



Exploring the properties of strongly interacting matter

theoretical background:

Quantum Chromo Dynamics (QCD)

experimental approach:

collision of ultra-relativistic heavy ions



## From matter to elementray particles

## Substructure of Matter

#### from atoms to nuclei to nucleons to quarks and gluons





# From matter to elementray particles ...to elementary particle matter





## Phase Transition from a Hadron Gas to the Quark-Gluon Plasma

temperatures in the early universe after  $10^{-6}~{
m sec:} \sim 10^{12}~{
m K}$ 

density of neutron stars:  $\sim$  (3-10)-times nuclear matter density





## Phase Transition from a Hadron Gas to the Quark-Gluon Plasma

temperatures in the early universe after  $10^{-6}~{
m sec:} \sim 10^{12}~{
m K}$ 

density of neutron stars:  $\sim$  (3-10)-times nuclear matter density

dense hadronic



#### quark gluon plasma

What are the properties of this new form of matter? (equation of state, screening)



### Phase Transition from a Hadron Gas to the Quark-Gluon Plasma

temperatures in the early universe after  $10^{-6}~{
m sec:} \sim 10^{12}~{
m K}$ 

density of neutron stars:  $\sim$  (3-10)-times nuclear matter density

### dense hadronic matter What are the properties of this hadron gas new form of matter? What are the critical parameter (equation of state, of the transition to the QGP? screening) $(T_c(\mu), \epsilon_c)$ increasing temperature

quark gluon plasma



## Phase Transition from a Hadron Gas to the Quark-Gluon Plasma

temperatures in the early universe after  $10^{-6}~{\rm sec:} \sim 10^{12}~{\rm K}$ 

density of neutron stars:  $\sim$  (3-10)-times nuclear matter density

#### quark gluon plasma



#### dense hadronic



## Phase diagram of strongly interacting matter

## From Hadron Gas to Quark Gluon Plasma



#### Highly Excited Nuclear Matter\*

G. F. Chapline, M. H. Johnson, E. Teller, and M. S. Weiss Lawrence Livermore Laboratory, University of California, Livermore, California 94550

(Received 4 September 1973)

It is suggested that very hot and dense nuclear matter may be formed in a transient state in "head-on" collisions of very energetic heavy ions with medium and heavy nuclei. A study of the particles emitted in these collisions should give clues as to the nature of dense hot nuclear matter. Some simple models regarding the effects of meson and  $N^*$  production on the properties of dense hot nuclear matter are discussed.

What will be the effect of higher resonances? Models of the strong interactions based on the "bootstrap" idea lead to a density of states that increases exponentially with mass. This results from the fact that each new resonant state can combine with particles of lower or equal mass to make more resonant states.<sup>8</sup> In particular, the statistical bootstrap model leads to a density of states of the form<sup>9,10</sup>

$$N(m) = Cm^{-3}e^{m/\theta_0}, (3)$$

where  $\theta_0$ , the "maximum temperature" of hadron matter, is about 174 MeV as determined from high-energy scattering experiments.<sup>5</sup> The param-

#### **Highly Excited Nuclear Matter\***

G. F. Chapline, M. H. Johnson, E. Teller, and M. S. Weiss Lawrence Livermore Laboratory, University of California, Livermore, California 94550

(Received 4 September 1973)

It is suggested that very hot and dense nuclear matter may be formed in a transient state in "head-on" collisions of very energetic heavy ions with medium and heavy nuclei. A study of the particles emitted in these collisions should give clues as to the nature of dense hot nuclear matter. Some simple models regarding the effects of meson and  $N^*$  production on the properties of dense hot nuclear matter are discussed.

What will be the effect of higher resonances? Models of the strong interactions based on the "bootstrap" idea lead to a density of states that increases exponentially with mass. This results from the fact that each new resonant state can combine with particles of lower or equal mass to make more resonant states.<sup>8</sup> In particular, the statistical bootstrap model leads to a density of states of the form<sup>9,10</sup>

$$N(m) = Cm^{-3}e^{m/\theta_0}, (3)$$

where  $\theta_0$ , the "maximum temperature" of hadron matter, is about 174 MeV as determined from high-energy scattering experiments.<sup>5</sup> The param-

#### **Highly Excited Nuclear Matter\***

G. F. Chapline, M. H. Johnson, E. Teller, and M. S. Weiss Lawrence Livermore Laboratory, University of California, Livermore, California 94550

(Received 4 September 1973)

It is suggested that very hot and dense nuclear matter may be formed in a transient state in "head-on" collisions of very energetic heavy ions with medium and heavy nuclei. A study of the particles emitted in these collisions should give clues as to the nature of dense hot nuclear matter. Some simple models regarding the effects of meson and  $N^*$  production on the properties of dense hot nuclear matter are discussed.

What will be the effect of higher resonances? Models of the strong interactions based on the "bootstrap" idea lead to a density of states that increases exponentially with mass. This results from the fact that each new resonant state can combine with particles of lower or equal mass to make more resonant states.<sup>8</sup> In particular, the statistical bootstrap model leads to a density of states of the form<sup>9,10</sup>

$$N(m) = Cm^{-3}e^{m/\theta_0},$$
 (3)

where  $\theta_0$ , the "maximum temperature" of hadron matter, is about 174 MeV as determined from high-energy scattering experiments.<sup>5</sup> The param-

#### resonance gas

#### **Highly Excited Nuclear Matter\***

G. F. Chapline, M. H. Johnson, E. Teller, and M. S. Weiss Lawrence Livermore Laboratory, University of California, Livermore, California 94550

(Received 4 September 1973)

It is suggested that very hot and dense nuclear matter may be formed in a transient state in "head-on" collisions of very energetic heavy ions with medium and heavy nuclei. A study of the particles emitted in these collisions should give clues as to the nature of dense hot nuclear matter. Some simple models regarding the effects of meson and  $N^*$  production on the properties of dense hot nuclear matter are discussed.

What will be the effect of higher resonances? Models of the strong interactions based on the "bootstrap" idea lead to a density of states that increases exponentially with mass. This results from the fact that each new resonant state can combine with particles of lower or equal mass to make more resonant states.<sup>8</sup> In particular, the statistical bootstrap model leads to a density of states of the form<sup>9,10</sup>

$$N(m) = Cm^{-3}e^{m/\theta_0}, (3)$$

where  $\theta_0$ , the "maximum temperature" of hadron matter, is about 174 MeV as determined from high-energy scattering experiments.<sup>5</sup> The param-

#### resonance gas

 $T_c \simeq 174 \; MeV \; \; (!!!)$ 





Collision of Au-Au and other ions in the Relativistic Heavy Ion Collider (RHIC) at Brookhaven Nat. Lab. (BNL)

Lattice Simulations of Finite Temperature QCD at Bielefeld University





## **APEmille**



## The machines

## <u>RHIC</u>

circumference:  $\sim 4 \text{ km}$ 

Au-Au collisions

beam energy: 200 GeV/A







## The events







#### LGT:

• equilibrium thermodynamics of QCD;

#### HIC:

evolution of a dense interacting medium described by QCD;







#### LGT:

- equilibrium thermodynamics of QCD;
- formulated in terms of basic degrees of freedom: quarks and gluons;

#### HIC:

- evolution of a dense interacting medium described by QCD;
- observable properties in terms of hadrons, leptons and photons;







#### LGT:

- equilibrium thermodynamics of QCD;
- formulated in terms of basic degrees of freedom: quarks and gluons;
- observables expressed in terms of temperature and chemical potential

#### HIC:

- evolution of a dense interacting medium described by QCD;
- observable properties in terms of hadrons, leptons and photons;
- observables parametrized in terms of energy and particle multiplicities







HIC:

- evolution of a dense interacting medium described by QCD;
- observable properties in terms of hadrons, leptons and photons;
- observables parametrized in terms of energy and particle multiplicities

## MODELS

#### LGT:

- equilibrium thermodynamics of QCD;
- formulated in terms of basic degrees of freedom: quarks and gluons;
- observables expressed in terms of temperature and chemical potential



# Analyzing hot and dense matter on the lattice: $N_{\sigma}^3 \times N_{\tau}$



partition function:  $Z(V, T, \mu) = \int \mathcal{D} \mathcal{A} \mathcal{D} \psi \mathcal{D} \bar{\psi} e^{-S_E}$ 

#### APEmille





# Analyzing hot and dense matter on the lattice: $N_{\sigma}^3 \times N_{\tau}$





Where lattice calculations do/will contribute to the development of theoretical concepts and the analysis of experimental observables



Where lattice calculations do/will contribute to the development of theoretical concepts and the analysis of experimental observables



Where lattice calculations do/will contribute to the development of theoretical concepts and the analysis of experimental observables

 $T_c, \ \epsilon_c$ phase diagram in the  $(T, \ \mu_B)$ -plane;  $\mu \simeq 0$  : RHIC (LHC)  $\mu > 0$  : SPS (GSI future) chiral critical point



Where lattice calculations do/will contribute to the development of theoretical concepts and the analysis of experimental observables

### EoS

energy density, pressure, velocity of sound,...; susceptibilities (baryon number fluctuations);

strangeness contribution



Where lattice calculations do/will contribute to the development of theoretical concepts and the analysis of

experimental observables

 $In - medium \\ hadron properties$ 

heavy quark potential, screening; charmonium spectroscopy; light quark bound states;

thermal dilepton rates



Where lattice calculations do/will contribute to the development of theoretical concepts and the analysis of

experimental observables

short vs. long distance physics

running coupling constant; transport coefficients



## crossover vs. phase transition







continuous transition for small chemical potential and small quark masses at

 $\begin{array}{l} T_c \simeq 170 \; MeV \\ \epsilon_c \simeq 0.7 \; GeV / fm^3 \end{array}$ 





continuous transition for small chemical potential and small quark masses at

 $T_c \simeq 170 \; MeV \\ \epsilon_c \simeq 0.7 \; GeV / fm^3$ 

2nd order phase transition; Ising universality class  $T_c(\mu)$  under investigation











## Critical temperature, equation of state









## Critical temperature, equation of state



- $a\simeq 0.2\,fm$  (continuum limit??)
- improved staggered fermions,
   ⇒ flavor symmetry breaking (need even better fermion actions)

 $\epsilon_c$ 

- $m_{PS}\simeq 770~MeV$  (!!!)
- ullet  $V\simeq (4\,{
  m fm})^3$  (thermodynamic limit)






QCD transition in a world with heavy pions:  $m_{\pi} \simeq 770~MeV$ not a "true phase transition"









. – p.13/31





. – p.13/31



## Particle ratios and freeze out conditions



### resonance gas: $Z(T, V, \mu_i) = \text{Tr}e^{-\beta(H-\sum_i \mu_i Q_i)}$

describes observed particle ratios and

freeze out conditions

P. Braun-Munzinger, D. Magestro, K. Redlich, J. Stachel, Phys. Lett. B518 (2001) 41



## Particle ratios and freeze out conditions



### resonance gas

describes observed particle ratios and freeze out conditions

- Is the freeze out temperature the critical temperature of the QCD transition?
- Which role do resonances play for the occurence of the transition to the QGP?
- ... and what about deconfinement and chiral symmetry restoration?

$$egin{array}{lll} \ln \ Z(T,V,\mu_B,..) \ = \ \sum_{m_i} \ln \ Z_i(T,V,\mu_B,..) \end{array}$$



## Critical temperature, equation of state and the resonance gas

Hagedorn spectrum :  $\rho(m_H) \sim c \ m_H^a \ e^{m_H/T_H}$ 

$$\ln Z(\mathbf{T}, \boldsymbol{\mu}_{B}) = \int \mathrm{d}m_{H} \ \rho(m_{H}) \ \ln Z_{m_{H}}(\mathbf{T}, \boldsymbol{\mu}_{B})$$

•  $\int \Rightarrow \sum \sim$  experimentally known resonances



## Critical temperature, equation of state and the resonance gas

Hagedorn spectrum :  $\rho(m_H) \sim c \ m_H^a \ e^{m_H/T_H}$ 

$$\ln Z(\boldsymbol{T},\boldsymbol{\mu_B}) = \int \mathrm{d}m_H \ \rho(m_H) \ \ln Z_{m_H}(\boldsymbol{T},\boldsymbol{\mu_B})$$



resonance gas: ~ 1500 d.o.f. from ~ 300 exp. known resonances VS. lattice calculation: (2+1)-flavor QCD,  $m_q/T = 0.4$ resonances give large contribution at  $T_c$ • explain eos for  $T \leq T_c$ ;



## Critical temperature, equation of state and the resonance gas

Hagedorn spectrum :  $\rho(m_H) \sim c \ m_H^a \ e^{m_H/T_H}$ 

$$\ln Z(\mathbf{T}, \boldsymbol{\mu}_{B}) = \int \mathrm{d}m_{H} \ \rho(m_{H}) \ \ln Z_{m_{H}}(\mathbf{T}, \boldsymbol{\mu}_{B})$$



resonance gas:

resonances give large contribution at  $T_c$ 

- explain eos for  $T \leq T_c$ ;
- explain eos for  $\mu_q > 0$  and  $T \leq T_c$ ;



## Fluctuations of the baryon number density ( $\mu > 0$ )

baryon number density fluctuations: (Bielefeld-Swansea, PRD68 (2003) 014507)





. – p.16/31



non-zero baryon number density:  $\mu > 0$ 

$$Z(\mathbf{V}, \mathbf{T}, \boldsymbol{\mu}) = \int \mathcal{D}\mathcal{A}\mathcal{D}\psi \mathcal{D}\bar{\psi} e^{-S_E(\mathbf{V}, \mathbf{T}, \boldsymbol{\mu})}$$
$$= \int \mathcal{D}\mathcal{A}\mathcal{D} \ det \ M(\boldsymbol{\mu}) e^{-S_E(\mathbf{V}, \mathbf{T})}$$
$$\uparrow \text{complex fermion determinant;}$$

long standing problem

 $\Rightarrow$  three (partial) solutions for large T, small  $\mu$ 



non-zero baryon number density:  $\mu > 0$ 

$$Z(\mathbf{V}, \mathbf{T}, \boldsymbol{\mu}) = \int \mathcal{D}\mathcal{A}\mathcal{D}\psi \mathcal{D}\bar{\psi} e^{-S_E(\mathbf{V}, \mathbf{T}, \boldsymbol{\mu})}$$
$$= \int \mathcal{D}\mathcal{A}\mathcal{D} \ det \ M(\boldsymbol{\mu}) e^{-S_E(\mathbf{V}, \mathbf{T})}$$
$$\uparrow \text{complex fermion determinant;}$$

long standing problem

- $\Rightarrow$  three (partial) solutions for large T, small  $\mu$
- exact evaluation of *det M*: works well on small lattices; requires reweighting Z. Fodor, S.D. Katz, JHEP 0203 (2002) 014
- Taylor expansion around  $\mu = 0$ : works well for small  $\mu$ ; requires reweighting C. R. Allton et al. (Bielefeld-Swansea), Phys. Rev. D66 (2002) 074507
- imaginary chemical potential: works well for small  $\mu$ ; requires analytic continuation Ph. deForcrand, O. Philipsen, Nucl. Phys. B642 (2002) 290



non-zero baryon number density:  $\mu > 0$ 

$$egin{aligned} Z(oldsymbol{V},oldsymbol{T},oldsymbol{\mu}) &= \int \mathcal{D}\mathcal{A}\mathcal{D}\psi\mathcal{D}ar{\psi} \ \mathrm{e}^{-S_E(oldsymbol{V},oldsymbol{T},oldsymbol{\mu})} \ &= \int \mathcal{D}\mathcal{A}\mathcal{D} \ det \ M(oldsymbol{\mu}) \ \mathrm{e}^{-S_E(oldsymbol{V},oldsymbol{T})} \end{aligned}$$



 $\begin{array}{ll} \frac{T_{c}(\mu)}{T_{c}(0)} &: & 1 - 0.0056(4)(\mu_{B}/T)^{2} \\ & \text{deForcrand, Philipsen (imag. } \mu) \\ & 1 - 0.0078(38)(\mu_{B}/T)^{2} \\ & \text{Bielefeld-Swansea} \end{array}$ 

 $(\mathcal{O}(\mu^2) \text{ reweighting})$ 



non-zero baryon number density:  $\mu > 0$ 

$$egin{aligned} Z(oldsymbol{V},oldsymbol{T},oldsymbol{\mu}) &= \int \mathcal{D}\mathcal{A}\mathcal{D}\psi\mathcal{D}ar{\psi} \ \mathrm{e}^{-S_E(oldsymbol{V},oldsymbol{T},oldsymbol{\mu})} \ &= \int \mathcal{D}\mathcal{A}\mathcal{D} \ det \ M(oldsymbol{\mu}) \ \mathrm{e}^{-S_E(oldsymbol{V},oldsymbol{T})} \end{aligned}$$





non-zero baryon number density:  $\mu > 0$ 

$$egin{aligned} Z(oldsymbol{V},oldsymbol{T},oldsymbol{\mu}) &= \int \mathcal{D}\mathcal{A}\mathcal{D}\psi\mathcal{D}ar{\psi} \ \mathrm{e}^{-S_E(oldsymbol{V},oldsymbol{T},oldsymbol{\mu})} \ &= \int \mathcal{D}\mathcal{A}\mathcal{D} \ det \ M(oldsymbol{\mu}) \ \mathrm{e}^{-S_E(oldsymbol{V},oldsymbol{T})} \end{aligned}$$





## In-medium properties of hadrons

 properties of light quark hadrons reflect chiral symmetry breaking: Goldstone pion, non-degenerate parity partners







## In-medium properties of hadrons

- properties of light quark hadrons reflect chiral symmetry breaking: Goldstone pion, non-degenerate parity partners
- properties of heavy quark hadrons reflect structure of the heavy quark potential: quarkonium spectra

#### Deconfinement





## In-medium properties of hadrons

- properties of light quark hadrons reflect chiral symmetry breaking: Goldstone pion, non-degenerate parity partners
- properties of heavy quark hadrons reflect structure of the heavy quark potential: quarkonium spectra

thermal modifications of chiral condensate and heavy quark potential will influence the hadron spectrum



observable consequences in dilepton spectra



## Running coupling at finite r and T... remnants of confinement in the QGP

up to  $T\simeq 3T_c$ :  $\alpha_{\rm qq}(r,T)\simeq \alpha_{\rm qq}(r,0)$  for r~<~0.1 fm



T = 0: S. Necco, R. Sommer, NP B622 (2002) 328. T > 0: O. Kaczmarek, FK, P. Petreczky and F. Zantow, hep-lat/0406036

# Thermal vector meson properties from dilepton rates in heavy ion collisions

dilepton rate vs. invariant mass of  $l^+l^-$  pair



 $\overline{\mathbf{q}}$   $\mu^+$  $\gamma^*$   $\mu^-$ 

dilepton pair ( $e^+e^-$ ,  $\mu^+\mu^-$ ) production through annihilation of "thermal"  $\bar{q}q$ -pairs in hot and dense matter

# Thermal vector meson properties from dilepton rates in heavy ion collisions

dilepton rate vs. invariant mass of  $l^+l^-$  pair



 $\overline{\mathbf{q}}$   $\mu^+$  $\gamma^*$   $\mu^-$ 

dilepton pair ( $e^+e^-$ ,  $\mu^+\mu^-$ ) production through annihilation of "thermal"  $\bar{q}q$ -pairs in hot and dense matter

rate  $\sim |q\bar{q} \rightarrow \gamma^*|^2 \cdot |l^+l^- \rightarrow \gamma^*|^2$ 



. – p.20/31

# Thermal vector meson properties from dilepton rates in heavy ion collisions

dilepton rate vs. invariant mass of  $l^+l^-$  pair



 $\overline{\mathbf{q}}$   $\mu^+$  $\gamma^*$   $\mu^-$ 

dilepton pair ( $e^+e^-$ ,  $\mu^+\mu^-$ ) production through annihilation of "thermal"  $\bar{q}q$ -pairs in hot and dense matter

rate  $\sim |q\bar{q} \rightarrow \gamma^*|^2 \cdot |l^+l^- \rightarrow \gamma^*|^2$ 



differential cross-section for  $\mu^+\mu^-$  pair production  $\Rightarrow$  thermal meson correlation function



## Thermal meson correlation functions and spectral functions

Thermal correlation functions: 2-point functions which describe propagation of a  $\bar{q}q$ -pair

spectral representation of correlator  $\Rightarrow$  dilepton and photon rates



spectral representation of

Euclidean correlation functions

spectral representation of thermal photon rate:  $\omega = |\vec{p}|$  $\omega \frac{\mathrm{d}^3 R^{\gamma}}{\mathrm{d}^3 p} = \frac{5\alpha}{6\pi^2} \frac{\sigma_V(\omega, \vec{p}, T)}{\omega^2(\mathrm{e}^{\omega/T} - 1)}$ 

spectral representation of thermal dilepton rate  $\frac{\mathrm{d}^4 W}{\mathrm{d}\omega \mathrm{d}^3 p} = \frac{5\alpha^2}{27\pi^2} \frac{\sigma_V(\omega, \vec{p}, T)}{\omega^2(\mathrm{e}^{\omega/T} - 1)}$ 

$$G_H^{eta}( au,ec{r}) = \int_0^\infty \mathrm{d}\omega \,\int rac{\mathrm{d}^3ec{p}}{(2\pi)^3} \, \sigma_H(\omega,ec{p},T) \, \mathrm{e}^{iec{p}ec{r}} \, rac{\mathrm{cosh}(\omega( au-1/2T))}{\mathrm{sinh}(\omega/2T)}$$



## Charmonium suppression in heavy ion collisions (SPS, CERN)

#### Suppression Pattern <sup>ntrl</sup> Mp/Nb combinatorial background 1.4 O-Cu. O-U ⊙ J/ U 1.2 $10^{4}$ Pb 1996 (min.bias) $(J/\psi)/(D-Y)_{29-45} = 17.1 \pm 0.3 \pm 0.1$ <sup>2</sup>h-Pb 1998 Fit for M > 2.85 $\gamma^{2}/dof = 1.03$ 10 S<sub>exp</sub> / S<sub>GI</sub> 0.8 $10^{2}$ Drell-Yan 0.6 open charm 10 0.4 u. p-W. p-U (200 GeV) 0.2 C, p-Al, p-Cu, p-W (450 GeV) 1 100 150 200 250 300 350 400 2 2 3 50 3 5 6 4 Number of participants $\sim$ energy density $M_{\mu\mu}$ (GeV/c<sup>2</sup>)

#### Charmonium Production: dilepton rate

invariant mass of the pair

differential cross-section for  $\mu^+\mu^-$  pair production

measured A-A rate normalized to rate expected from known p-A collisions

M.C.Abreu (NA50), Phys.Lett. B477 (2000) 28



## Charmonium suppression in heavy ion collisions (SPS, CERN)

#### Charmonium Production: dilepton rate



#### invariant mass of the pair

differential cross-section for  $\mu^+\mu^-$  pair production

#### Suppression Pattern



use Bjorken formula to convert participants to energy density

 $\mathrm{d}E_T$ 



## Heavy quark spectral functions and correlation functions

#### reconstructed correlation functions



reconstructed spectral functions using the Maximum Entropy Method





## Heavy quark spectral functions and correlation functions



reconstructed spectral functions using the Maximum Entropy Method





## Heavy quark spectral functions and correlation functions



reconstructed spectral functions using the Maximum Entropy Method





## Outlook: Next generation lattice calculations

- Thermodynamics of pure gauge theory has been "solved" on (1-10)GFlops computers (1996)
- Thermodynamics of QCD with "still too heavy" quarks has been studied on (10-100) GFlops computers
- Analysis of "continuum and thermodynamic limit" of QCD thermodynamics with light quarks, requires computers with  $\sim 10$  TFlops peak speed.(LatFor, 2003)

#### QCDOC and apeNEXT:

scaling to 10's of Teraflops

with \$1/MFlops Cost/performance





## Progress in lattice calculations... depends on...

- development of (special purpose) computer hardware
- ٩





## Progress in lattice calculations... depends on...

### special purpose computer hardware

contribution of research done on special purpose computer to 10 top cited papers in LGT (1999-2004)





## Outlook: Next generation computers for lattice gauge theory



today: APEmille

so far the only dedicated large-scale computer installation used predominantly for QCD thermodynamics exists in Bielefeld: 120 GFlops



## Outlook: Next generation computers for lattice gauge theory

## **QCDOC** and apeNEXT



QCD thermodynamics on the next generation of special purpose dedicated QCD computers

installations with (10-20) TFlops peak speed are planned in the USA and Europe



## apeNEXT:

### Next generation of APE computers



BackPlane



## apeNEXT:

### Next generation of APE computers



#### BackPlane



### apeNEXT:

## Next generation of APE computers



BackPlane


# QCDOC: Next generation of Columbia-RIKEN computer

#### $Columbia-RIKEN-UKQCD\ Collaboration$



- 2 node daughter card
- prototypes exist since 07/2003



64 - node mother board



# QCDOC: Next generation of Columbia-RIKEN computer

### $Columbia-RIKEN-UKQCD\ Collaboration$





512 - node machine: (360 - 450) GFlops

currently being debugged prototype (05/2004):
0.25 Tbyte memory; 6 Gbit/sec Ethernet I/O bandwidth

### $QCDOC\ computing\ center\ at\ BNL:$

- 10 TFlops machine for RBRC:  $\sim$  autumn 2004
- $\bullet$  10 TFlops machine for american LGT community:  $\sim$  early 2005
- •... larger installations possible and needed!



 $\bullet$  LGT calculations contributed a lot to the understanding/interpretation of HIC



- $\bullet$  LGT calculations contributed a lot to the understanding/interpretation of HIC
- a major effort is still needed to provide results from calculations with the physical quark mass spectrum close to the continuum limit



- $\bullet$  LGT calculations contributed a lot to the understanding/interpretation of HIC
- a major effort is still needed to provide results from calculations with the physical quark mass spectrum close to the continuum limit
- Wish list for 2004/05 is ready: QCD thermodynamics on Teraflops computers



- $\bullet$  LGT calculations contributed a lot to the understanding/interpretation of HIC
- a major effort is still needed to provide results from calculations with the physical quark mass spectrum close to the continuum limit
- Wish list for 2004/05 is ready: QCD thermodynamics on Teraflops computers



Also these calculations have to face the well-known lattice beasts: ultra-violet and infra-red problems



- $\bullet$  LGT calculations contributed a lot to the understanding/interpretation of HIC
- a major effort is still needed to provide results from calculations with the physical quark mass spectrum close to the continuum limit
- Wish list for 2004/05 is ready: QCD thermodynamics on Teraflops computers



Also these calculations have to face the well-known lattice beasts: ultra-violet and infra-red problems