

Muon Cooling and the Muon Collider

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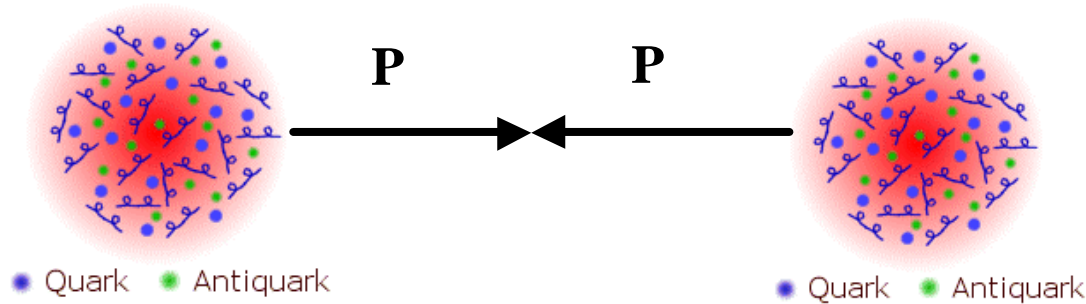
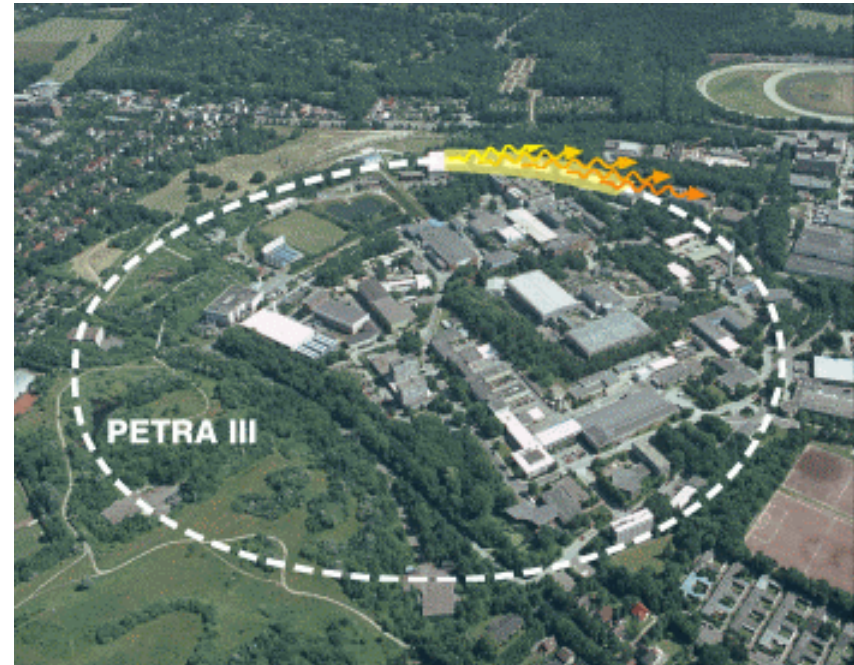
- Motivation
- Difficulties
- Focus on Cooling (frictional cooling)

Why a Muon Collider ?

No synchrotron radiation
problem (cf electron)

$$P \propto (E/m)^4$$

Can build a high energy circular
accelerator.

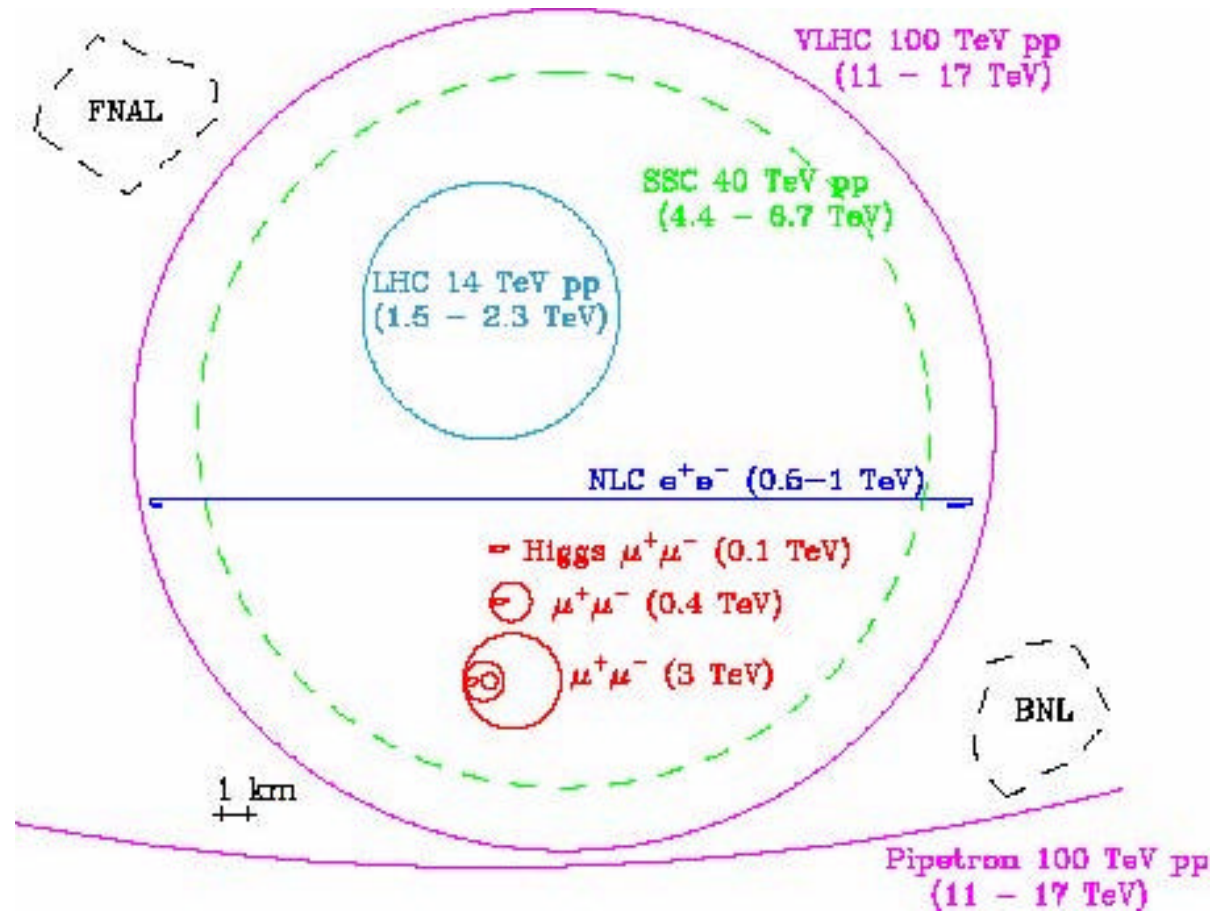


Collide point
particles rather than
complex objects

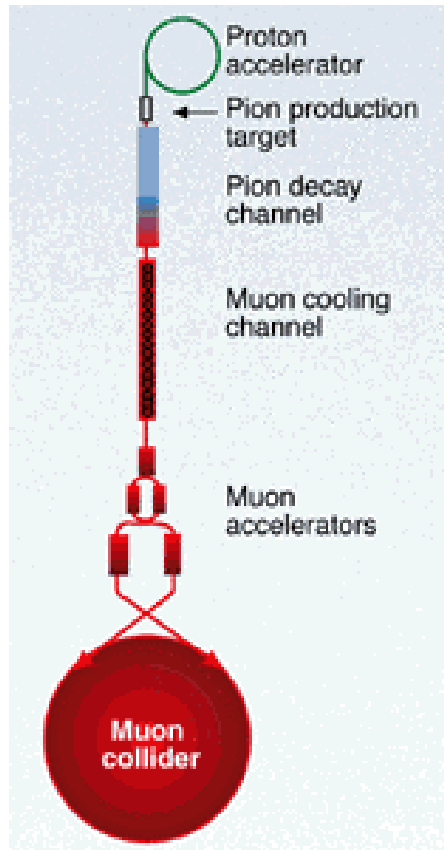
Some Difficulties

- Muons decay, so are not readily available – need multi MW source. Large starting cost.
- Muons decay, so time available for cooling, bunching, acceleration is very limited. Need to develop new techniques, technologies.
- Large experimental backgrounds from muon decays (for a collider). Not the usual clean electron collider environment.
- High energy colliders with high muon flux will face critical limitation from neutrino induced radiation.

Dimensions of Some Colliders discussed



Physics at a Muon Collider



Muon Collider Complex:

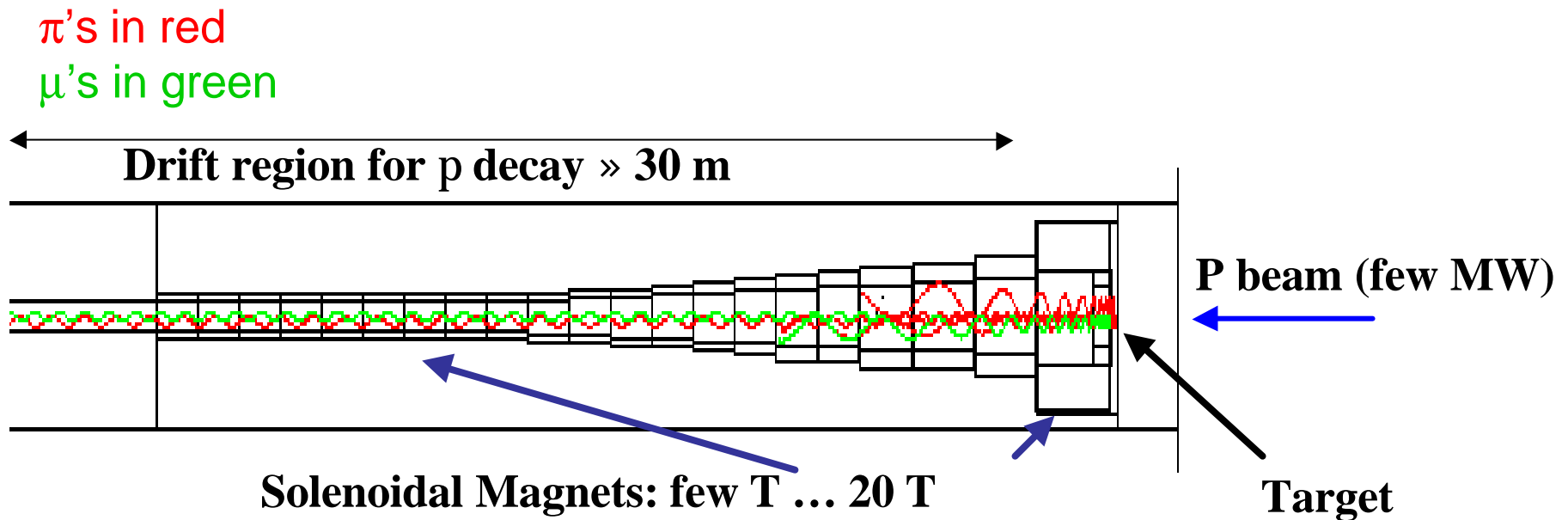
- Proton Driver 2-16GeV; 1-4MW leading to 10^{22} p/year
- π production target & Strong Field Capture
- COOLING resultant μ beam
- μ acceleration
- Storage & collisions

- Stopped μ physics
- ν physics From target, stored μ
- Higgs Factory $\sigma_{\mu\mu \rightarrow H^-} = 40000 \sigma_{ee \rightarrow H}$
- Higher Energy Frontier

HIGH ENERGY MUON COLLIDER PARAMETERS

Baseline parameters for high energy muon colliders. From "Status of Muon Collider Research and Development and Future Plans," Muon Collider Collaboration, C. M. Ankenbrandt *et al.*, *Phys. Rev. ST Accel. Beams* **2**, 081001 (1999).

COM energy (TeV)	0.4	3.0
ρ energy (GeV)	16	16
ρ 's/bunch	2.5×10^{13}	2.5×10^{13}
Bunches/fill	4	4
Rep. rate (Hz)	15	15
ρ power (MW)	4	4
μ / bunch	2×10^{12}	2×10^{12}
μ power (MW)	4	28
Wall power (MW)	120	204
Collider circum. (m)	1000	6000
Ave bending field (T)	4.7	5.2
rms $\delta\rho/\rho$ (%)	0.14	0.16
$6D \varepsilon_{6,N} (\pi\text{m})^3$	1.7×10^{-10}	1.7×10^{-10}
rms $\varepsilon_n (\pi \text{ mm mrad})$	50	50
β^* (cm)	2.6	0.3
σ_z (cm)	2.6	0.3
σ_r spot (μm)	2.6	3.2
σ_θ IP (mrad)	1.0	1.1
Tune shift	0.044	0.044
n_{turns} (effective)	700	785
Luminosity ($\text{cm}^{-2} \text{ s}^{-1}$)	10^{33}	7×10^{34}



Simplified emittance estimate:

At end of drift, rms x,y,z approx 0.05,0.05,10 m

$$P_x, P_y, P_z \text{ approx } 50, 50, 100 \text{ MeV}/c$$

Normalized 6D emittance is product divided by $(m_\mu c)^3$

$$\epsilon_{6D,N}^{\text{drift}} \approx 1.7 \cdot 10^{-4} (\pi m)^3$$

Emittance needed for Muon Collider

$$\epsilon_{6D,N}^{\text{collider}} \approx 1.7 \cdot 10^{-10} (\pi m)^3$$

This reduction of 6 orders of magnitude must be done with reasonable efficiency !

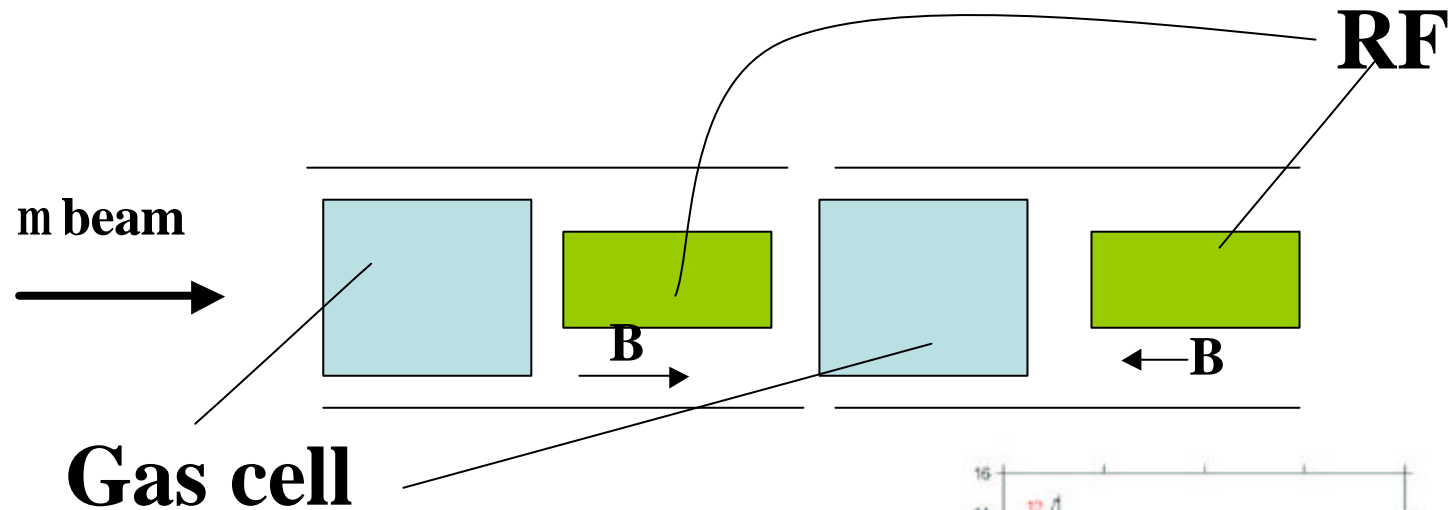
Muon Cooling

Muon Cooling is the signature challenge of a Muon Collider

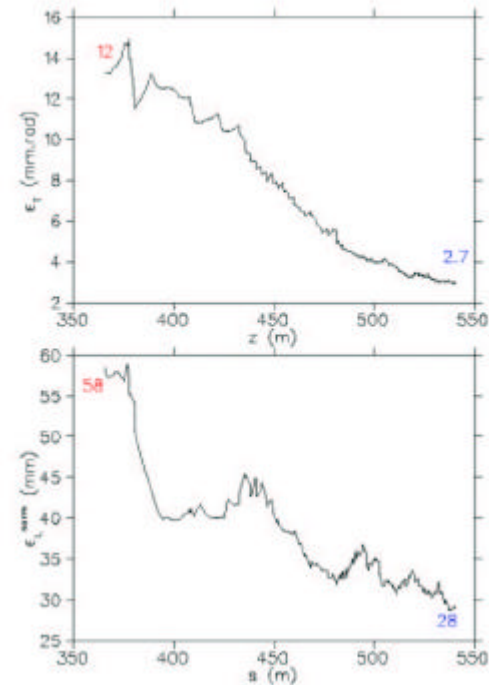
Cooler beams would allow fewer muons for a given luminosity, thereby

- Reducing the experimental background
- Reducing the radiation from muon decays
- Reducing the radiation from neutrino interactions
- Allowing for smaller apertures in machine elements, and so driving the cost down

Cooling Ideas

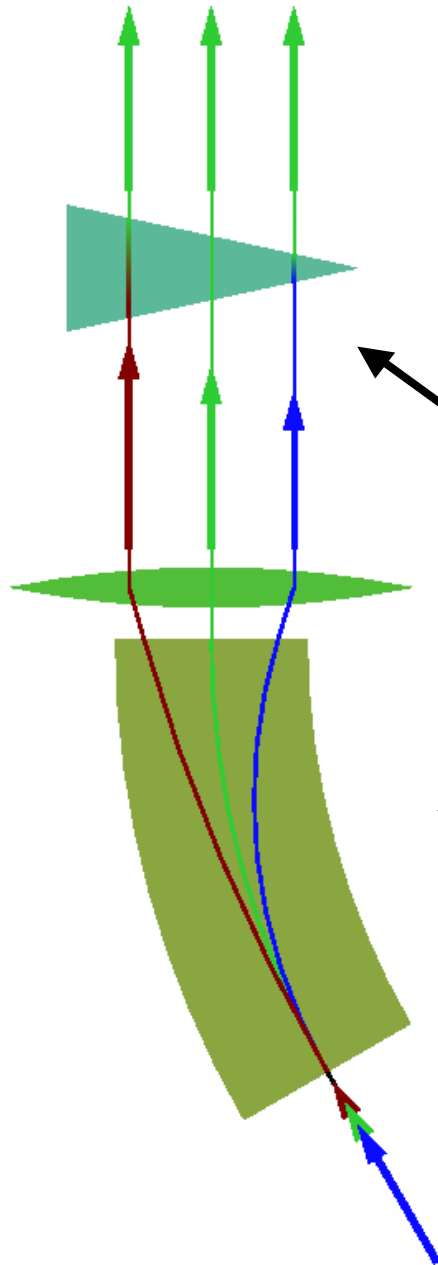


Ionization cooling (Skrinsky, Neuffer, Palmer, ...): muons are maintained at ca. 200 MeV. Transverse cooling of order x20 seems feasible (see v-factory feasibility studies 1-2). Longitudinal cooling not solved.



Longitudinal Cooling via Emittance Exchange

Transform longitudinal phase space into transverse (know how to cool transverse)



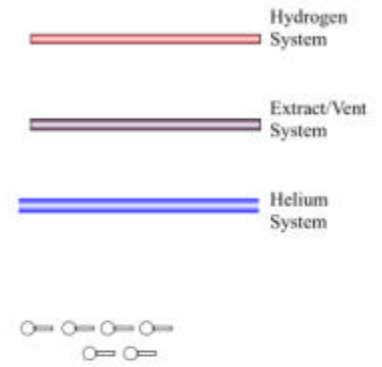
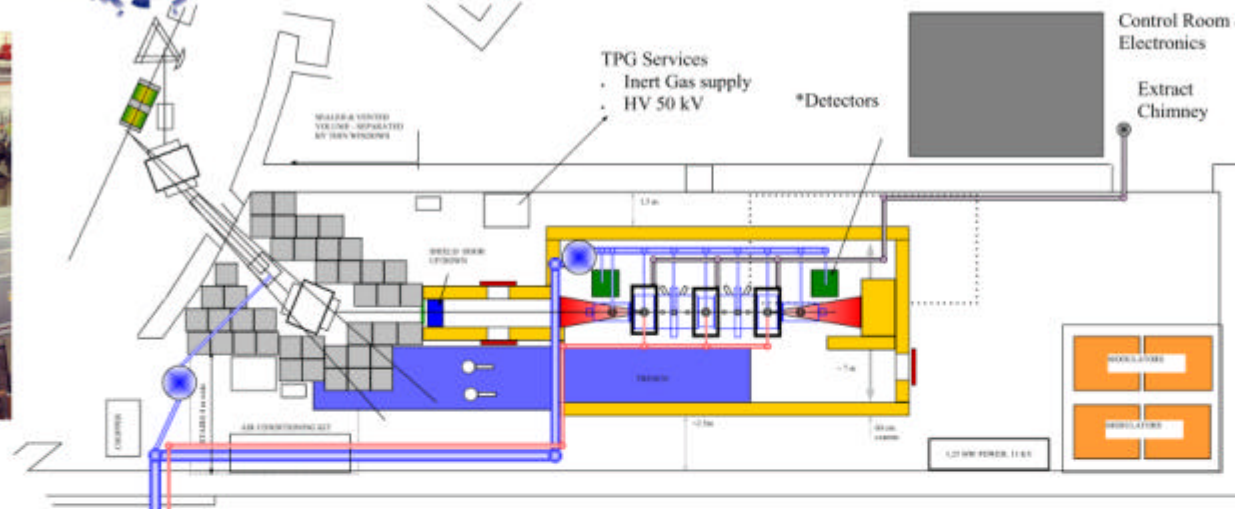
Wedge shaped absorber

Bent solenoid produces dispersion

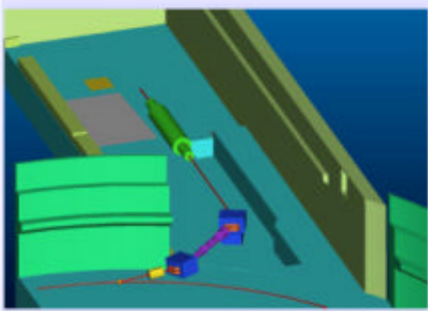
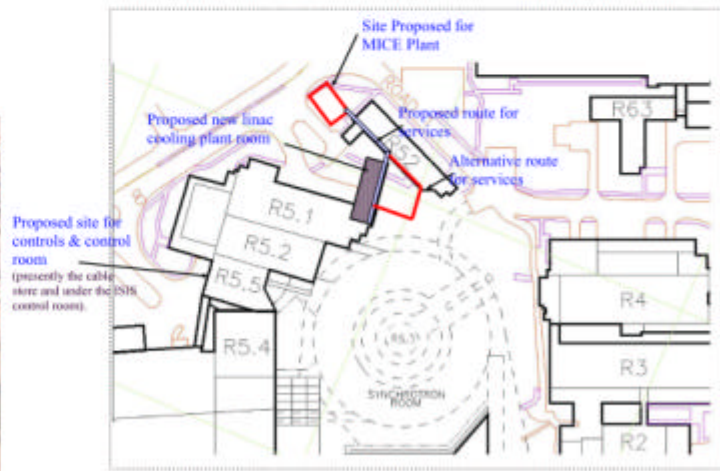
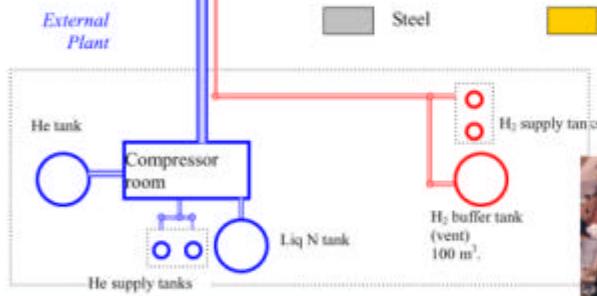
ISIS



Muon Ionization Cooling Experiment

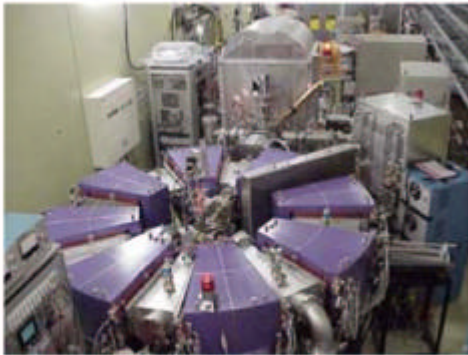


*TPG: Electronics
 *SCIFI: Electronics & Cryostat



at the Rutherford Appleton Laboratory

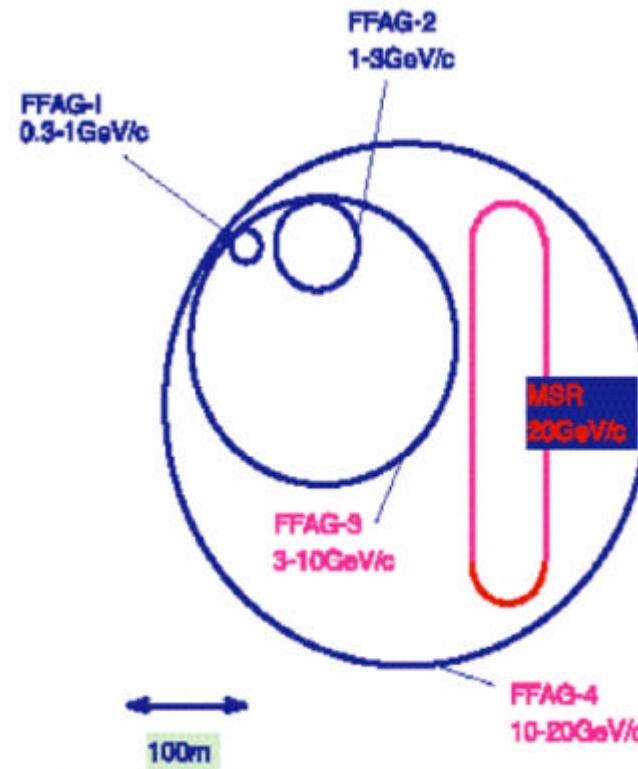
KEK – focus on FFAG acceleration



0.5-MeV Proton FFAG
POP at KEK



150-MeV Proton FFAG
Under construction at KEK



Series of FFAG acceleration

Frictional Cooling

Studies at Columbia University/Nevis Labs

Allen Caldwell, Raphael Galea

+

Stefan Schlenstedt (DESY/Zeuthen)

Halina Abramowicz (Tel Aviv University)



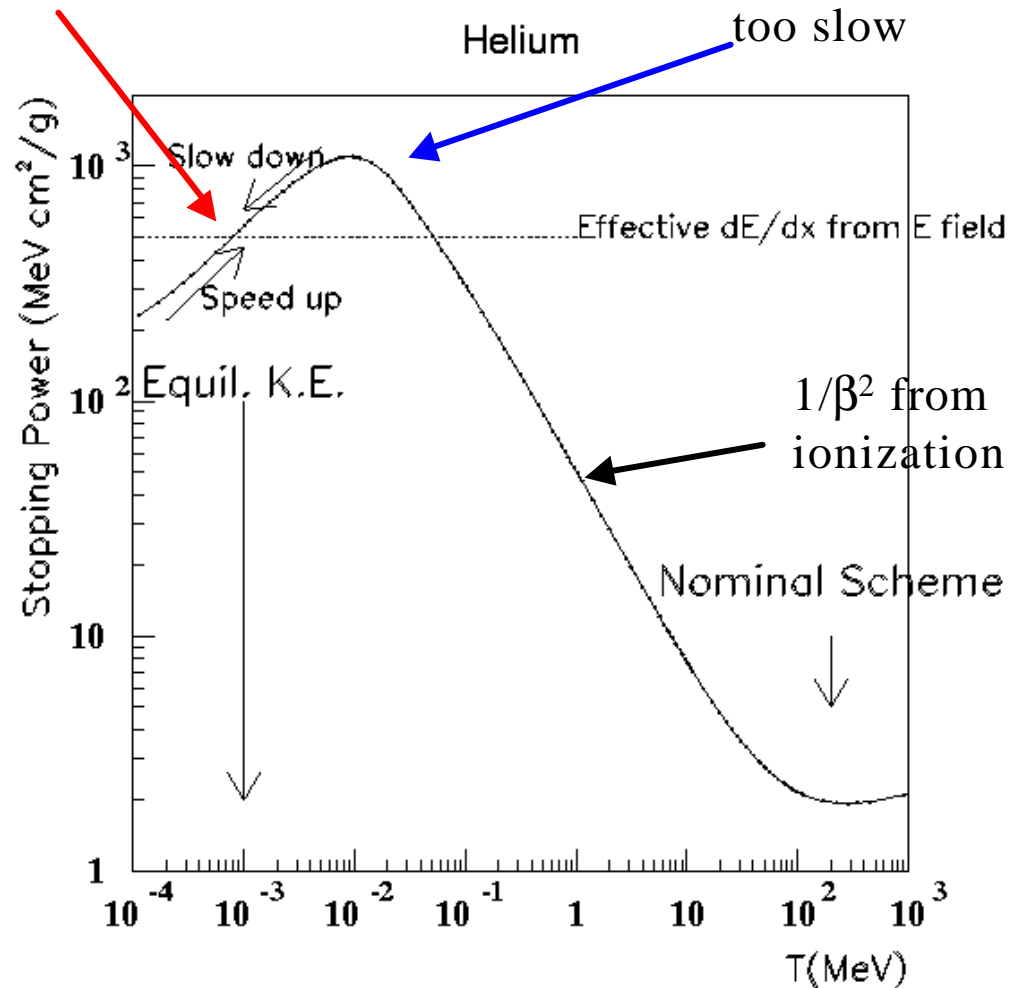
Summer Students: Emily Alden
Christos Georgiou
Daniel Greenwald
Laura Newburgh
Yujin Ning
Inna Shpiro
Will Serber

Frictional Cooling

Nuclear scattering, excitation,
charge exchange, ionization

Ionization
stops, muon
too slow

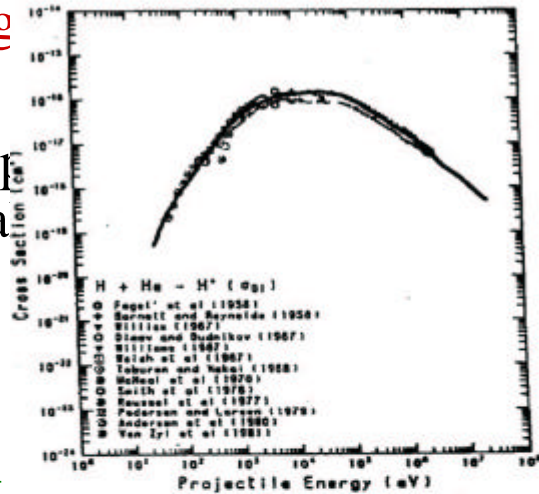
- Bring muons to a kinetic energy (T) where dE/dx increases with T
- Constant **E-field** applied to muons resulting in **equilibrium energy**
- Big issue – how to maintain efficiency
- Similar idea first studied by Kottmann et al., PSI



Problems/Comments

Y. NAKAI et al. Charge T

GRAPH IV. Cross Section vs Energy. σ_{01}
See page 74 for Explanation



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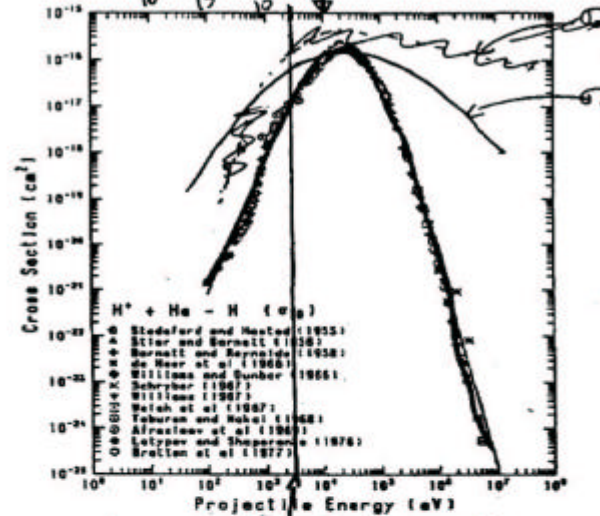
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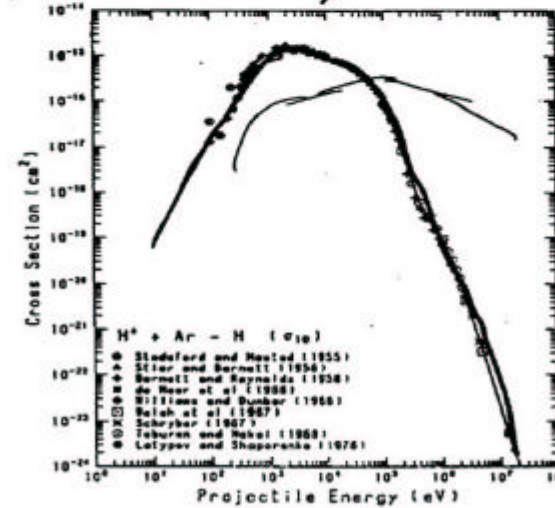
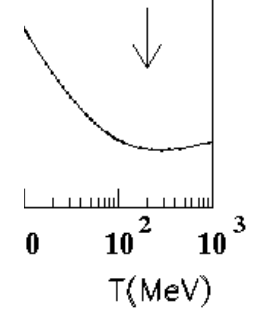
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GRAPH I. Cross Section vs Ener
See page 74 for E

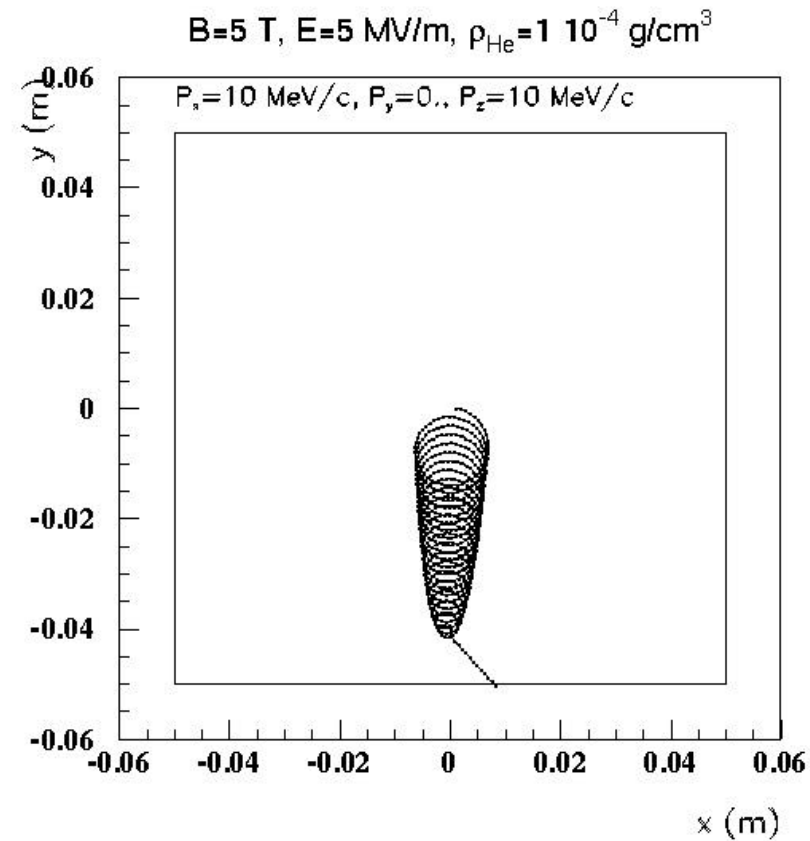
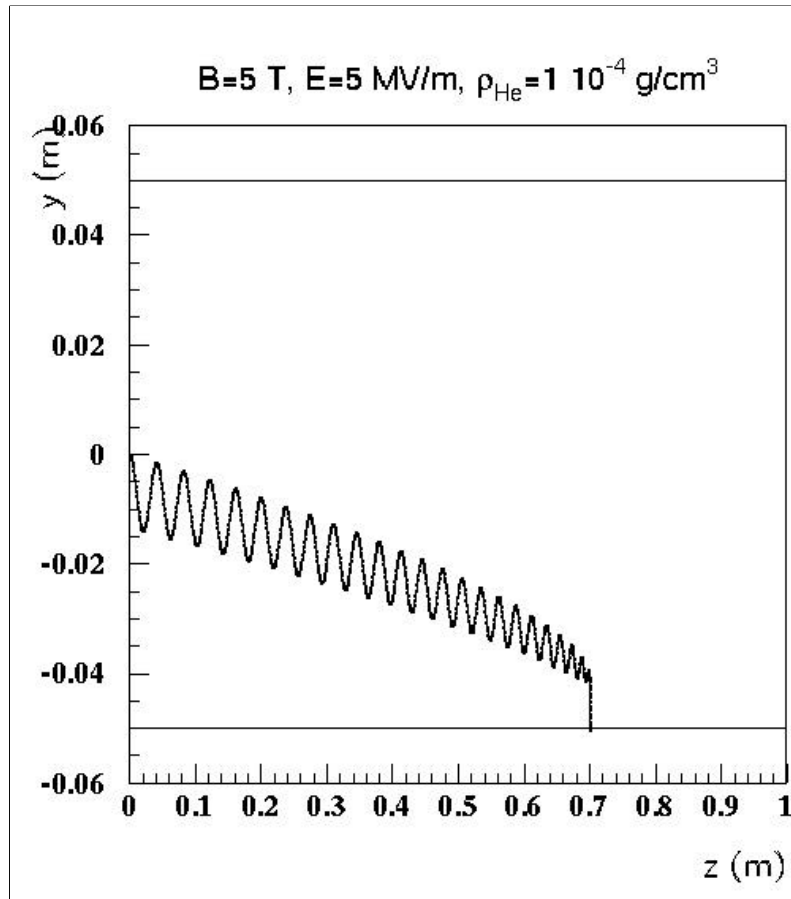


re dE/dx from E field

ominal Scheme



Frictional Cooling: particle trajectory

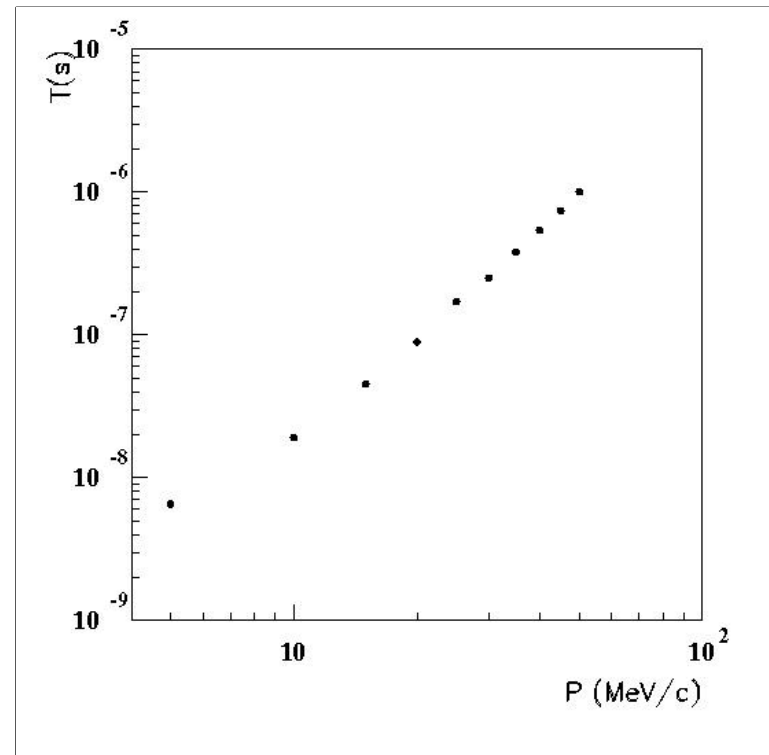
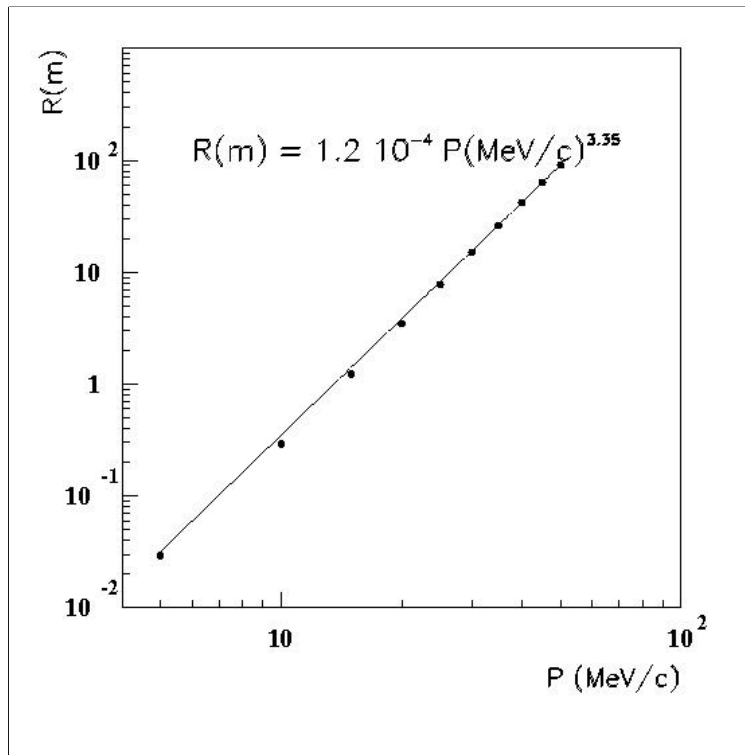


**** Using continuous energy loss**

Frictional Cooling: stop the m

- High energy μ 's travel a long distance to stop
- High energy μ 's take a long time to stop

Plots for
1. $10^{-4} \text{ g/cm}^3 \text{ He}$

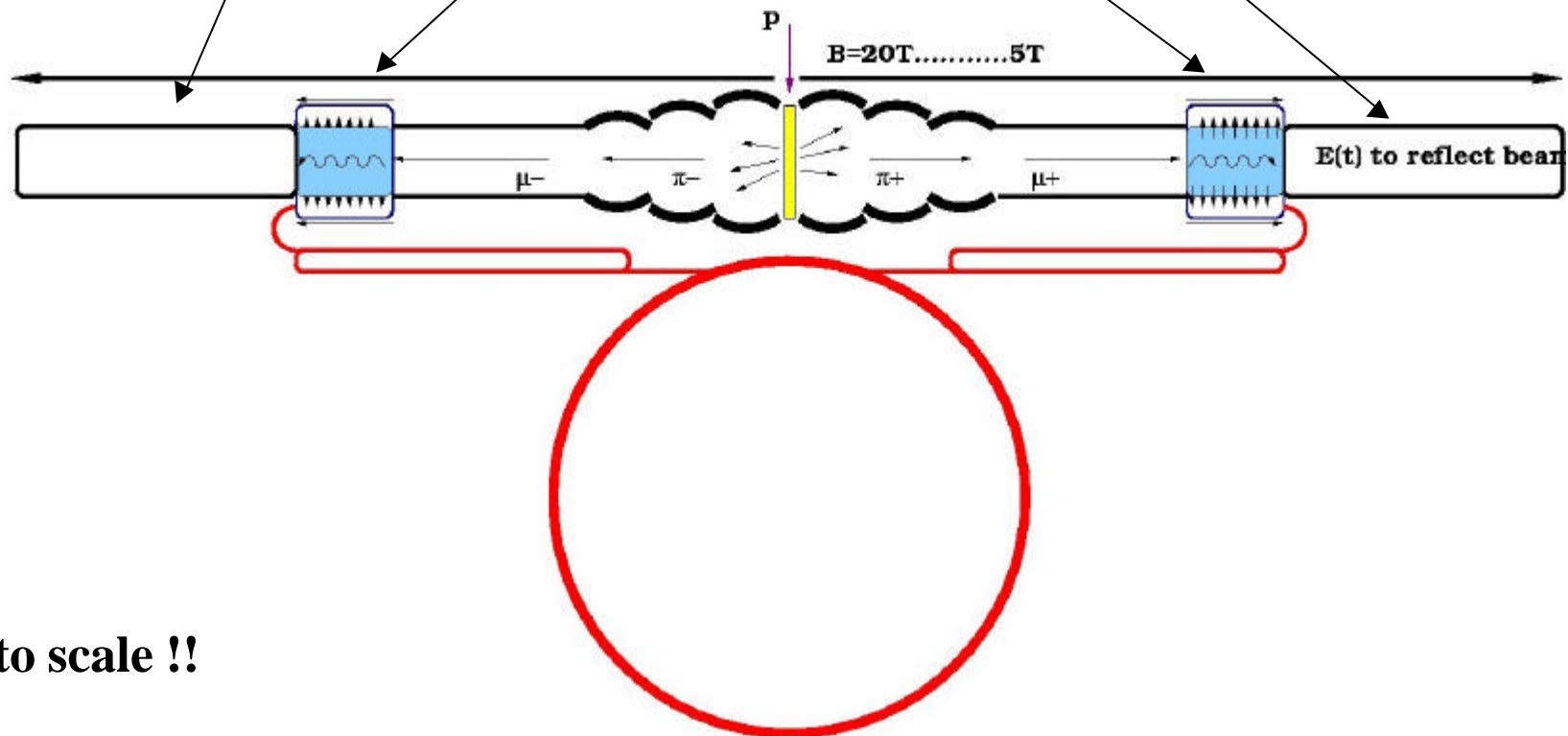


Optimize for low initial muon momentum,
+ phase rotation

Schematic Layout for m Collider Scheme Studied

Phase rotation sections

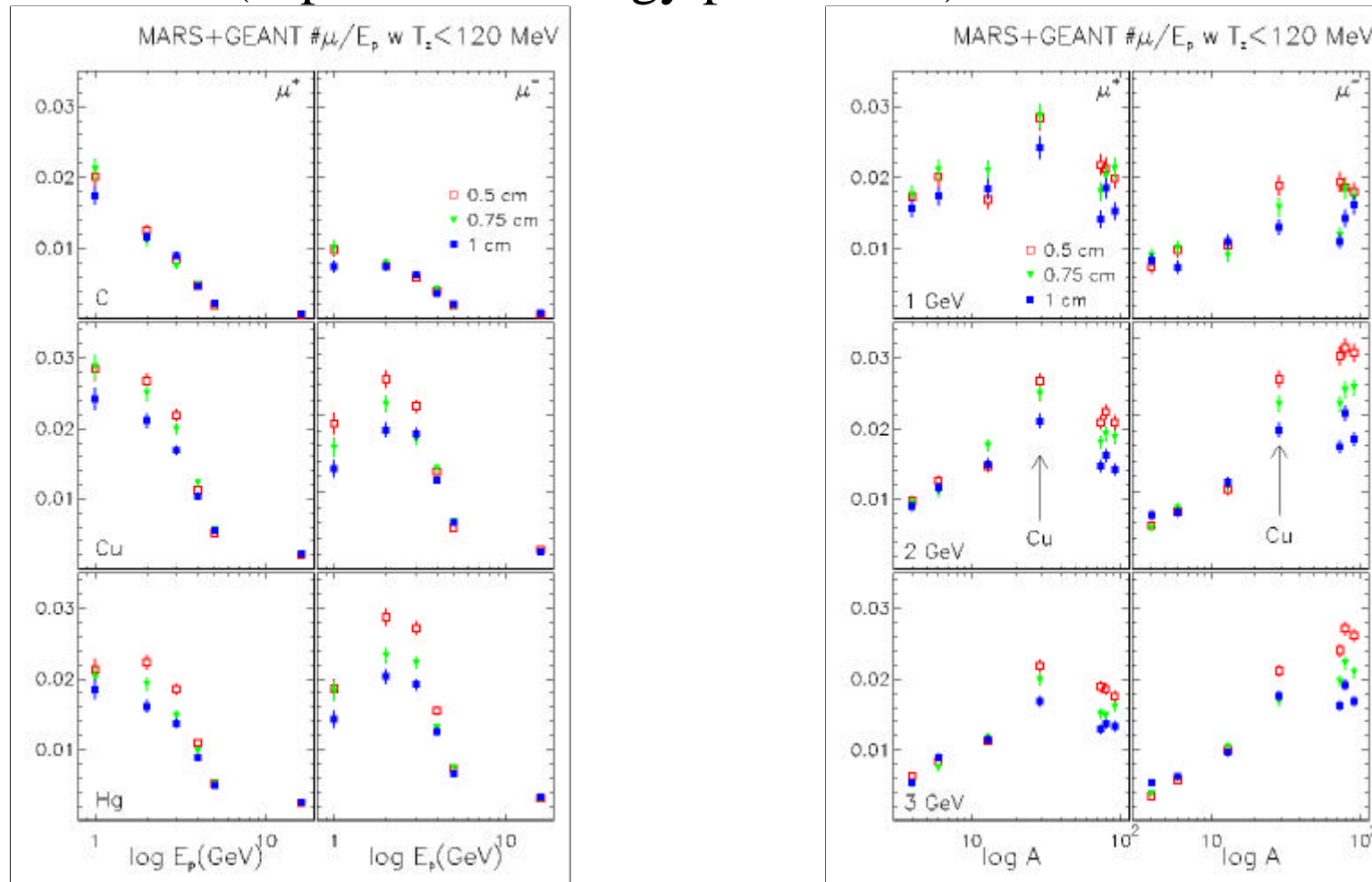
Cooling cells

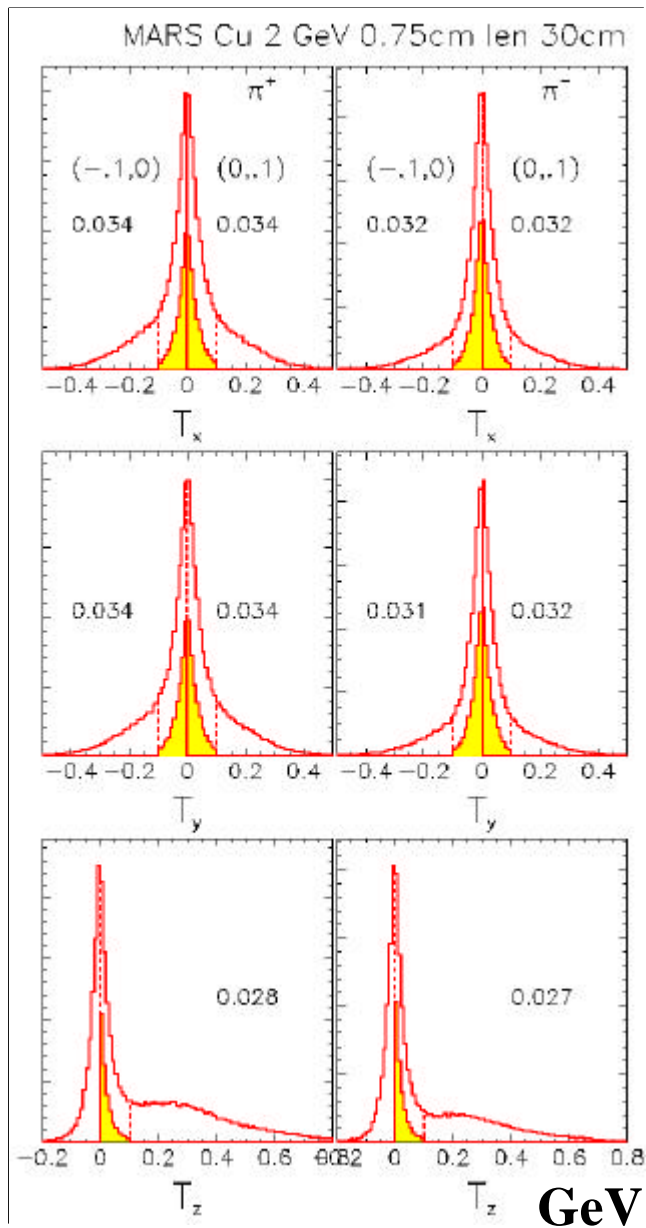


Not to scale !!

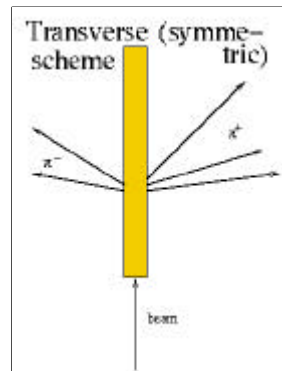
Detailed Simulation

Full MARS target simulation, optimized for low energy muon yield: **2 GeV protons** on **Cu** with proton beam transverse to solenoids (capture low energy pion cloud).

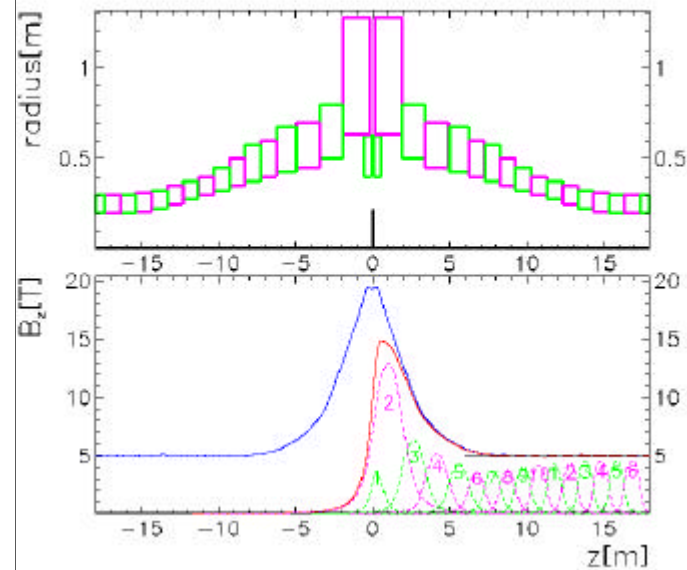




Target System

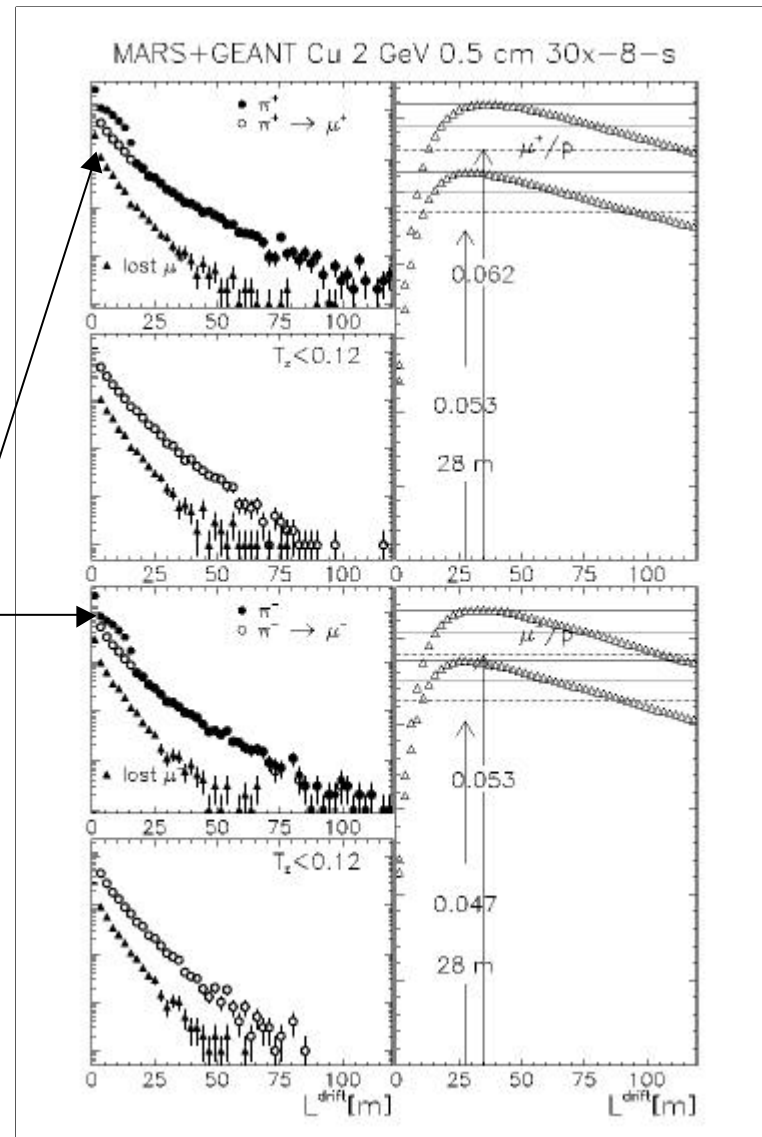


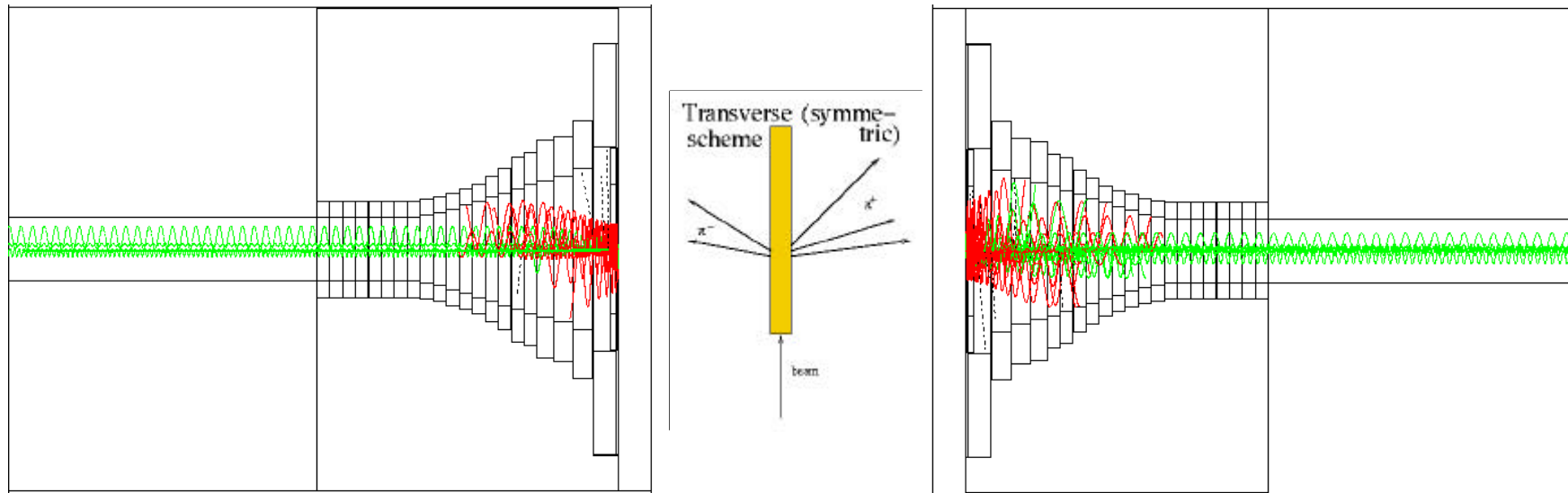
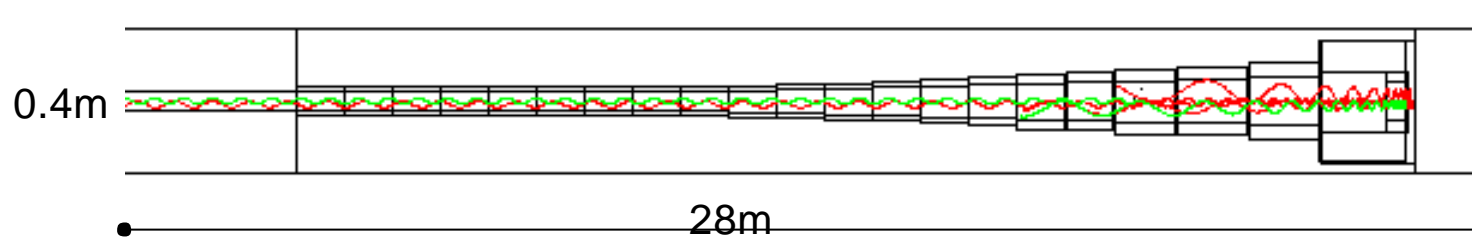
- cool μ^+ & μ^- at the same time
- calculated new symmetric magnet with gap for target



Target & Drift Optimize yield

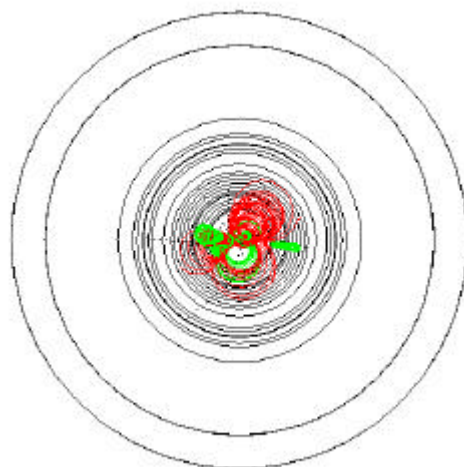
- Optimize drift length for μ yield
- Some π 's lost in Magnet aperture





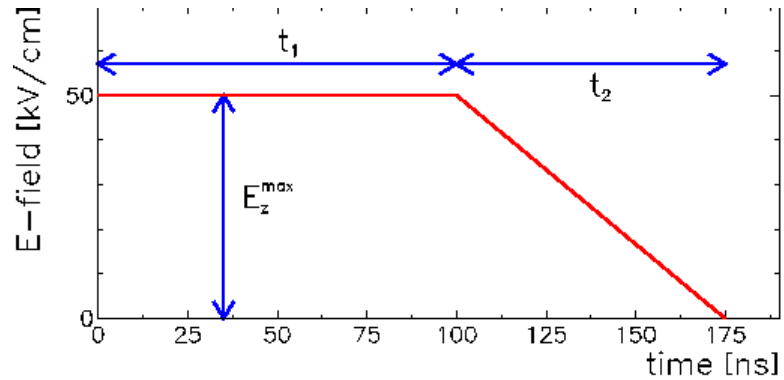
π 's in red
 μ 's in green

GEANT simulation

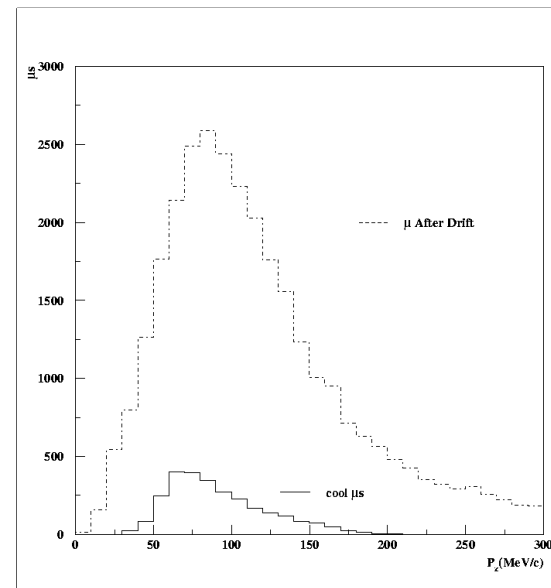
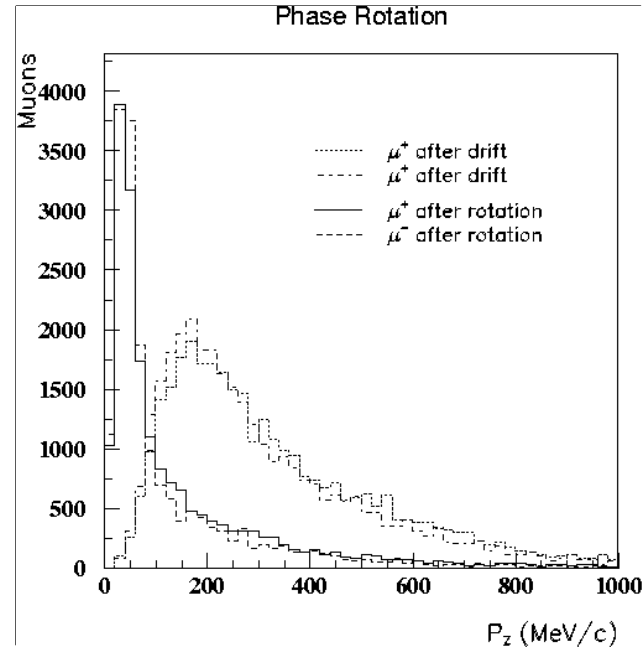


View into beam

Phase Rotation



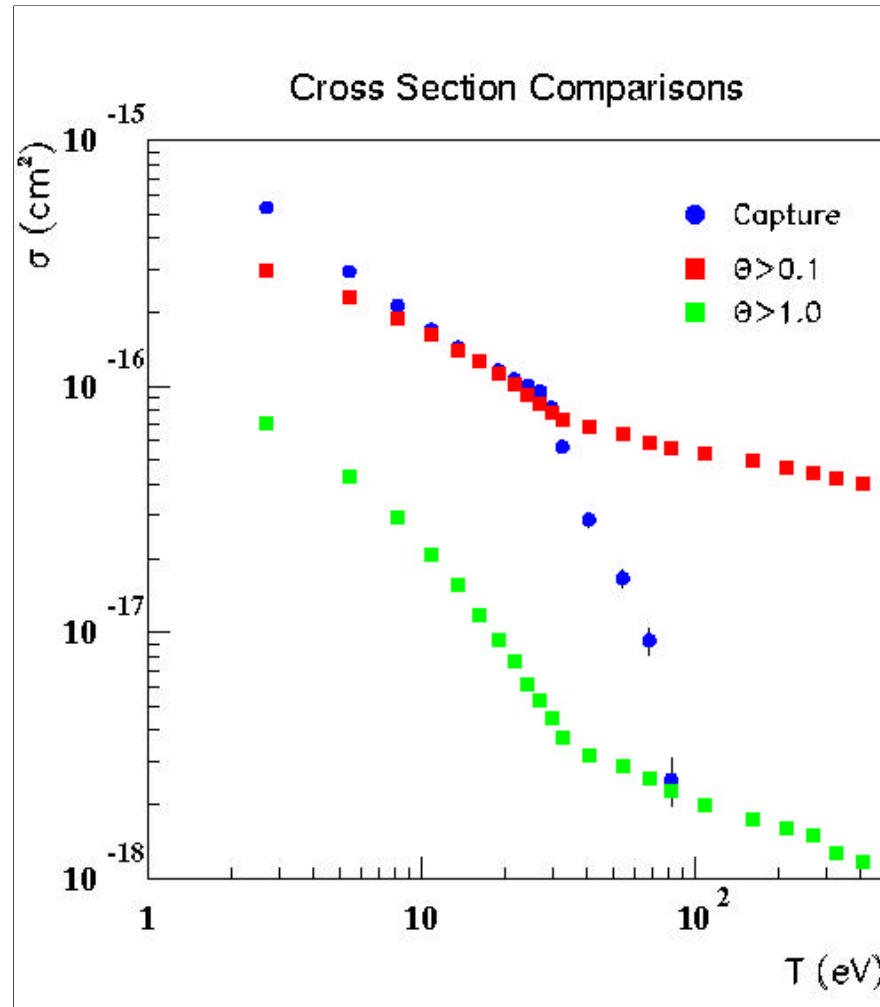
- First attempt simple form
- Vary t_1, t_2 & E_{\max} for maximum low energy yield



Cooling cell simulation

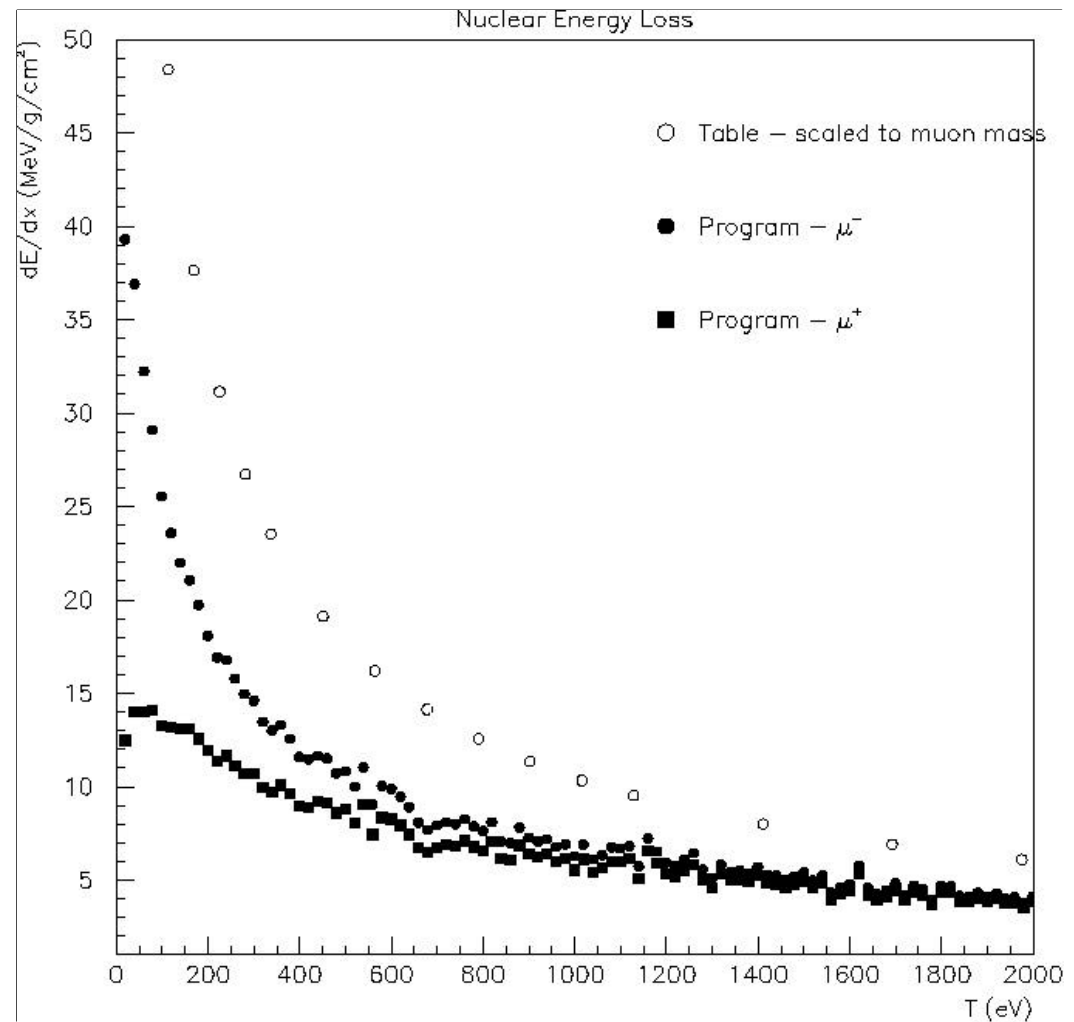
He gas is used for μ^+ , H_2 for μ^- .

- Individual nuclear scatters are simulated – crucial in determining final phase space, survival probability.
- Incorporate scattering cross sections into the cooling program
- Include μ^- capture cross section using calculations of Cohen (Phys. Rev. A. Vol 62 022512-1)
- Electronic energy loss treated as continuous

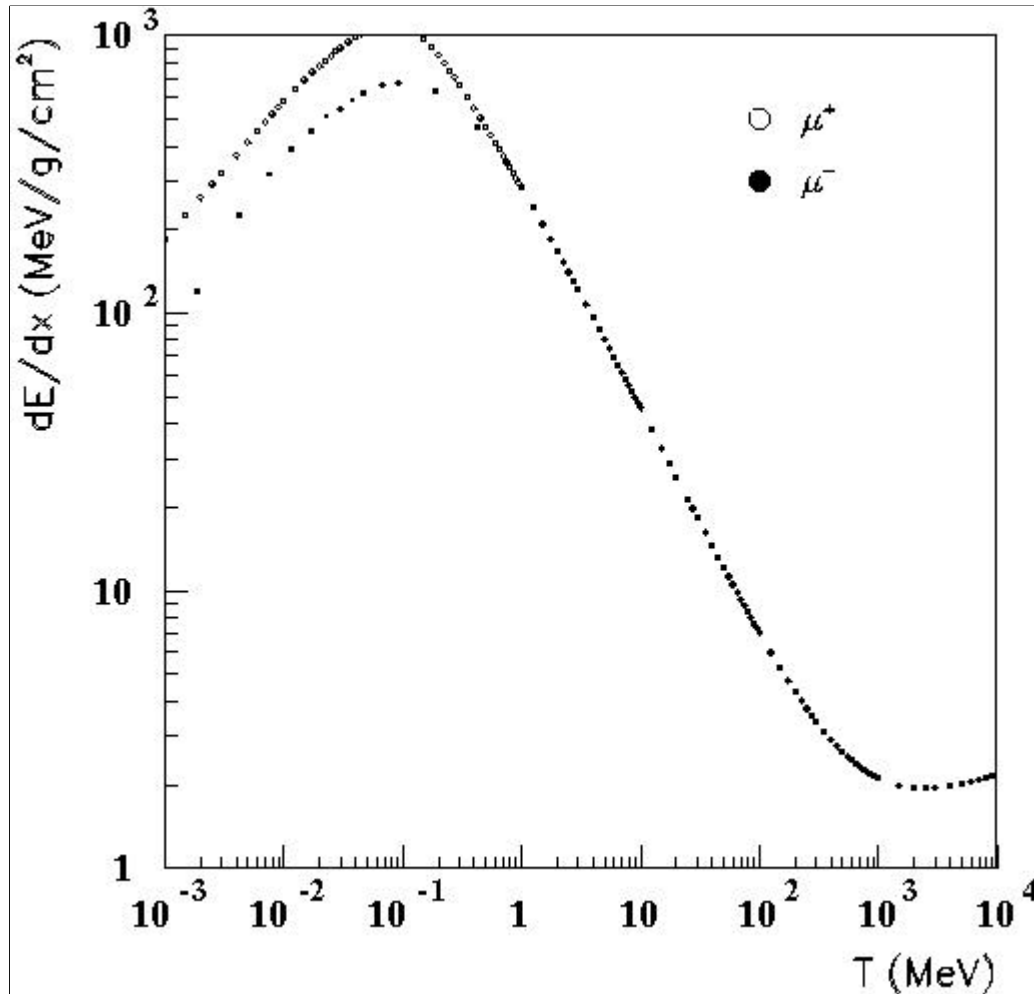


Scattering Cross Sections

- Scan impact parameter and calculate $\theta(b)$, $d\sigma/d\theta$ from which one can get $\lambda_{\text{mean free path}}$
- Use screened Coulomb Potential (Everhart et. al. Phys. Rev. 99 (1955) 1287)
- Simulate all scatters $\theta > 0.05$ rad
- Simulation accurately reproduces ICRU tables



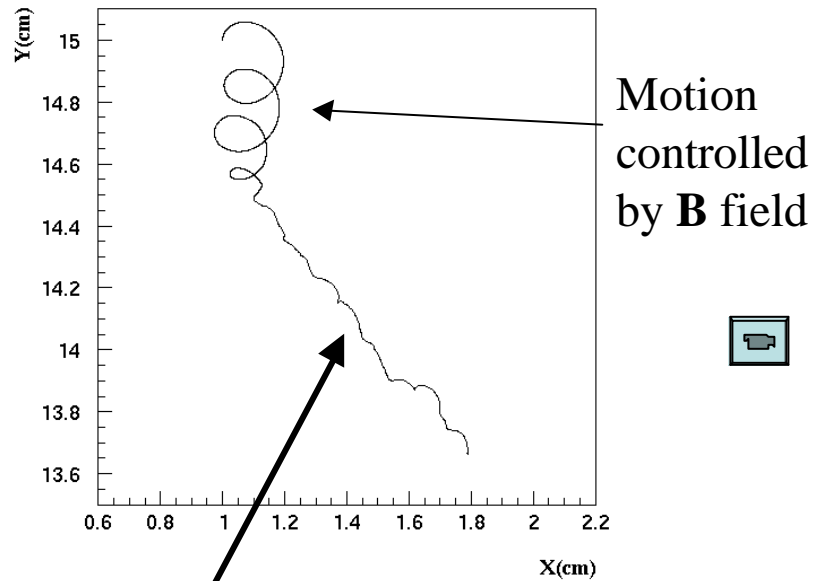
Barkas Effect



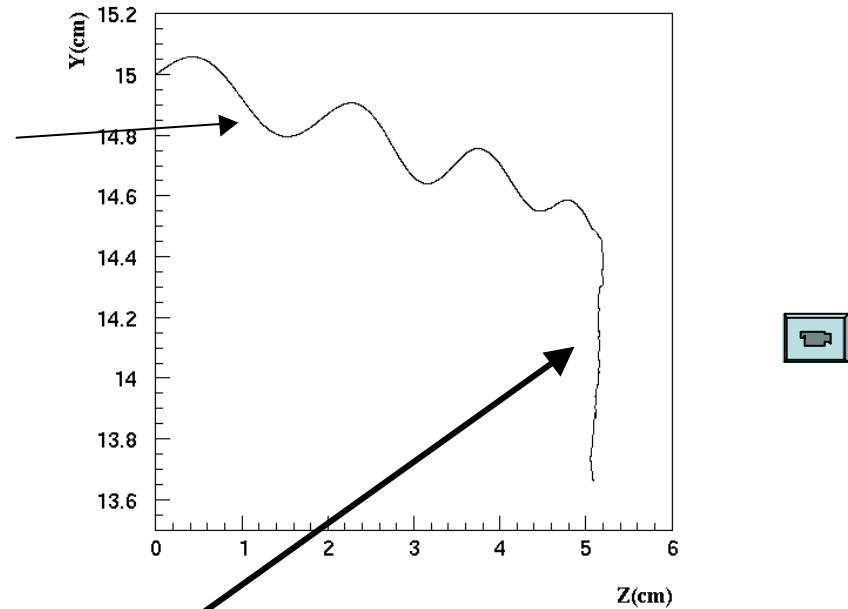
- Difference in μ^+ & μ^- energy loss rates at dE/dx peak
- Due to charge exchange for μ^+
- parameterized data from Agnello et. al. (*Phys. Rev. Lett.* 74 (1995) 371)
- Only used for the electronic part of dE/dx

Trajectories in detailed simulation

Transverse motion



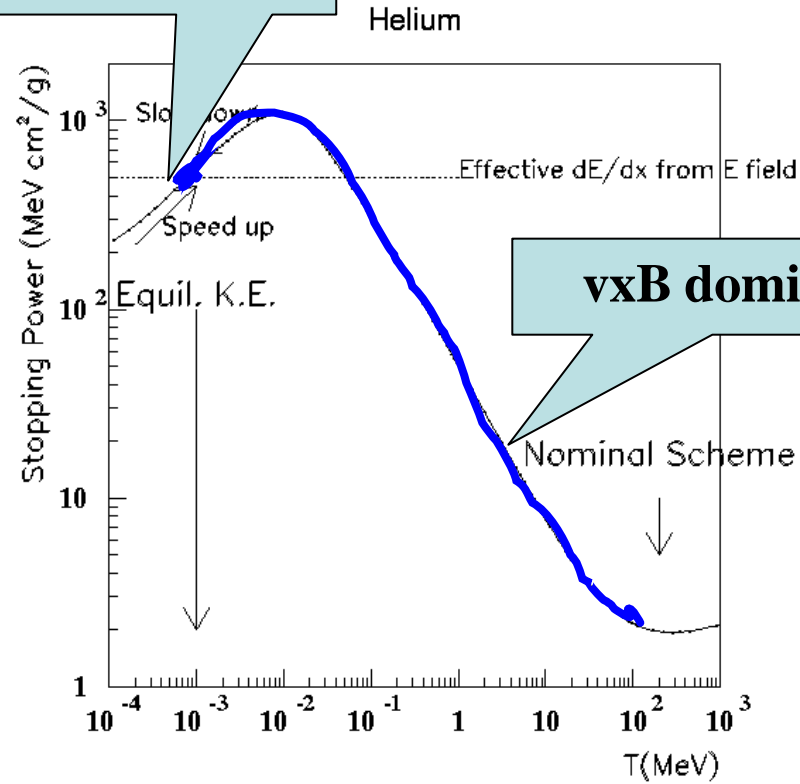
Longitudinal motion



Lorentz angle drift, with nuclear scattering

Final stages of muon trajectory in gas cell

E dominates

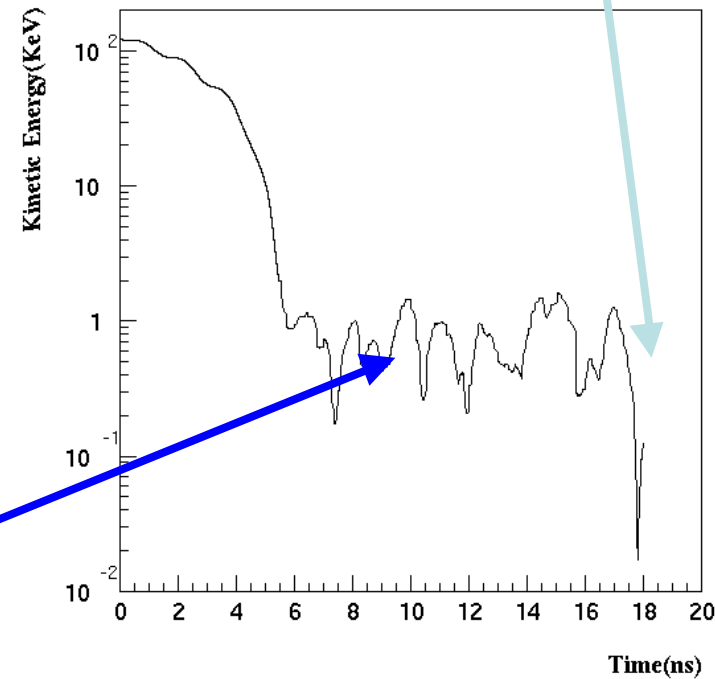


$$\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B}) - d\mathbf{T}/dx$$

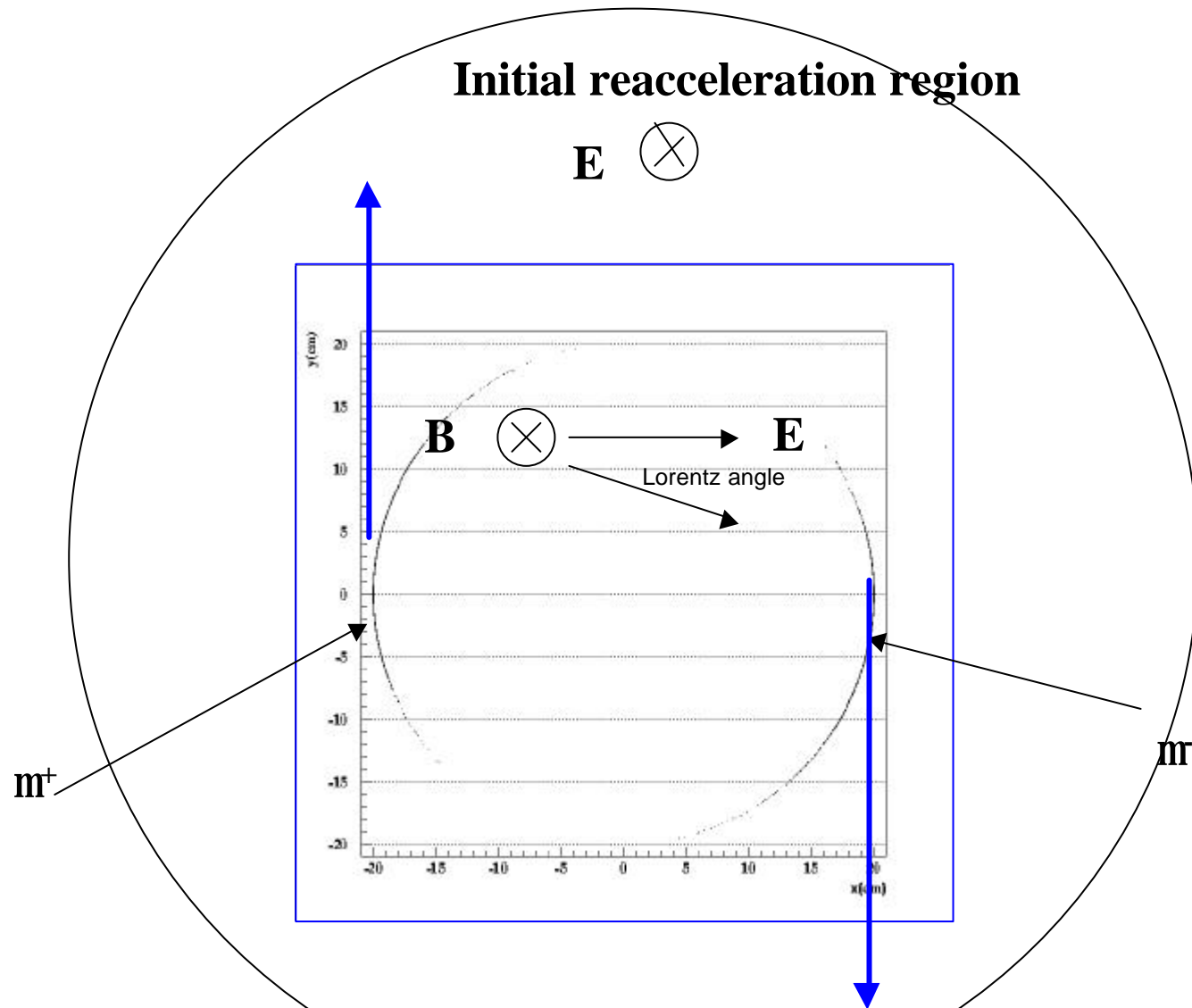
vxB dominates

Dangerous because of μ^- capture possibility

Oscillations about equilibrium define emittance.



After the gas cell



Yields & Emittance

Look at muons coming out of 11m cooling cell region after initial reacceleration.

Yield: approx 0.002 μ per 2GeV proton after cooling cell.
Need to improve yield by factor 3 or more.

Emittance: rms

$x = 0.015$ m
$y = 0.036$ m
$z = 30$ m (actually βct)
$P_x = 0.18$ MeV
$P_y = 0.18$ MeV
$P_z = 4.0$ MeV

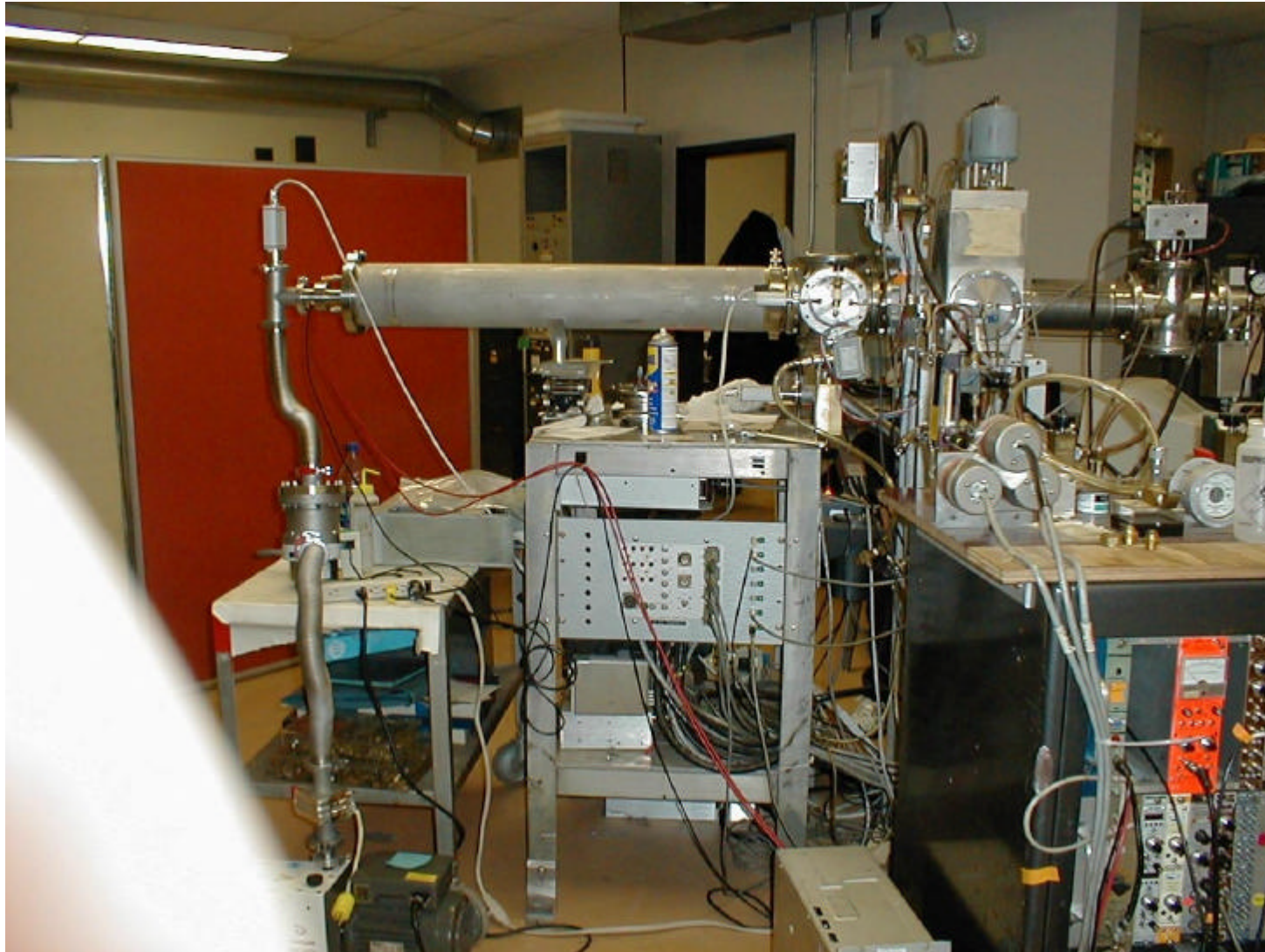
$$\epsilon_{6D,N} = 5.7 \cdot 10^{-11} (\pi m)^3$$

Problems/Things to investigate...

- Extraction of μs through window in gas cell
 - Must be very thin to pass low energy μs
 - Must be reasonably gas tight
- Can we apply high electric fields in gas cell without breakdown (large number of free electrons, ions) ?
Plasma generation \rightarrow screening of field.
- Reacceleration & bunch compression for injection into storage ring
- The μ^- capture cross section depends very sensitively on kinetic energy & falls off sharply for kinetic energies greater than e^- binding energy. NO DATA – simulations use theoretical calculation
- +...

First try to demonstrate frictional cooling with protons.

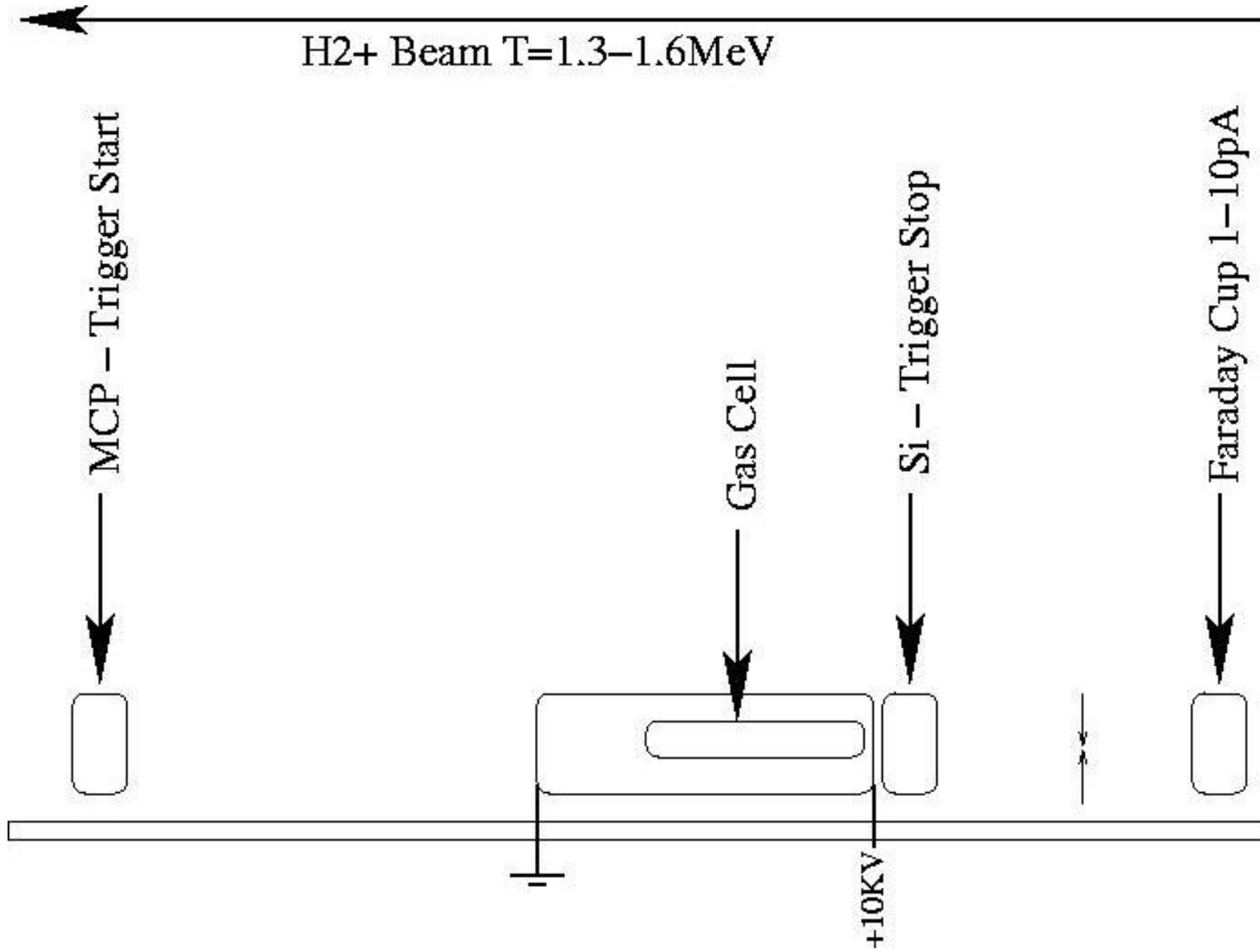
RARAF Facility – Nevis Lab/Columbia University

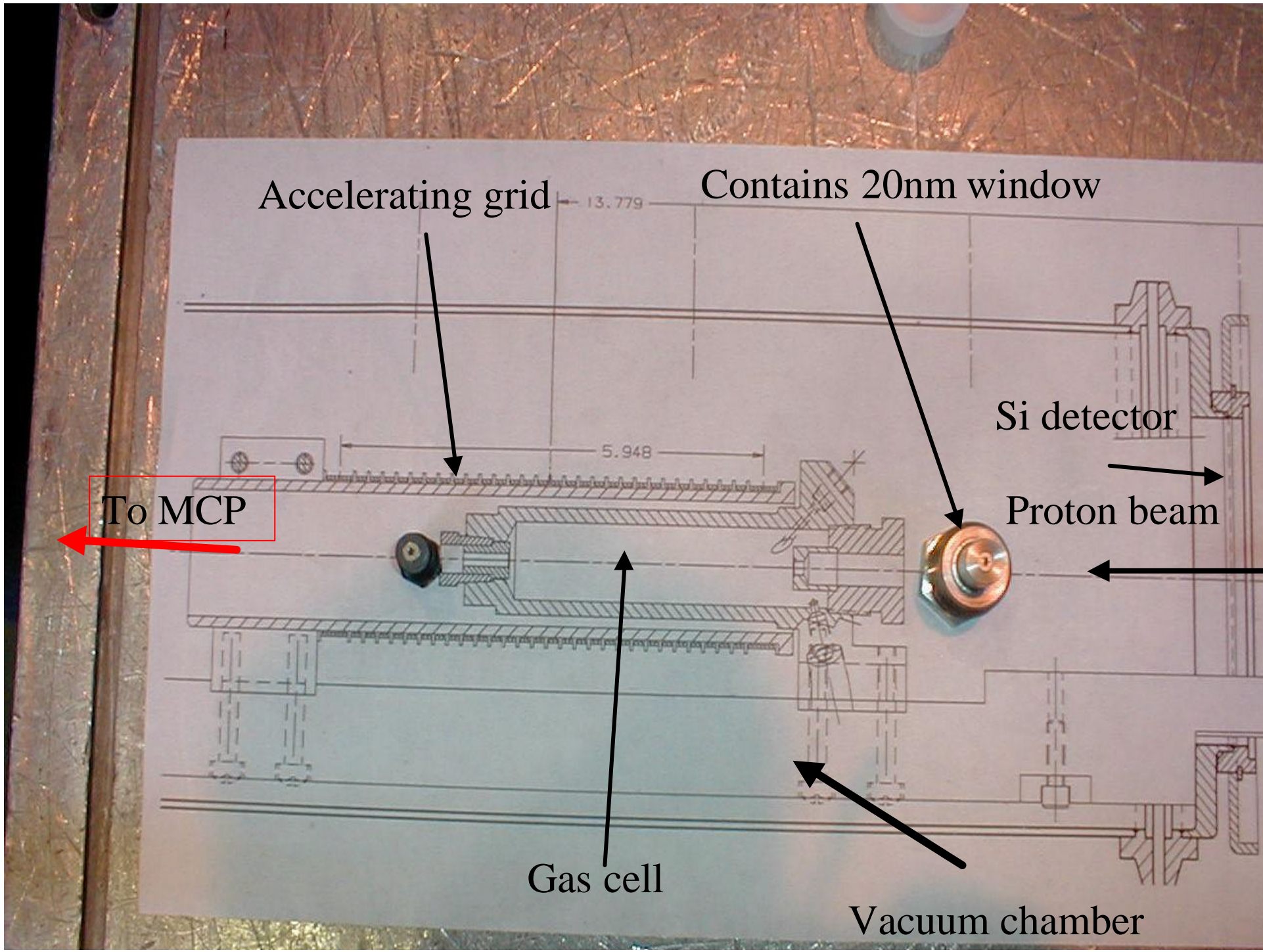


£ 4 MeV p



VandeGraaf
accelerator for
biomedical
research





Accelerating grid

Contains 20nm window

Si detector

Proton beam

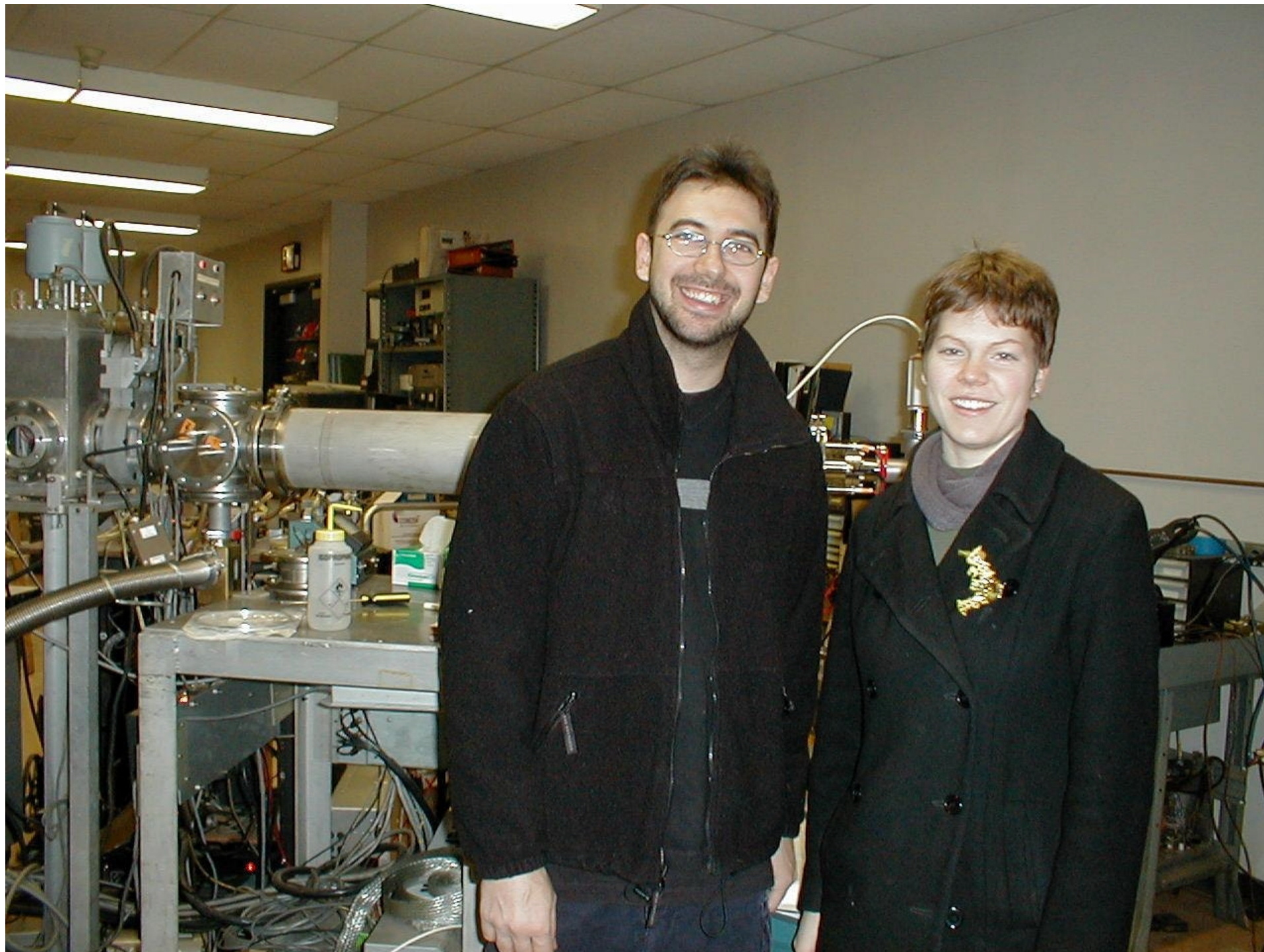
Gas cell

Vacuum chamber

To MCP

13.779

5.948

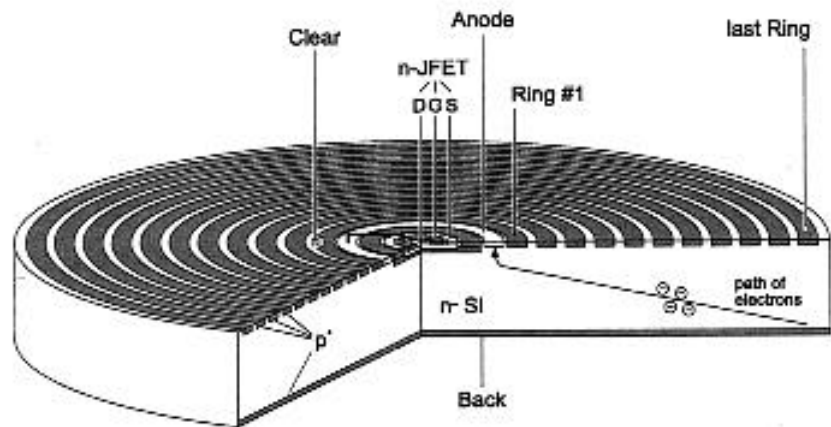


Future Plans

- Frictional cooling tests at MPI with 5T Solenoid, alpha source
- Study gas breakdown in high E,B fields
- R&D on thin windows
- Beam tests with muons to measure μ capture cross section

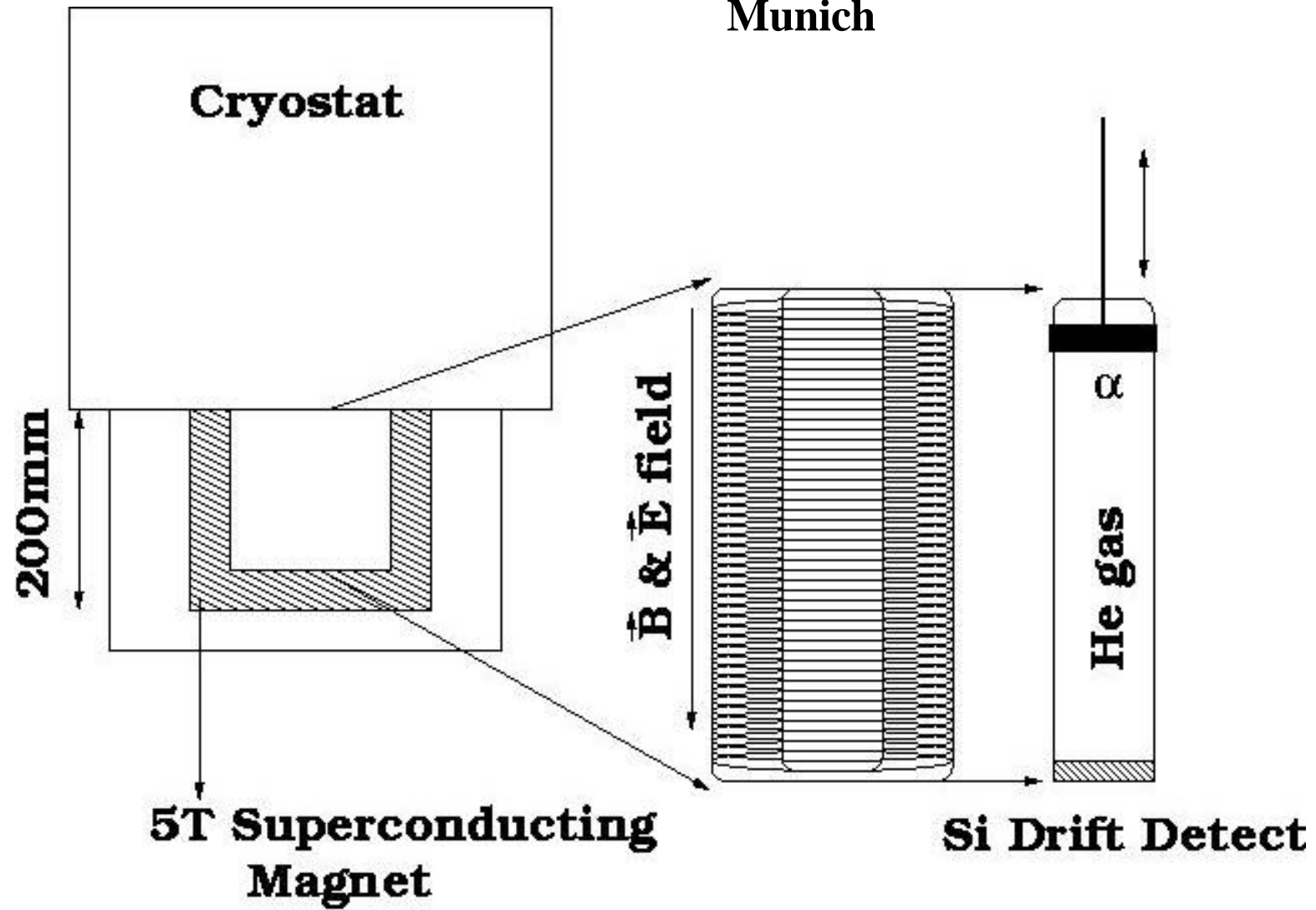


muon initially captured in large n orbit, then cascades down to n=1.
Transition n=2→n=1 releases few KeV x-ray.



Si drift detector
Developed my MPI
HLL

Lab situated at MPI-WHI in
Munich



Conclusions

- Muon Collider complex would be a boon for physics
- We need to solve the muon cooling problem
- Different schemes should be investigated
- We are doing some simulation and experimental studies of **frictional cooling**.
So far, so good, but a long way to go !