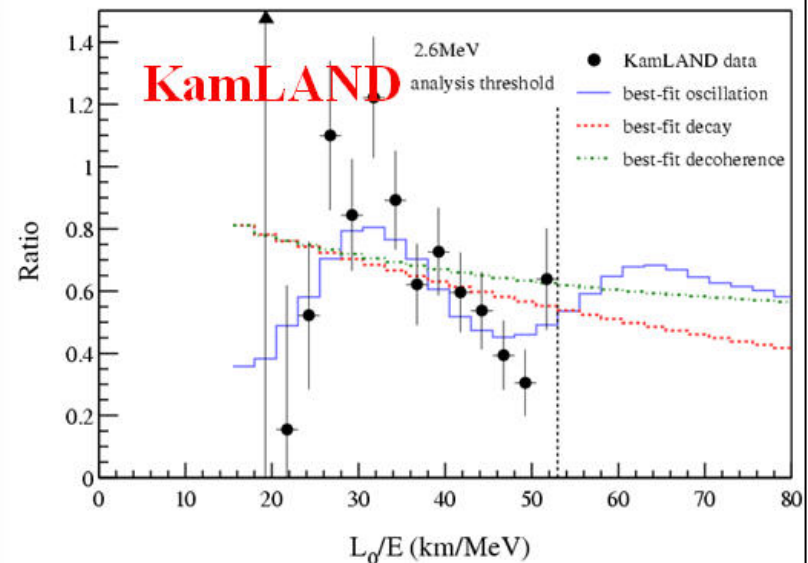
* *V. Barger et al. Phys. Rev. Lett. 82 (1999) 2640*

† *E. Lisi et al., Phys. Rev. Lett. 85 (2000) 1166*



Decay rejected at 3.4σ

Decoherence rejected at 3.8σ

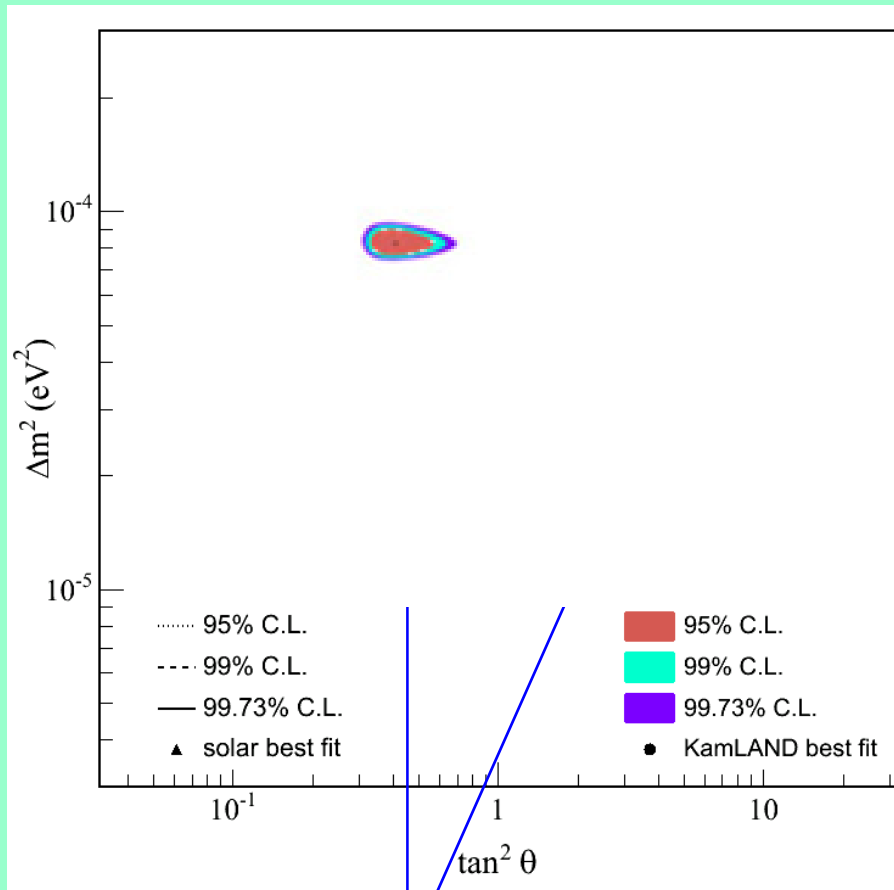


Decay* excluded at 95% CL

Decoherence† excluded at 94% CL

L/E analyses from atmospheric (Θ_{23}) and reactor (Θ_{12}): oscillation favored over other exotic explanations

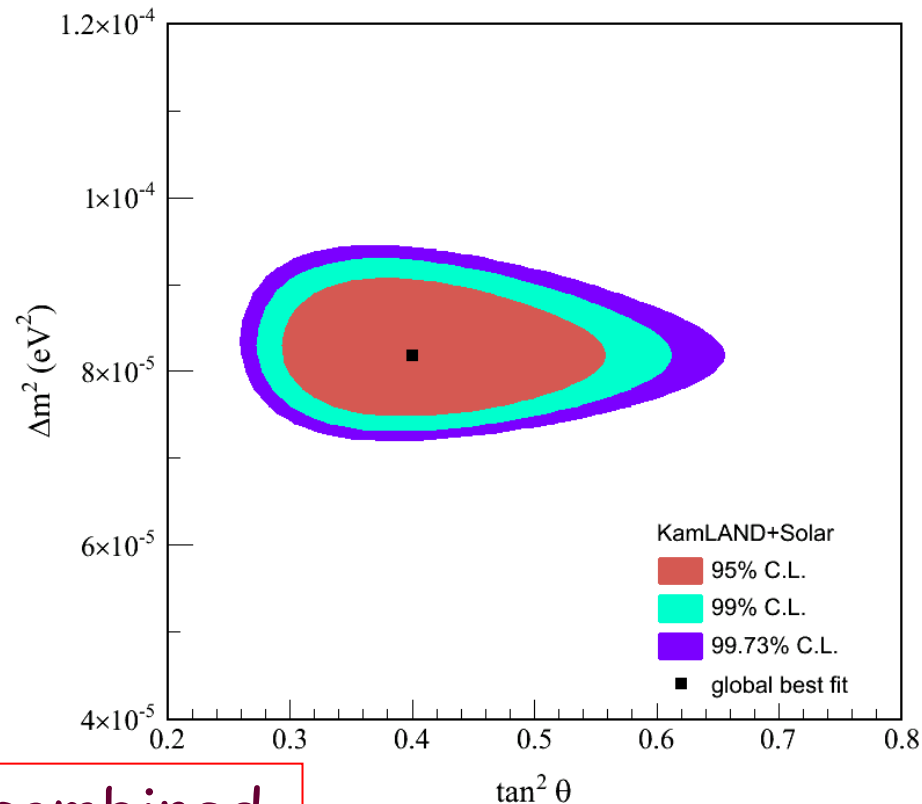
Combined solar ν - KamLAND 2-flavor analysis



Includes (small) matter effects

$$\Delta m_{12}^2 = 8.2^{+0.6}_{-0.5} \times 10^{-5} \text{ eV}^2$$

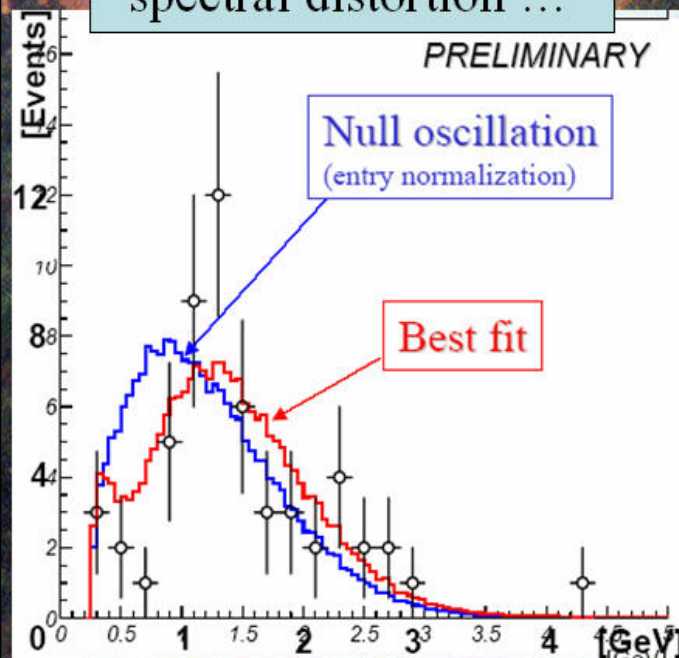
$$\tan^2 \theta_{12} = 0.40^{+0.09}_{-0.07}$$



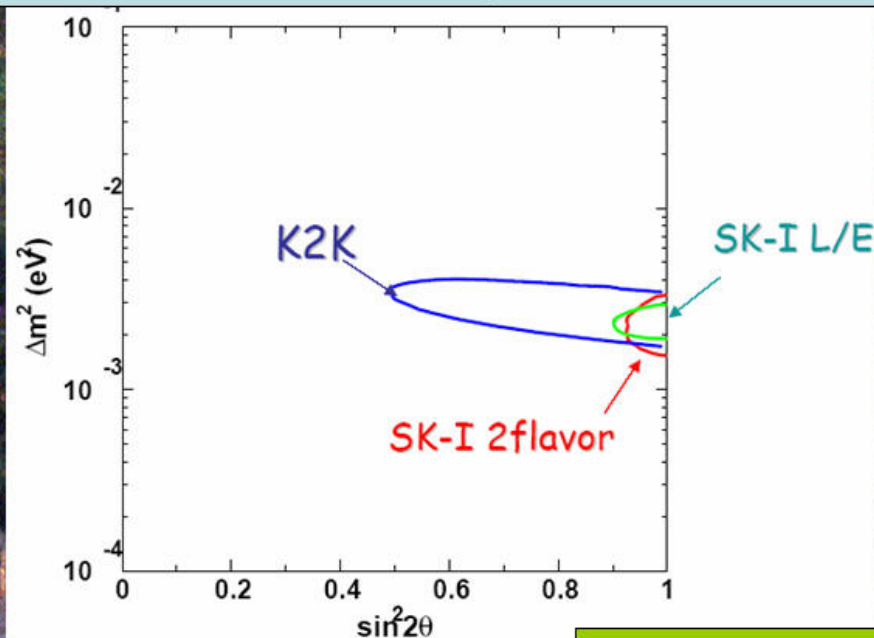
Solar(SK,SNO) and KamLAND combined

K2K confirms Super-Kamiokande

K2K Rate suppression
and
spectral distortion ...



... agree with SK azimuthal distributions
and
L/E analysis



McGrew, Wang

Atmospheric and accelerator-based (Θ_{23}) consistent

Summary

- **Neutrino oscillations established**

- Solar: $\Delta m_{12}^2 = (8.2^{+0.6}_{-0.5}) \times 10^{-3} \text{ eV}^2$

$\tan^2 \theta_{12} = 0.40^{+0.09}_{-0.07}$ **large but not maximal**

- Atmospheric: $\Delta m_{23}^2 = (2.4 \pm 0.4) \times 10^{-3} \text{ eV}^2$

$\sin^2 2\theta_{23} > 0.92$ @ 90%CL

- From global fit: $\sin^2 2\theta_{13} < 0.09$ @ 90%CL

- **Our next goal:**

- Absolute neutrino masses β -decay experiments

- Dirac or majorana ? $0\nu\beta\beta$ -decay experiments

- θ_{13} , CP and Mass hierarchy

Reactor, Long baseline accelerator based experiments

➔ **A great success in the past**

A long way to go in the future

Current limit: 2.2 eV
soon down to 0.5 eV?

Sensitivity of
next exp: 0.01 eV

Sensitivity of
proposed exp: ~0.01

MiniBooNE should clarify LSDN result (2005?)

MINOS start in 2005

CNGS(OPERA) start 2006

T2K, Nova on the horizon

Electro-weak fits: Improved top mass from the Tevatron



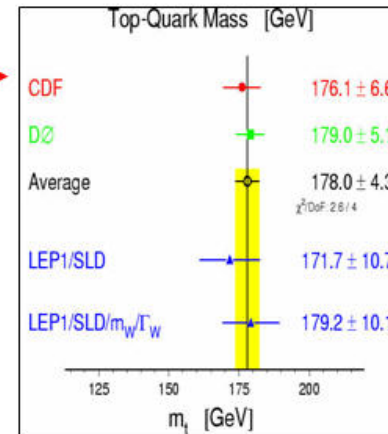
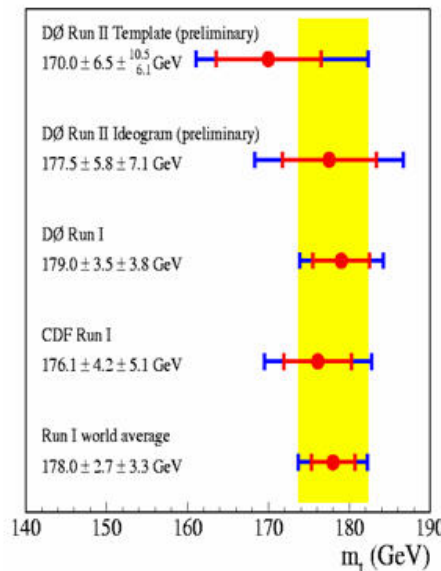
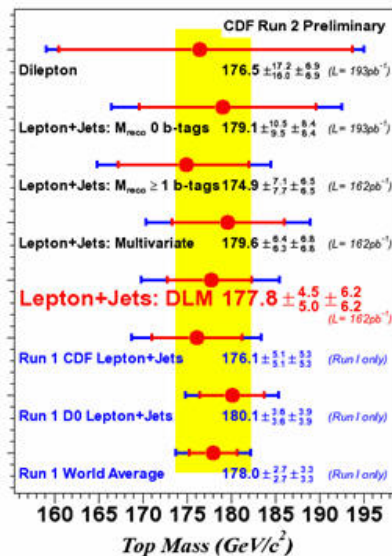
Tevatron Top Quark Mass Measurement



New combined Run I result →
(Was $m_t = 174.3 \pm 5.1$ GeV)

$$m_t = 178.0 \pm 4.3 \text{ GeV}$$

Run II top quark mass results from both detectors are available



Systematic error
(mainly jet energy scale)
is becoming limiting accuracy
factor

TeV EWWG is working on combining Run II top mass measurement from CDF and DØ



Experimental Limits on Higgs Mass



Available experimental limits

→ direct searches at LEP

$M_H > 114 \text{ GeV}$ at 95% C.L.

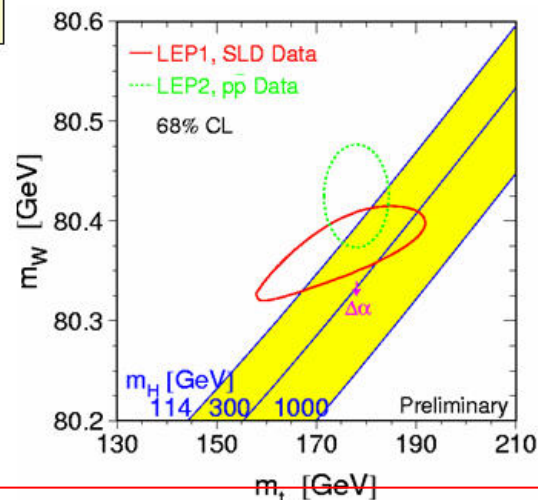
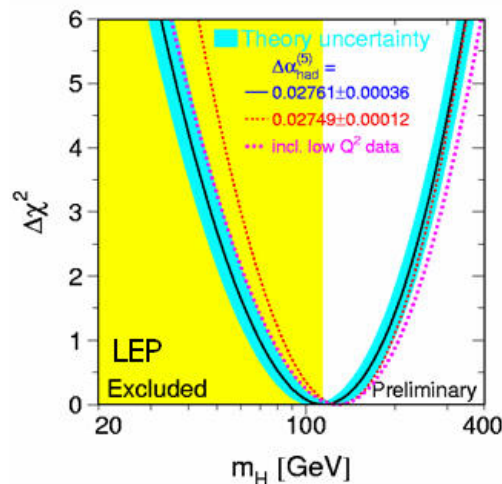
→ precision EW fits

$$m_H = 114^{+69}_{-45} \text{ GeV}$$

$m_H \leq 260 \text{ GeV}$ at 95% C.L.

Light Higgs favored

Tevatron provides:
Precision m_t and M_W measurements
Direct searches
→ SM Higgs
→ non-SM Higgs

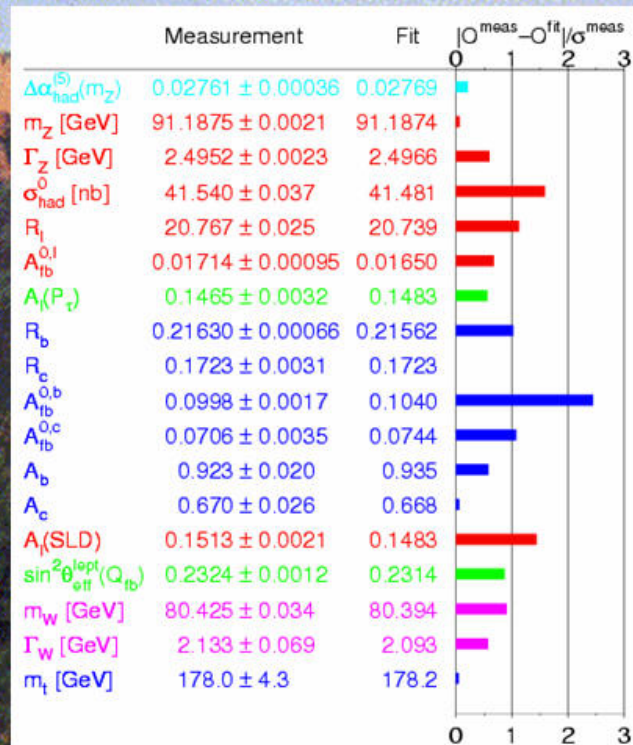


Leads to a Higgs mass which is the same as that of the limit from the direct search as LEP...

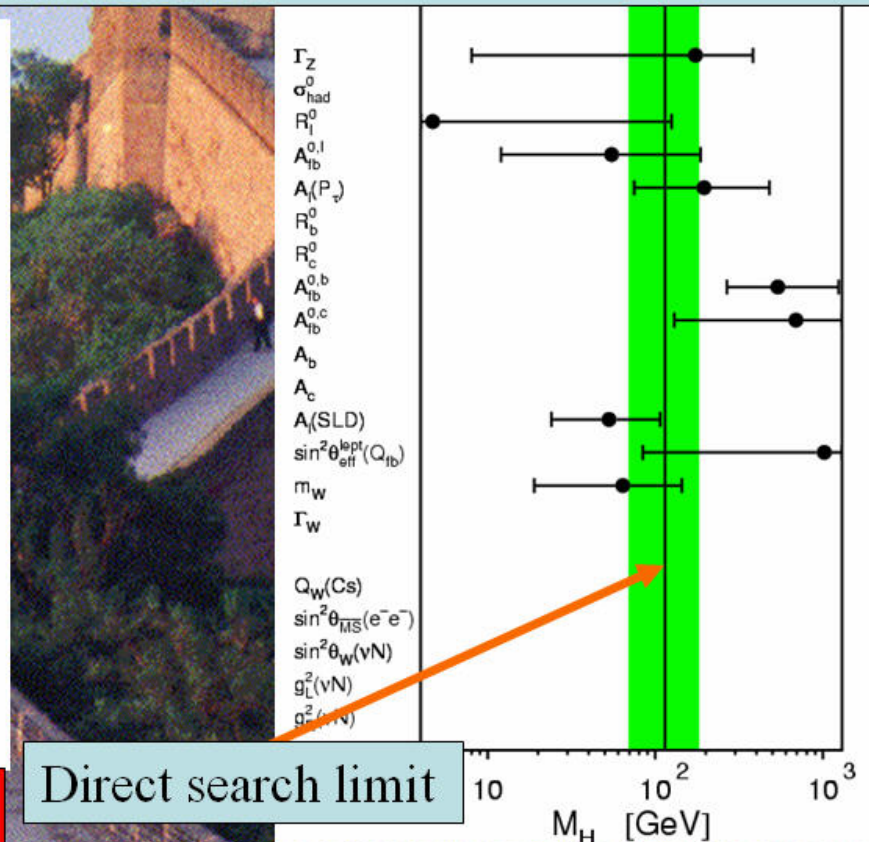
How Good is the Global Electroweak Fit?

Teubert

Pulls on Global Fit



Higgs mass from individual measurements



Heavy flavours $\neq$ leptons, m_W

Electroweak fits give satisfactory results:
 $g-2$, $\sin^2\theta$ from NuTeV (theory issues: see S-O. M. talk)

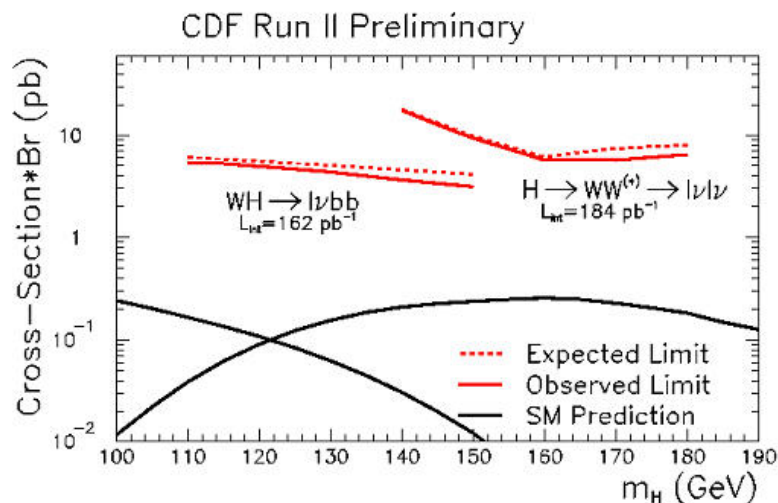
Higgs search at the Tevatron



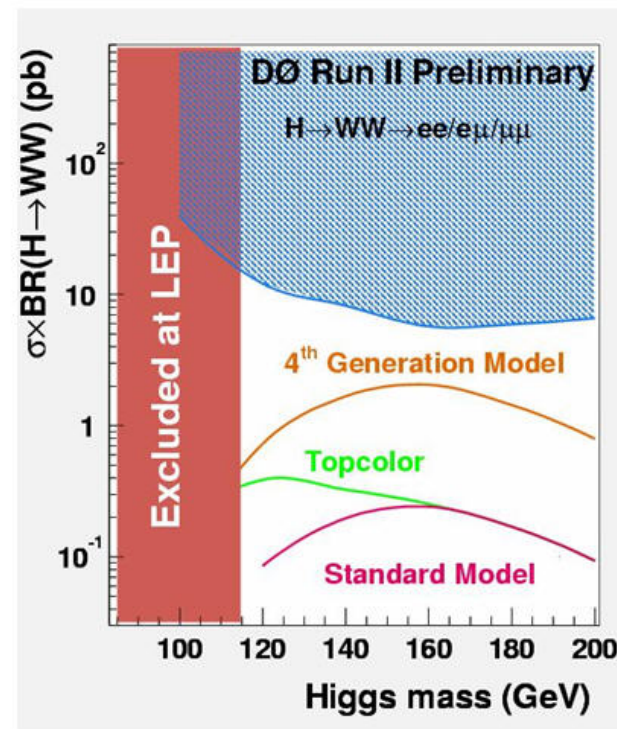
Current Limits on SM Higgs Search



Both experiments set 95% C.L. on SM Higgs cross section $\times$ Br



Limits already exceeding Run I results



D0 light (115 GeV) Higgs search limit
 $\sigma(WH) \times BR(H \rightarrow bb) < 12.4 \text{ pb}^{-1}$ at 95% C.L.



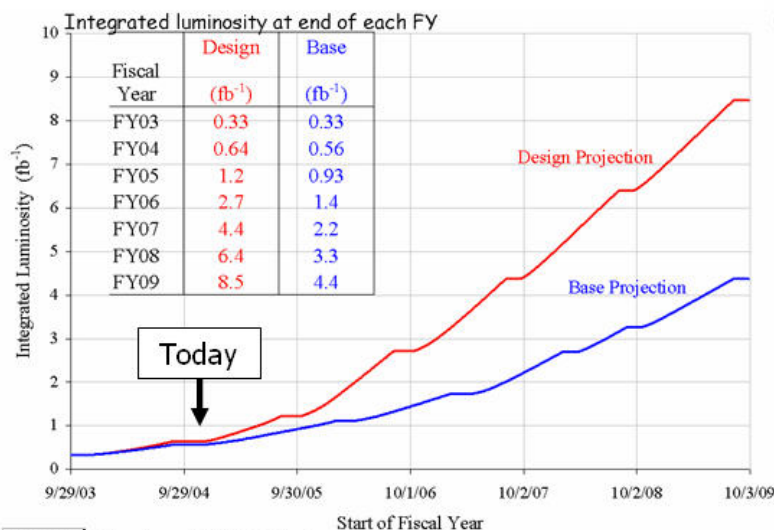
Tevatron Long Term Luminosity Plan



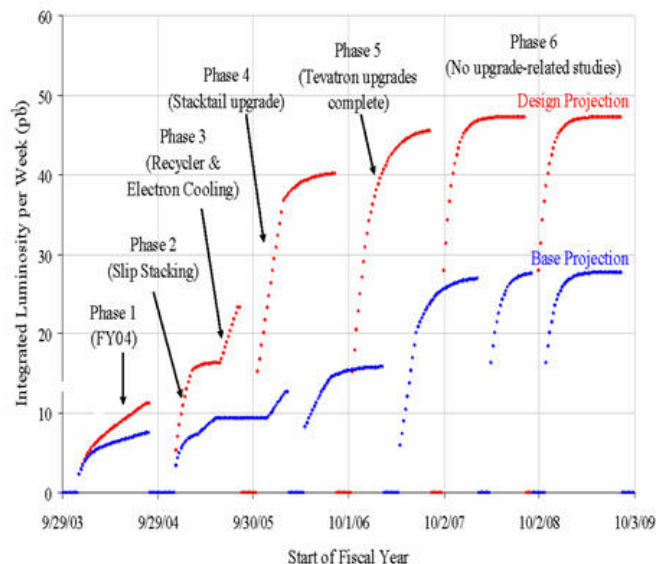
To reach higher masses with the same energy
→ higher luminosity

Increase in number of antiprotons
→ the key for higher luminosity

Expected peak luminosity
→ $3 \cdot 10^{32} \text{ cm}^{-2} \text{ sec}^{-1}$ by 2007



Dmitri Denisov, ICHEP04, August 21 2004



Currently expecting delivered
luminosity to each experiment
→ $4\text{-}8 \text{ fb}^{-1}$
by the end of 2009

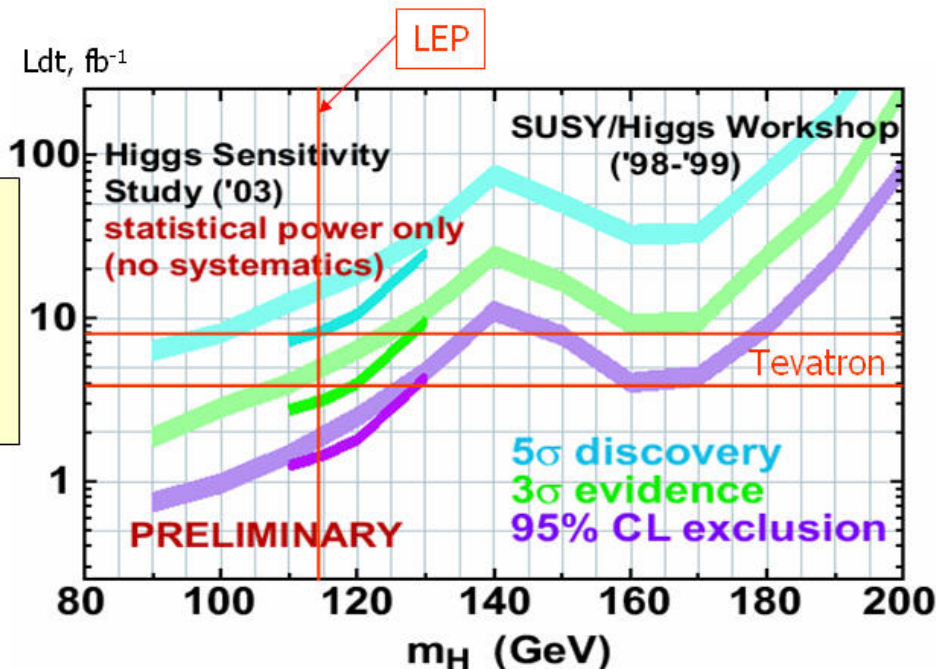
Tevatron projection is $4\text{-}8 \text{ fb}^{-1}$ by 2009



Tevatron SM Higgs Search: Outlook



Updated in 2003 in the low Higgs mass region
 $W(Z)H \rightarrow l\nu(\nu\nu, ll)bb$ to include
→ better detector understanding
→ optimization of analysis



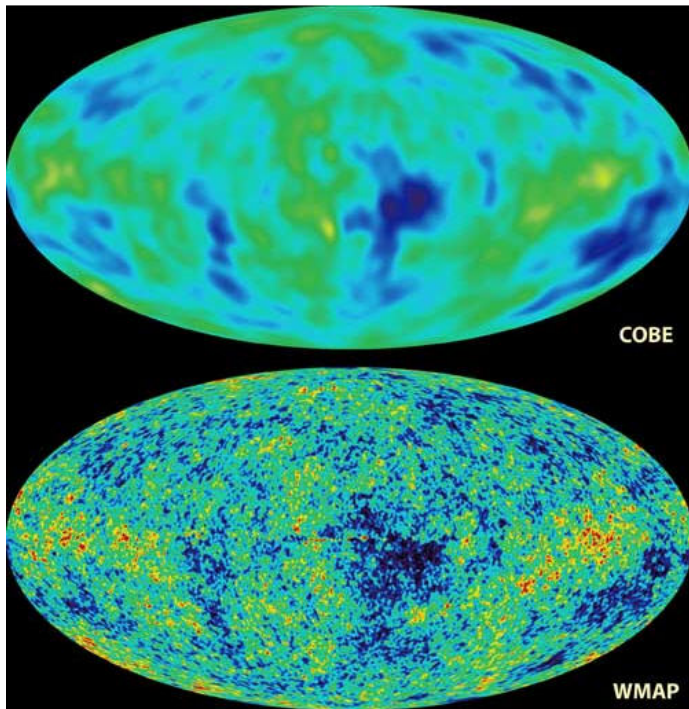
Sensitivity in the mass region above LEP limit starts at $\sim 2 \text{ fb}^{-1}$

Meanwhile

- optimizing analysis techniques
- understanding detectors better
- searching for non-SM Higgs with higher production cross sections or enhanced branching into modes with lower backgrounds

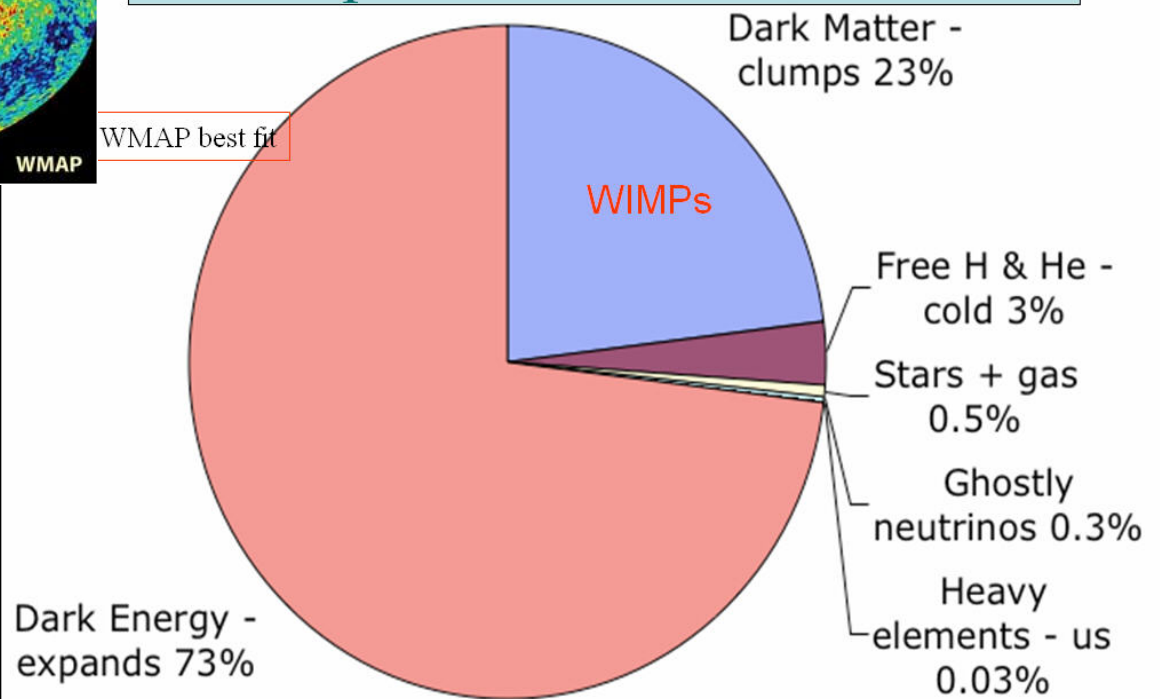
Higgs prospects at the Tevatron

"New" standard cosmology

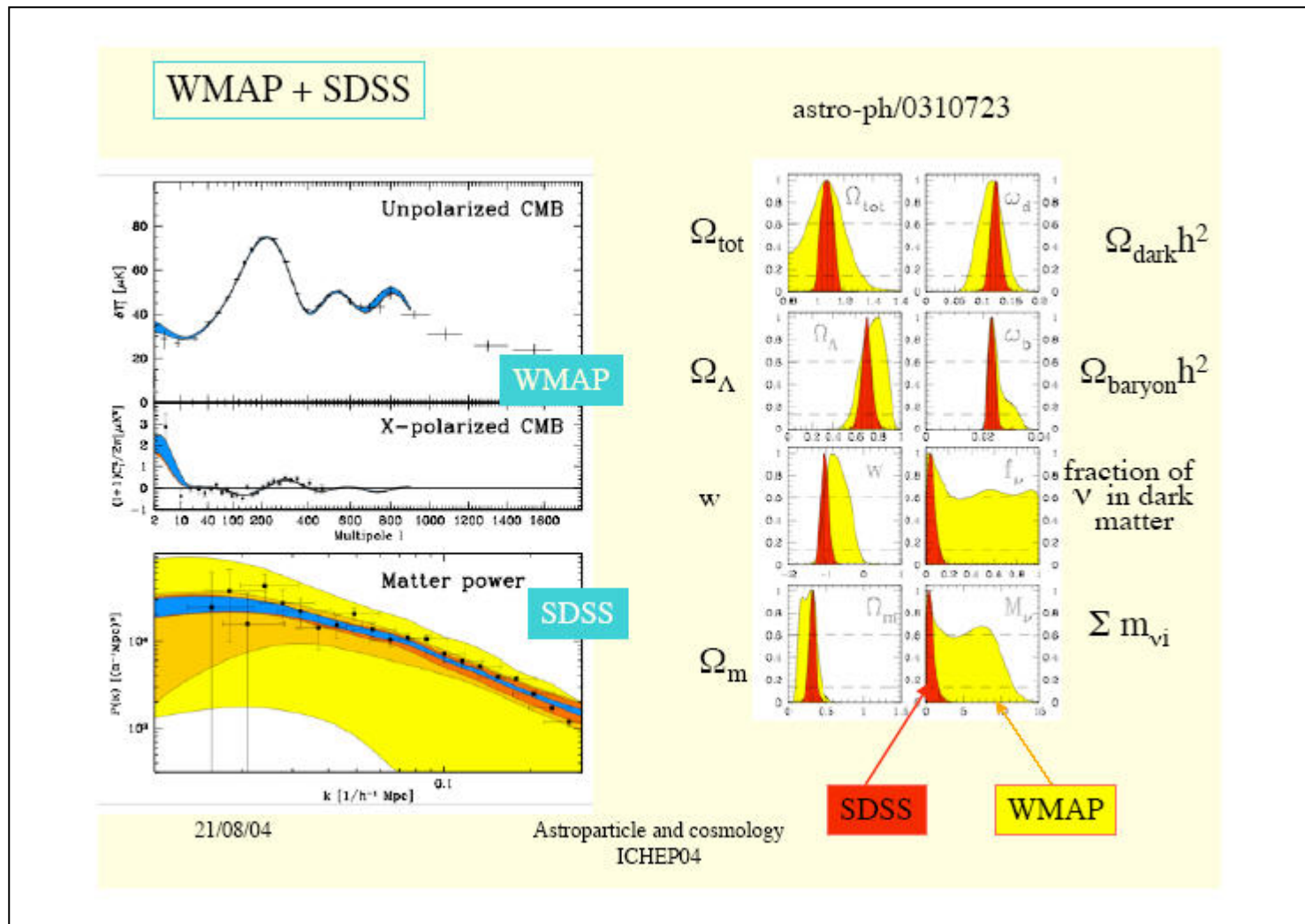


CMB anisotropy
measurements
from WMAP

Composition of the Cosmos



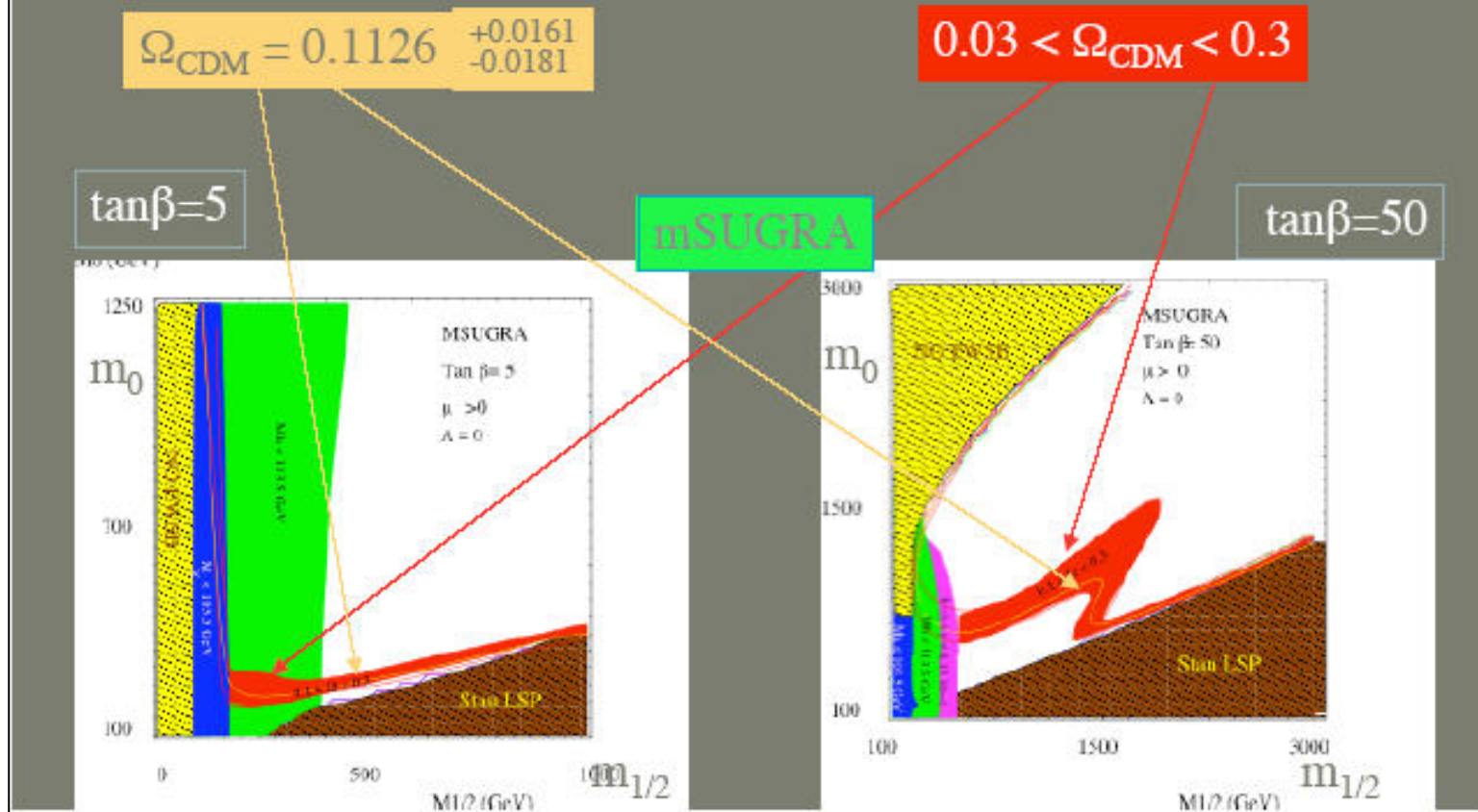
SDSS (Sloan Digital Sky Survey): clustering of galaxies



Strong constraints on cosmological parameters

Which, in turn, puts constraints on, e.g. SUSY

In the context of supersymmetric models, the constraint on dark matter severely constrains the parameter space :



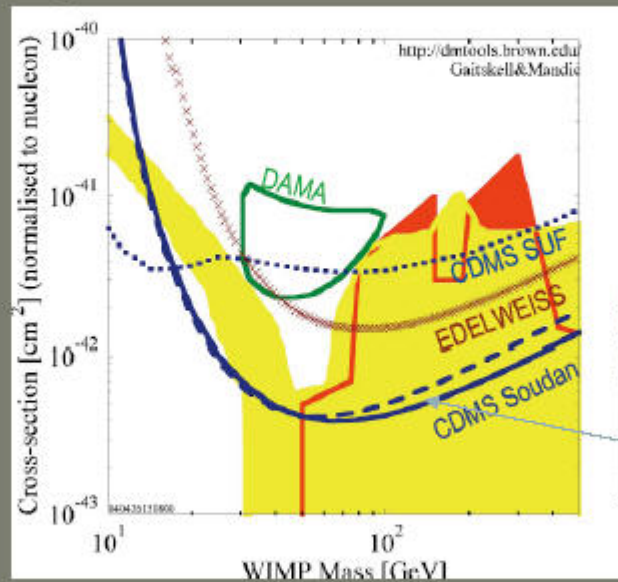
Binetruy

Experimental search for dark matter

Dark matter : direct detection



New results presented by CDMS of the first data run in the Soudan mine



P. Cushman



C. Sander

indirect detection

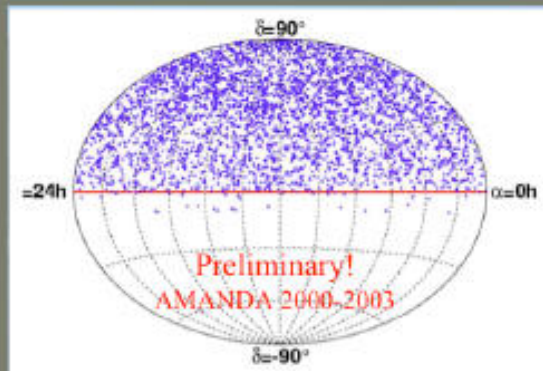
Connect EGRET excess in GeV cosmic γ rays with the presence of a halo of susy dark matter.

Binetruy

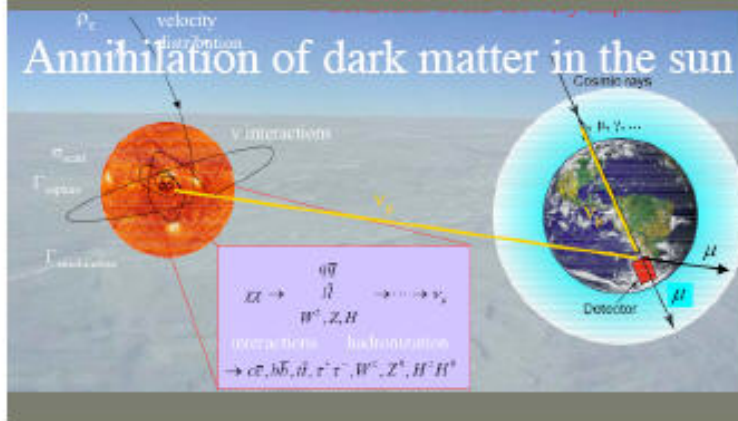
CDMS limit now contradicts DAMA positive result

Amanda

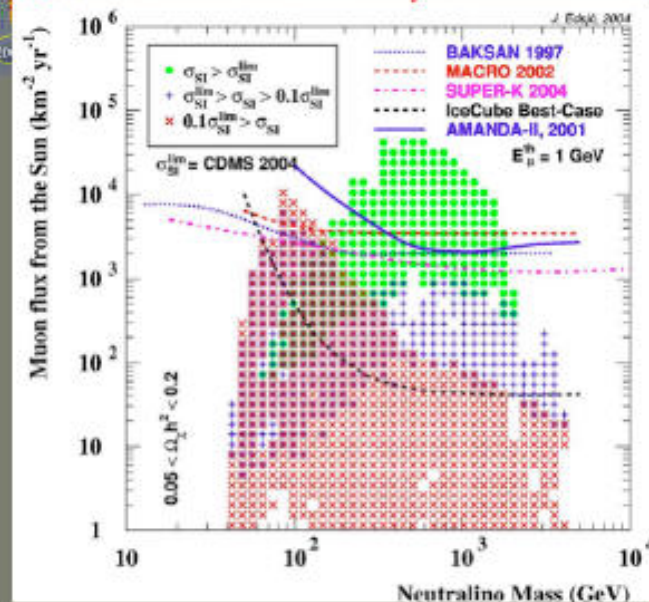
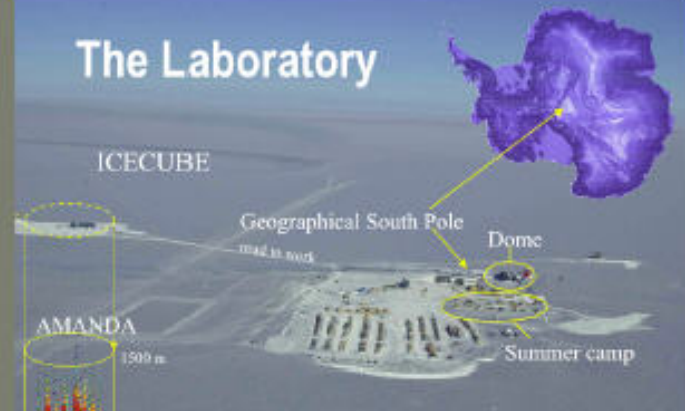
P. Olbrechts



No evidence for point sources



The Laboratory



Dark matter search at neutrino observatories

Binetruy

Final Remarks

- Many confirmations of Standard Model.
 - QCD becoming ever more precise. Deeper understanding of QCD essential in understanding many other results.
 - CKM description is now becoming very solid; no room for big deviations.
 - Neutrino oscillation also seems solid.
 - EW fits (despite some small deviations) work well -> light Higgs
 - SM + neutrino mixing appears to be a very good description of nature.
- On the other hand, as the precision of B and neutrino sector measurements increase, one is stuck more and more by the large number of apparently random parameters in the SM.
 - SM looks more and more like a description than an explanation.
 - E.g. Why does the quark and neutrino mixings look so different?
 - Several speakers commented on $\Theta_{12} + \Theta_{\text{cabbibo}} \approx \pi/4$
- Many reasons to be optimistic
 - B factories
 - Many new neutrino experiments on horizon
 - Strong connection with astroparticle physics and cosmology
 - LHC on the way
 - Linear Collider technology decision taken
 - Ideas for future (>30 years) accelerators (talk by Yokoya)